

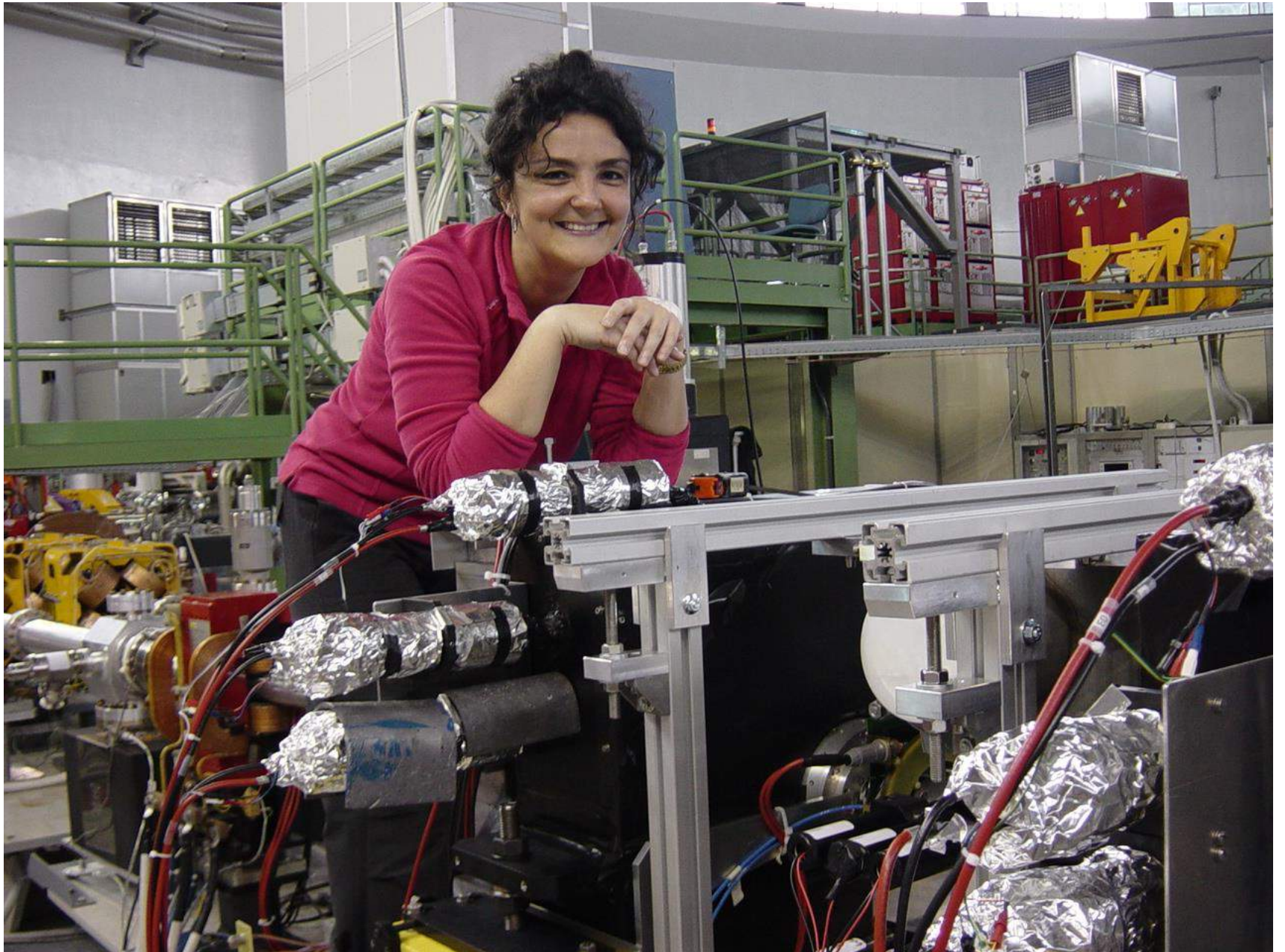


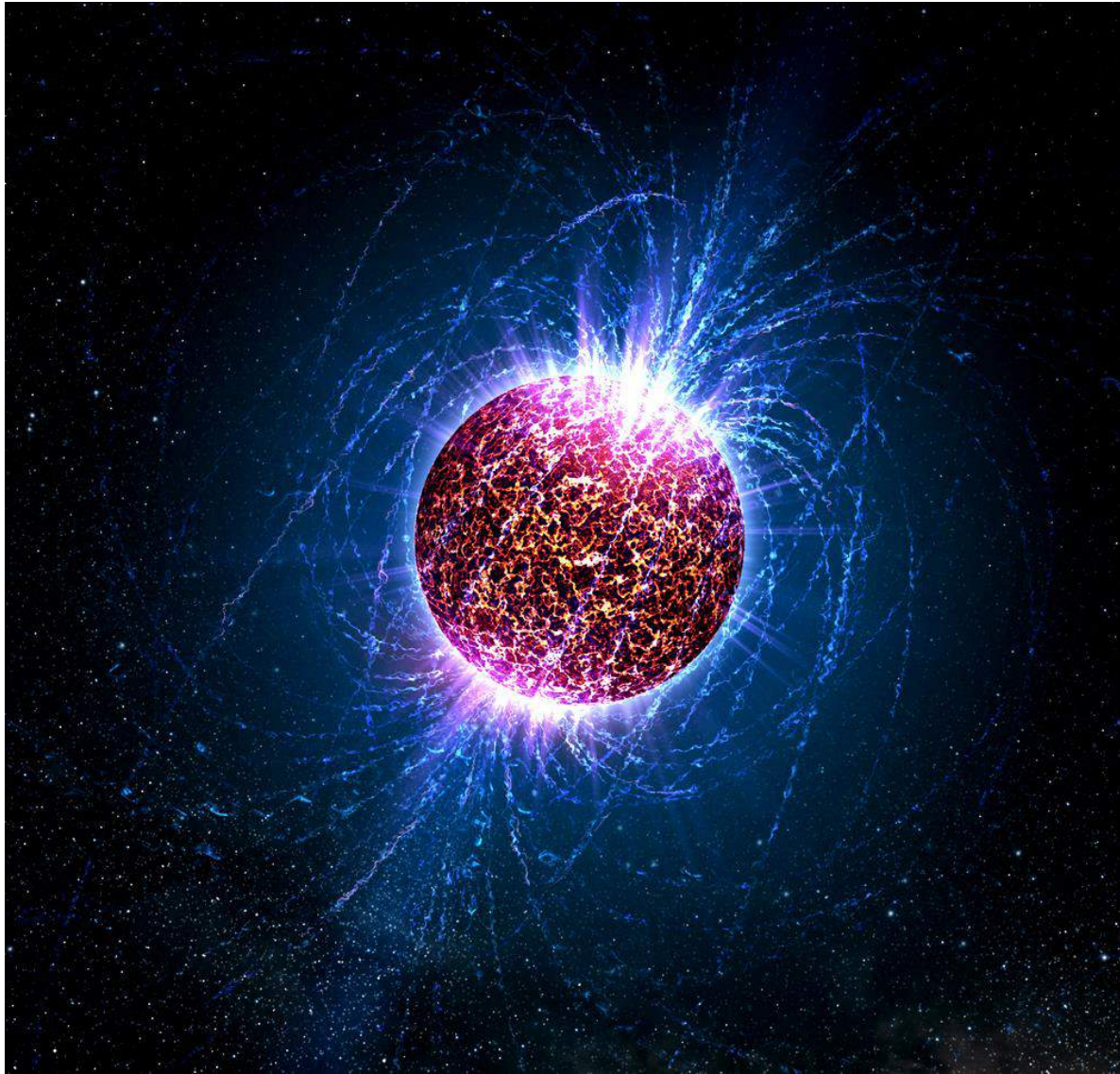
# **Dal Big Bag ai buchi neri Le meraviglie dell'Universo!**

**Catalina Curceanu, INFN-LENS**

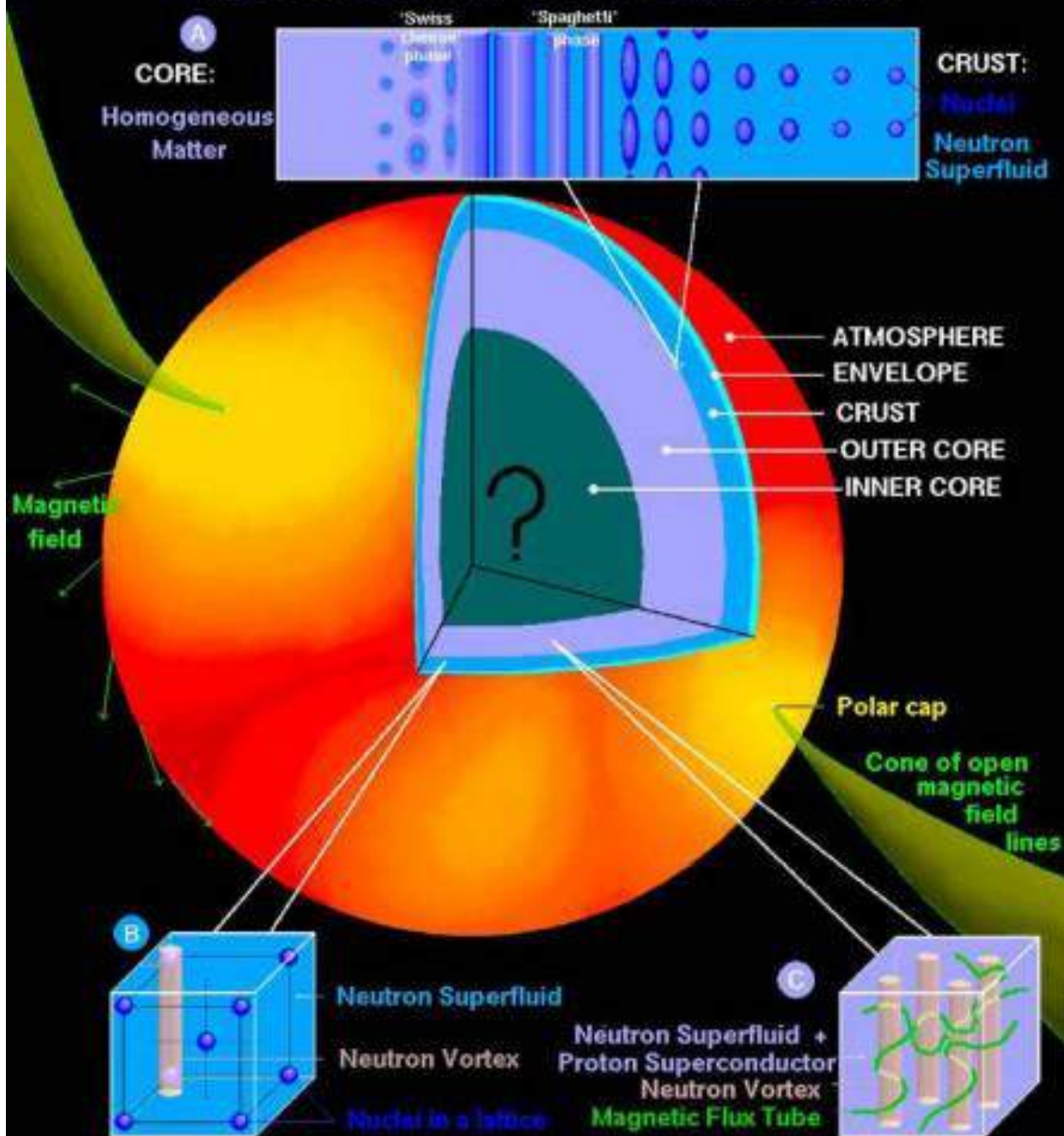
**Liceo Marconi, Pesaro, 3 maggio 2016**







# A NEUTRON STAR: SURFACE and INTERIOR





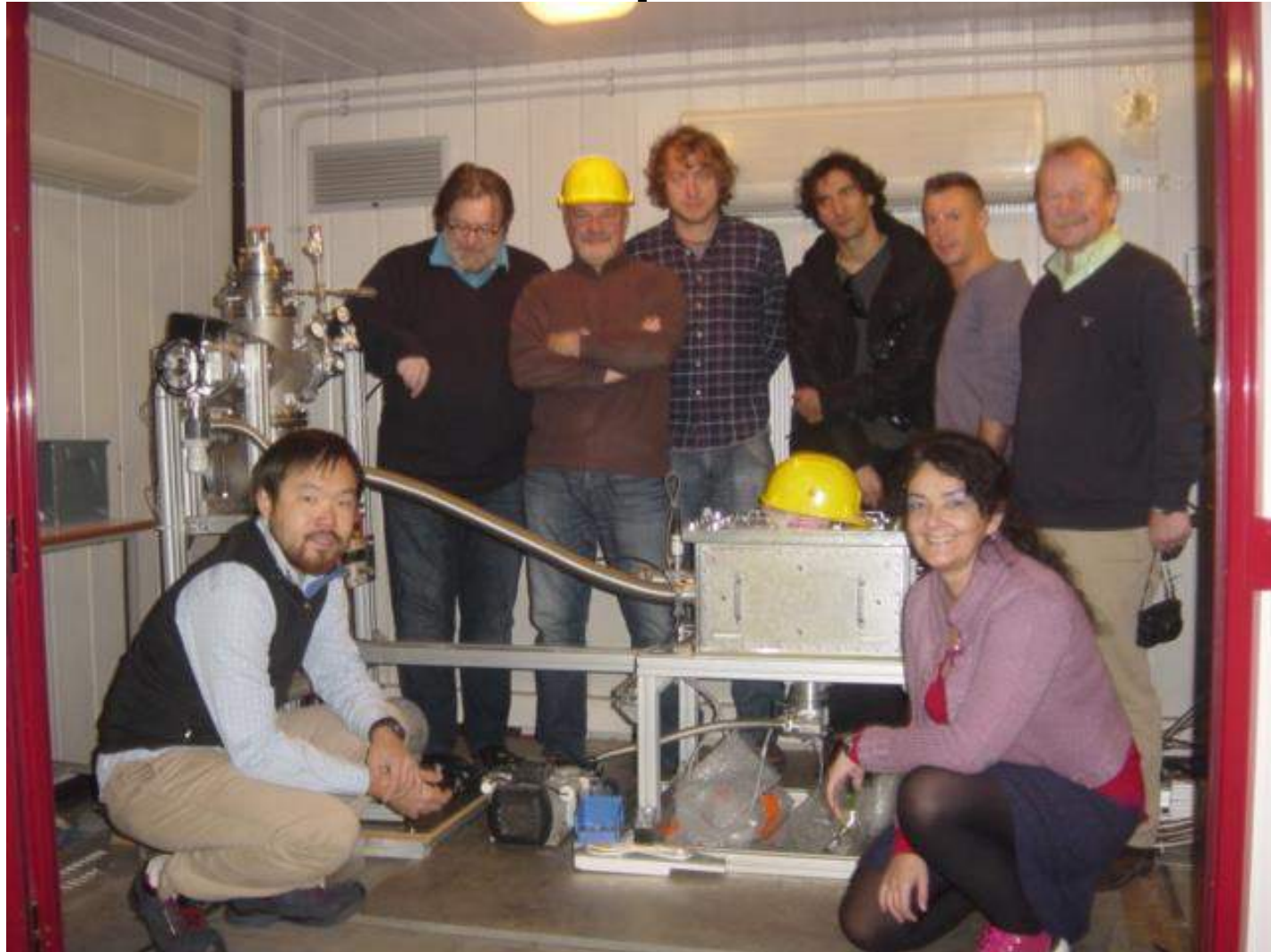


# VIP setup at LNGS





# Final setup at LNGS



# Principio di sovrapposizione quantistica

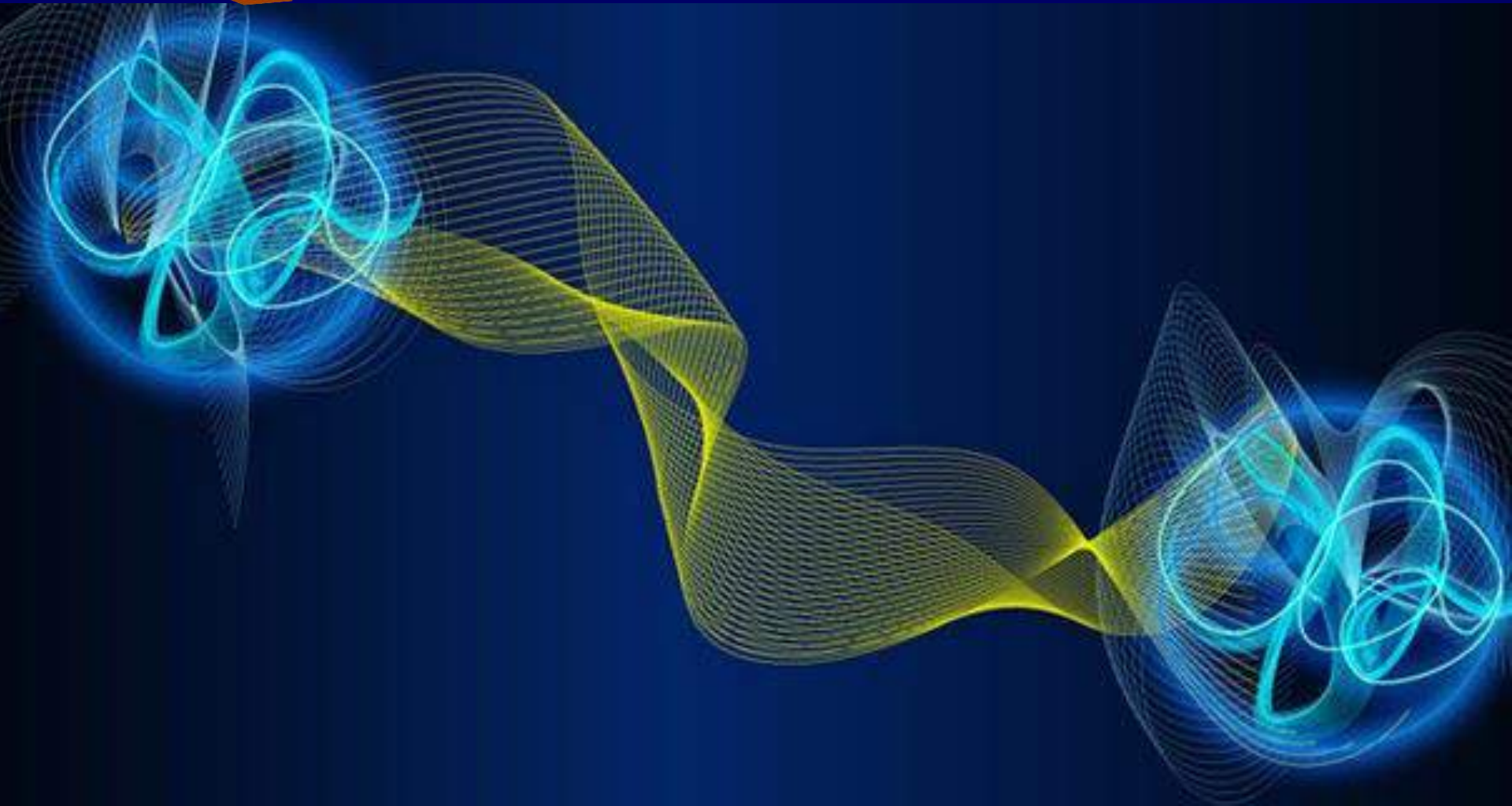
$$\Psi = c_1 \psi_1 + c_2 \psi_2$$

$$|\Psi\rangle = \frac{|\text{alive}\rangle + |\text{dead}\rangle}{\sqrt{2}}$$





# L'entanglement



# After 100 years of General Relativity...

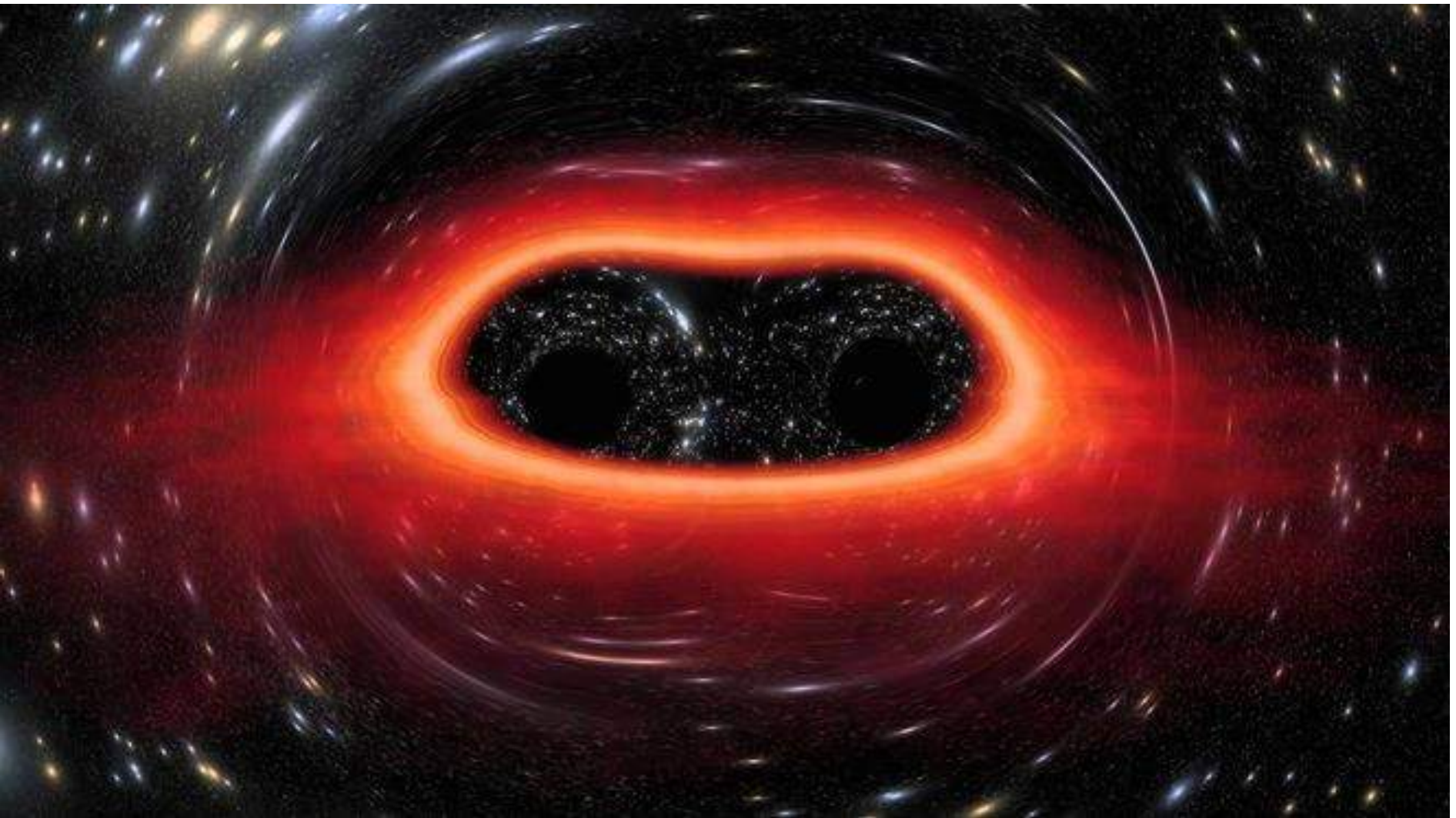
Imagine travelling through space on a beam of light at the speed of light.



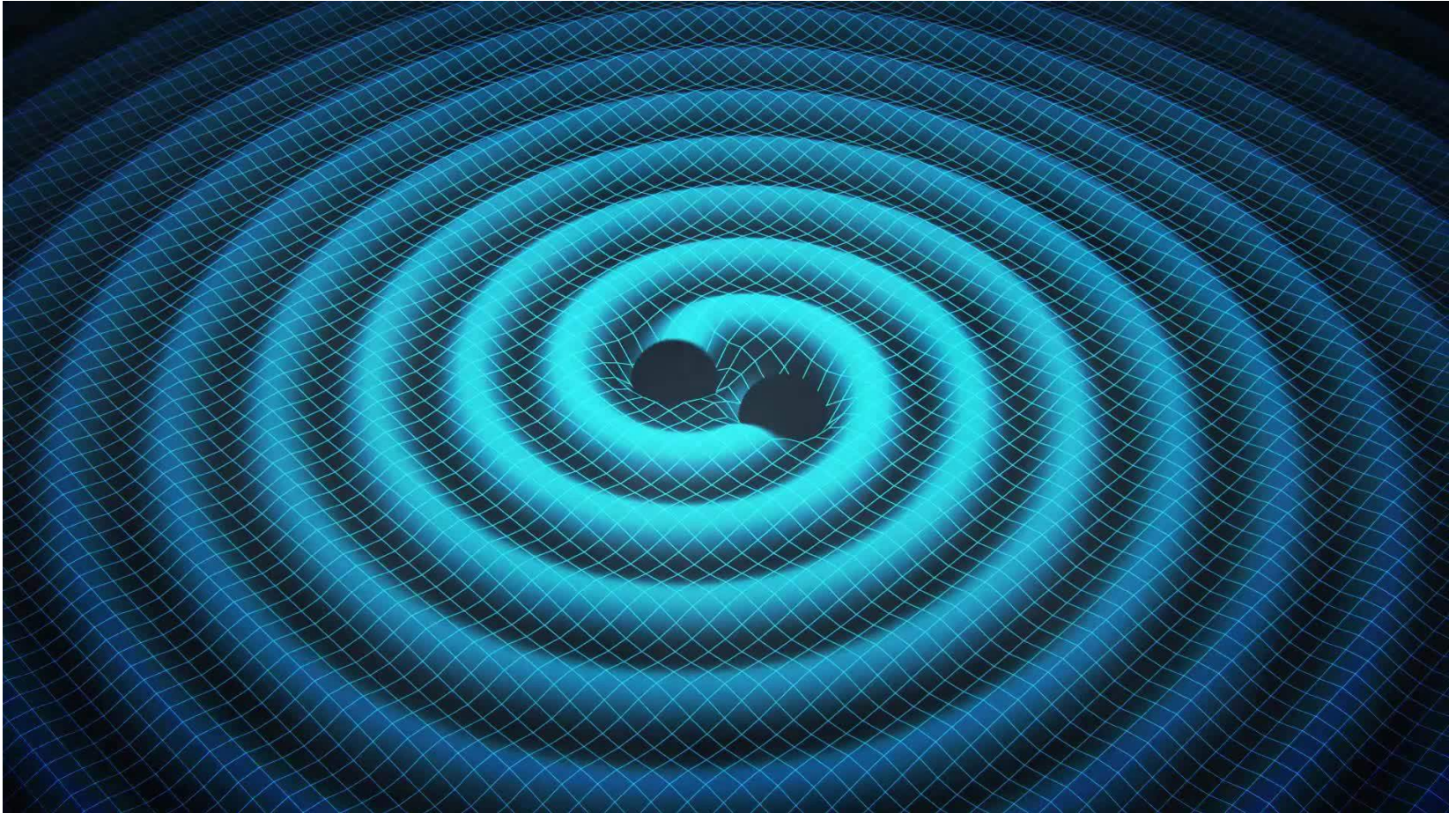
Albert Einstein, theory of relativity, gravity, velocity, energy, mass, speed, time,  $E=mc^2$  Albert Ein

Bobonart

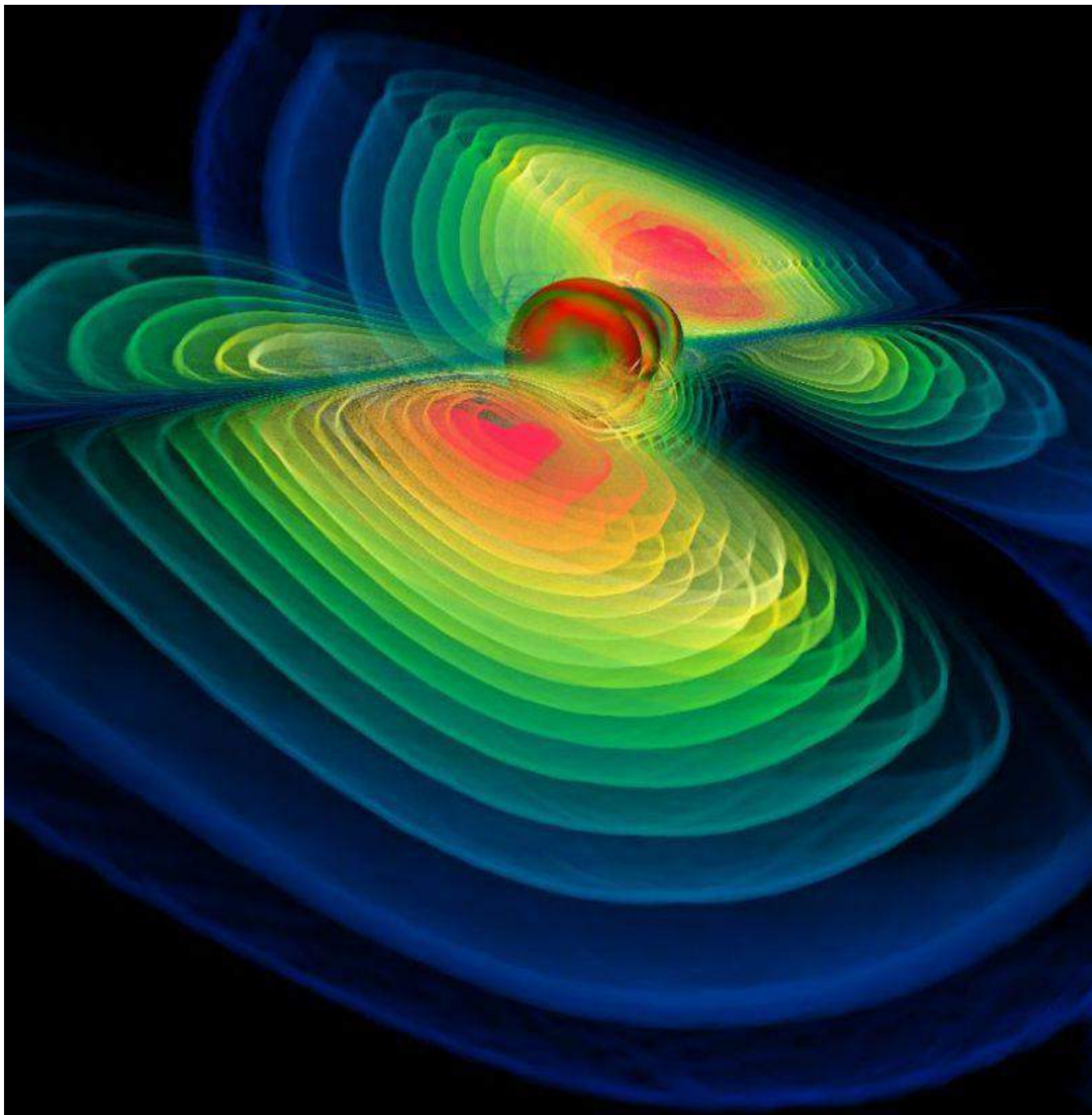
**Scoperta delle onde gravitazionali (14 Sept.  
2015 -> 11 Feb 2016)**



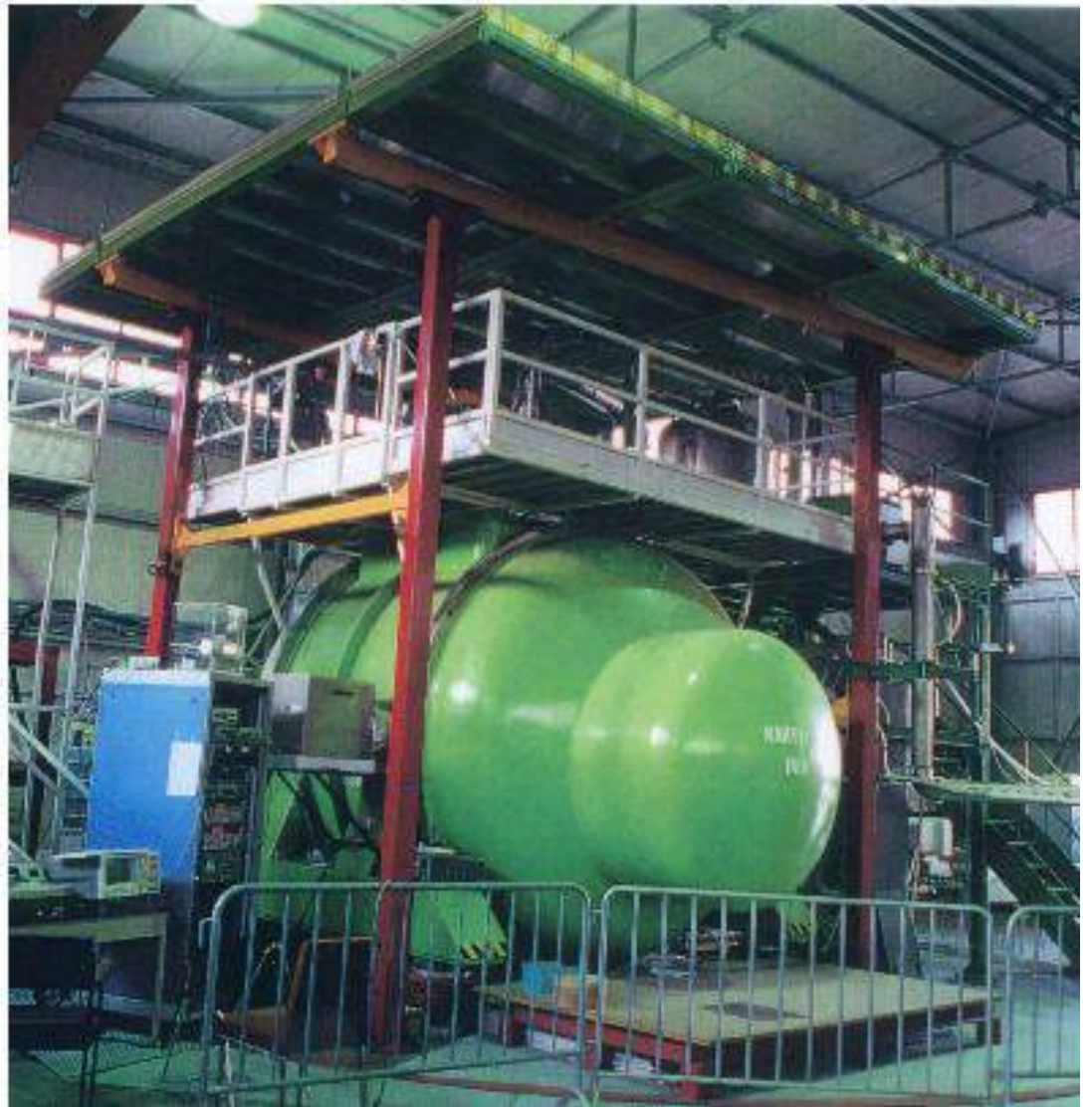
**Onde gravitazionali (14 Sept. 2015 -> 11 Feb 2016)**







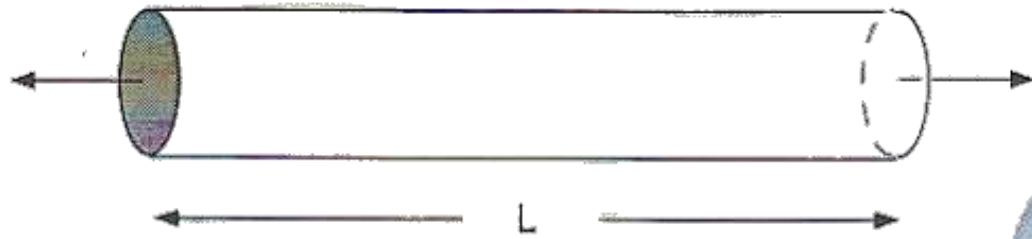
## ***NAUTILUS ai LNF-INFN***



# *Search for gravitational waves:*

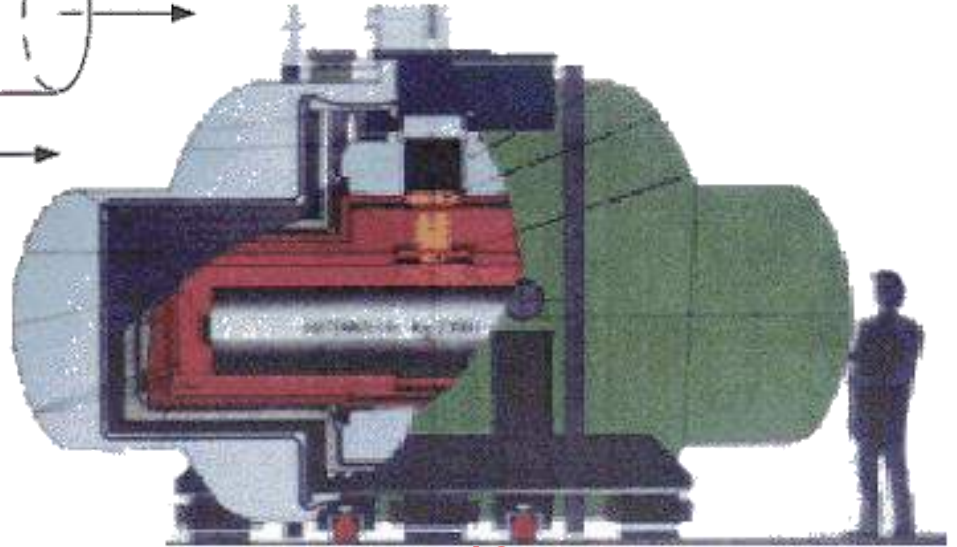
↓ gw

## *NAUTILUS*



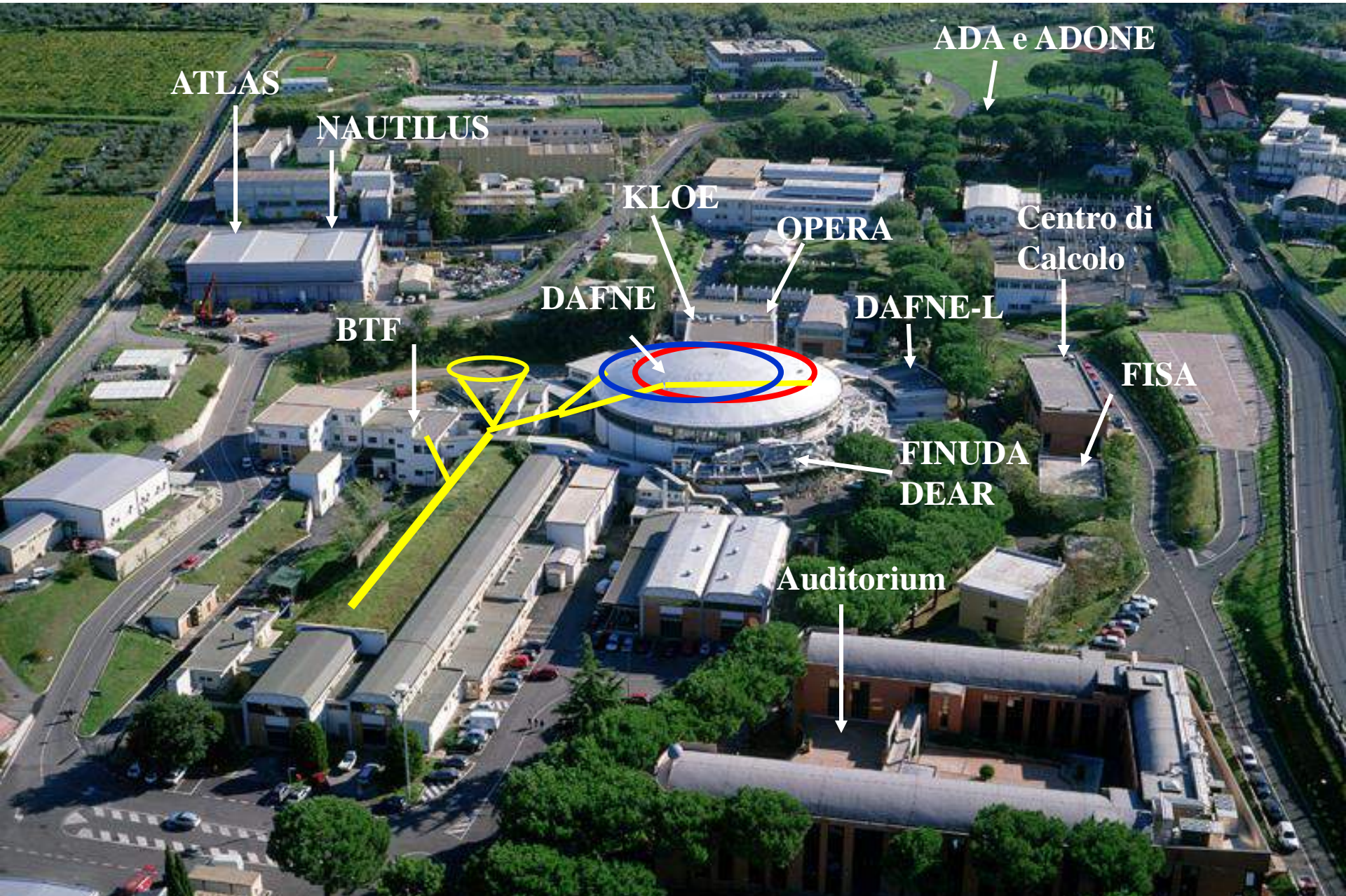
vibrations in the bar  
(for Al, L=3 m, f=915 Hz)

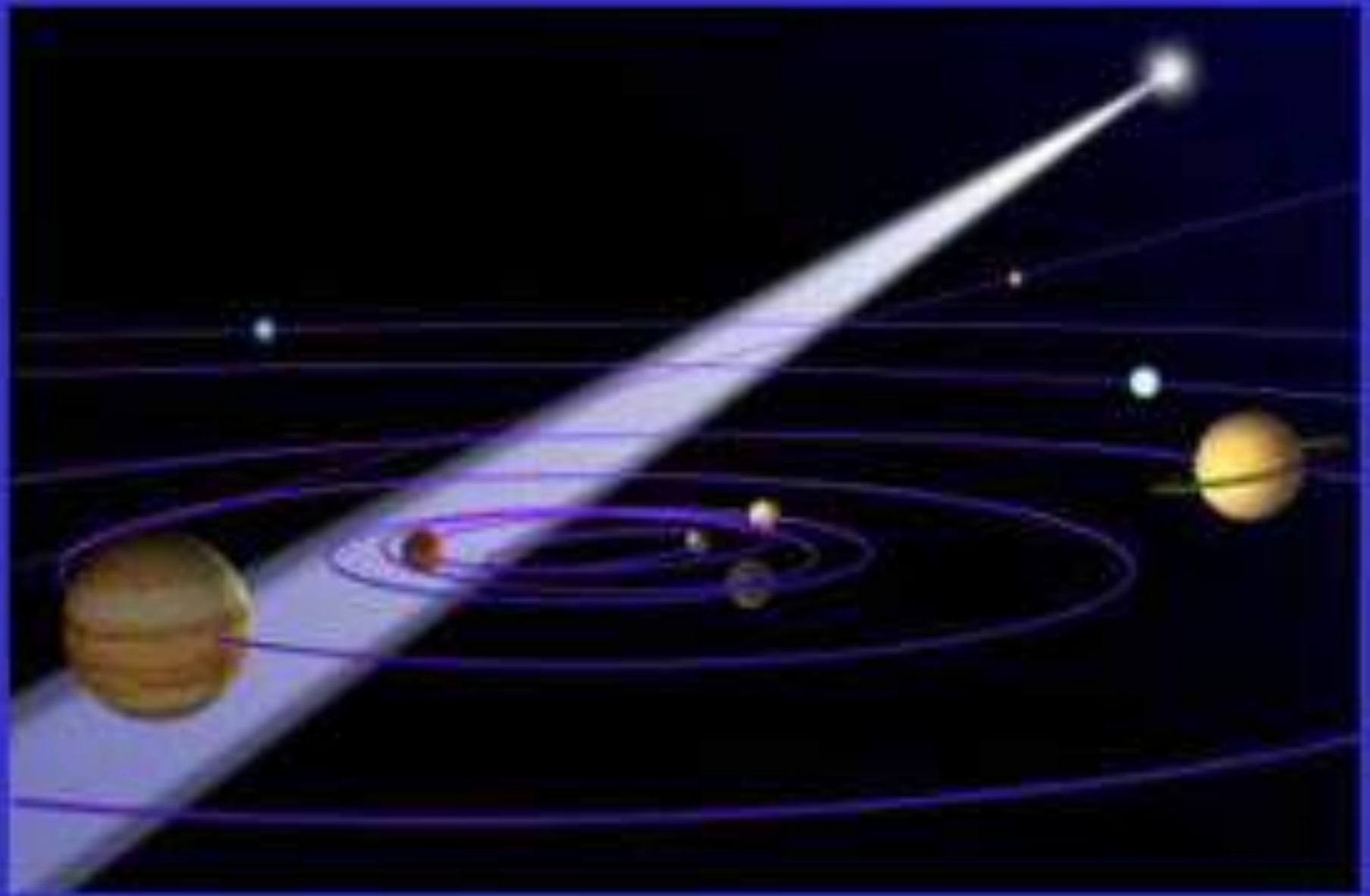
$$\frac{\Delta L}{L} \approx h$$

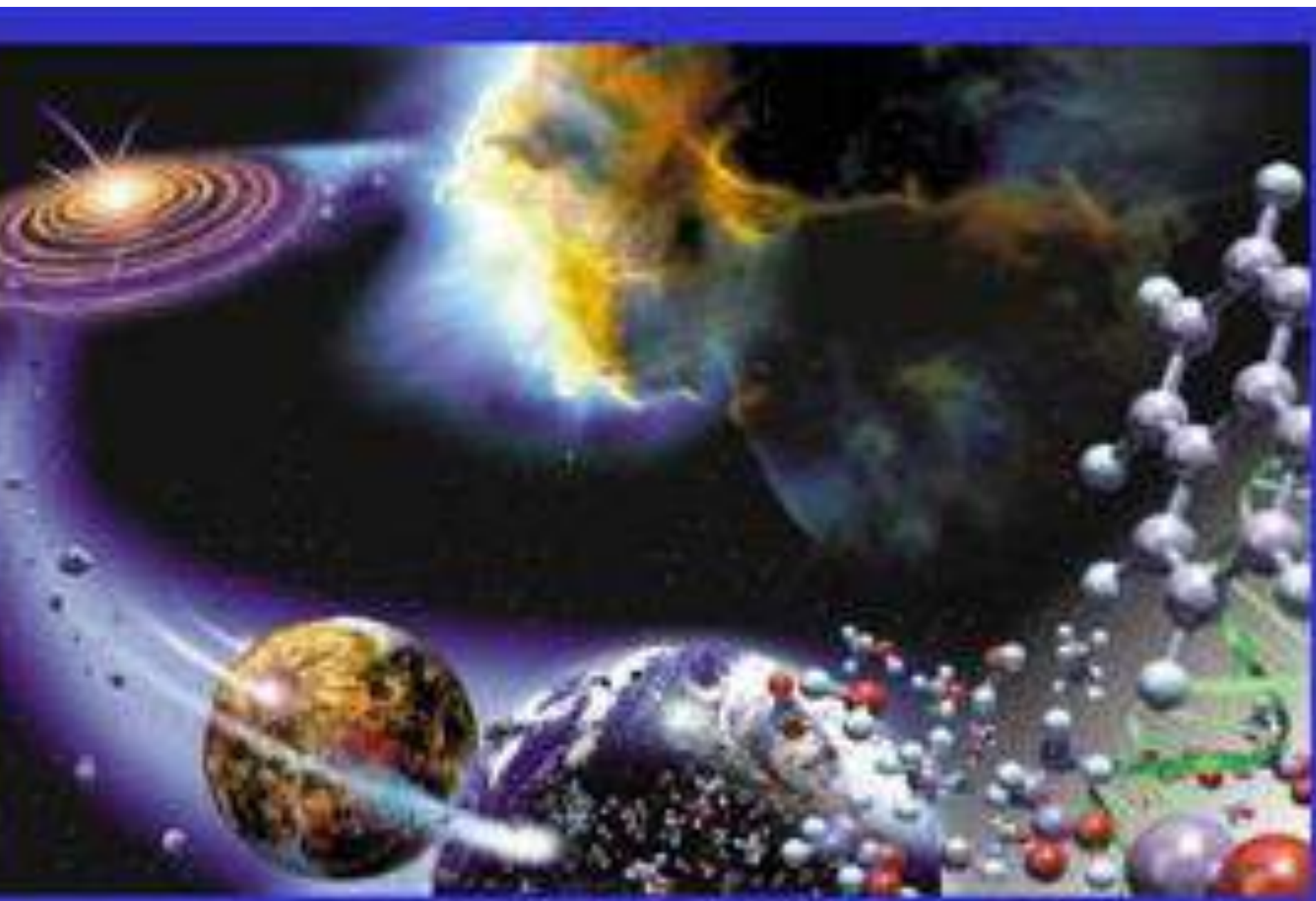


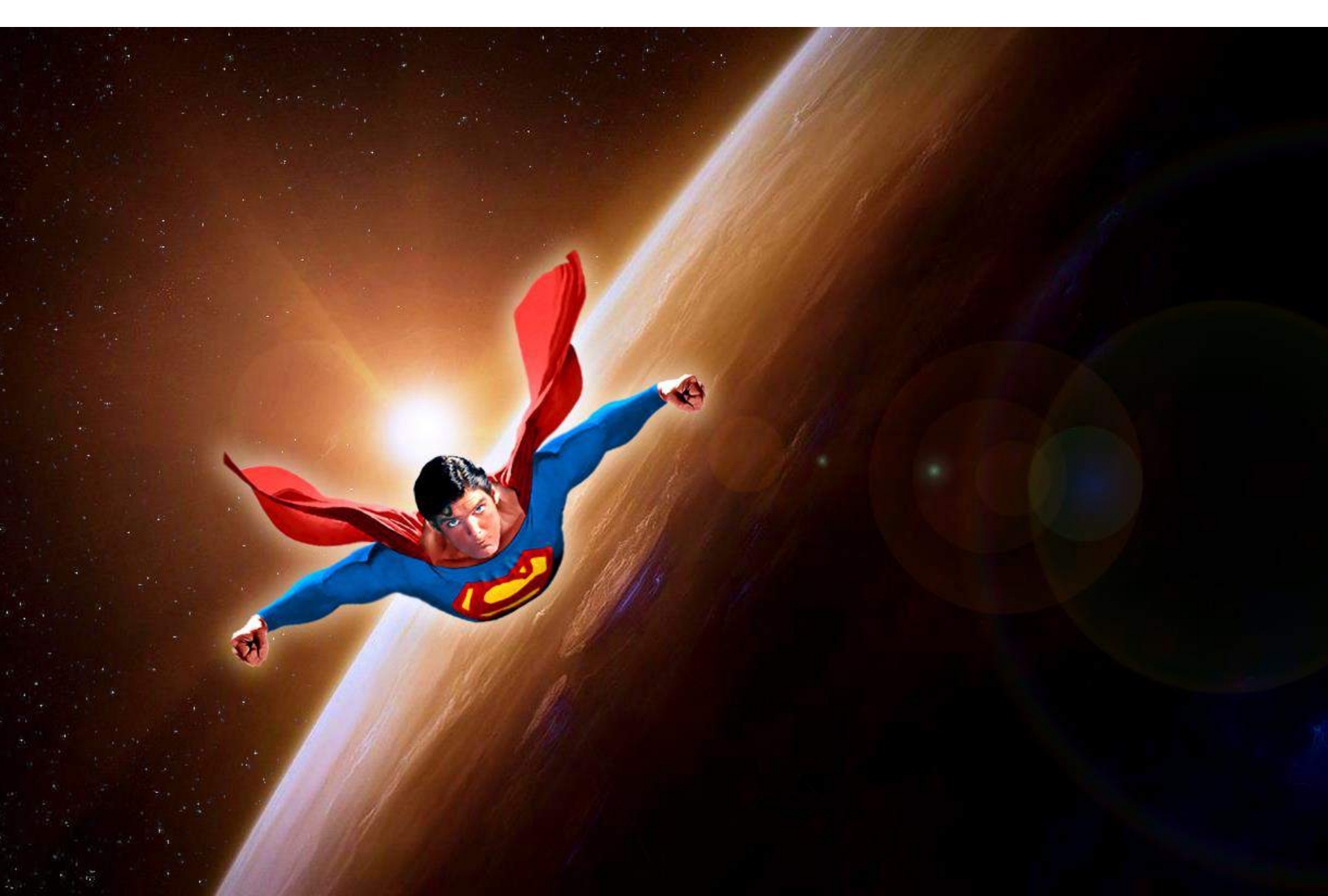
- **Supernova in our Galassia  $h=10^{-18}$**
- **Supernova in Virgo  $h=10^{-21}$**
- **Thermal noise @  $T=300$  K,  $\Delta L=10^{-16}$  m**
- **Thermal noise @  $T=3$  K,  $\Delta L=10^{-17}$  m**
- **Thermal noise @  $T=300$  mK  $\rightarrow \Delta L=10^{-18}$  m**

*Laboratori Nazionali di Frascati, info: <http://www.lnf.infn.it/sis/>*



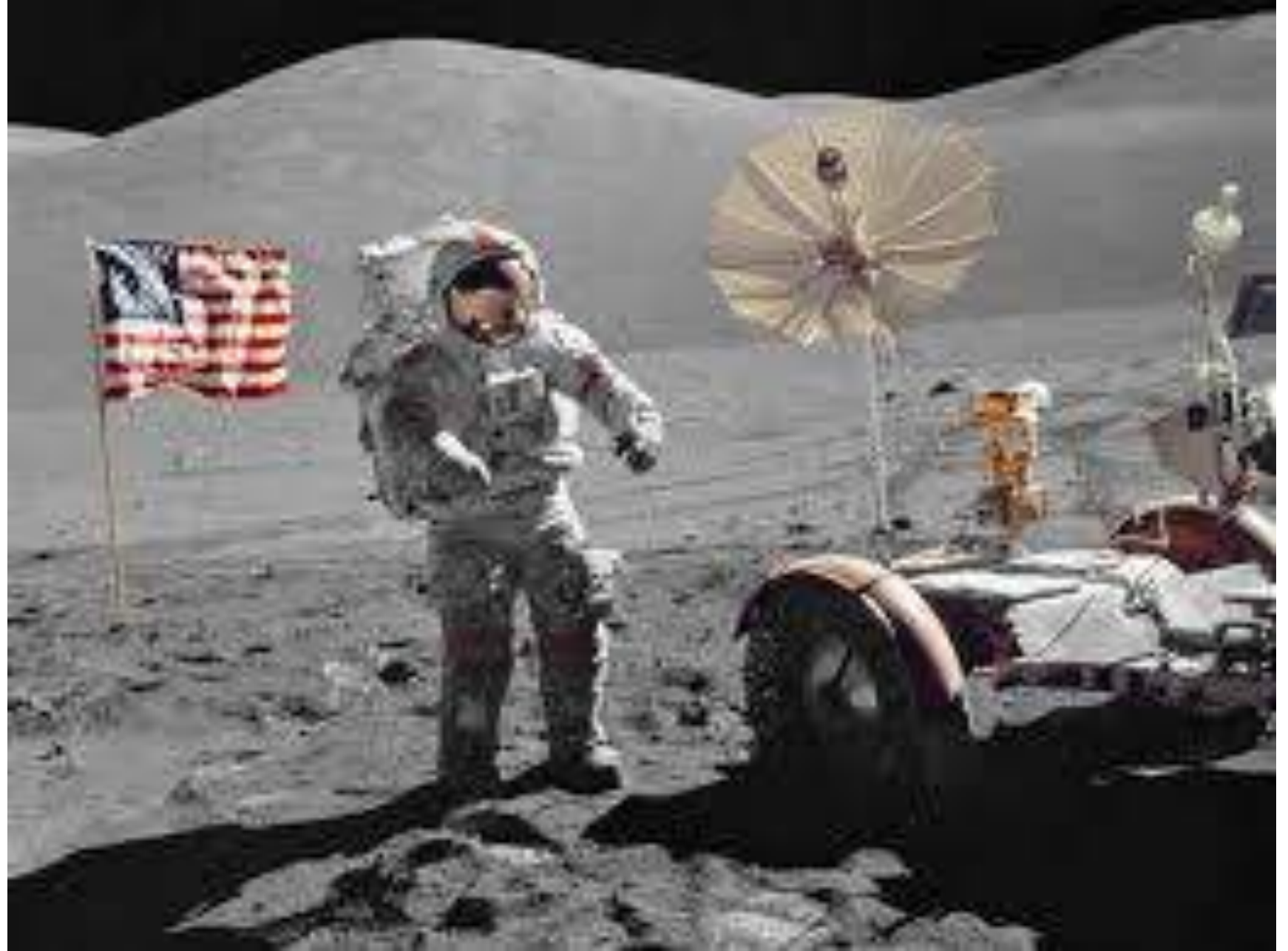






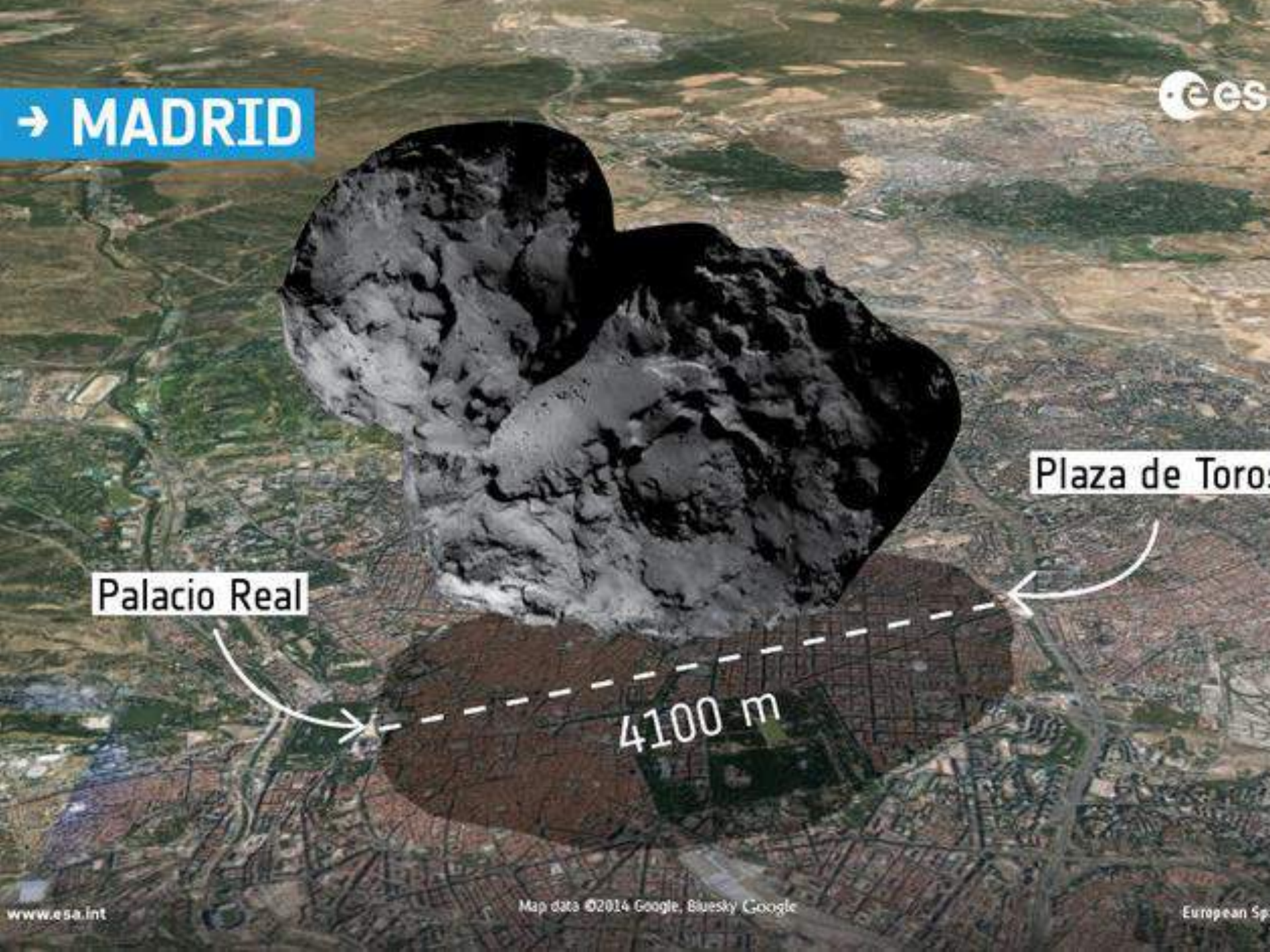








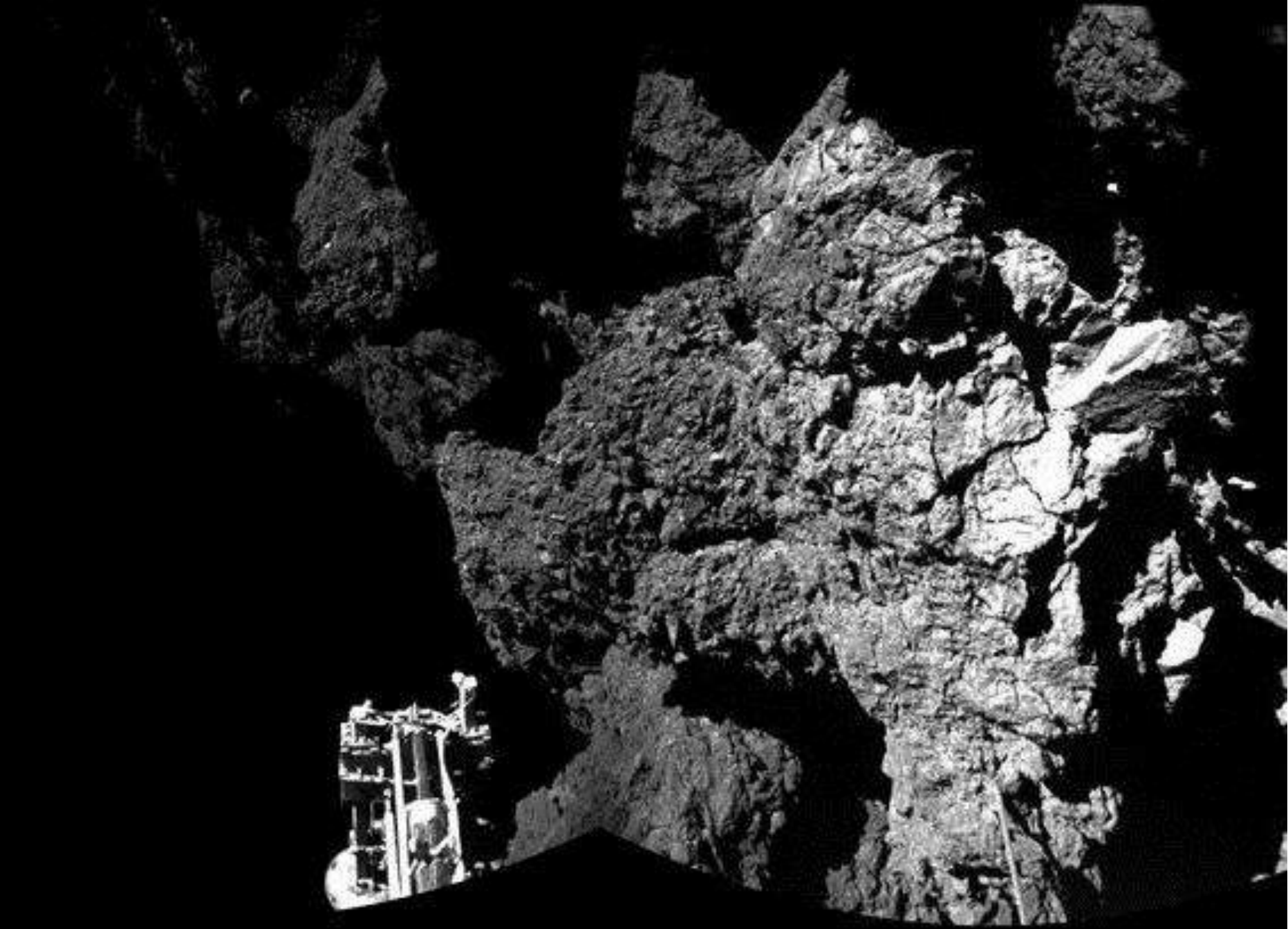
→ MADRID



Palacio Real

Plaza de Toros

4100 m



Philae carries 10 research tools and a transmitter that can communicate data to Rosetta to be relayed home. Philae can swivel on its base to sample different areas.

Lander

A monitor studies the magnetic field and solar wind

Ovens analyze comet material

When they're on opposite sides of the comet, Rosetta and Philae send each other radio signals to map 67P's internal structure

Solar cells gather weak sunlight to power the craft after the main battery dies

Seismographs in Philae's feet detect activity in 67P's core as

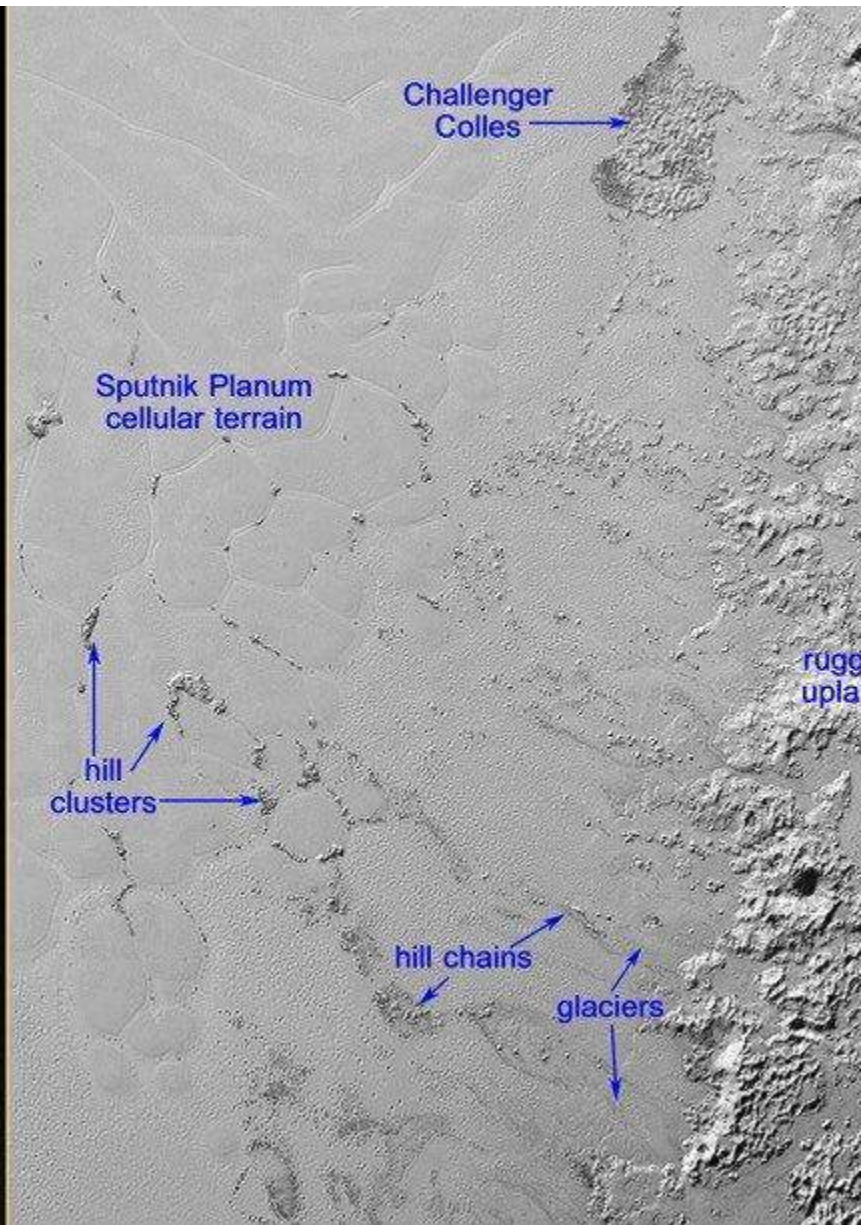
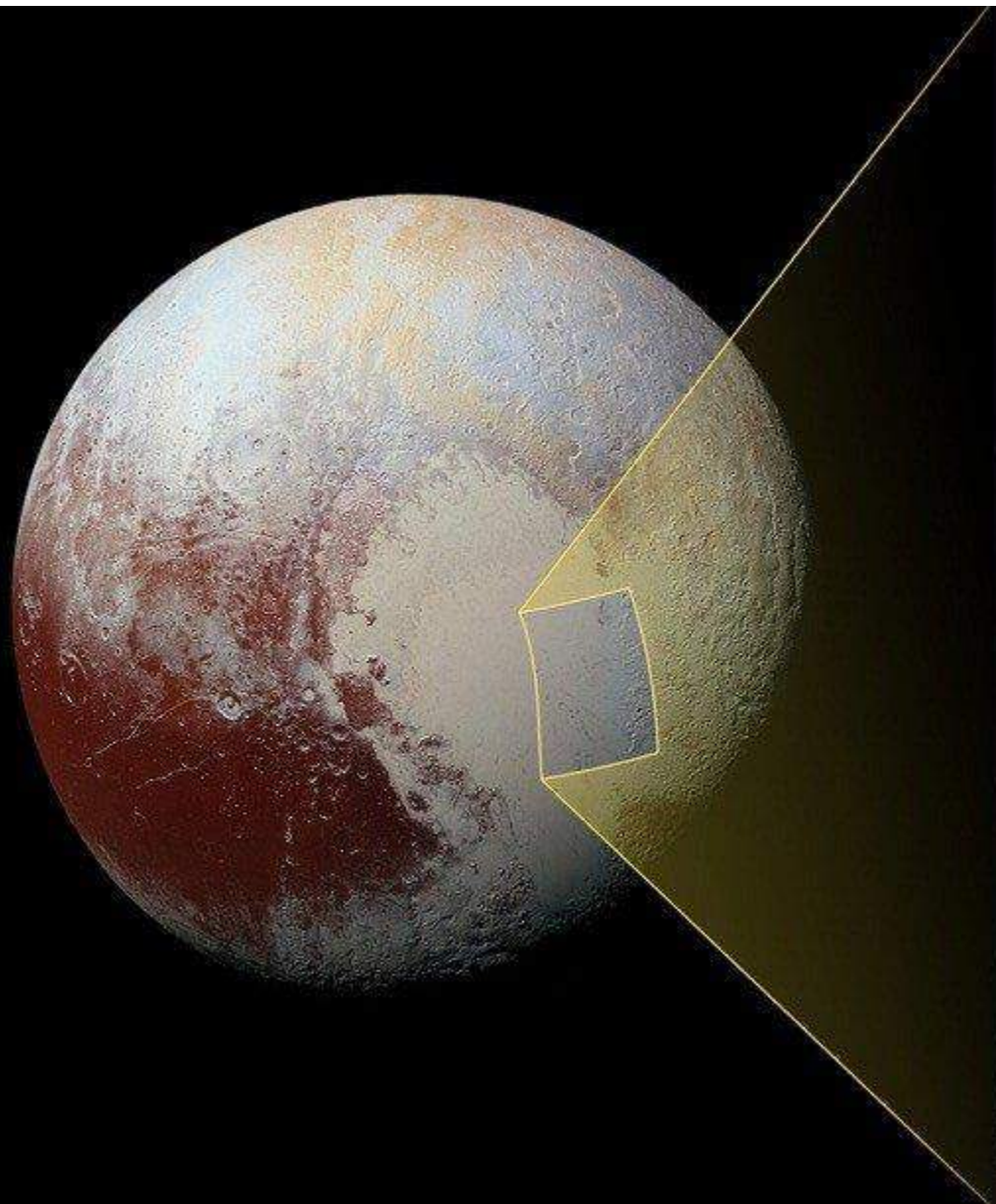
#### INSIDE THE COMET

A probe will penetrate the comet if the surface layers are porous

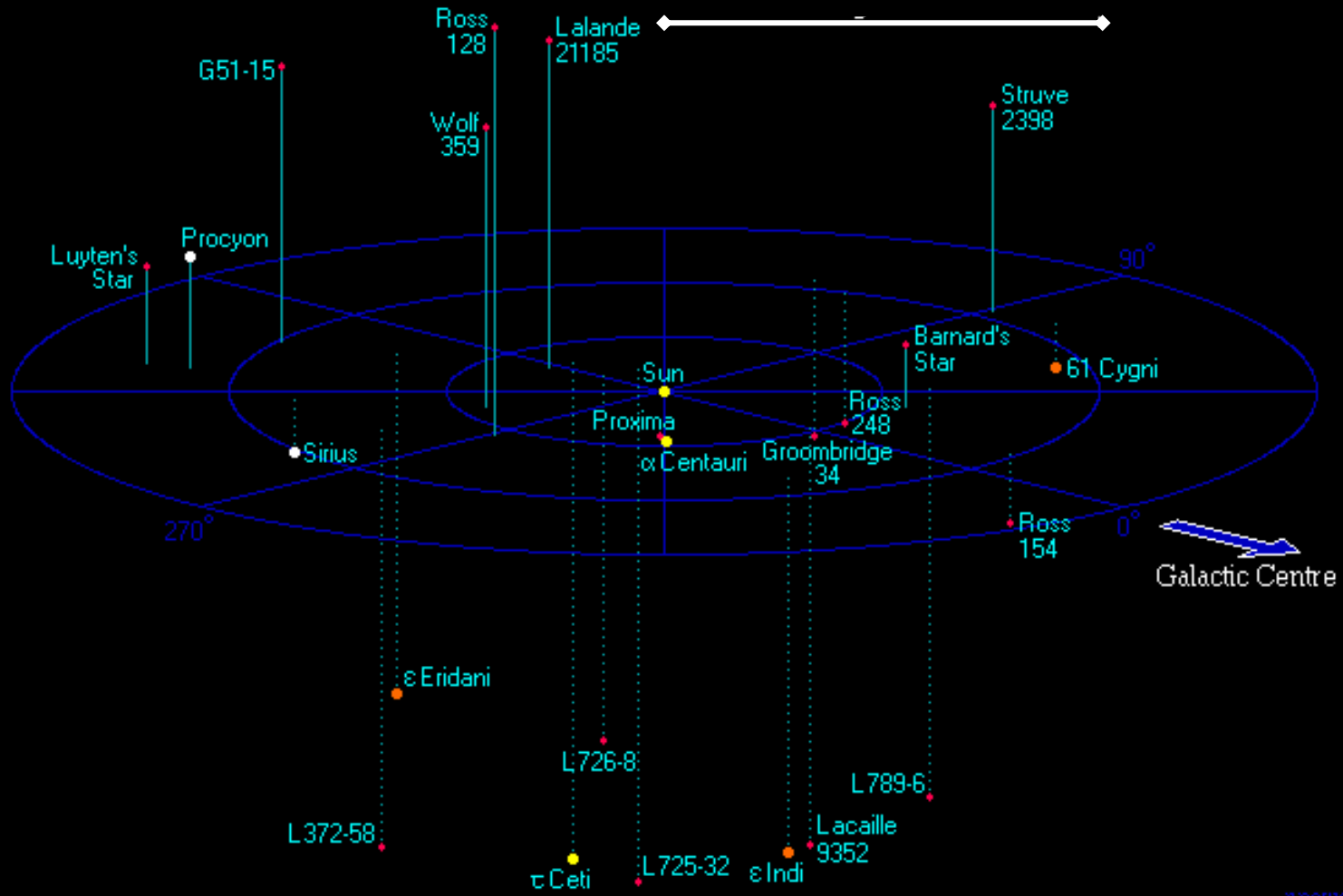
11 in. deep

Measurement will include the temperature and density of the comet's interior

**10** YEARS TO REACH

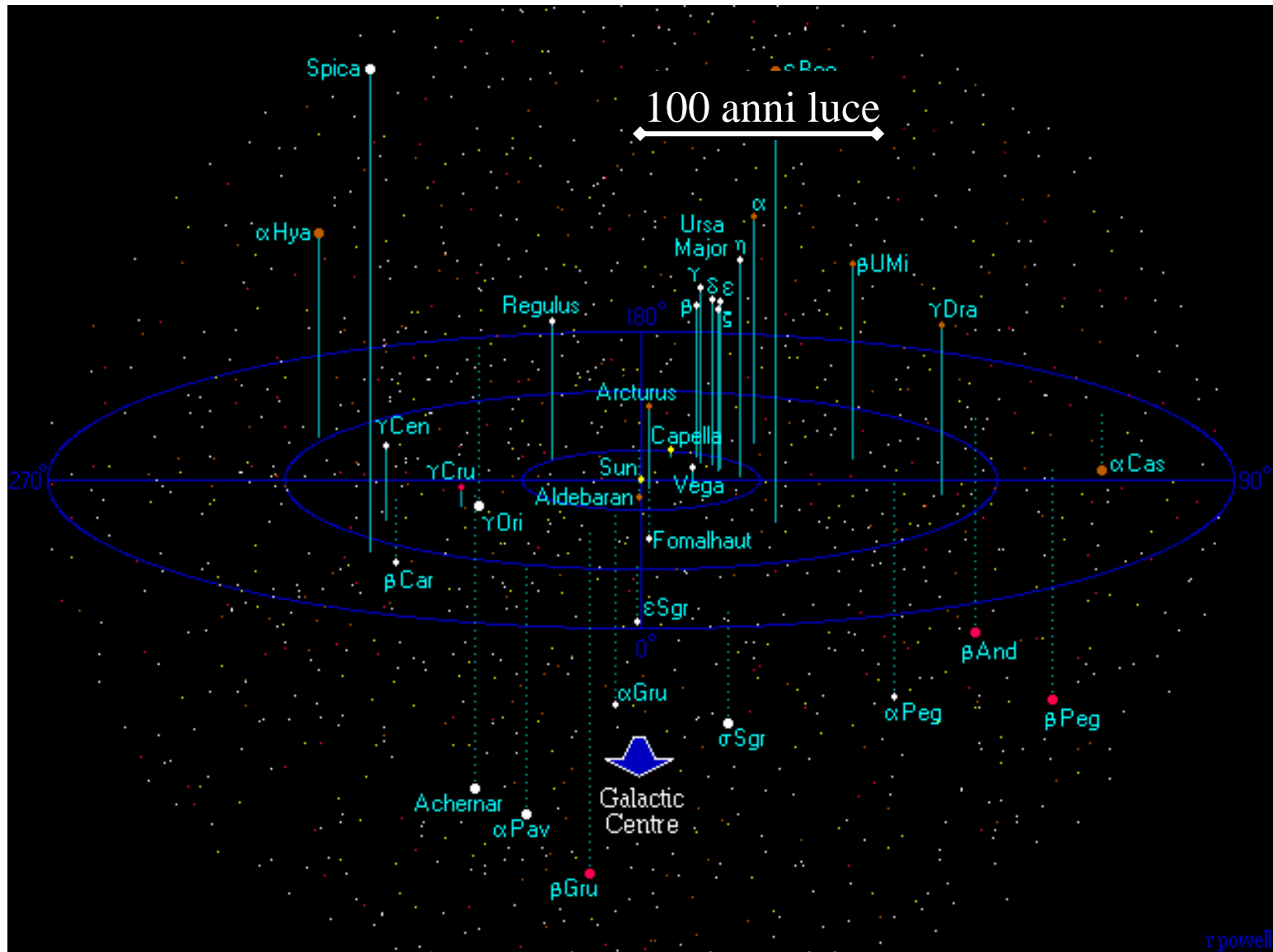


10 anni luce



r powell

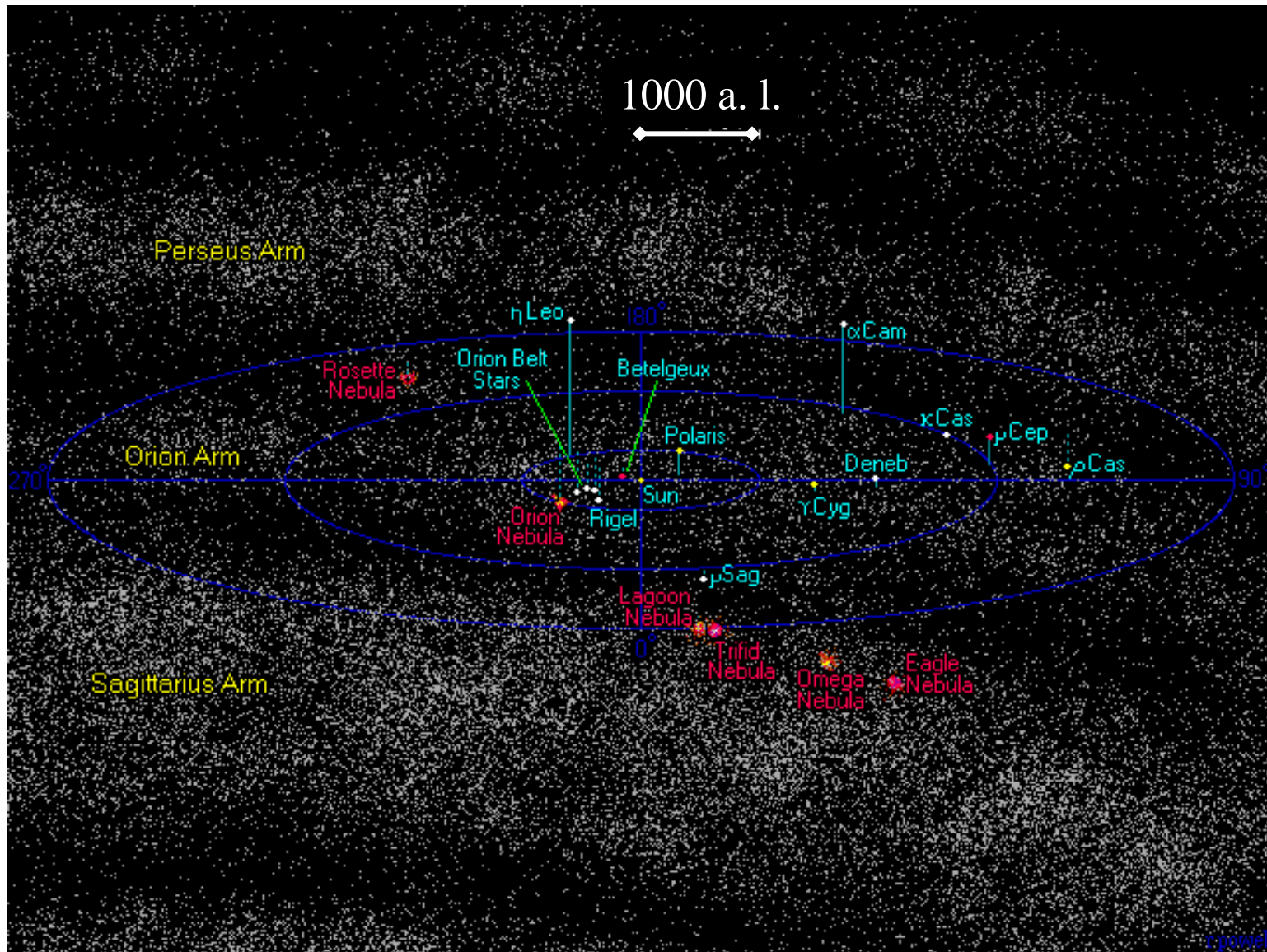
Zoom Out x20



Zoom In x20

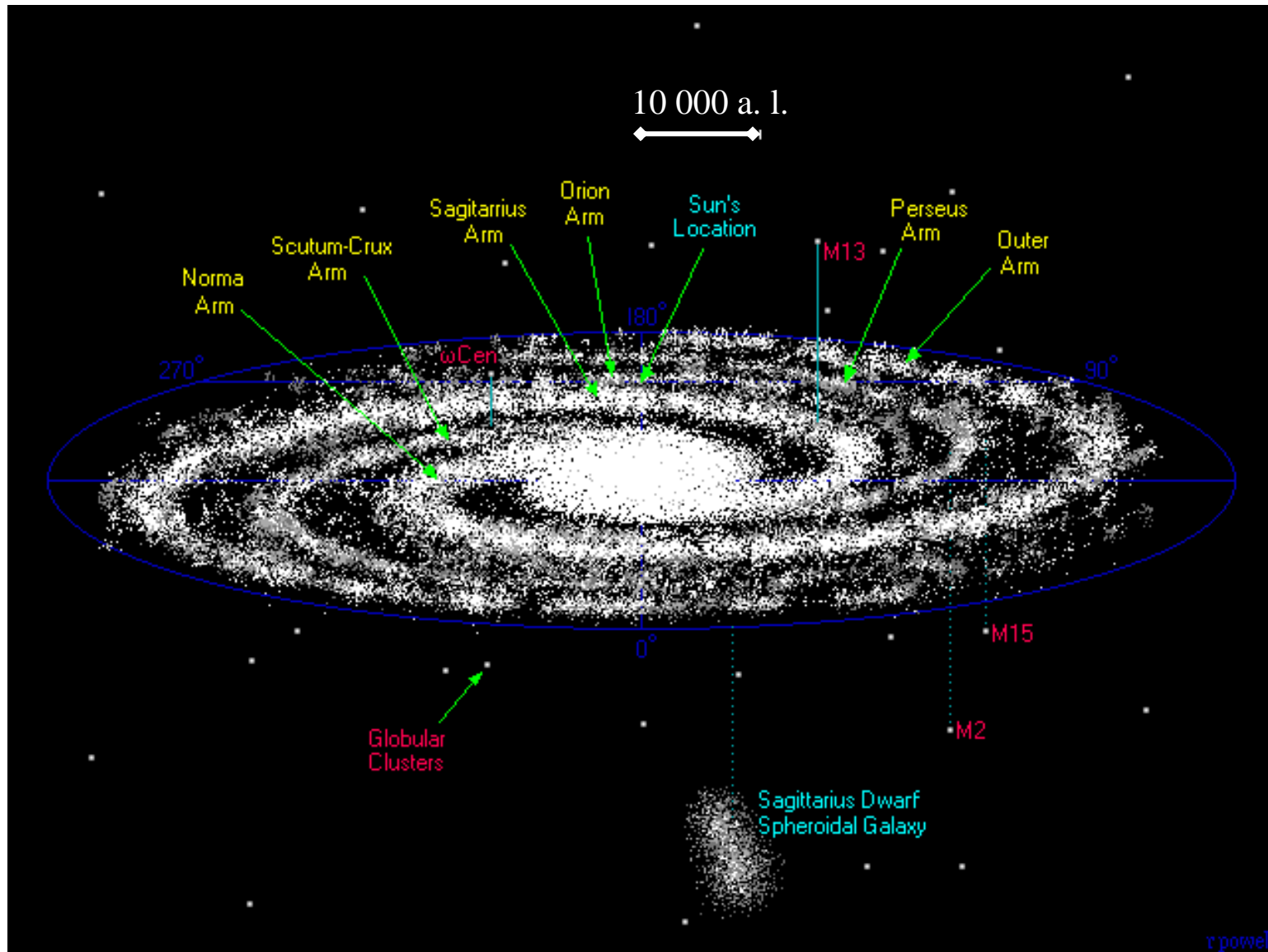
Zoom In x20





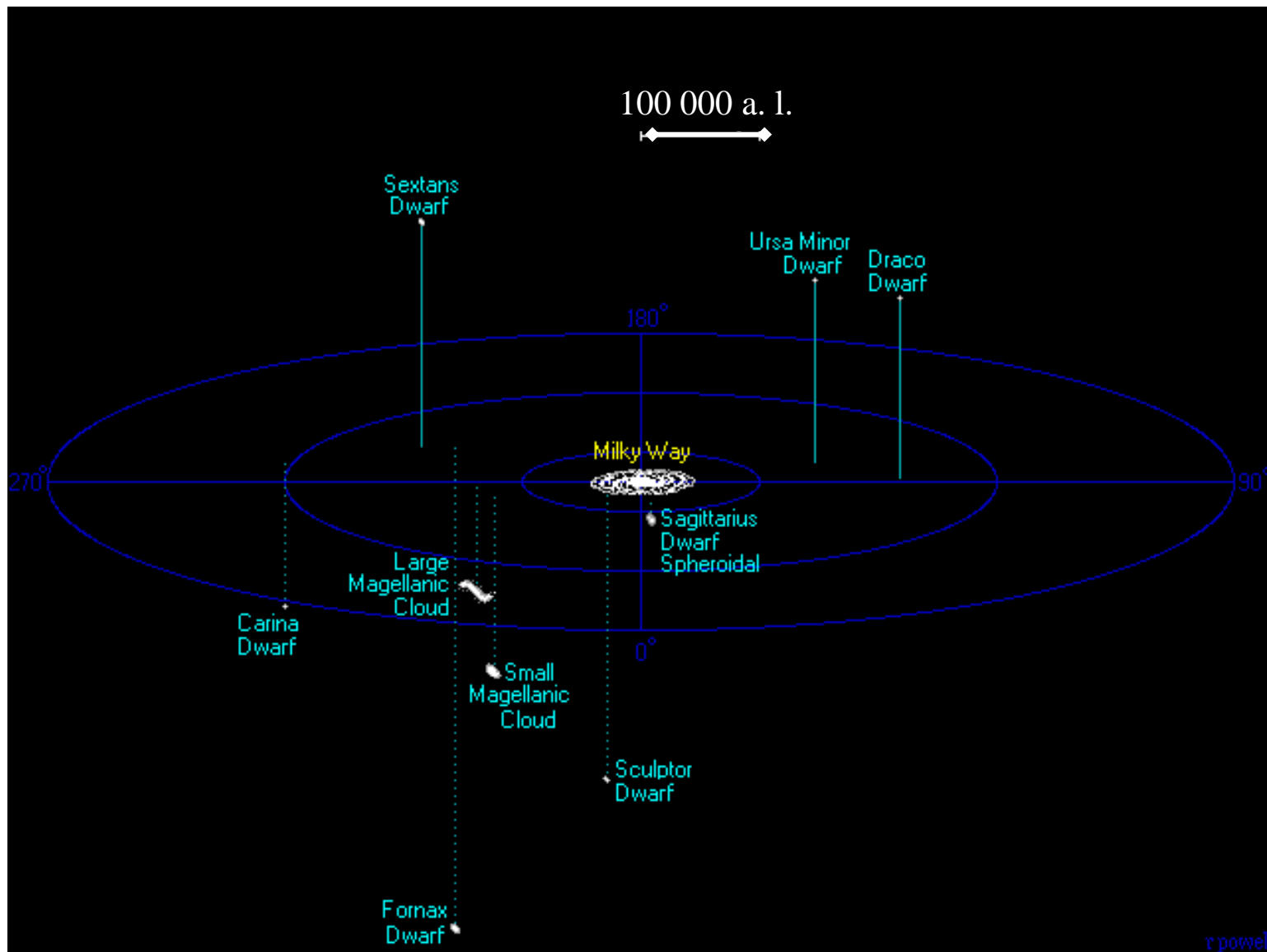
Zoom In x20

Zoom Out x10



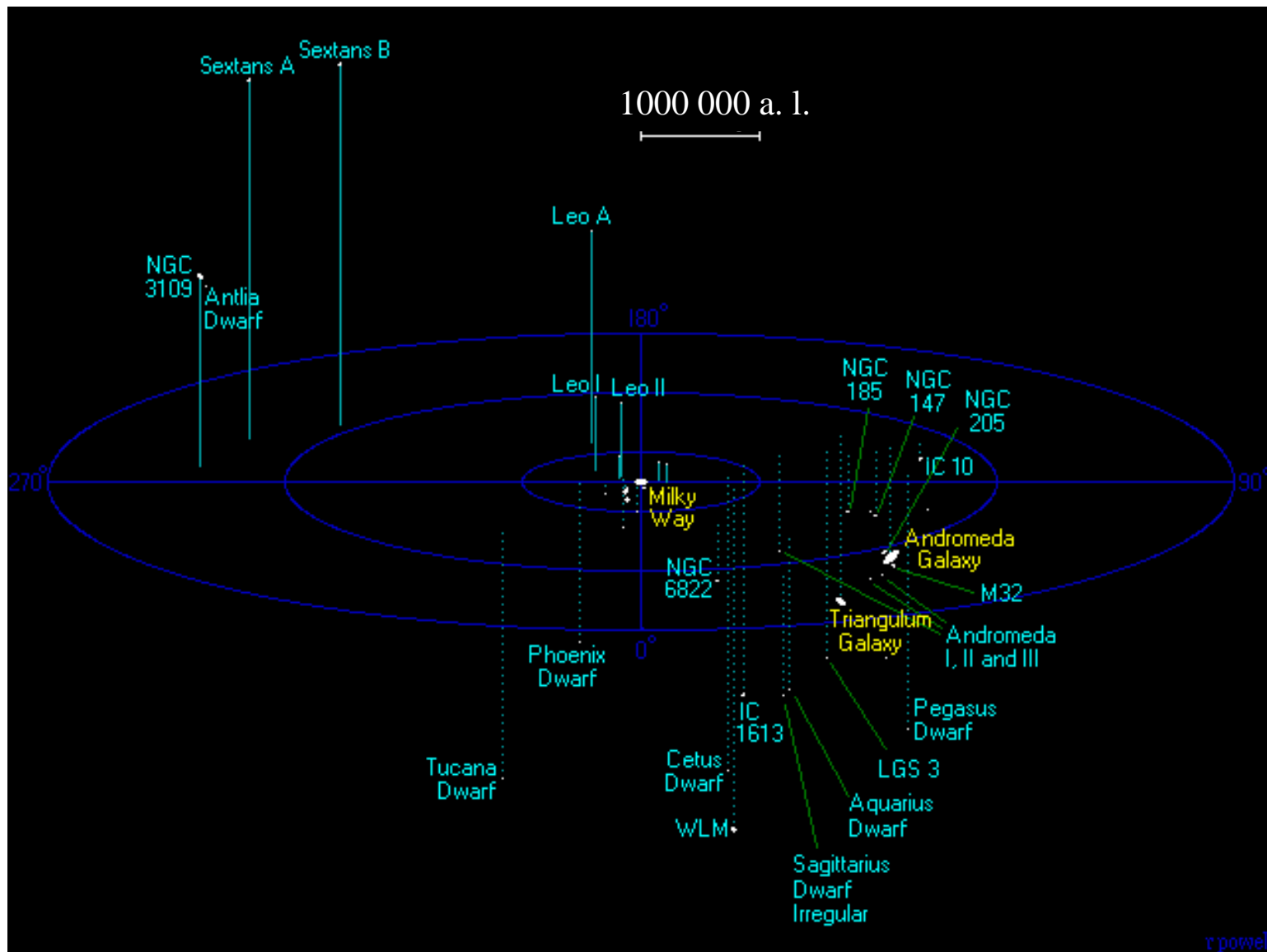
Zoom In x10

Zoom Out x10



Zoom In x10

Zoom Out x10

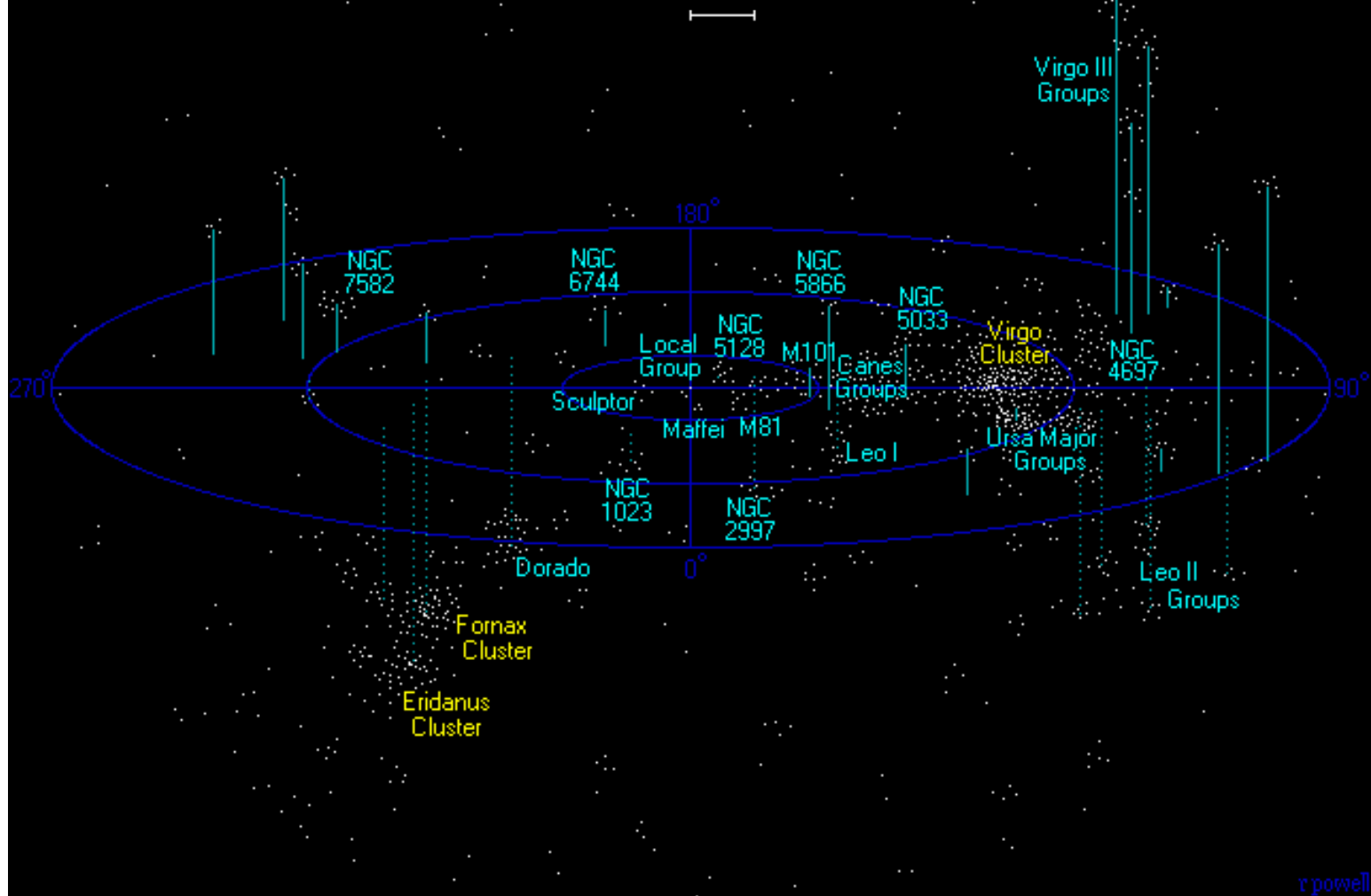


r powell

Zoom In x10

Zoom Out x20

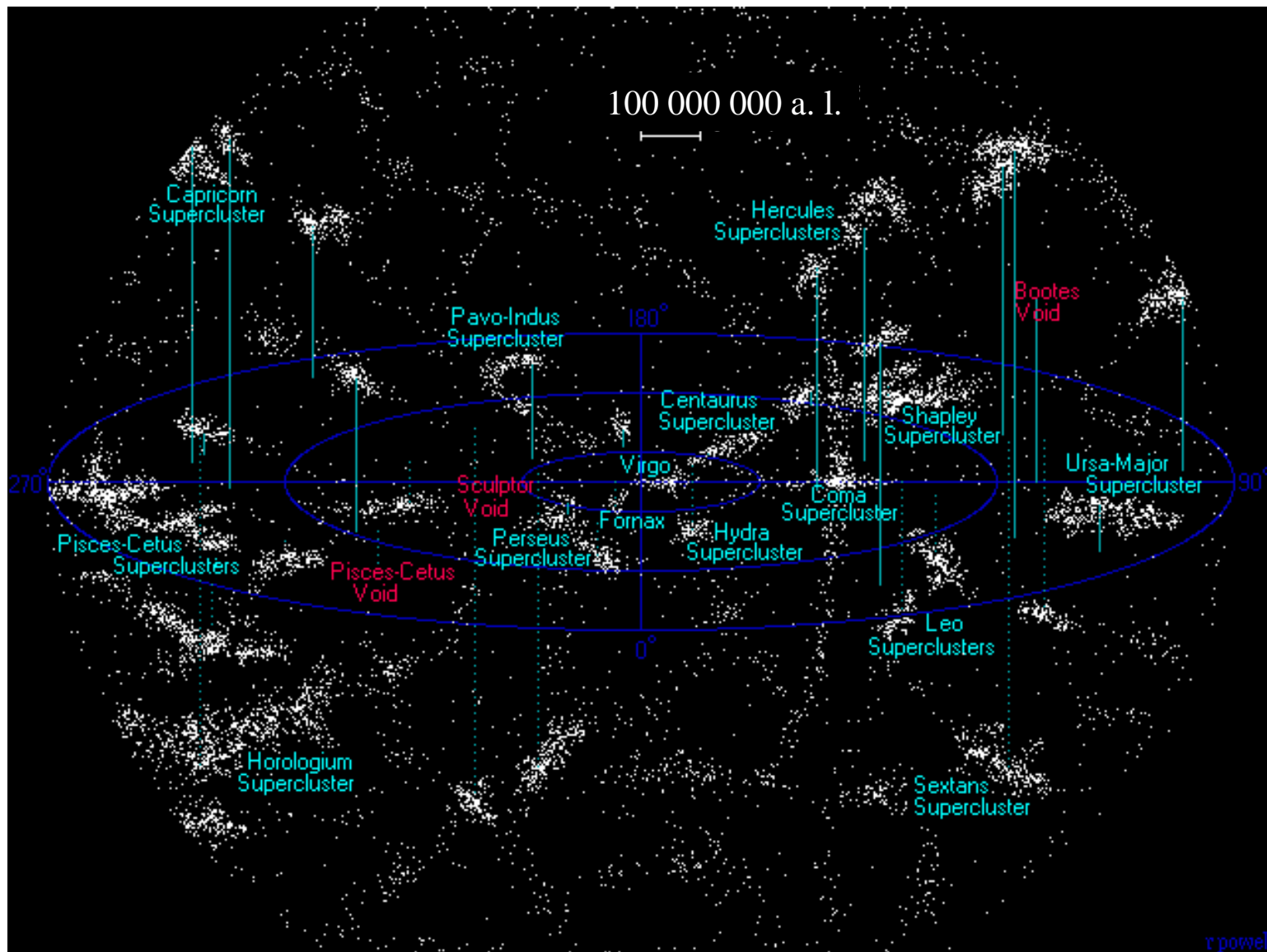
10 000 000 a. l.



r powell

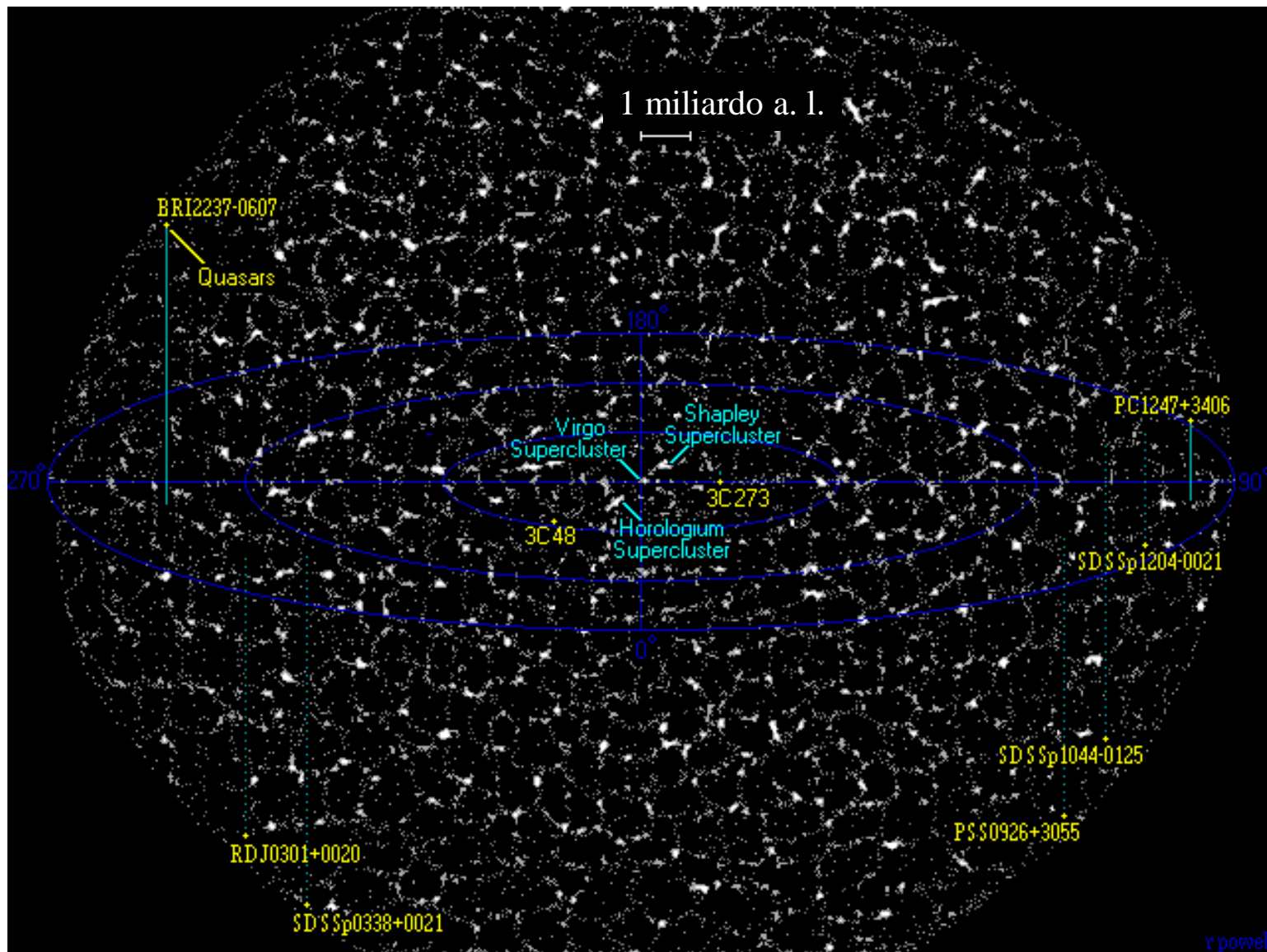
Zoom In x20

Zoom Out x10



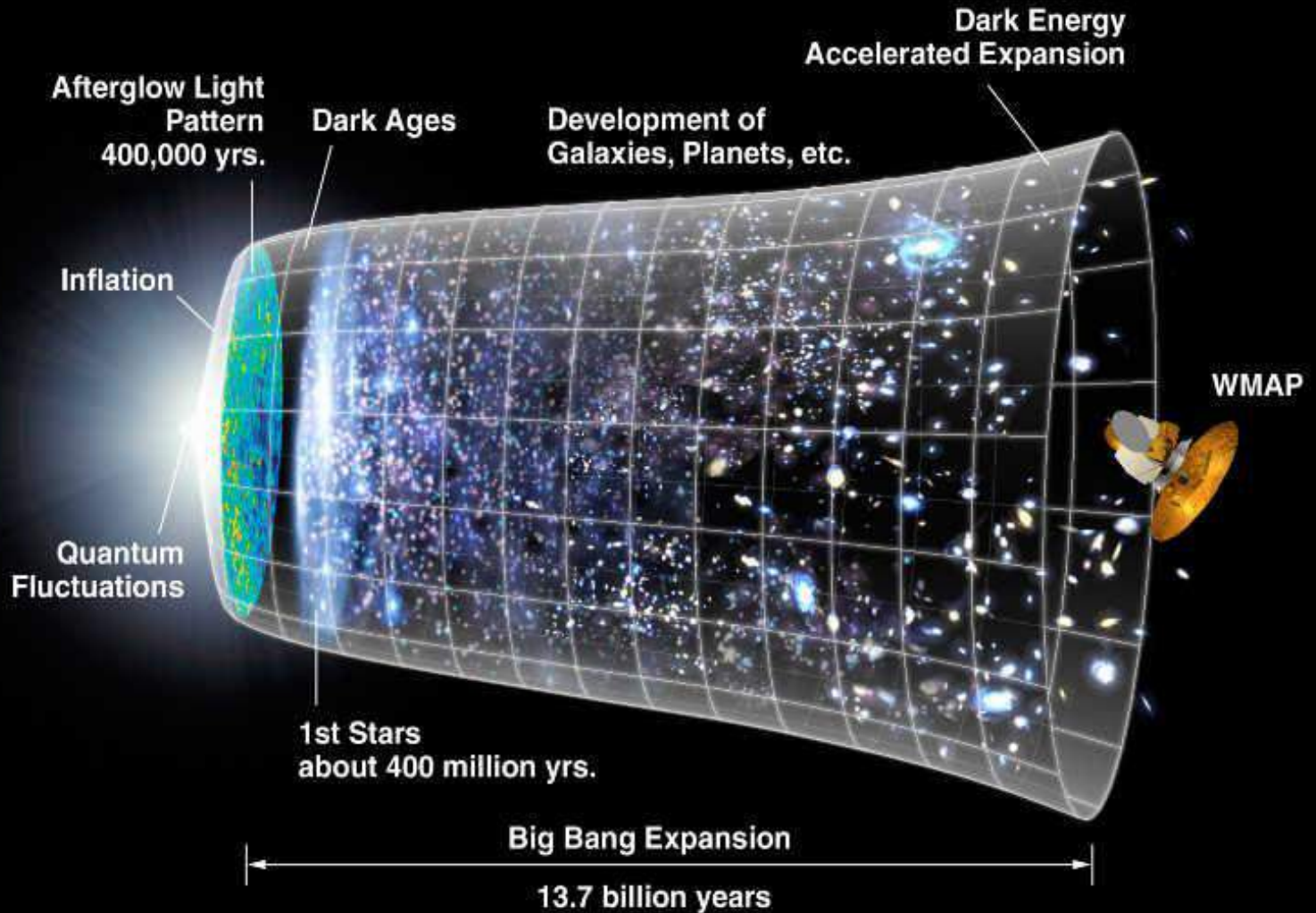
Zoom In x10

Zoom Out x15



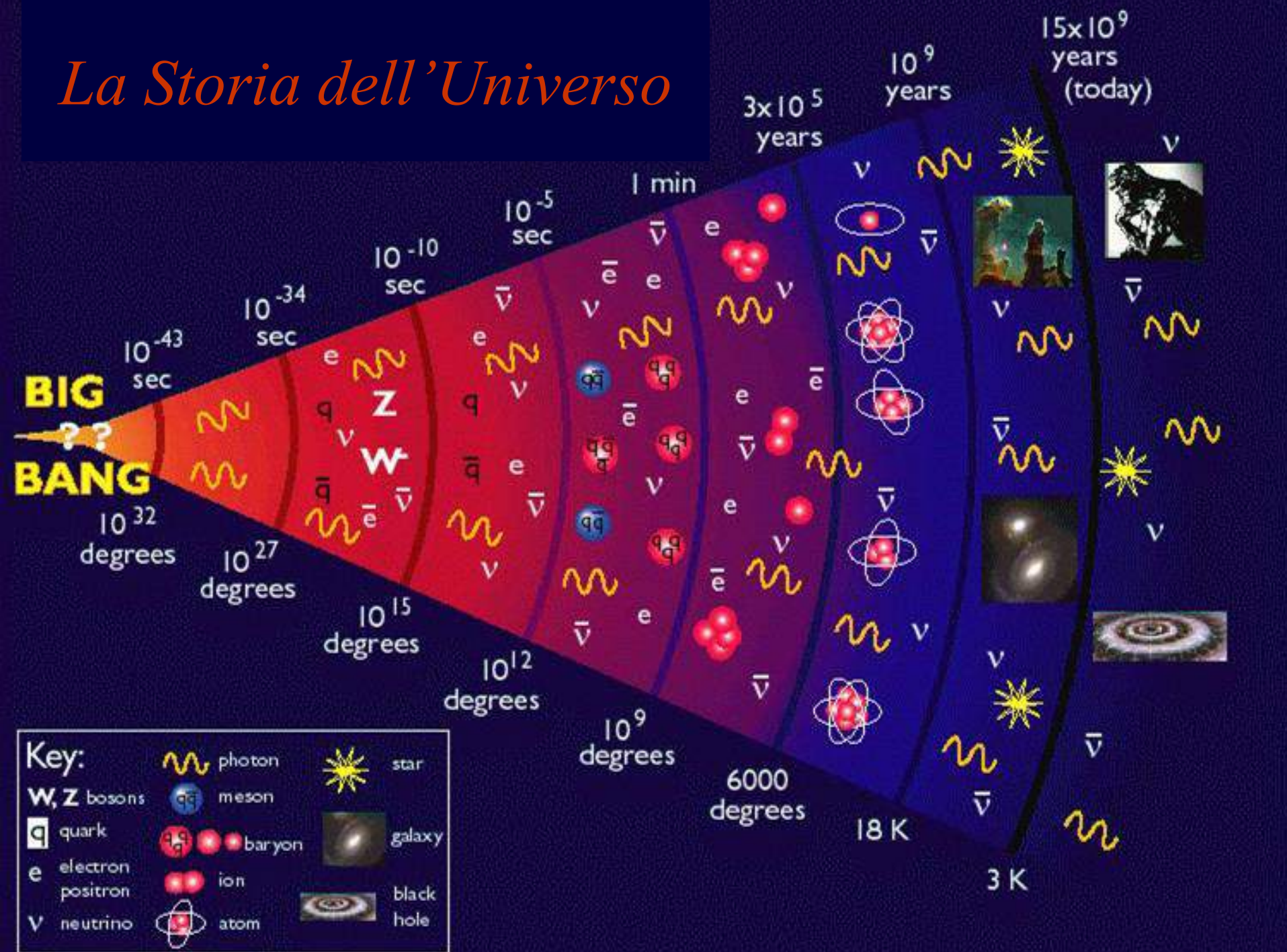
Zoom In x15

# The Big Bang Model



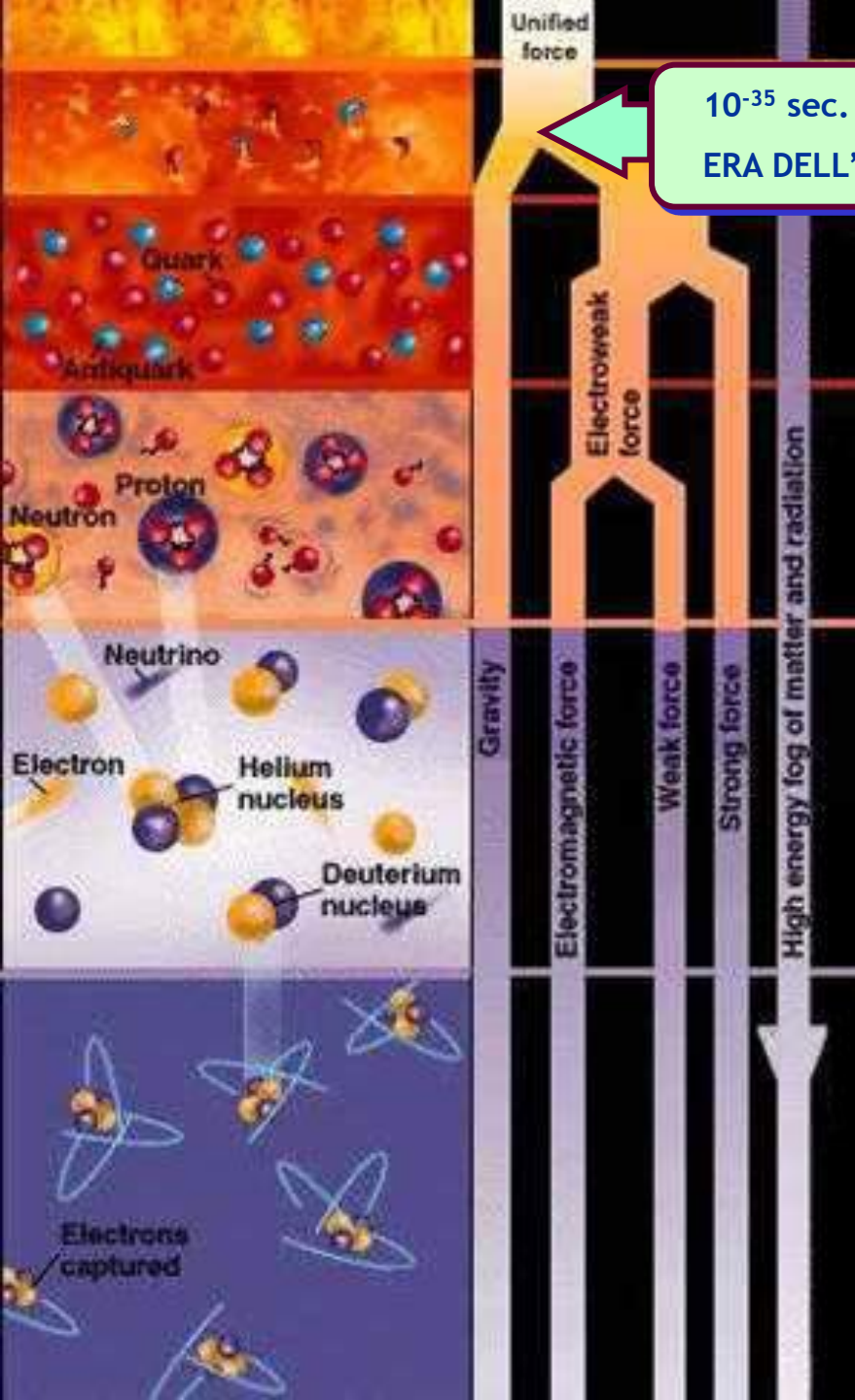


# La Storia dell'Universo



# Nascita del tempo e dello spazio

- ❖ Un evento che diede inizio alla scala del tempo e dello spazio. Di conseguenza in origine tutto doveva essere concentrato in un minuscolo “punto”, dalla **densità e gravità infinite**, dove il **tempo e lo spazio** erano pari a zero e la **temperatura** dell'ordine di miliardi di miliardi di gradi.
- ❖ Cosa ci fosse prima rimane per ora un mistero



$10^{-35}$  sec.

ERA DELL'INFLAZIONE

# Il Big Bang

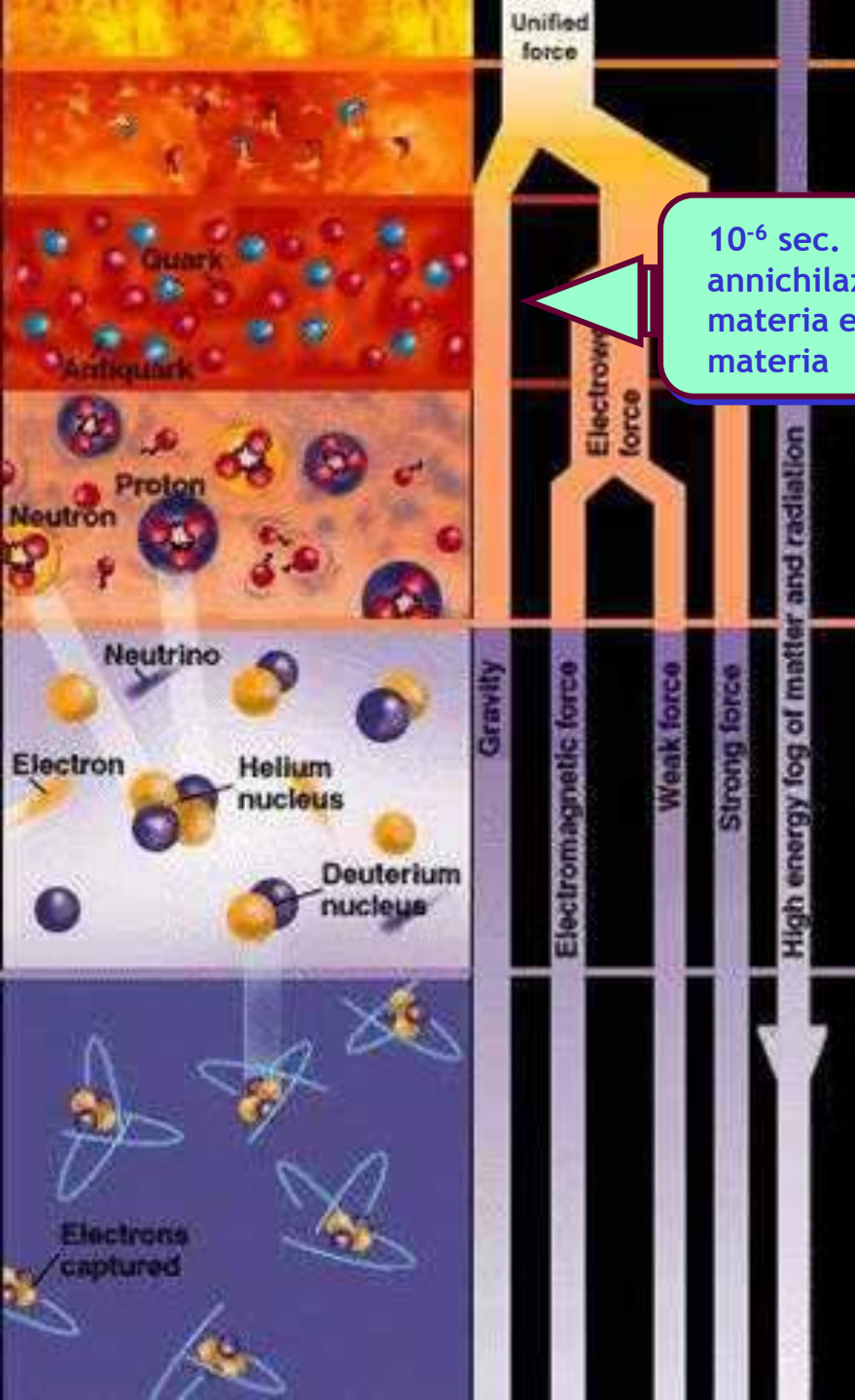
- ❖ Il Big Bang non è stato un'esplosione nello spazio, ma l'espansione dello spazio.
- ❖ Durante l'era dell'inflazione ( $10^{-35}$  sec.) l'universo si è espanso in modo esponenziale (con velocità molto più grande di quella della luce....)
- ❖ ? Commento: BICEP2.

# Il Big Bang

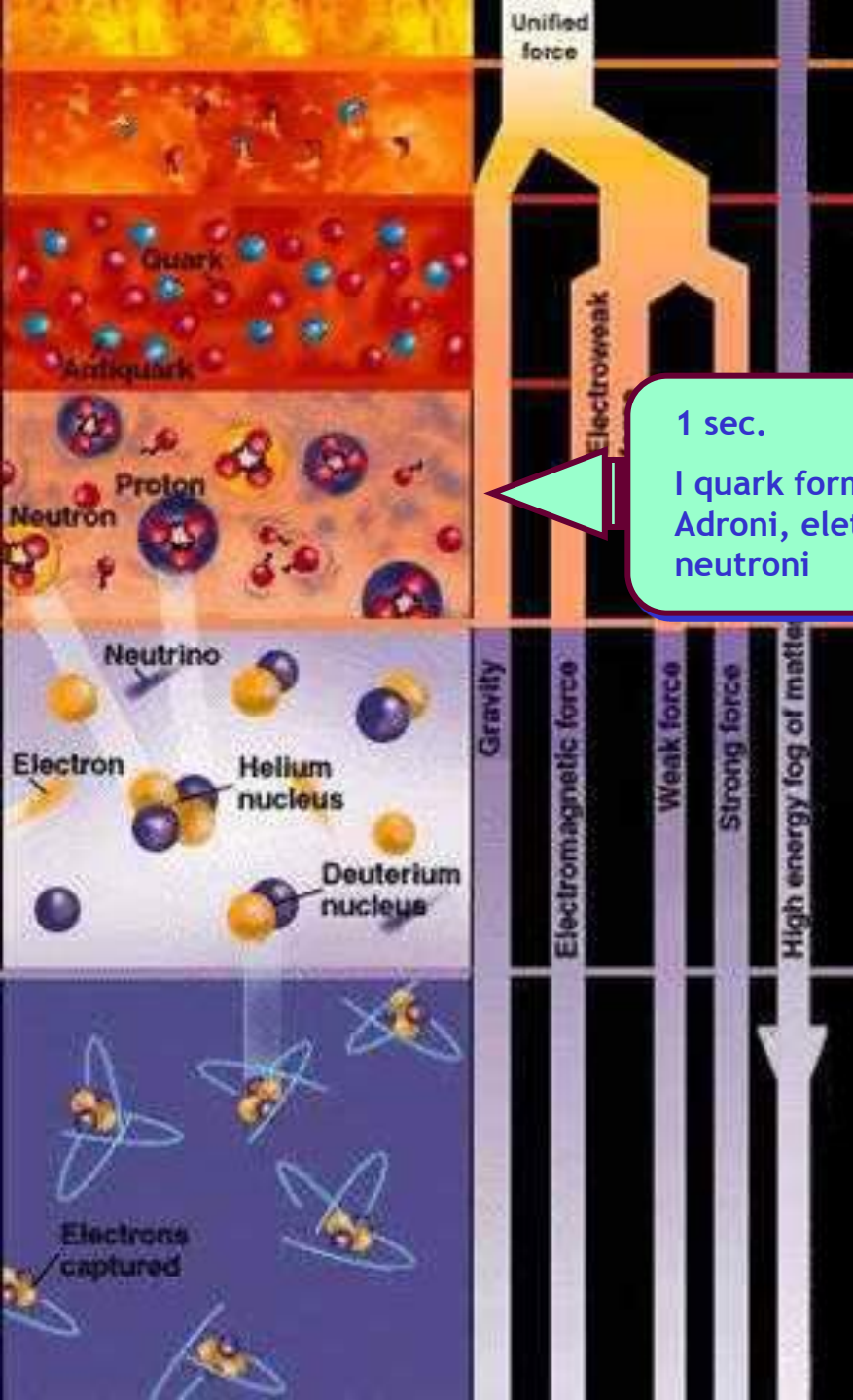
10<sup>-6</sup> sec.  
annichilazione di  
materia e anti-  
materia

## ERA DEI QUARK

- ❖ In questa fase si sono formate dall'energia moltissime copie di quark e antiquark, che si annichilivano ridiventando energia.
- ❖ Cosa succede all'antimateria?



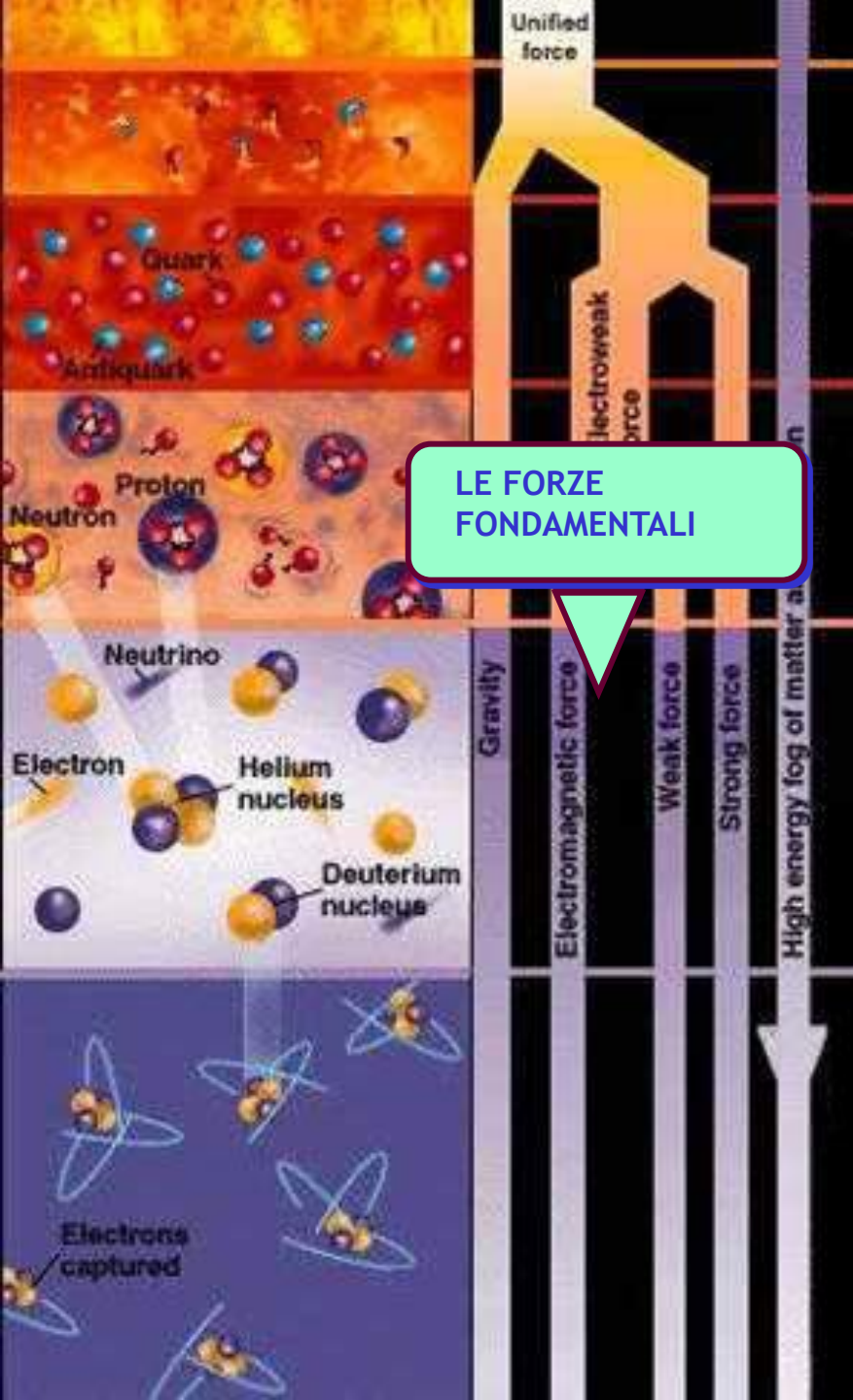
# Il Big Bang



## I PRIMI PROTONI E NEUTRONI

Dopo  $1 \mu\text{s}$  l'universo era abbastanza freddo perché i quark potessero combinarsi e formare particelle più massicce: protoni e neutroni

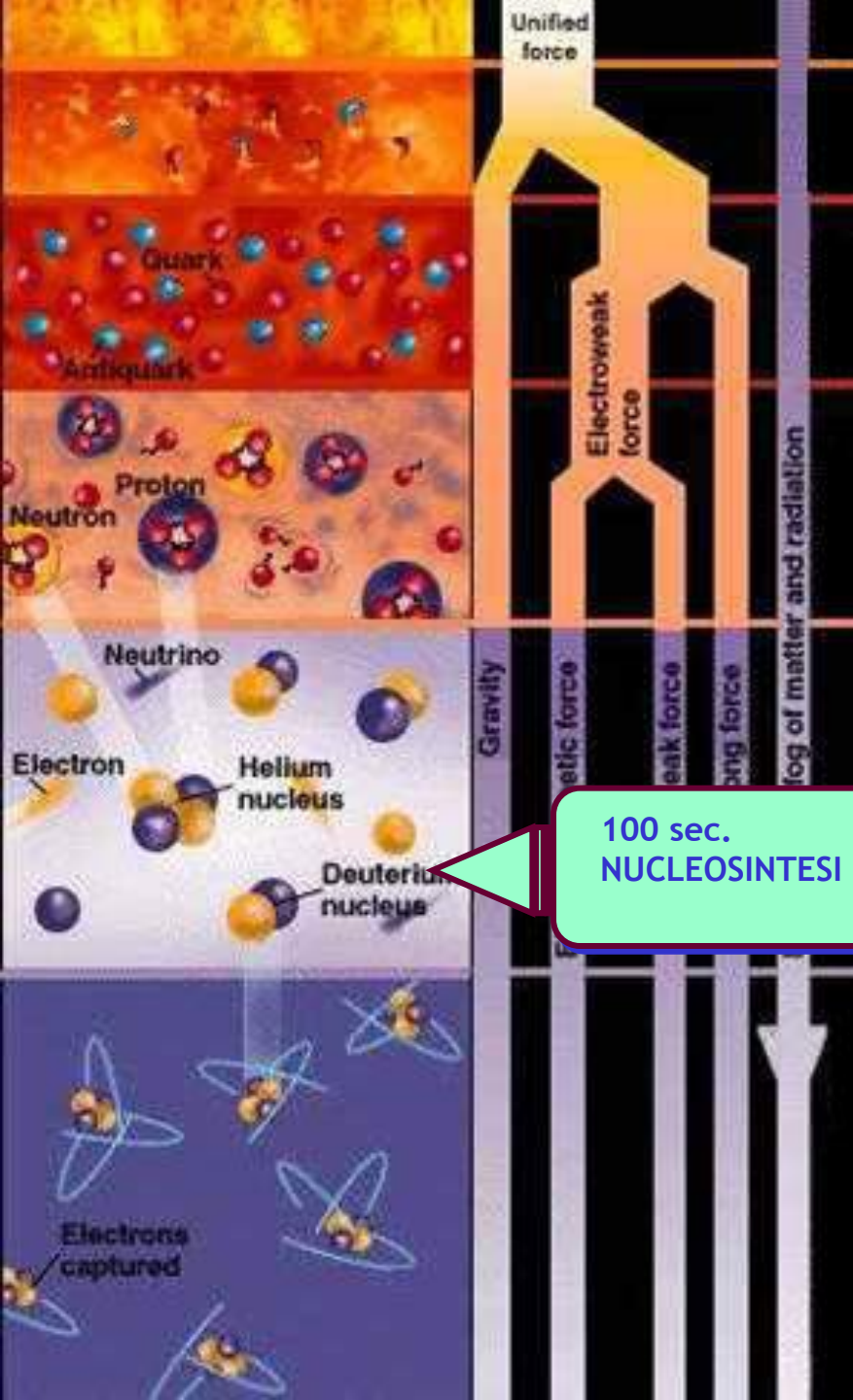
# Il Big Bang



- Subito dopo entrarono in gioco le forze fondamentali dell'universo, ed alla già esistente forza gravitazionale, che regola l'attrazione fra le masse, si aggiunsero le altre tre che insieme a questa governano l'universo:

- **la forza debole**, che agisce a livello atomico,
- **la forza forte** che governa i nuclei atomici,
- **la forza elettromagnetica** responsabile di tutti i fenomeni elettromagnetici quali la luce, le onde radio, ecc...

# Il Big Bang

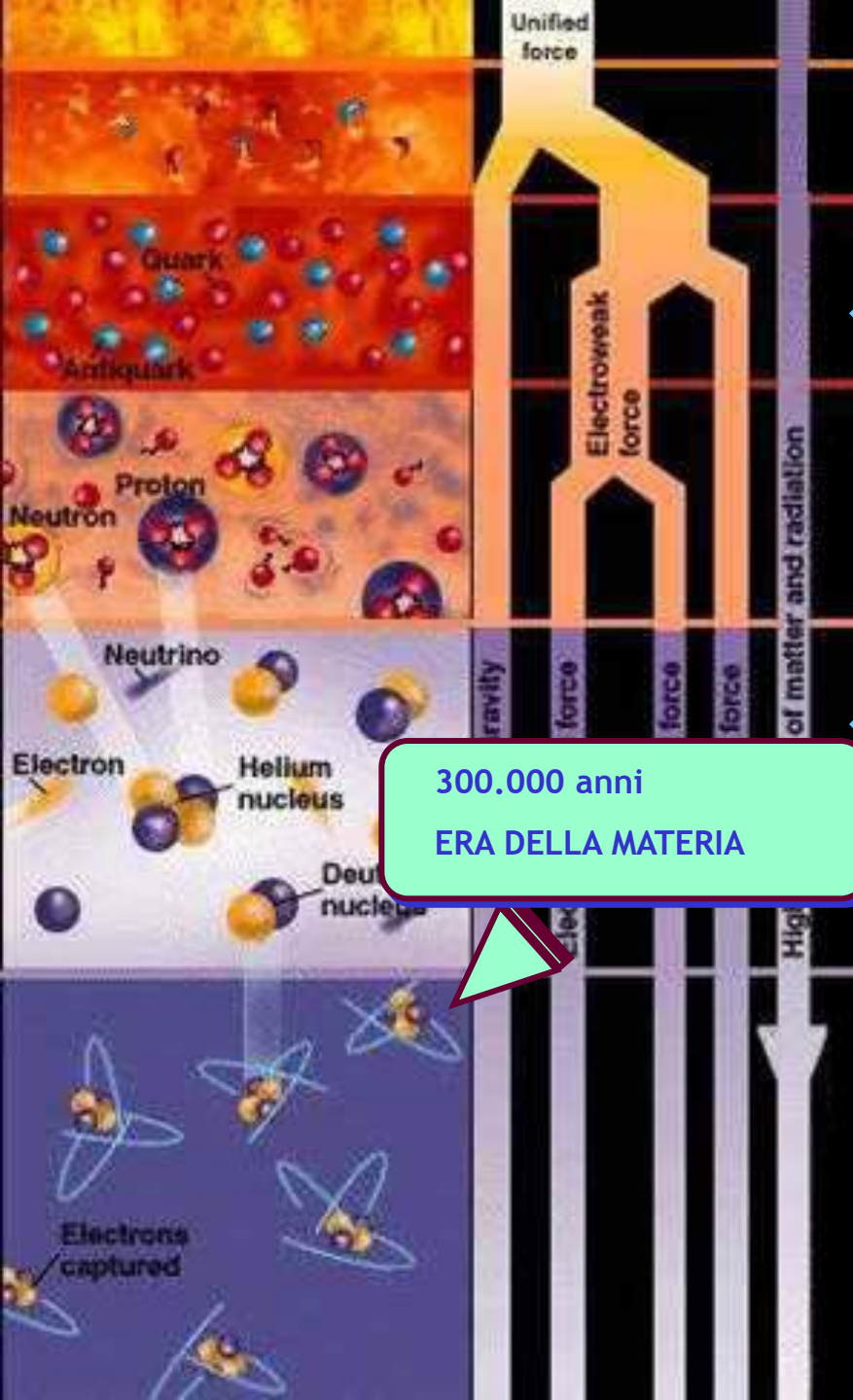


- ERA DELLA NUCLEOSINTESI
- In questa fase i neutroni si sono gradualmente trasformati in protoni. Mentre la temperatura dell'universo si abbassava, gradualmente i neutroni si combinavano con i protoni per formare i primi nuclei di elio ognuno formato da 2 protoni e 2 neutroni.

# Il Big Bang

❖ Dopo alcune centinaia di migliaia di anni, l'universo era divenuto ancora meno denso e più freddo, avveniva la **"ricombinazione"**, i protoni e gli elettroni si combinavano per creare i primi atomi di idrogeno.

❖ Finiva a quel punto la prima parte della storia dell'universo, quella dominata dalla radiazione, ed iniziava **"l'era della materia"** che vedeva l'aggregazione delle particelle nelle prime forme atomiche, mentre i fotoni, liberi ormai da ogni vincolo, potevano così irradiarsi in tutte le direzioni sotto forma di radiazione cosmica di fondo.





# Cronologia delle scoperte

## Tappe fondamentali

1915

Einstein

Teoria della relatività  
Universo statico

1922

Friedmann

Teoria espansione  
universo

1929

Hubble

Red Shift  
Le galassie si allontanano

1946

Gamow

Teoria del Big Bang

1965

Penzias e Wilson

Radiazione di fondo

1981

Guth- Sato

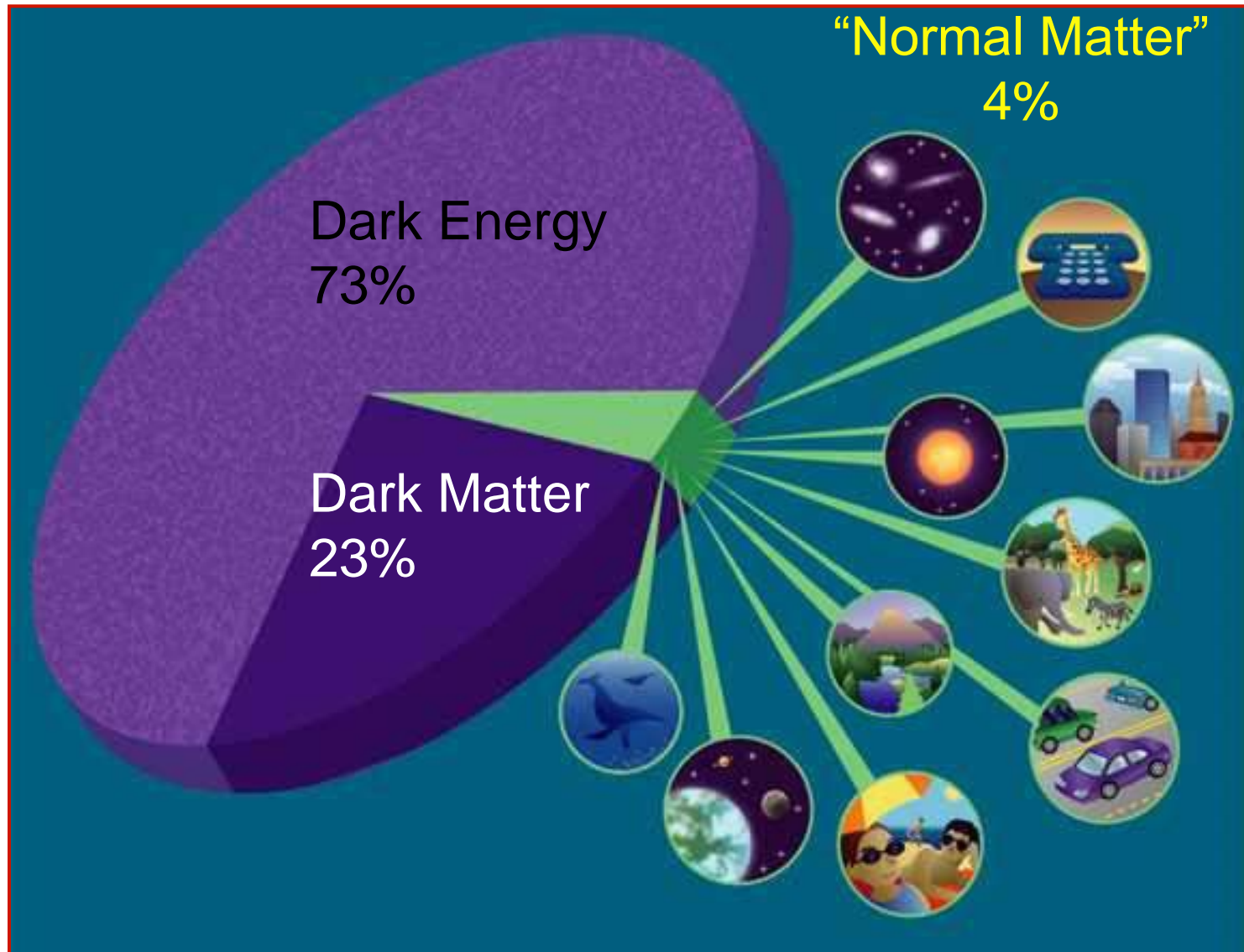
Universo inflazionario

1998

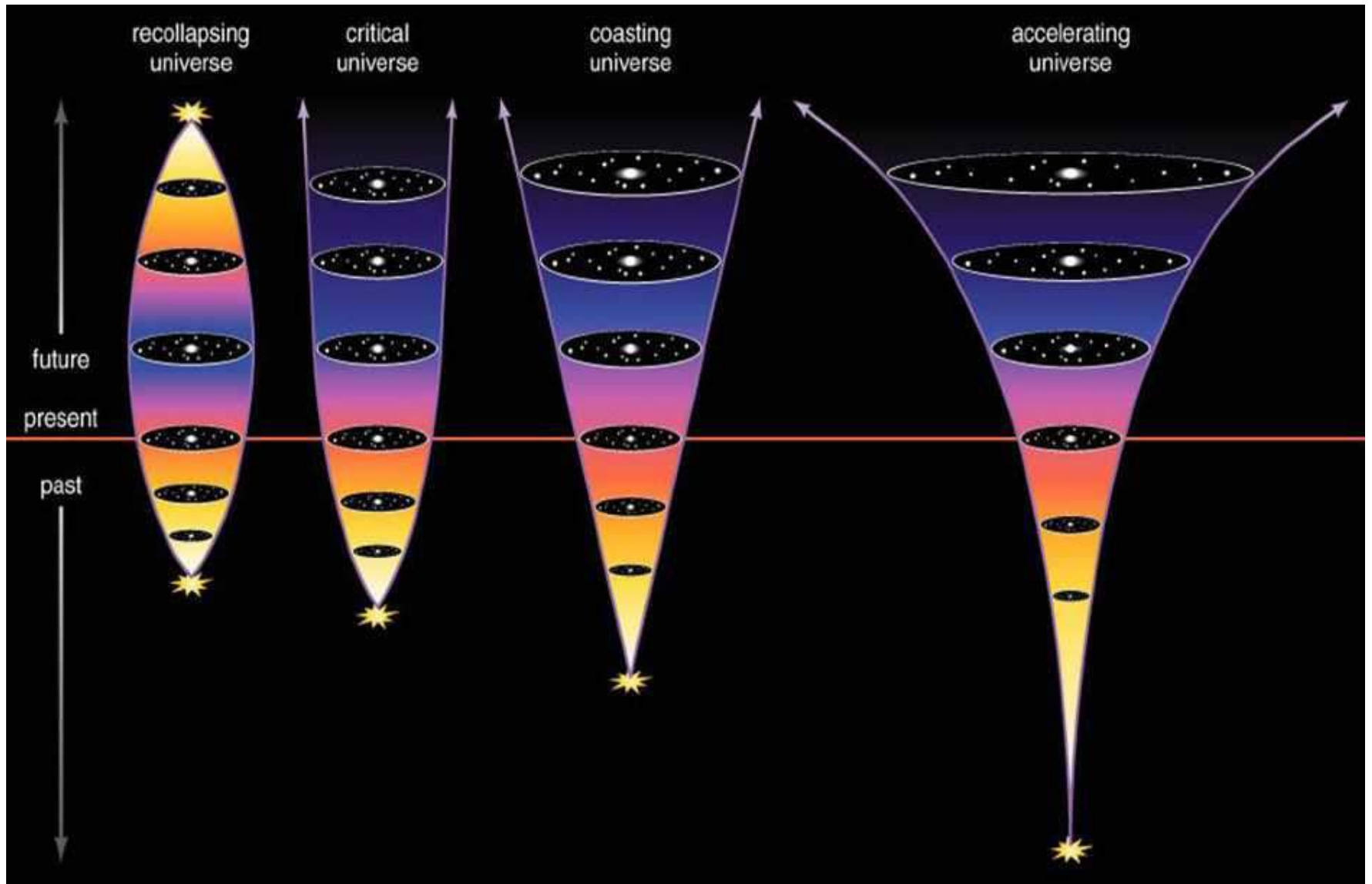
Ricerche sulle supernovae

Espansione dell'universo  
accelerata

# Materia ed energia



# I modelli Cosmologici



**Da cosa è costituita  
la materia oscura?**

**Non è fatta di materia barionica  
ma di particelle sconosciute**

**(Forse LHC?)**

Una forza che

si oppone alla gravità

Antigravità

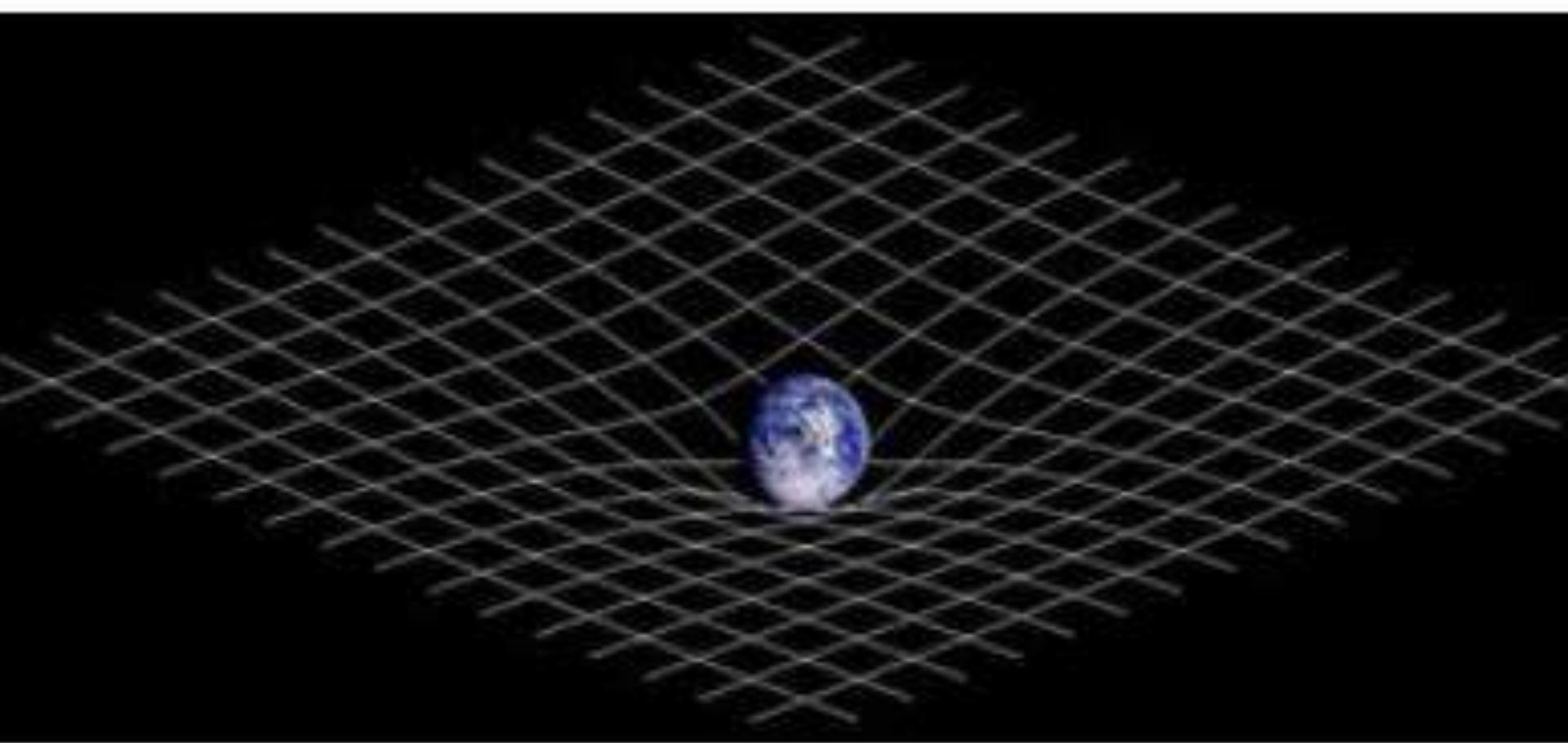
Quintessenza

Costante

cosmologica

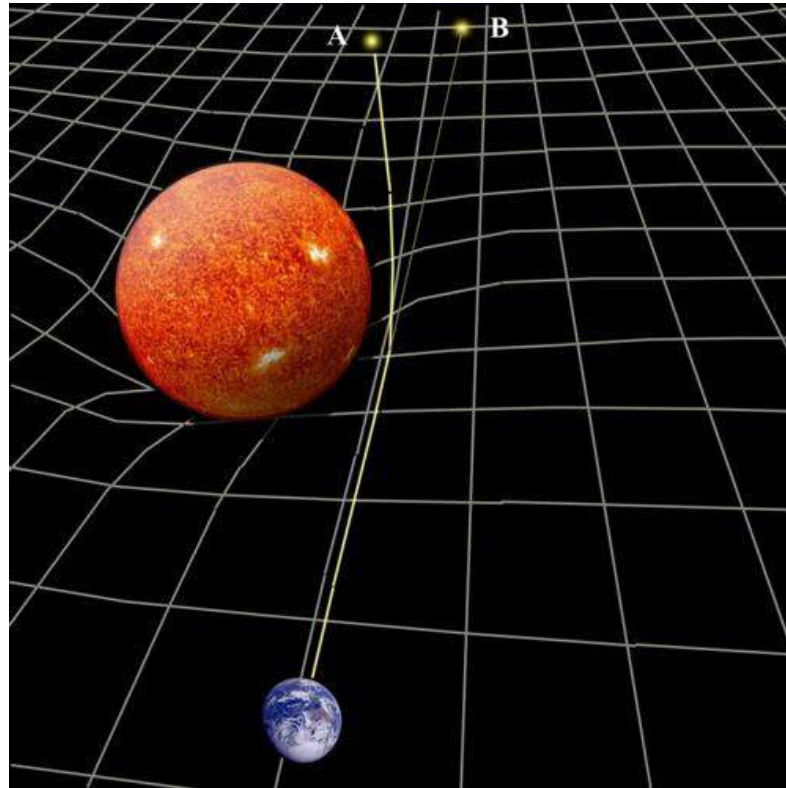
Energia oscura







# Relativita' generale

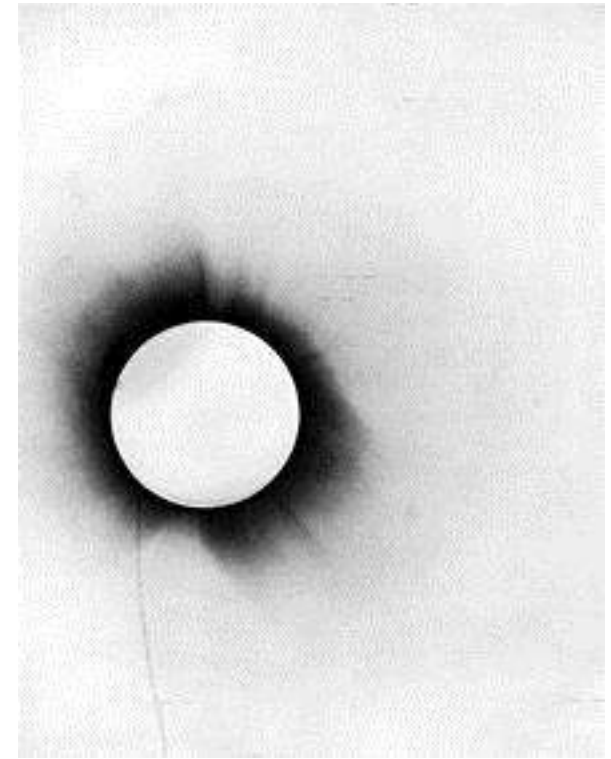
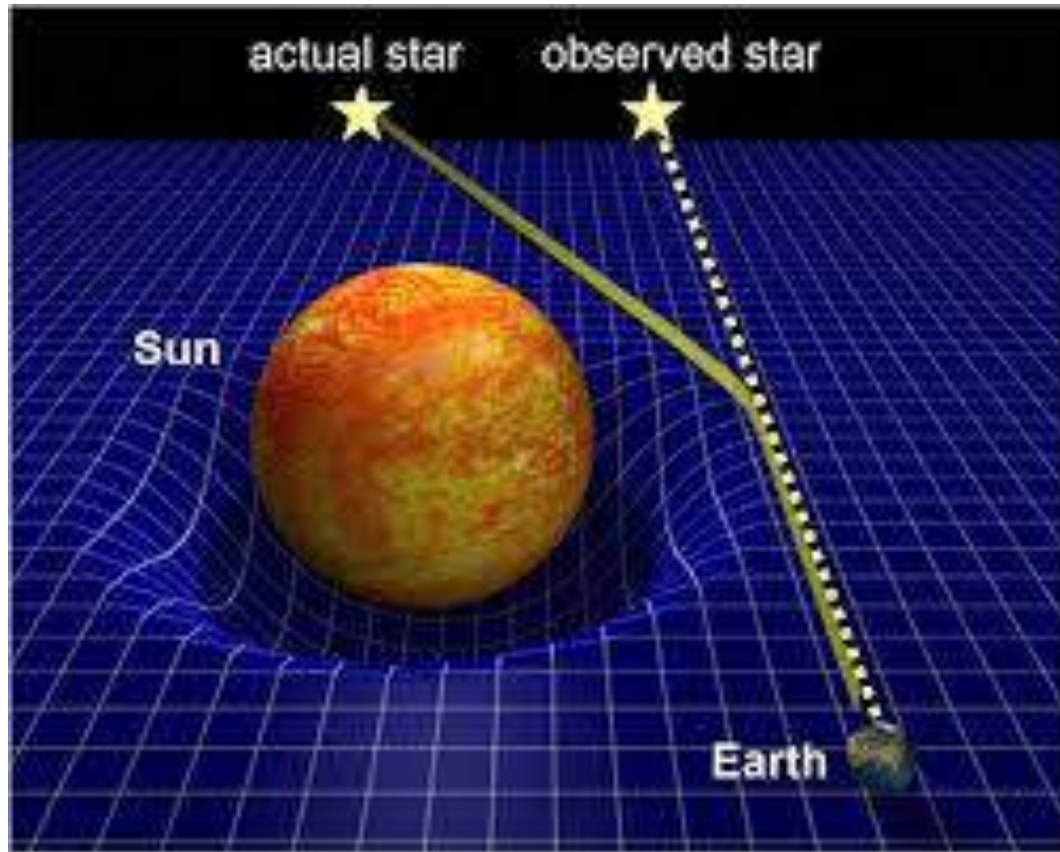


La materia dice allo spazio come curvarsi, *lo spazio dice alla materia* come muoversi.

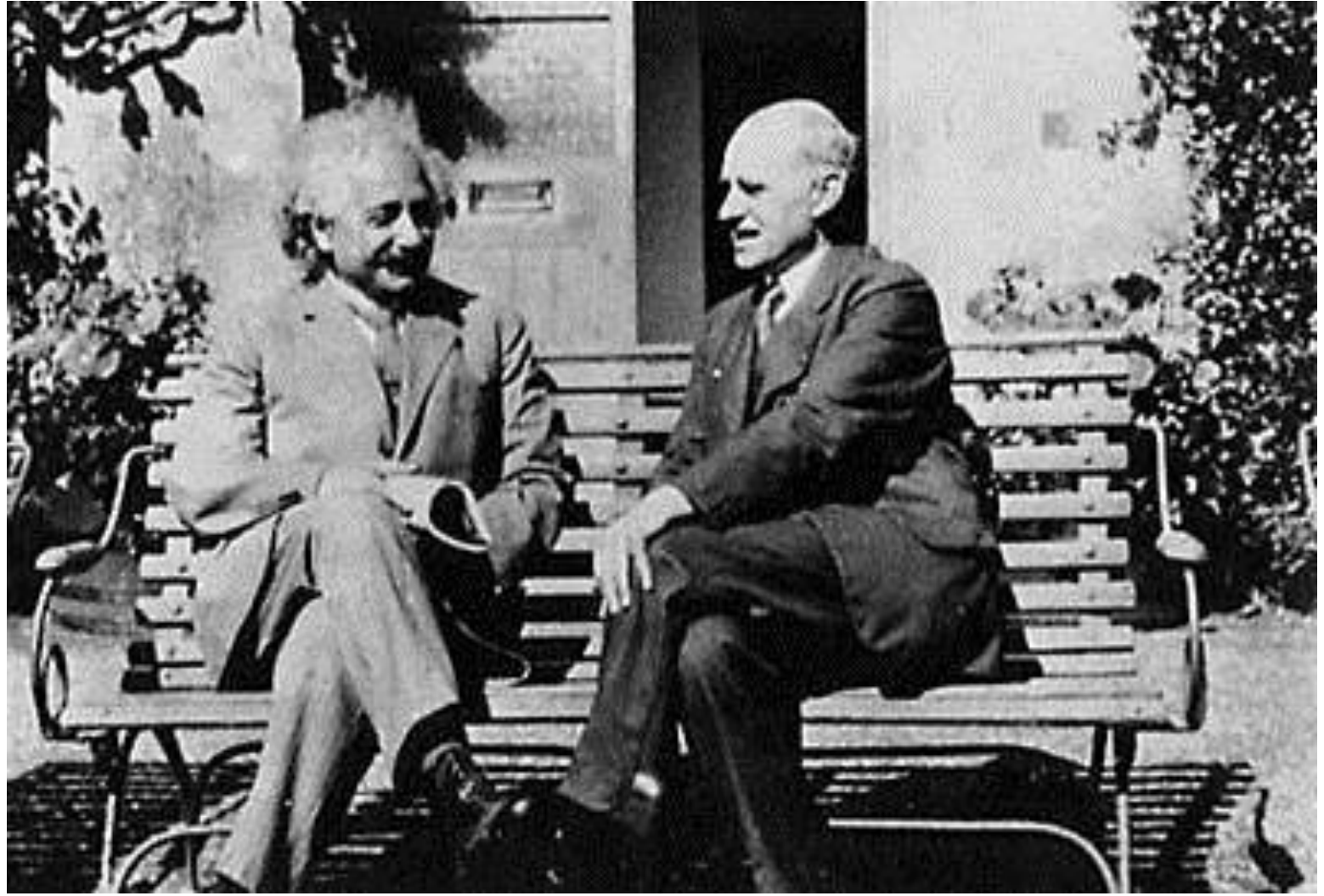
# Relativita' generale

$$R_{\mu\nu} - \frac{1}{2}R g_{\mu\nu} + \underbrace{\Lambda}_{?} g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

# Relativita' generale



**1919 – conferma relativita' generale: Sir Arthur Eddington (foto durante eclissi...)**

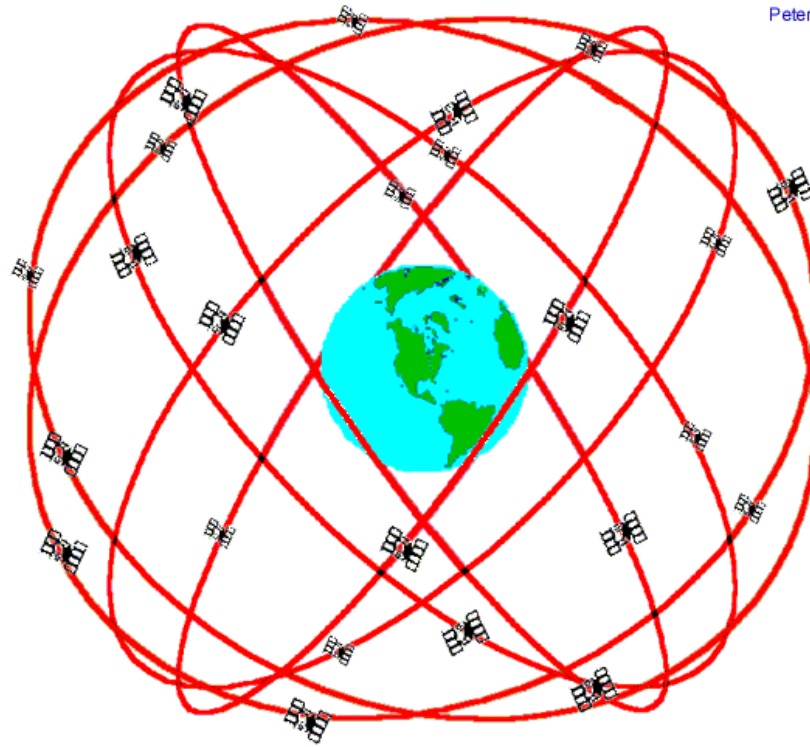


# DAΦNE Collider at LNF-INFN



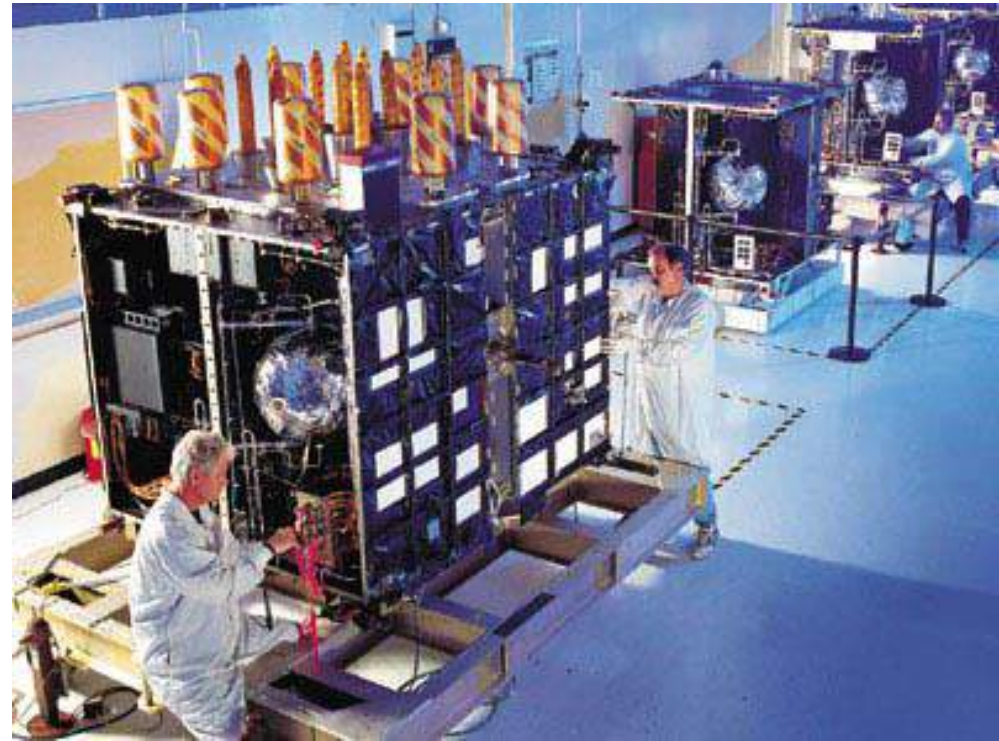
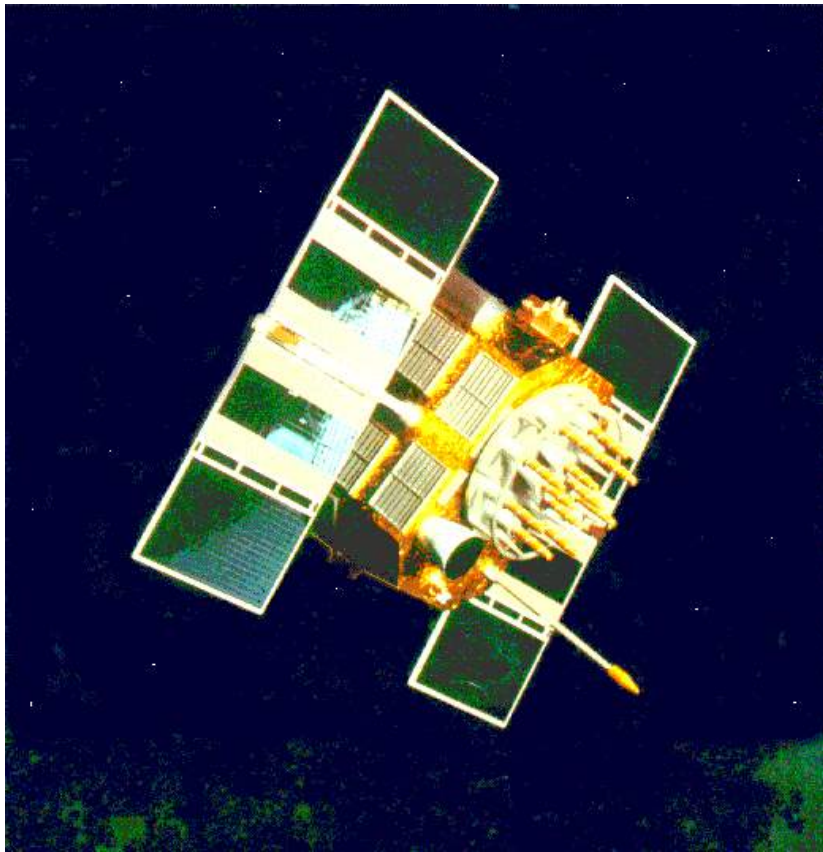
# The Global Position System (GPS)

Peter H. Dana 9/22/98

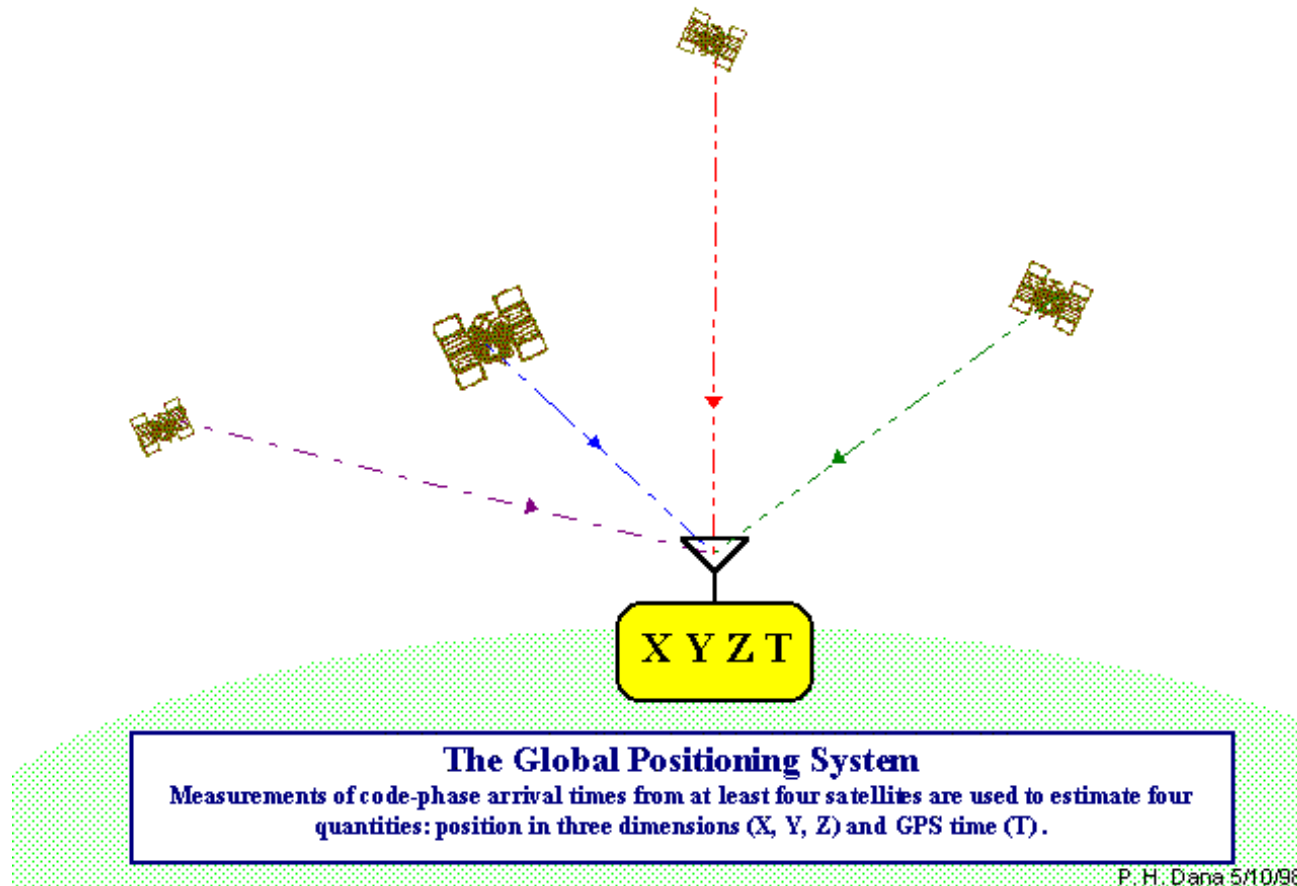


**GPS Nominal Constellation**  
**24 Satellites in 6 Orbital Planes**  
**4 Satellites in each Plane**  
**20,200 km Altitudes, 55 Degree Inclination**

# The Global Position System (GPS)

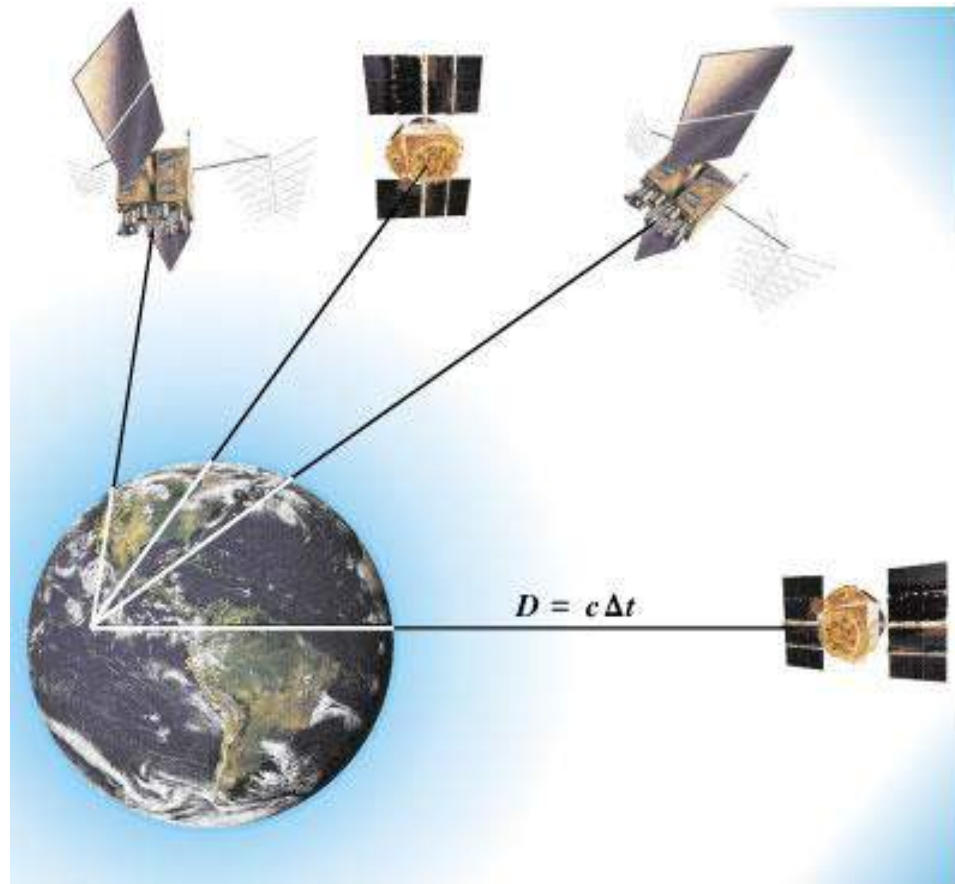


# The Global Position System (GPS)

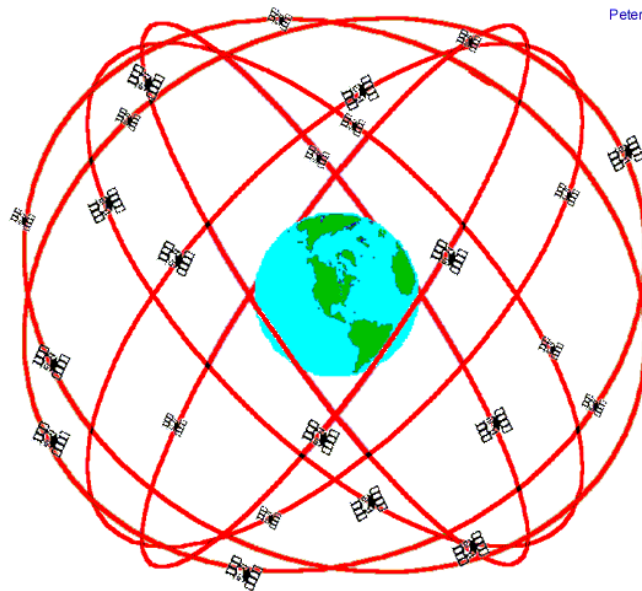




# The Global Position System (GPS)



# The Global Position System (GPS) e la relativita'!



Peter H. Dana 9/22/98

GPS Nominal Constellation  
24 Satellites in 6 Orbital Planes  
4 Satellites in each Plane  
20,200 km Altitudes, 55 Degree Inclination

$$v = 3.87 \text{ Km/s}$$

$$t_{01} - t_{01}' = (1 - \gamma) t_{01}$$

1 orbit ~ 12 hours

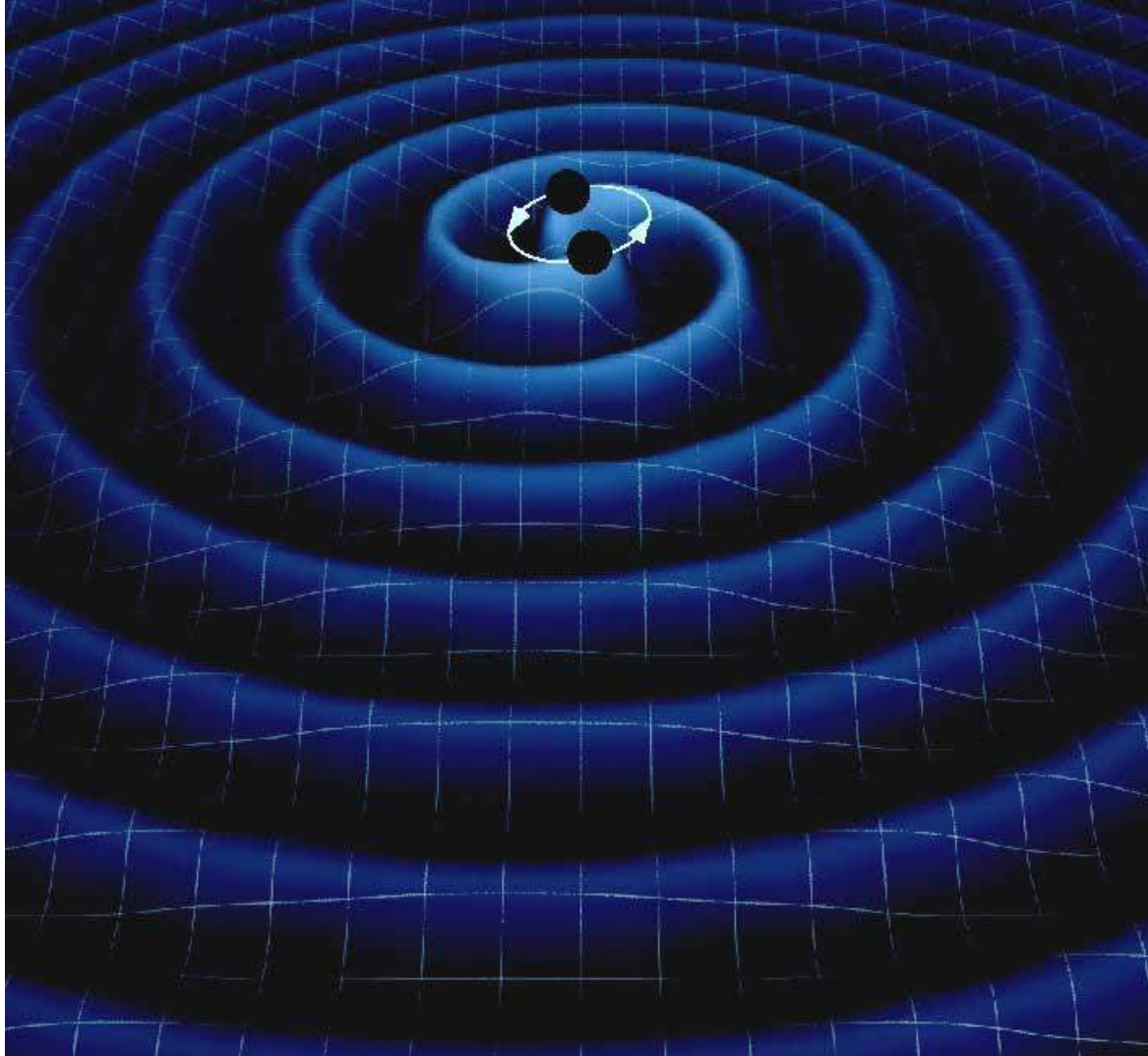
$$\rightarrow t_{01} - t_{01}' \sim 6 \cdot 10^{-8}$$

s

$$\rightarrow \Delta D \sim 18 \text{ m}$$

Due to other effects connected to general relativity  $\Delta D \sim 100 \text{ m}$

# Onde gravitazionali





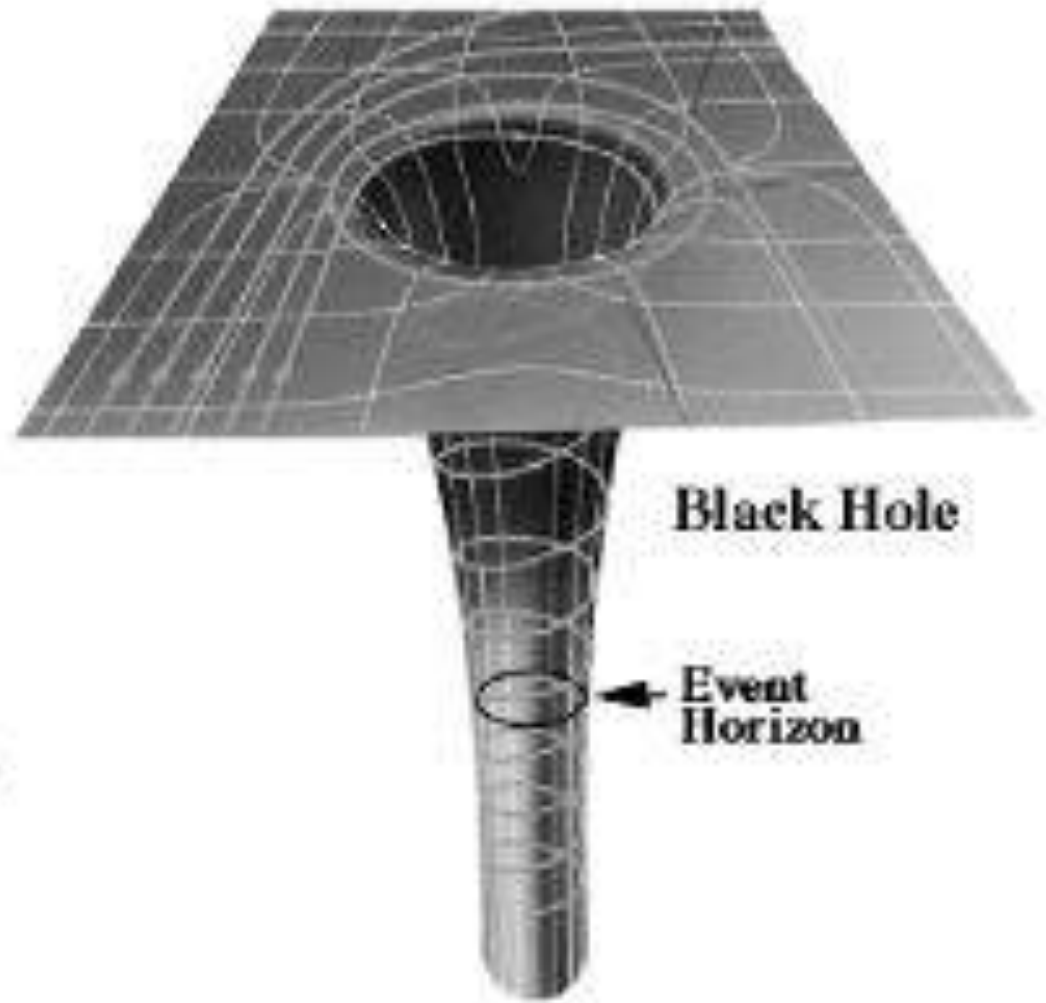
**Sun**



**White Dwarf**



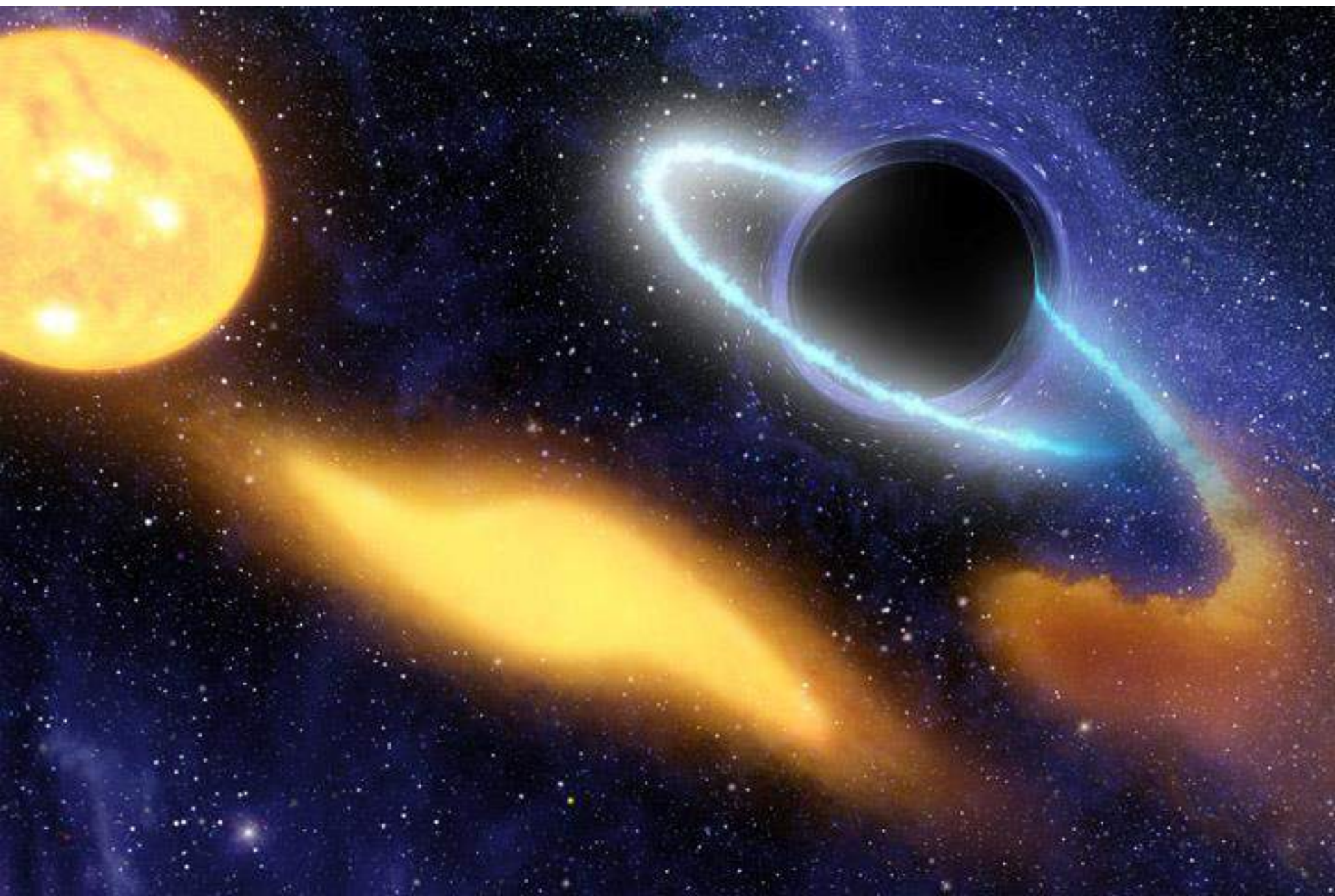
**Neutron Star**



**Black Hole**

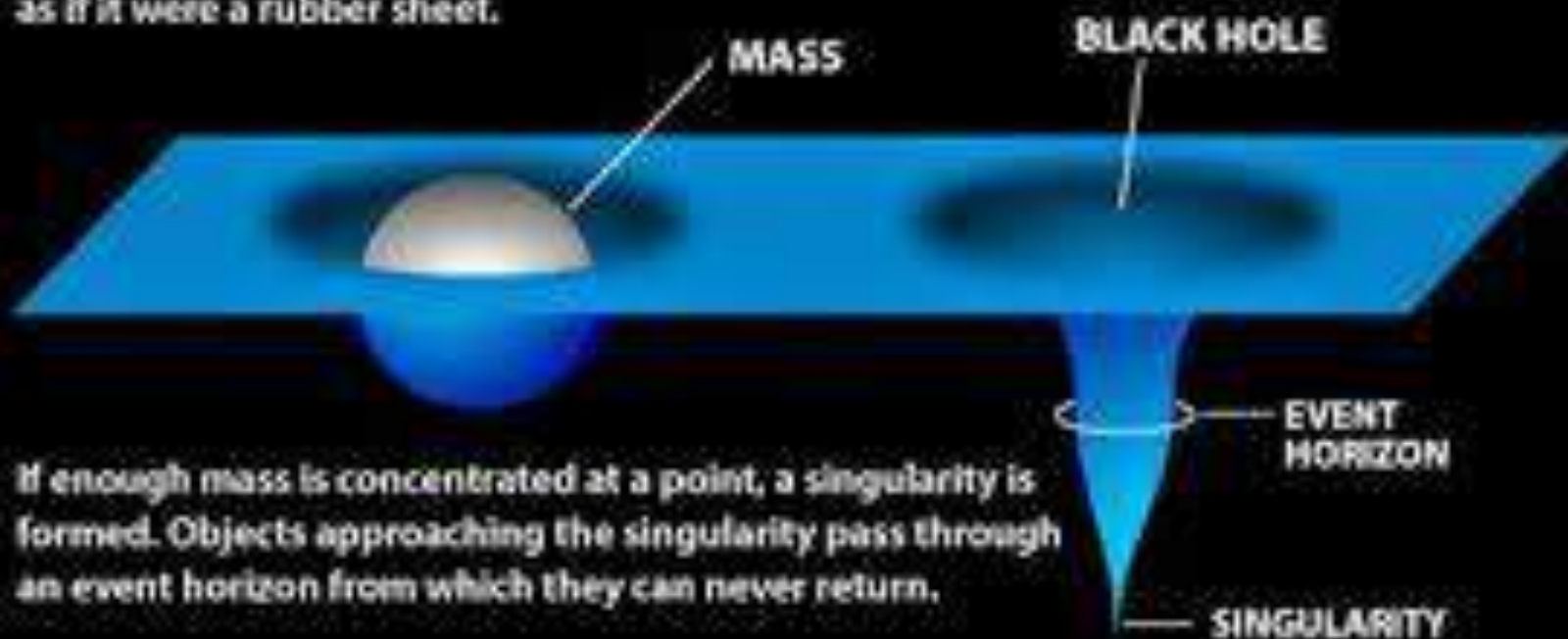
**Event Horizon**

Credit: Adam Apollo

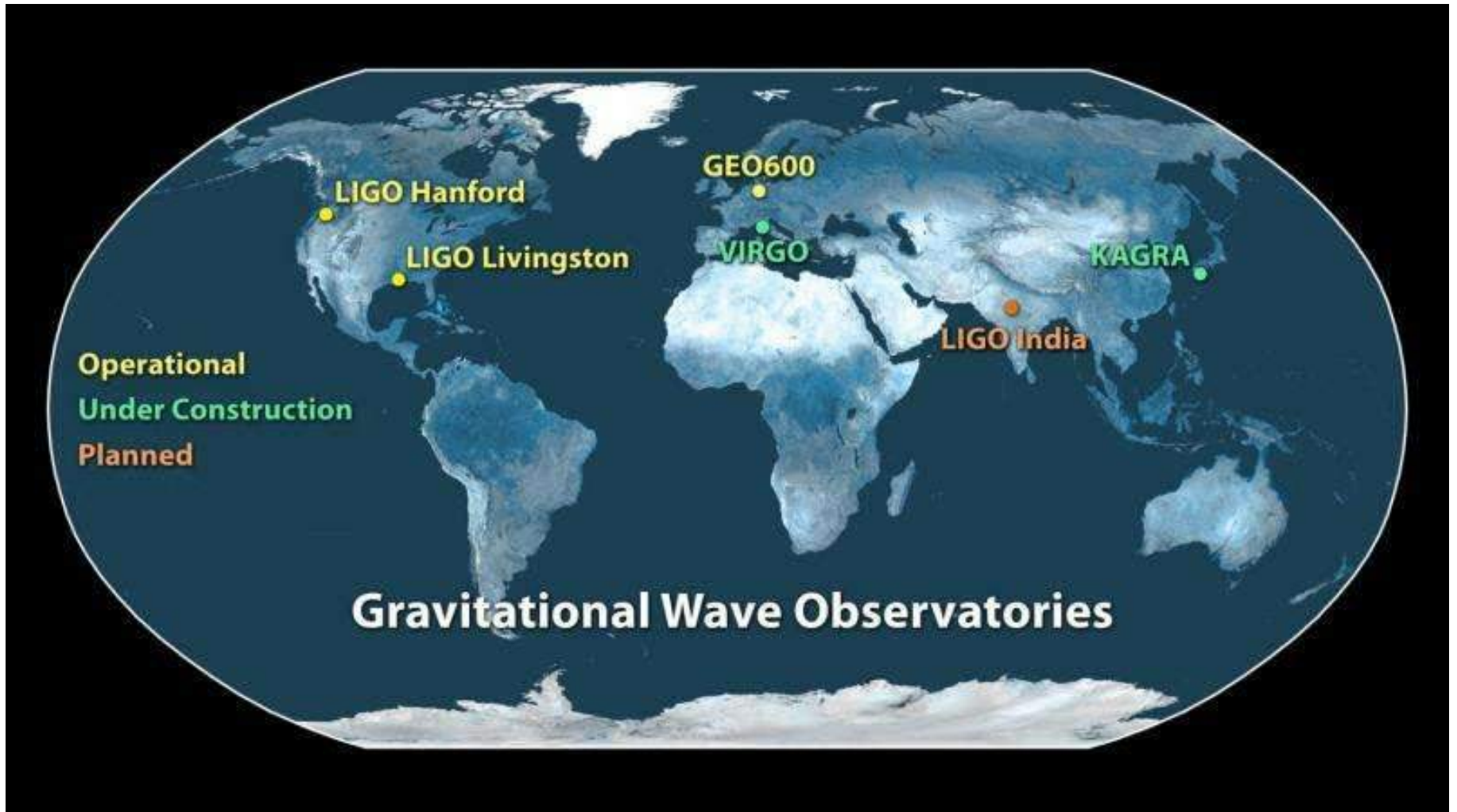


# HOW TO MAKE A BLACK HOLE

The presence of mass distorts the local space-time as if it were a rubber sheet.

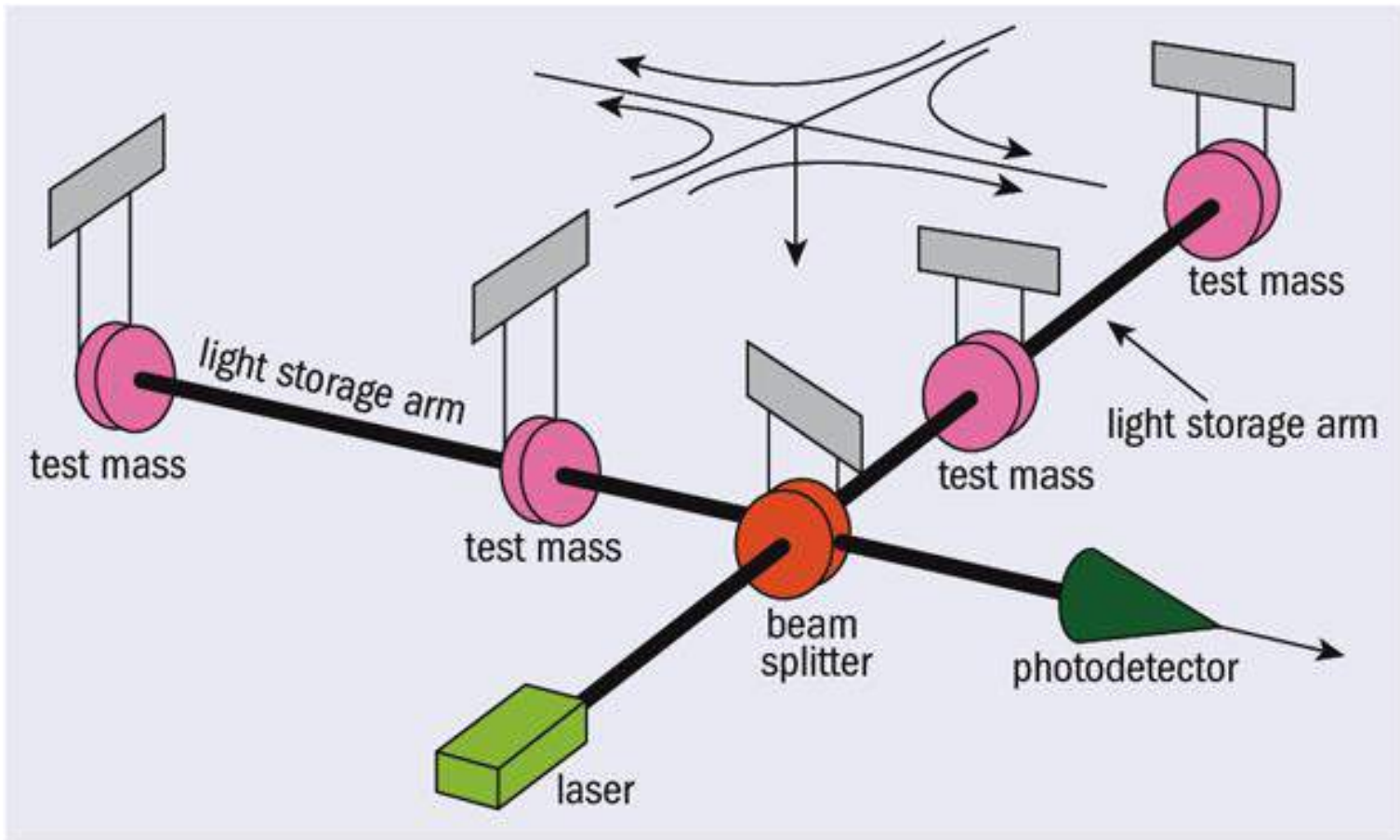


If enough mass is concentrated at a point, a singularity is formed. Objects approaching the singularity pass through an event horizon from which they can never return.





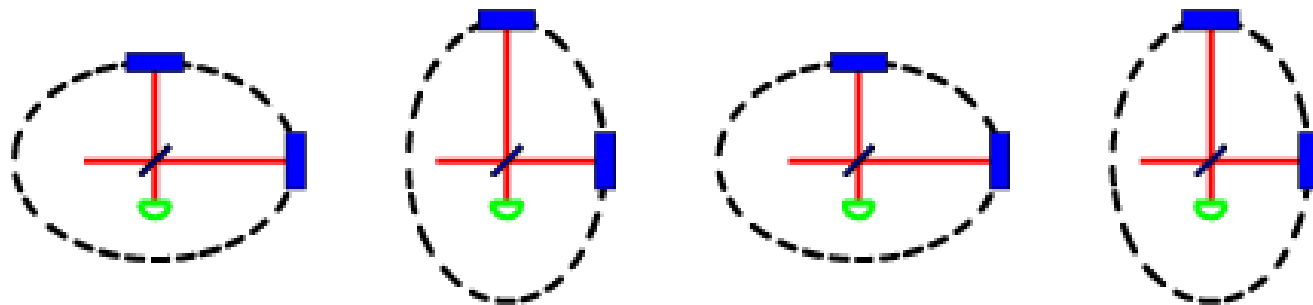
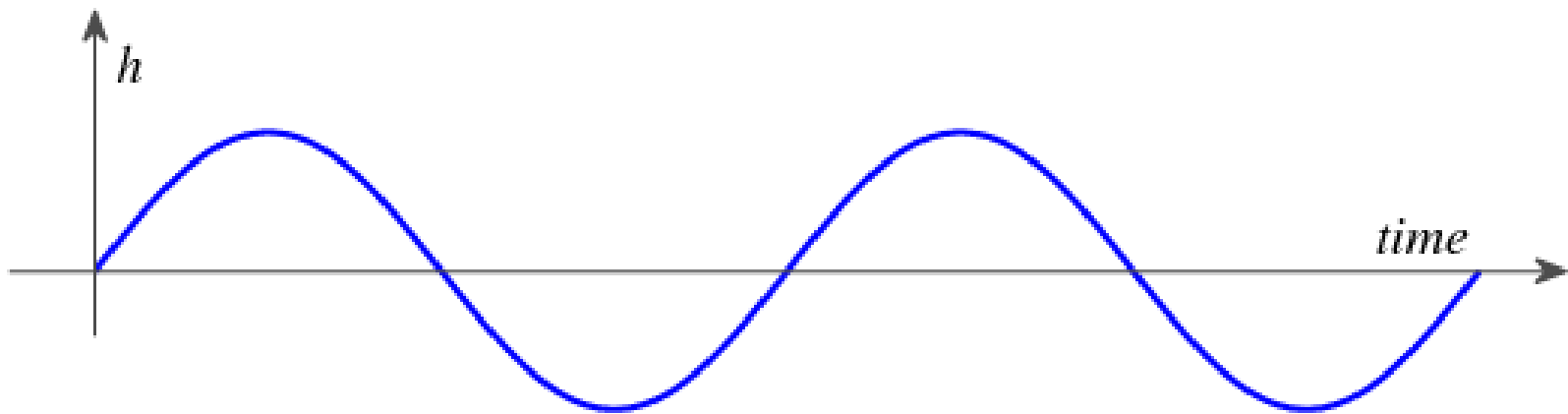




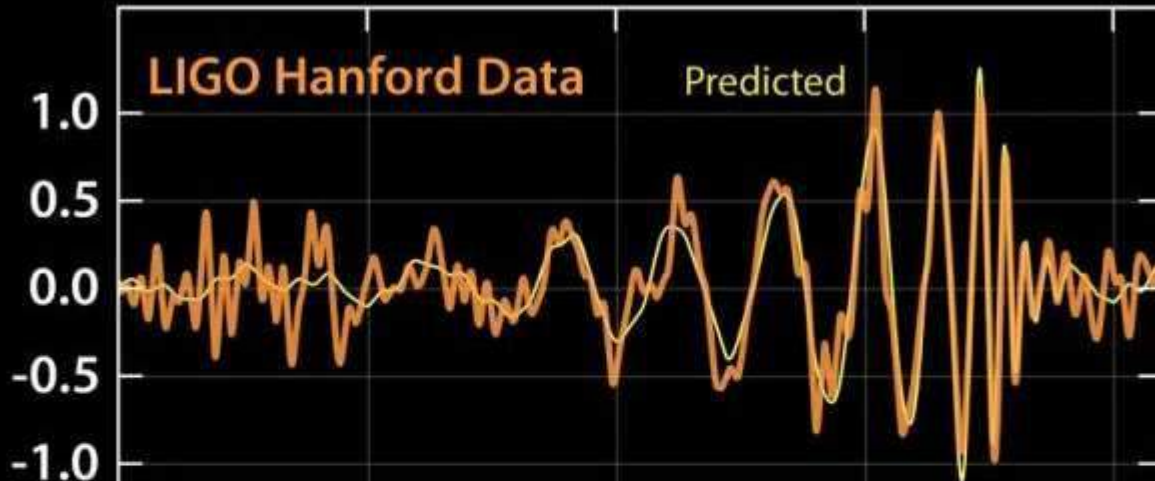


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05/20/2012

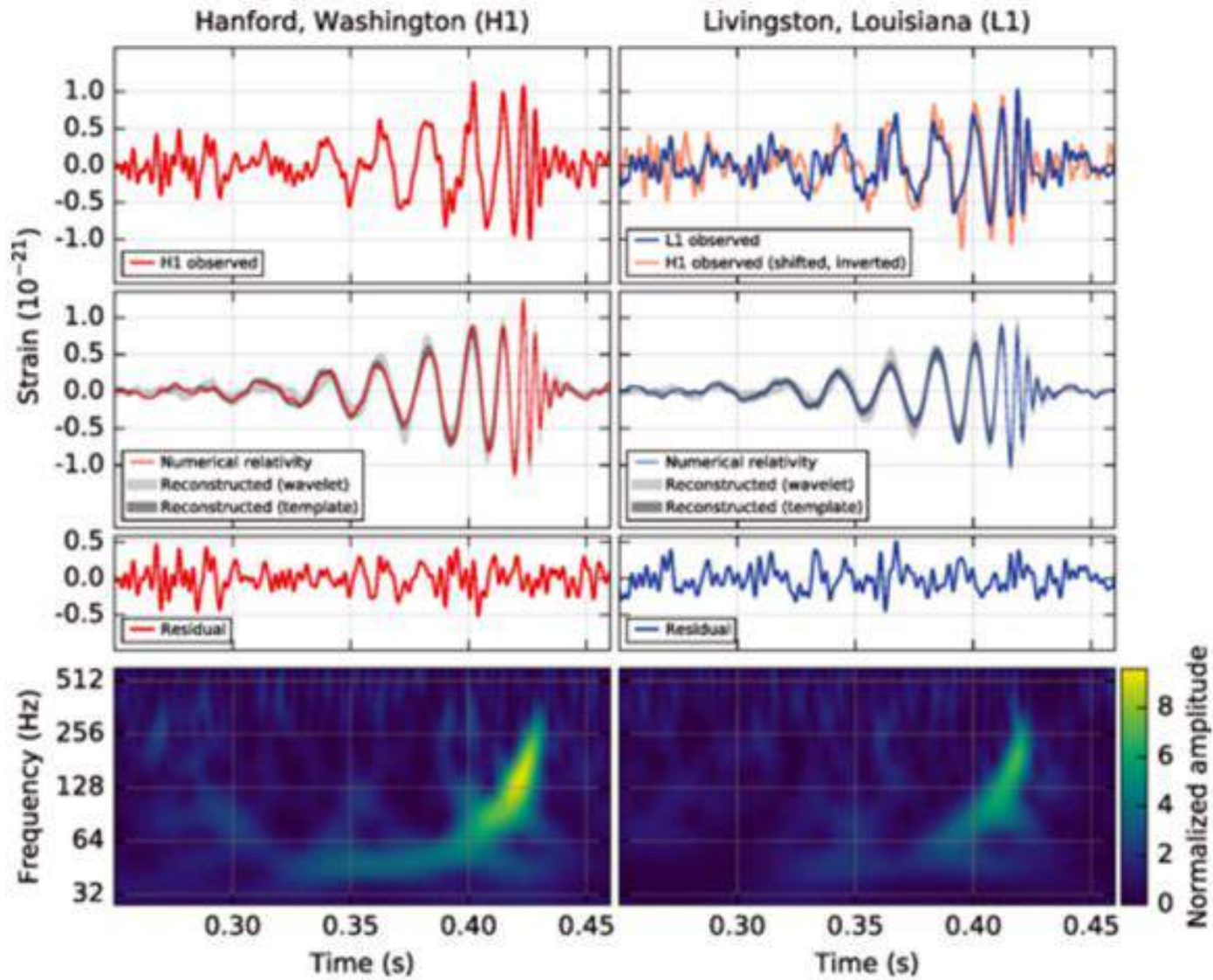


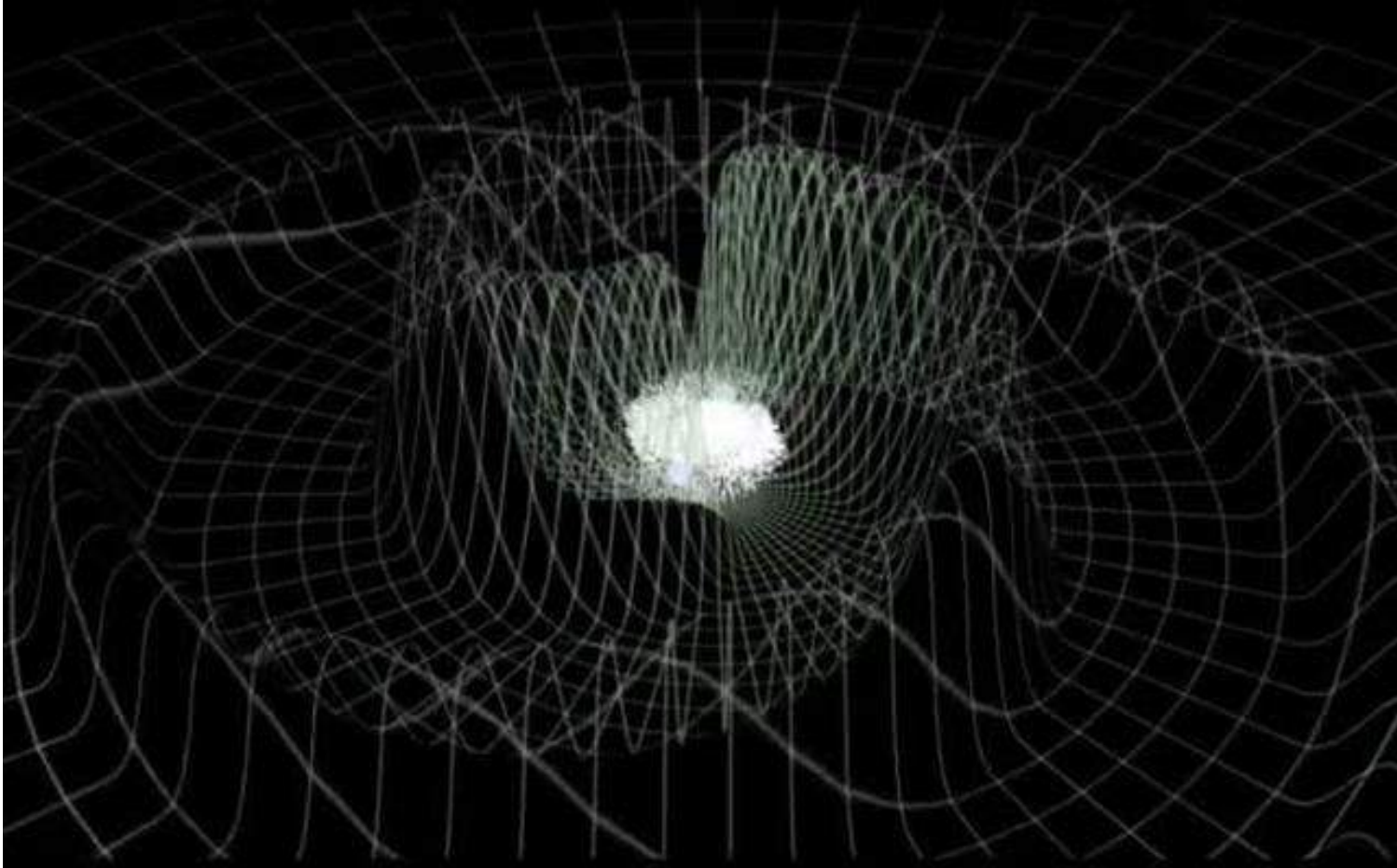
Strain ( $10^{-21}$ )



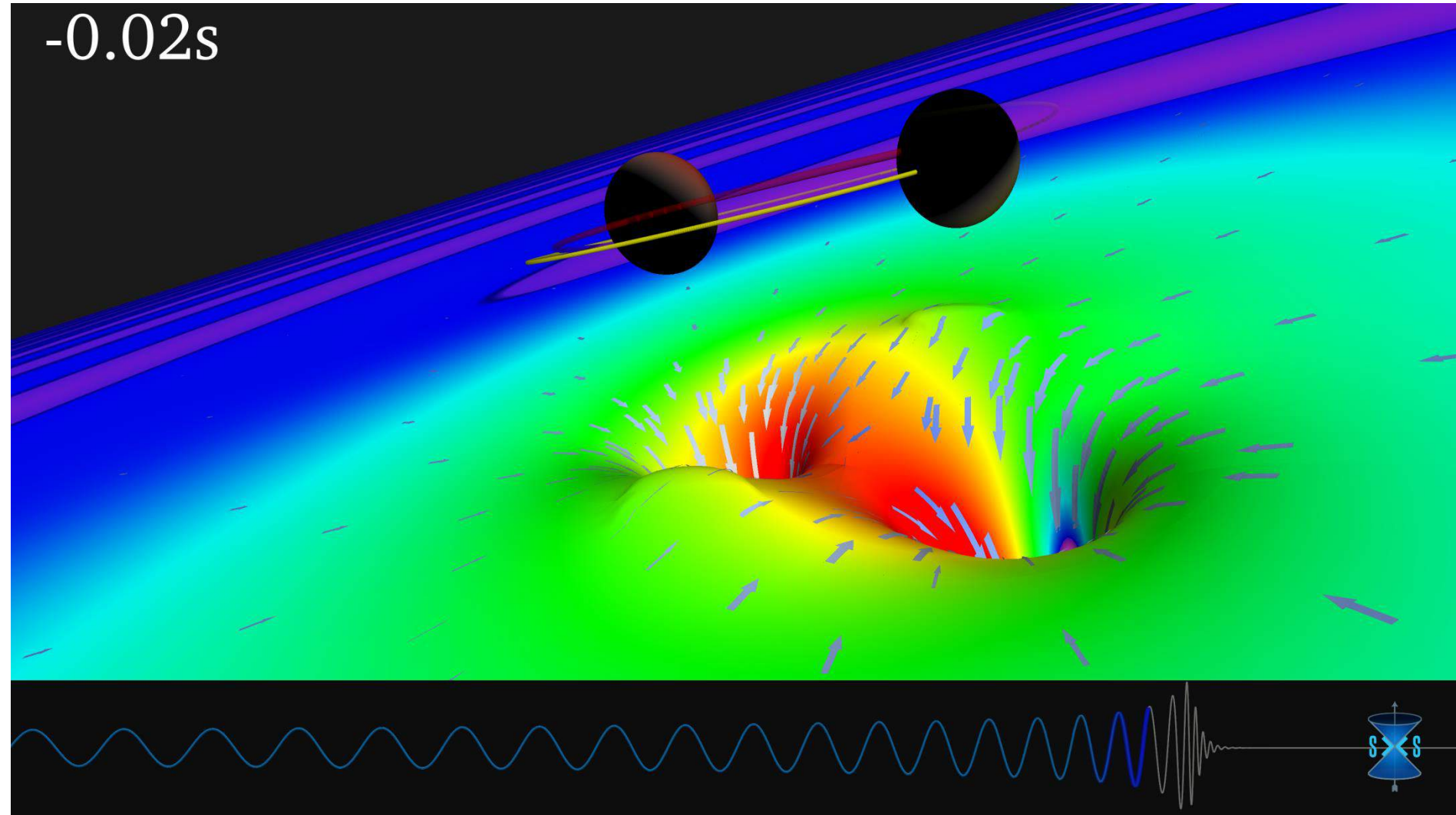
n ( $10^{-21}$ )



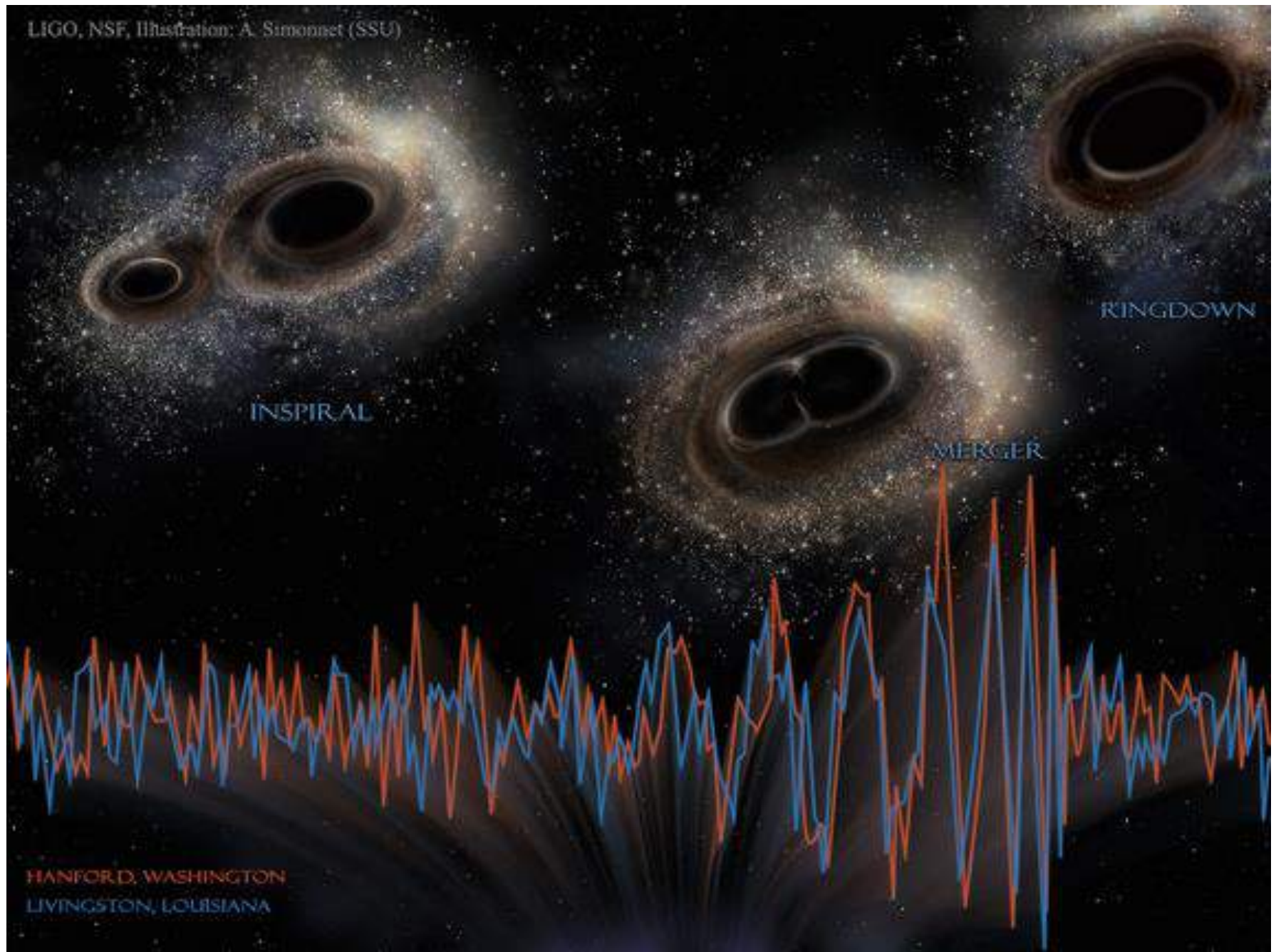




-0.02s



LIGO, NSF, Illustration: A. Simonnet (SSU)



HANFORD, WASHINGTON  
LIVINGSTON, LOUISIANA

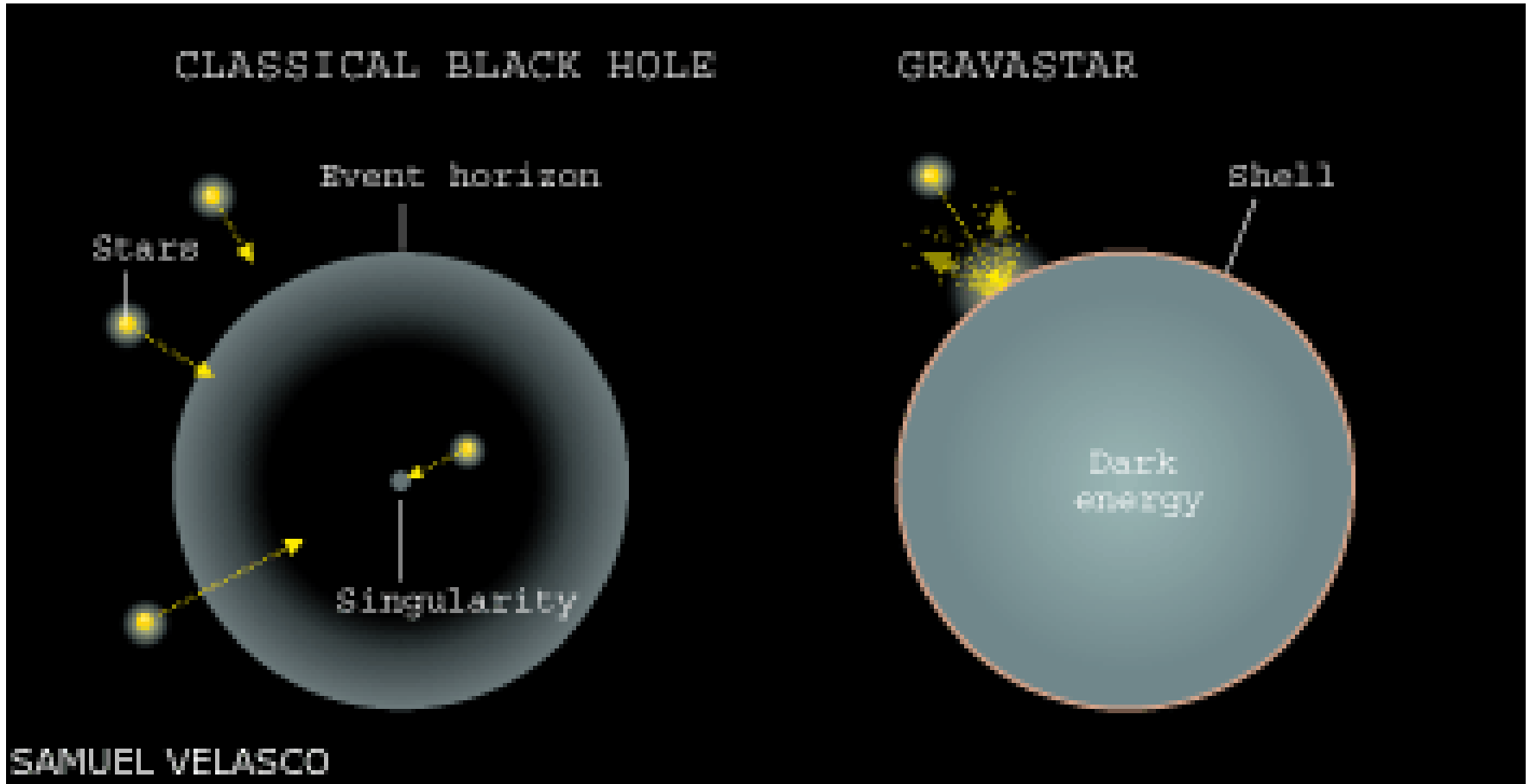


# Gravastar???

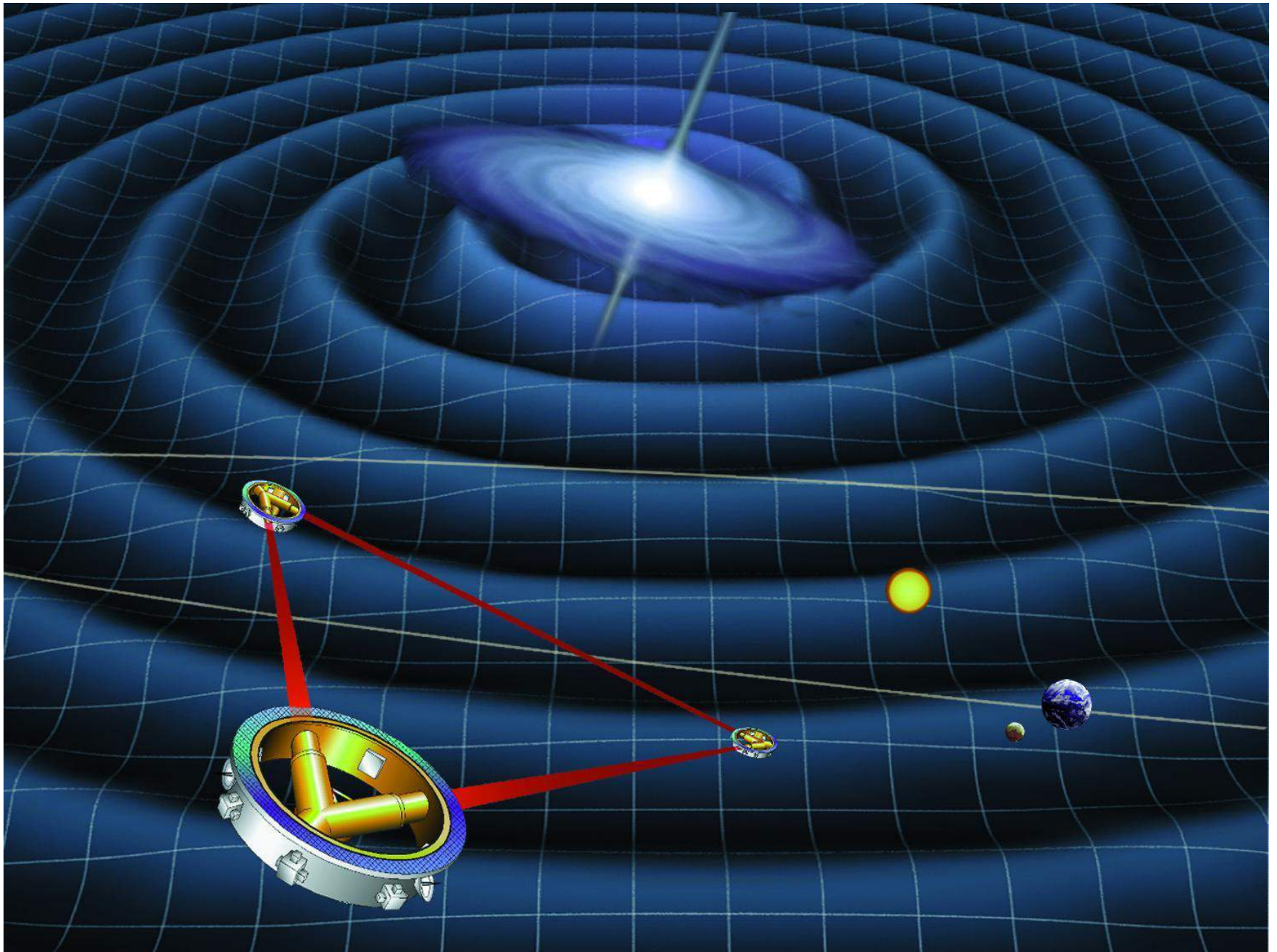
## **Una gravastar? Che cos'è?**

«Le gravastars e i wormholes fanno parte di quelli che in gergo vengono chiamati “*black-hole mimickers*”, ossia oggetti che possono essere tanto compatti quanto un buco nero ma che non possiedono l'orizzonte degli eventi. Una gravastar è una stella esotica, la cui forza gravitazionale è bilanciata da un nucleo interno fatto di energia oscura: in pratica, di materia con una pressione negativa. I wormholes sono invece una sorta di tunnel spaziotemporale che connette due regioni distanti del nostro universo o addirittura due universi diversi. Il punto più stretto del tunnel, detto “gola”, è anch'esso, formato da materia esotica, simile a quella delle gravastar.

# Gravastar???







# Lanciarsi in un buco nero?







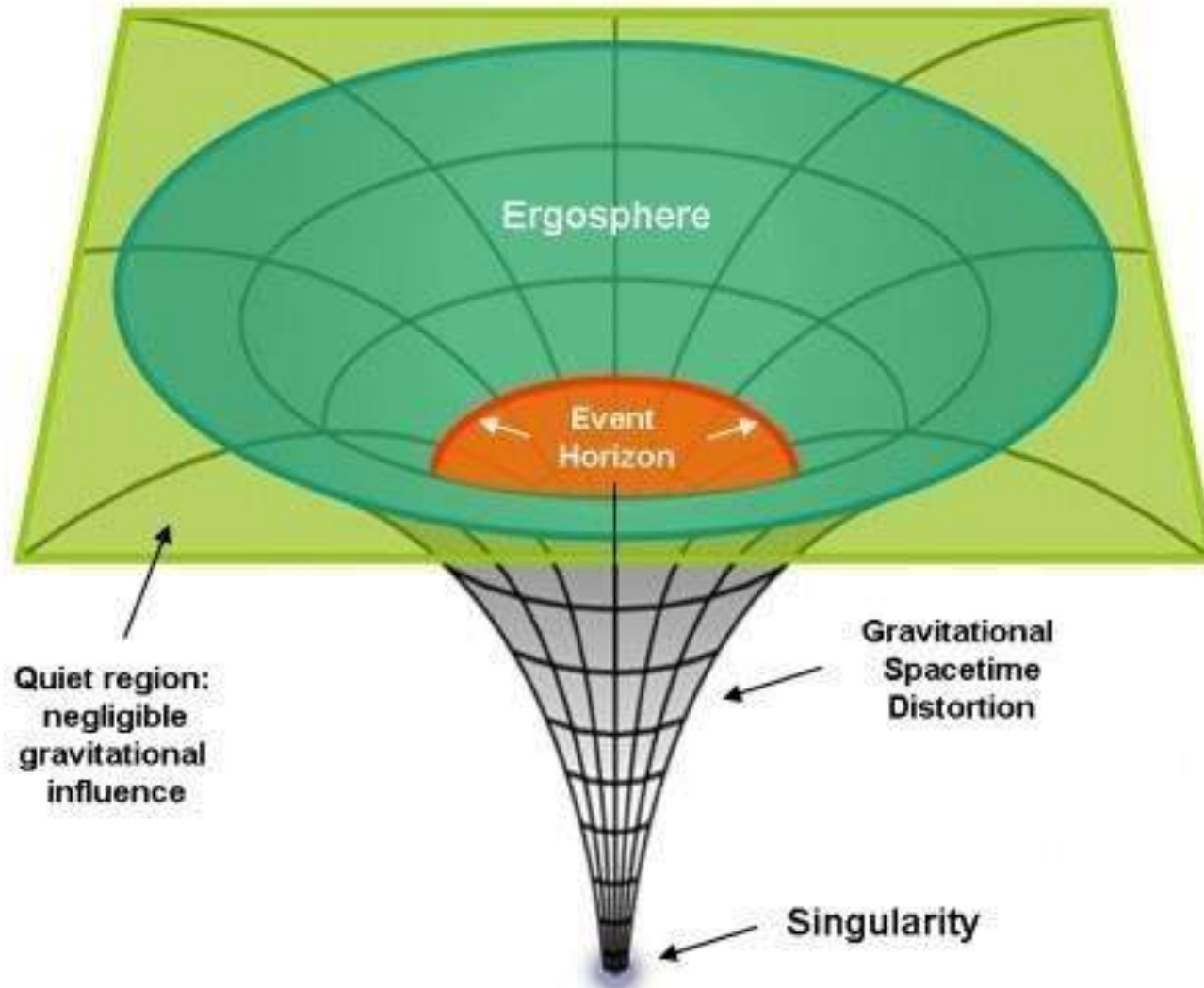








# Black Hole Regions

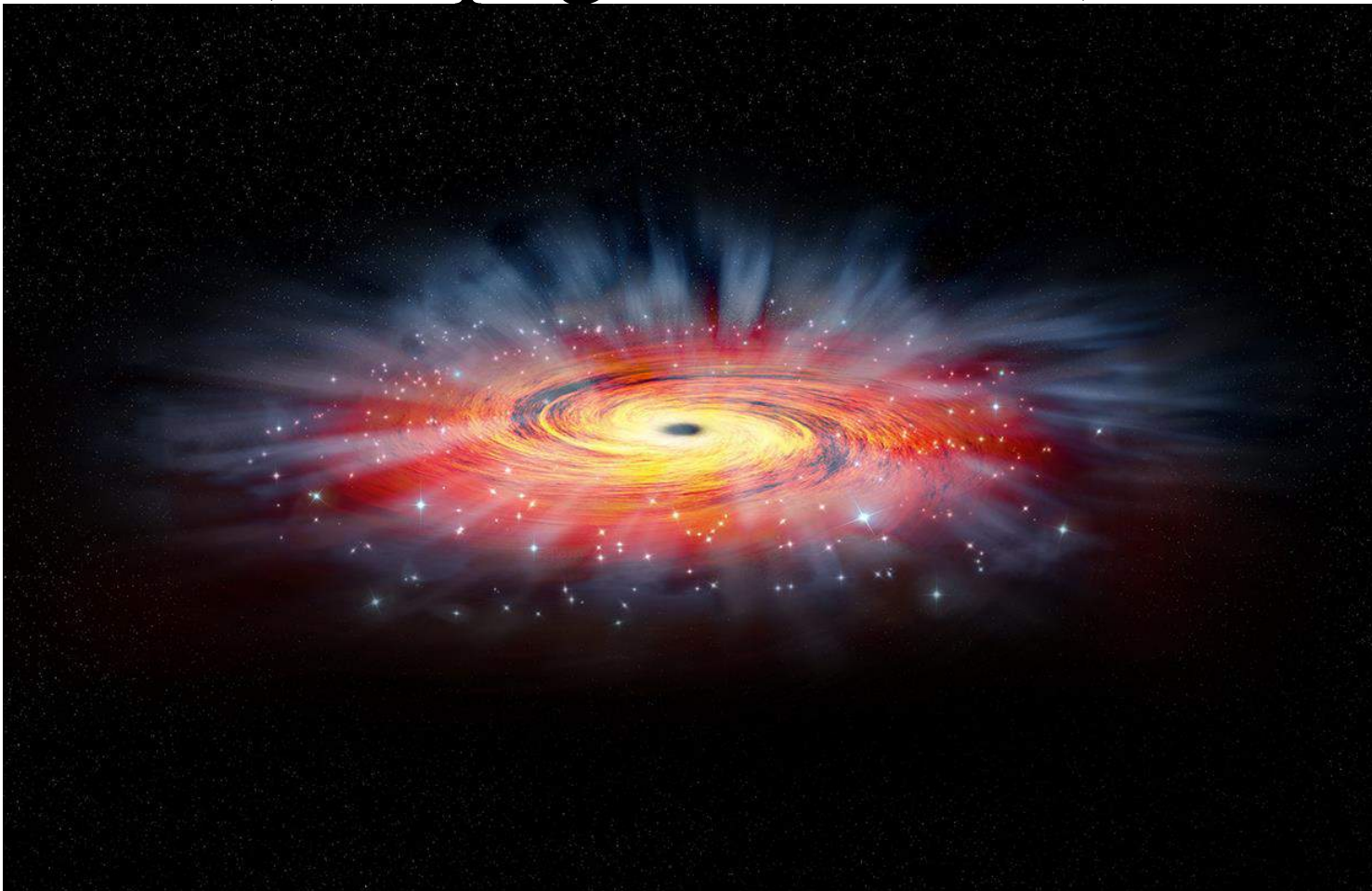


# Spaghettification!

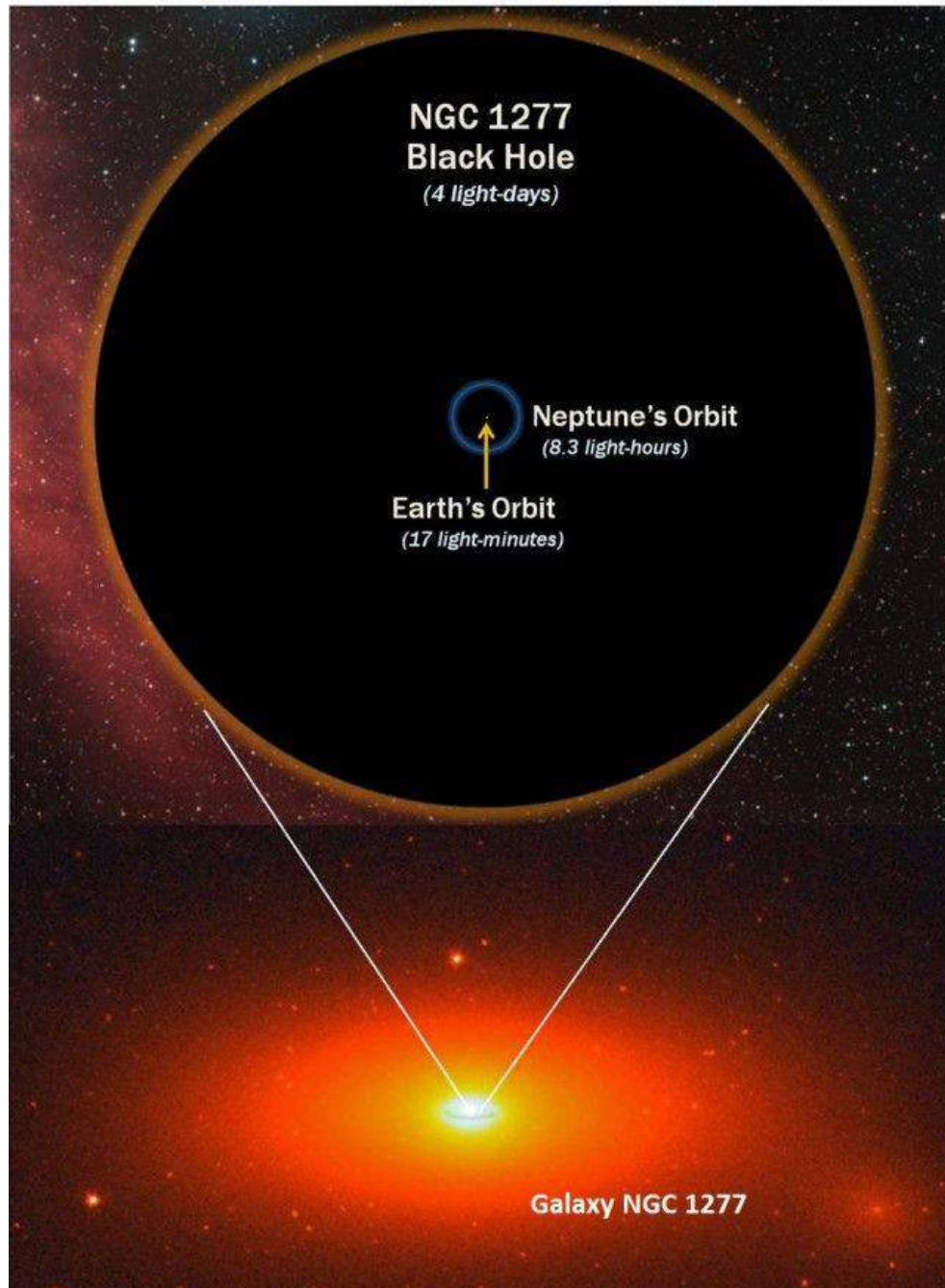


To Black Hole

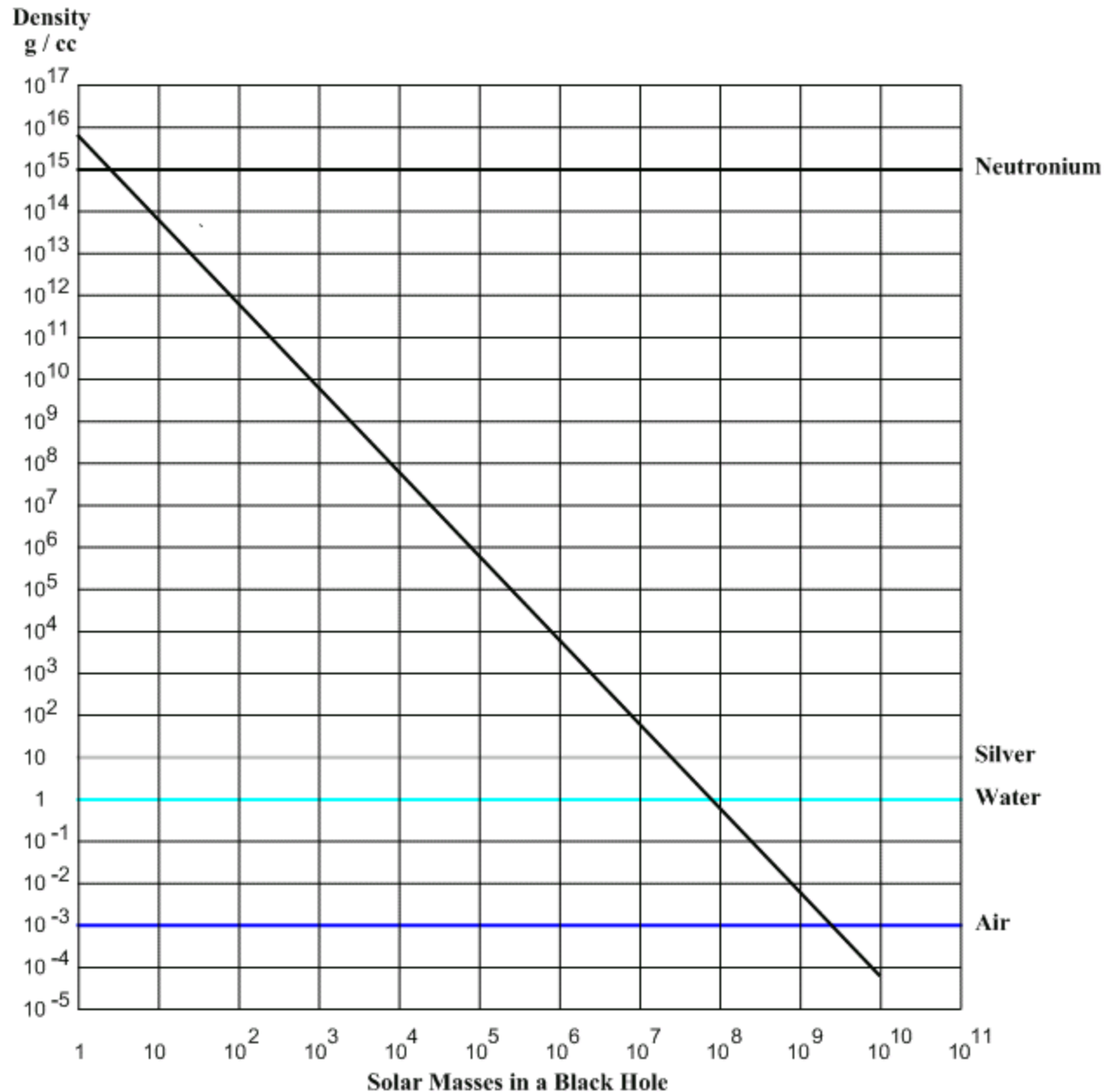
# Buchi neri supermassicci densita' bassa (no spaghetificaton)



# Black Hole – mass about 17 miliardi di volte massa Sole

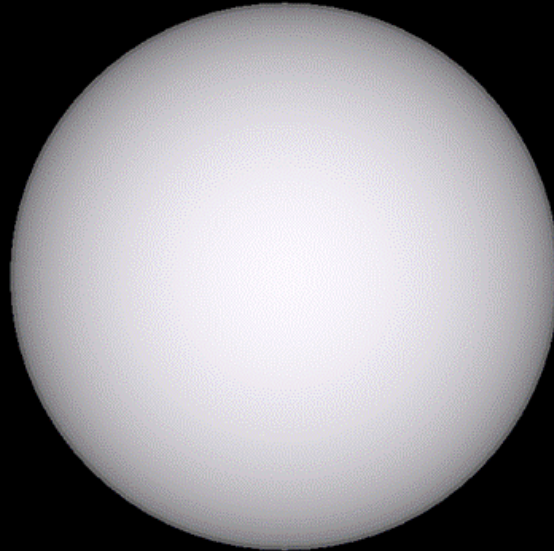


# Densita' buchi neri

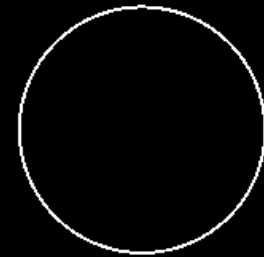




Manhattan  
(spaceimaging.com)



Neutron Star  
 $M = 1.5 M_{\text{sun}}$   
 $R \approx 10 \text{ km}$



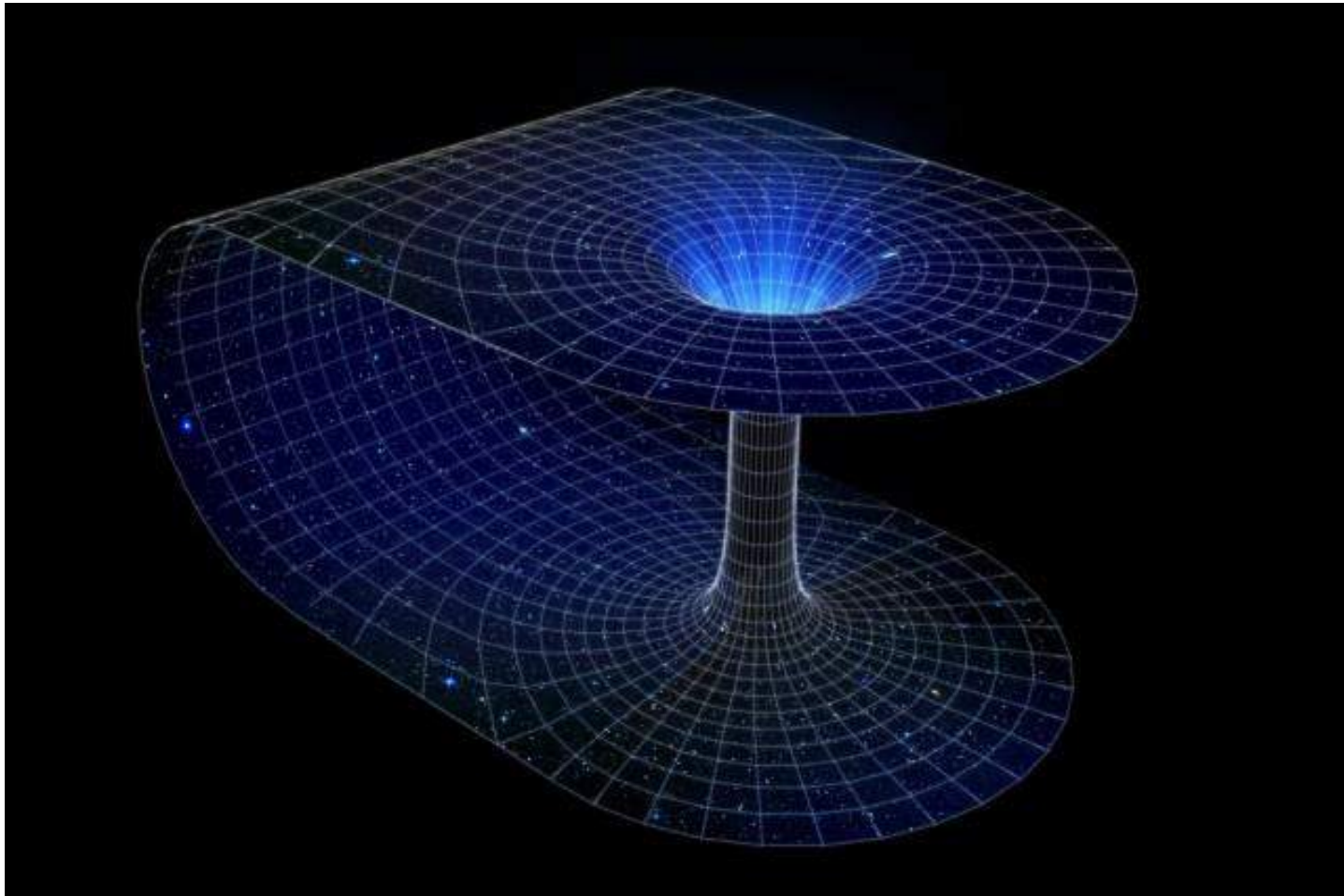
Black Hole  
 $M = 1.5 M_{\text{sun}}$   
 $R_S = 4.5 \text{ km}$



# Viaggi interstellati?



# Viaggi interstellati?

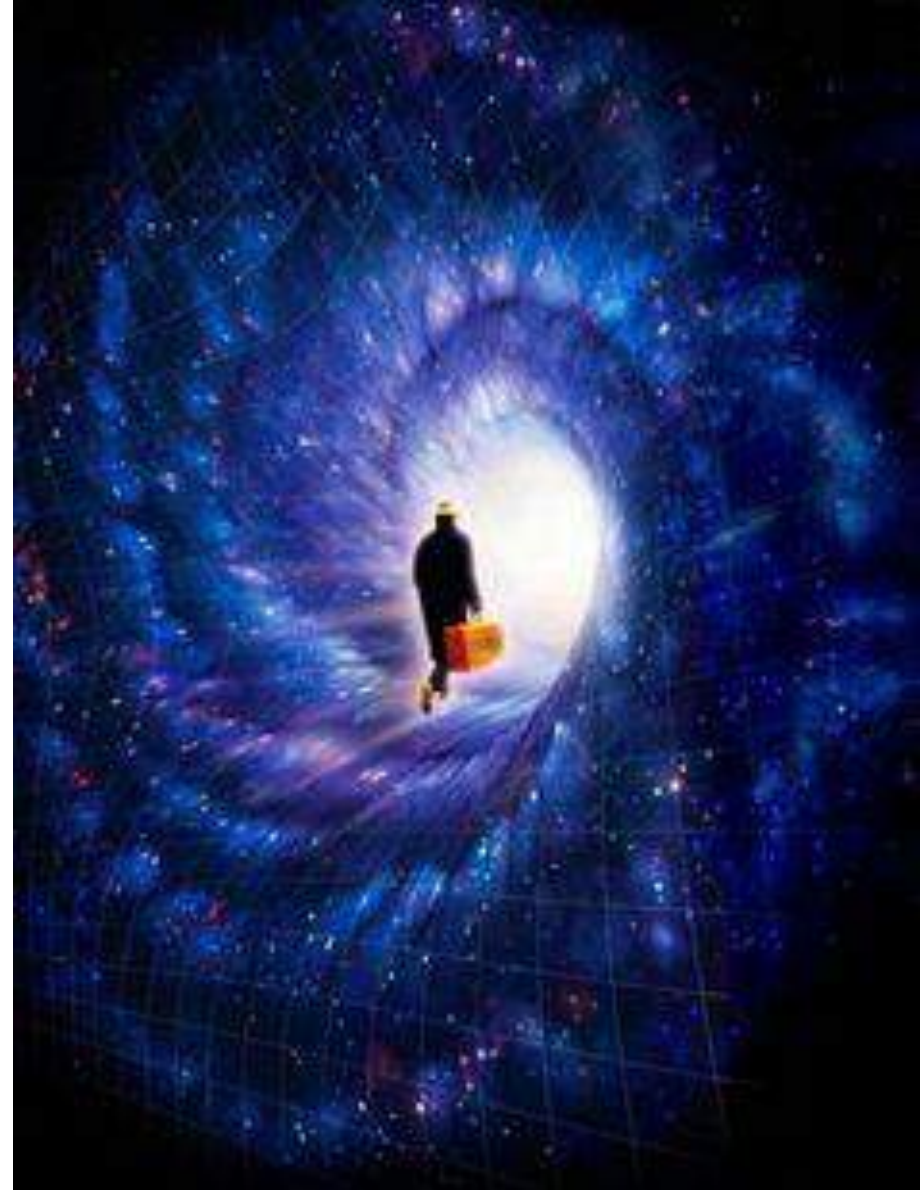


Viaggi in altri Universi?



SCIENCEPHOTOLIBRARY

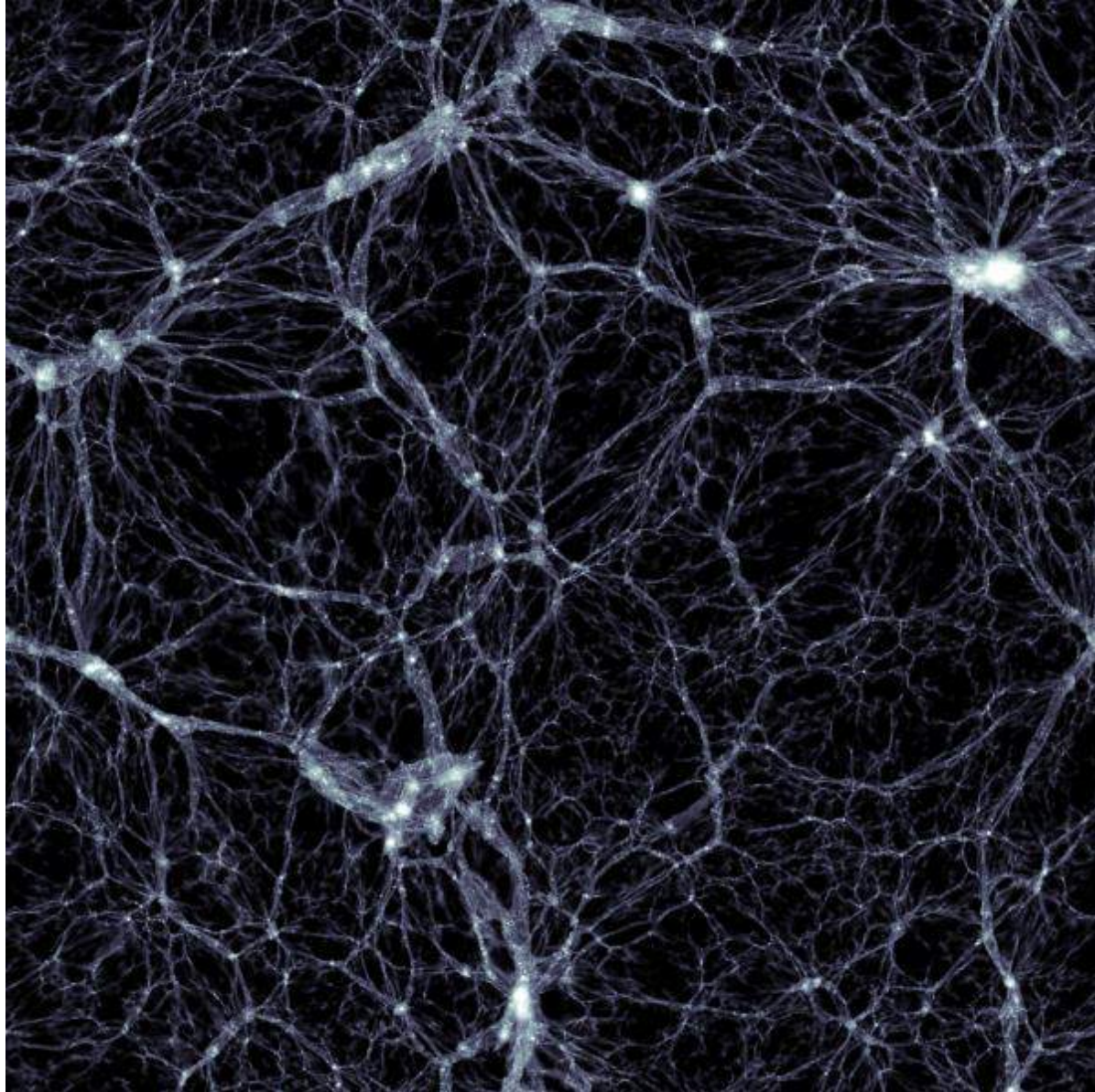
**Viaggi in altri Universi?**



Per capire i buchi neri:  
gravita' quantistica!



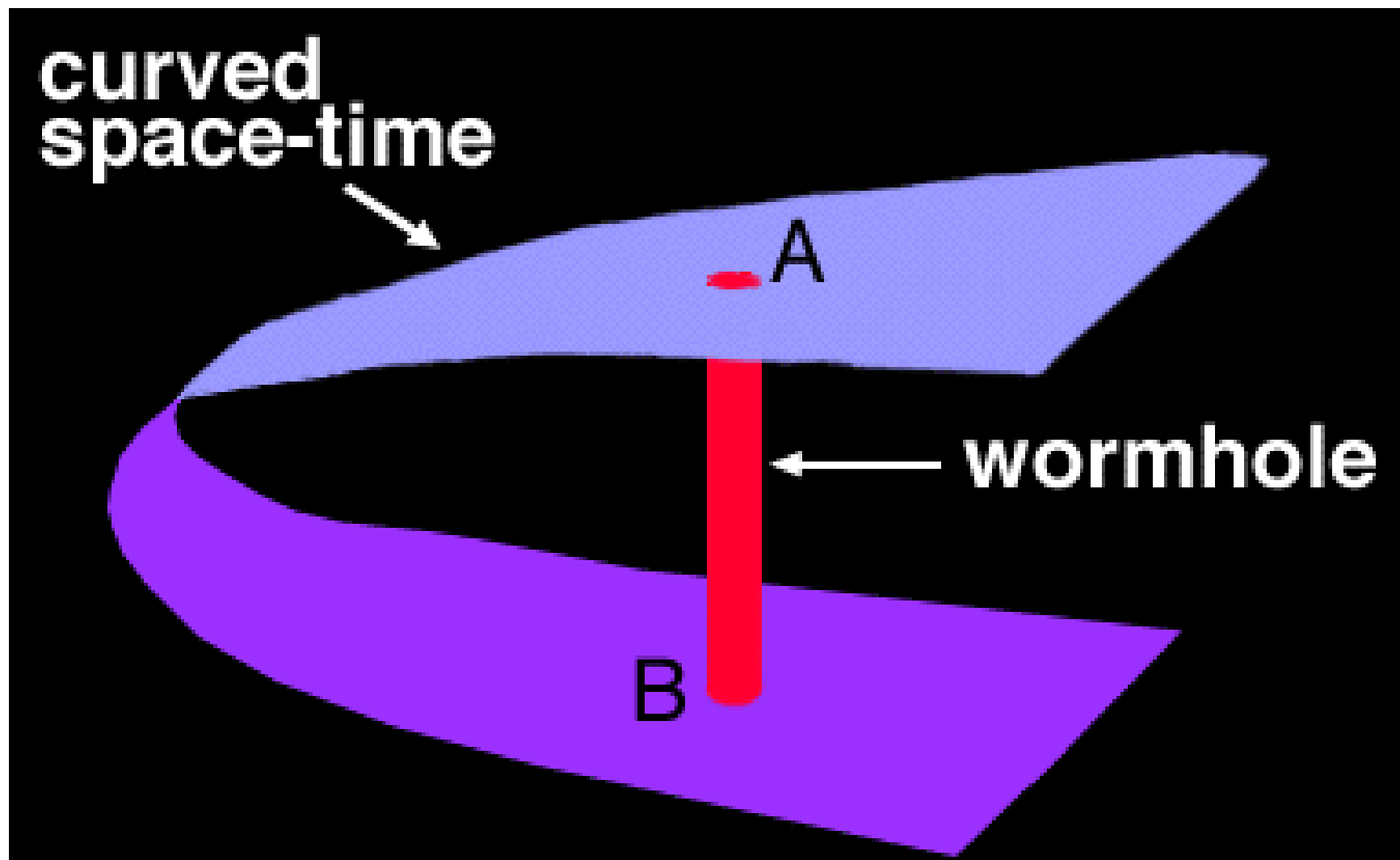
# Struttura dello spazio - tempo



# Viaggi interstellati?

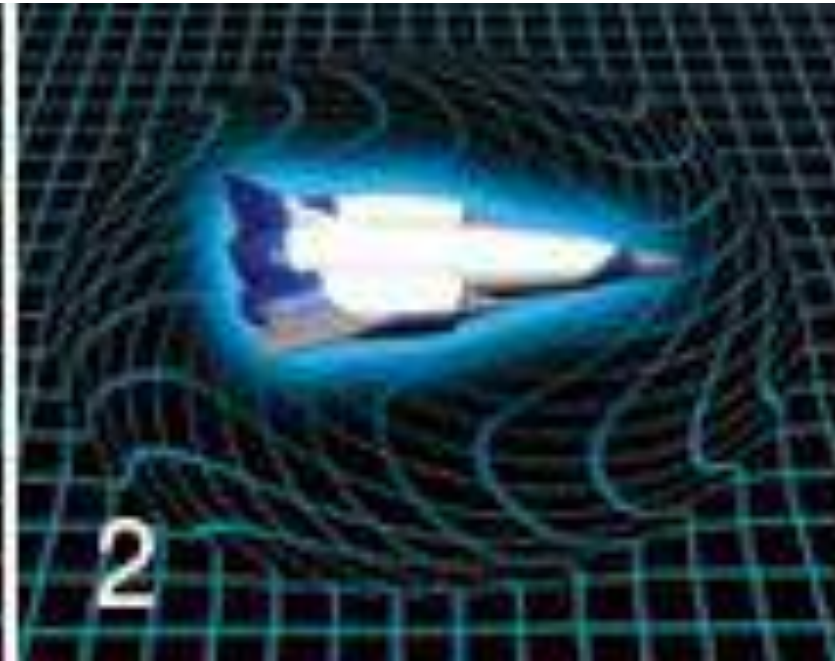
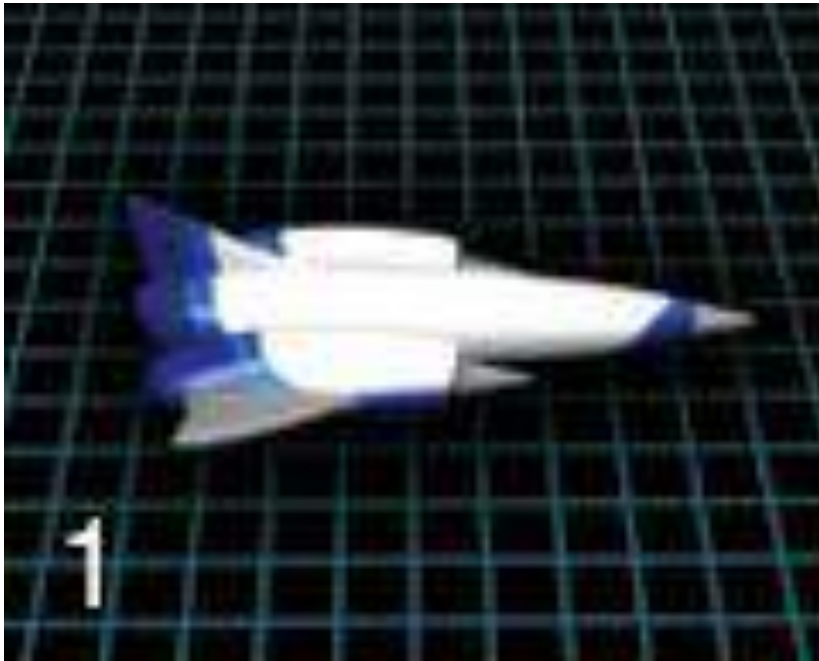


# Scorciatoie spazio-temporali

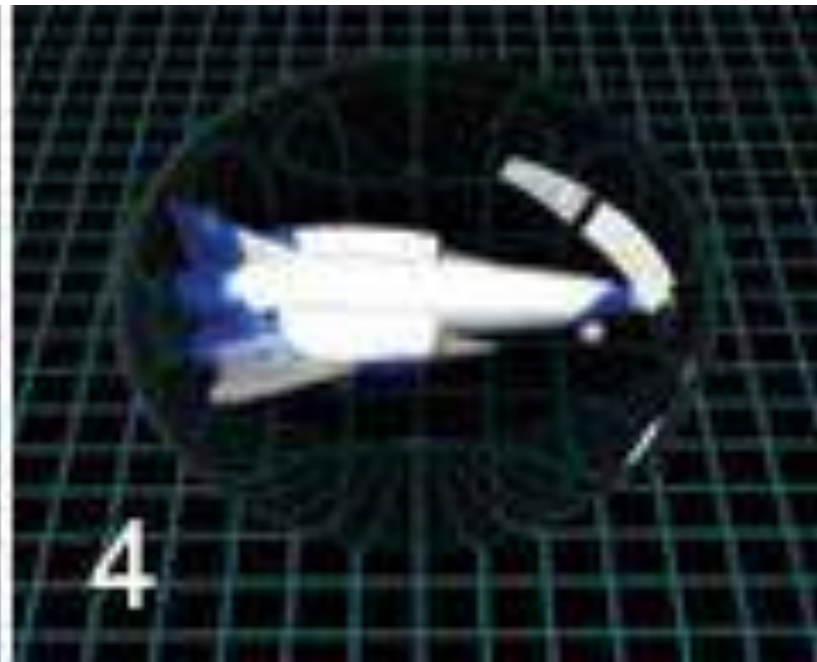




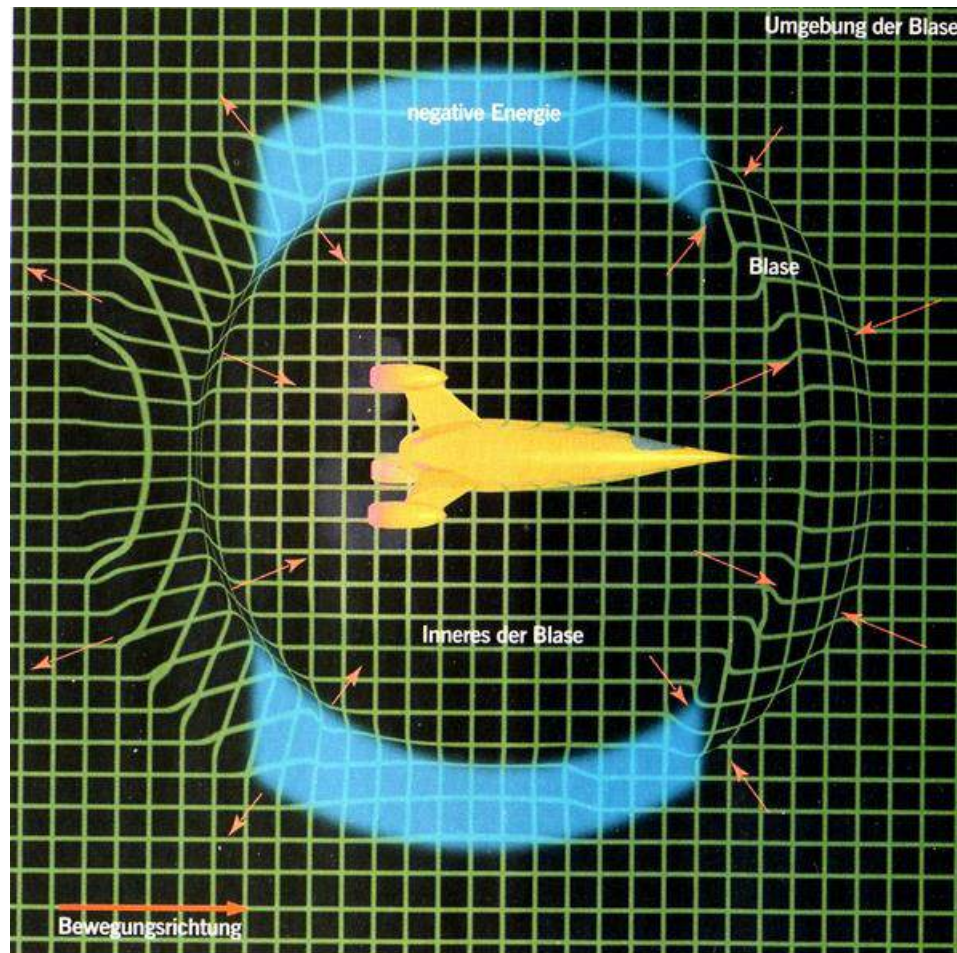
# Motori che manipolano lo spazio-tempo



# Motori che manipolano lo spazio-tempo



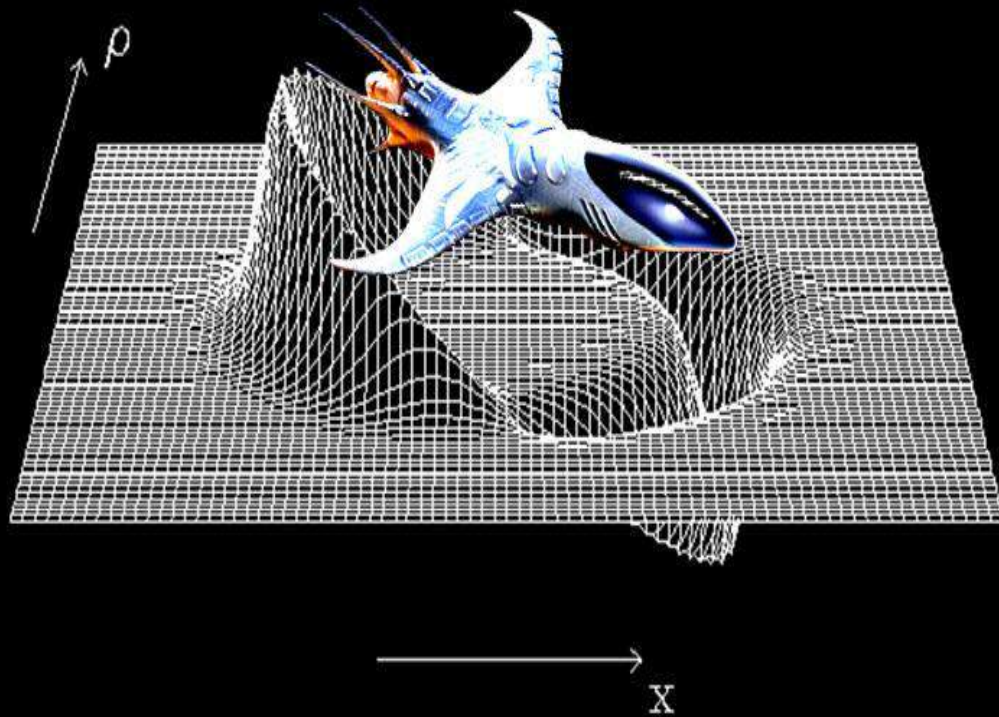
# Motori che manipolano lo spazio-tempo: Alcubierre – surf sulle onde spazio- temporali





# Alcubierre Warp Drive

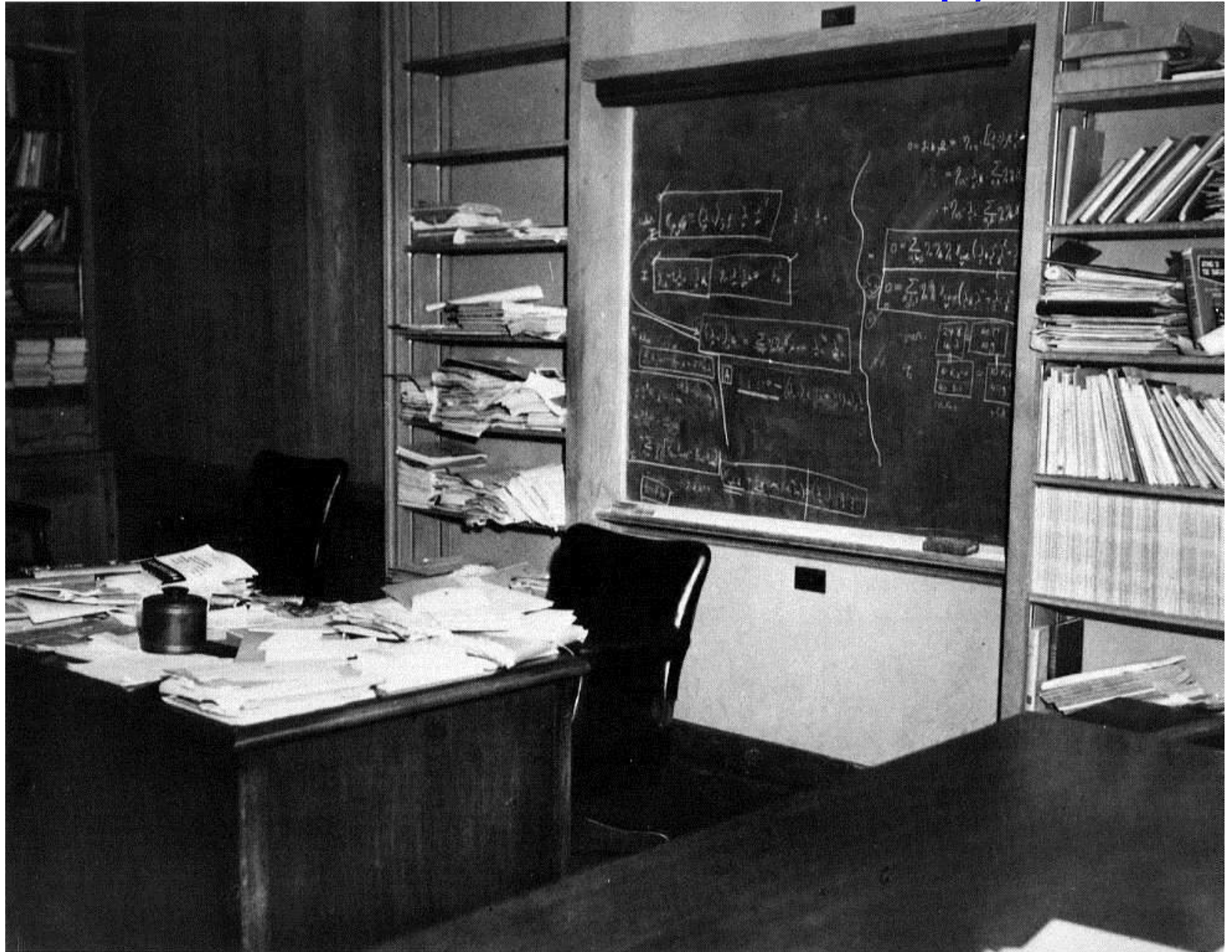
$$\psi = -\alpha \text{Tr}(K)$$



Alcubierre Warp Drive: stretches spacetime in a wave causing the fabric of space ahead of a spacecraft to contract and the space behind it to expand.

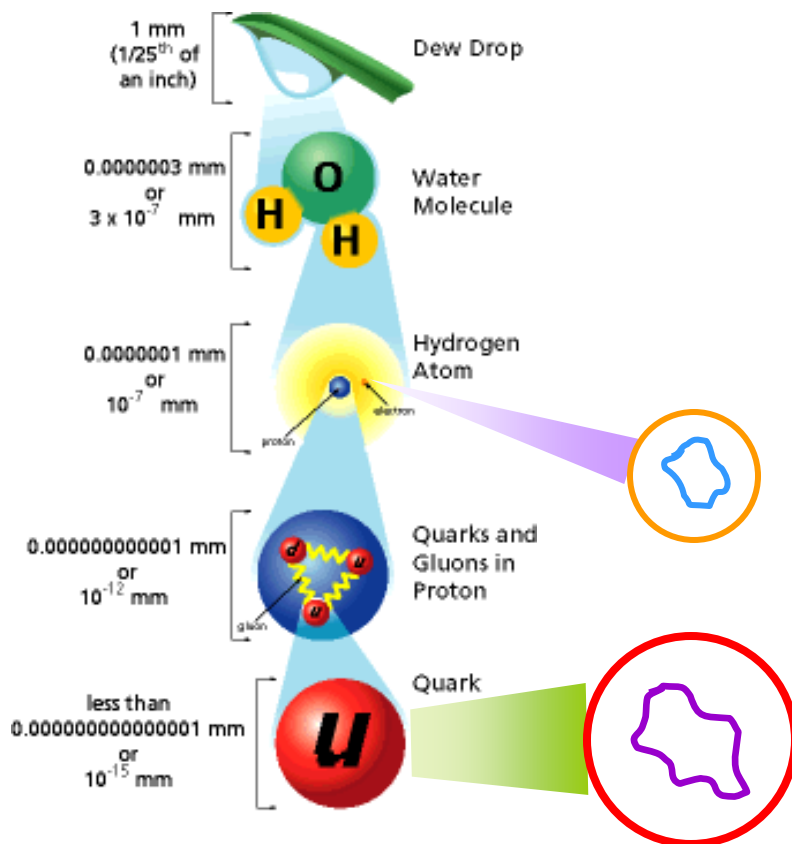
The ship can ride the wave to accelerate to high speeds and time travel.

# *Einstein – l'ultima lavagna*



# Questioni Aperte

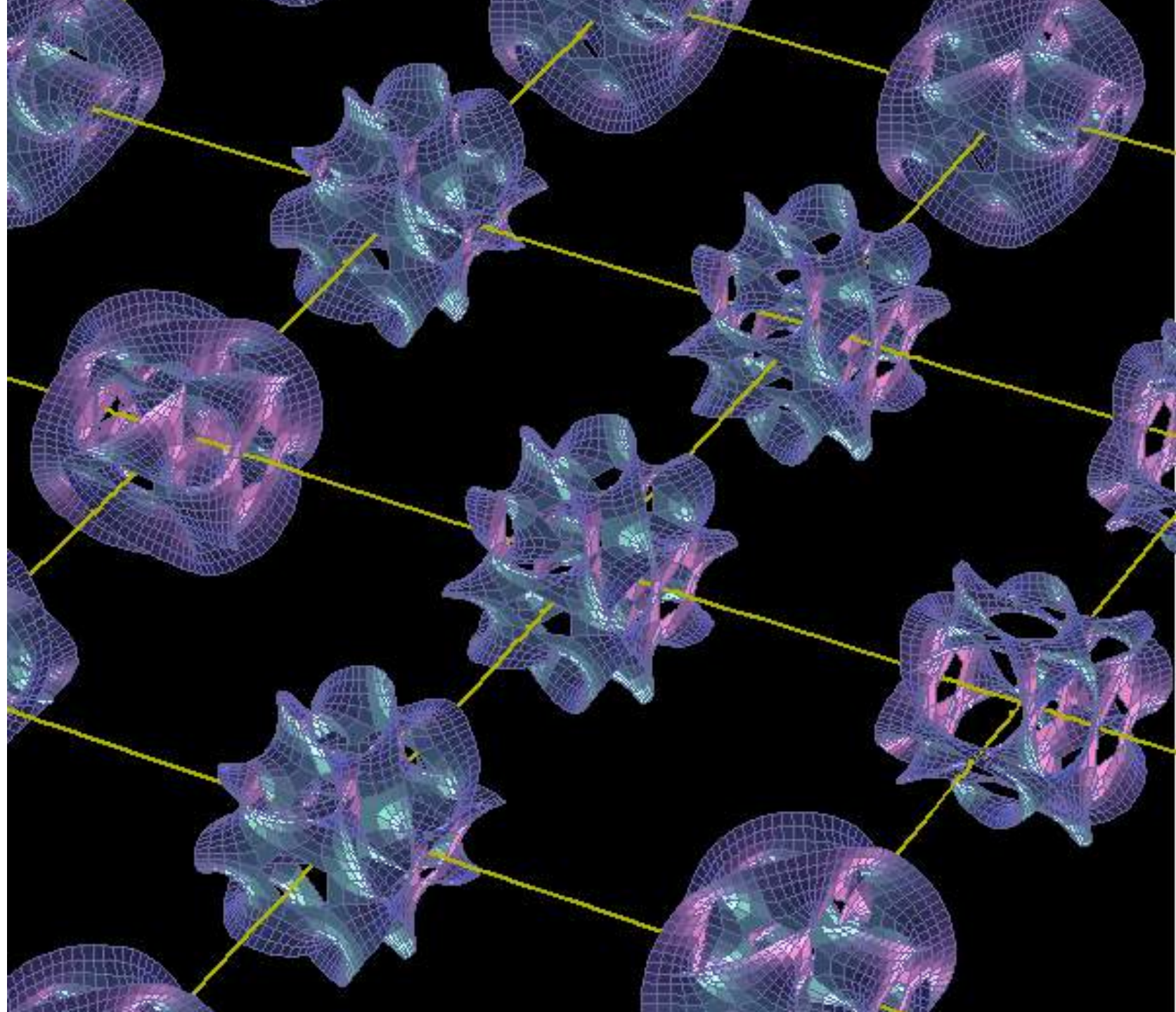
- Le particelle sono veramente puntiformi ?

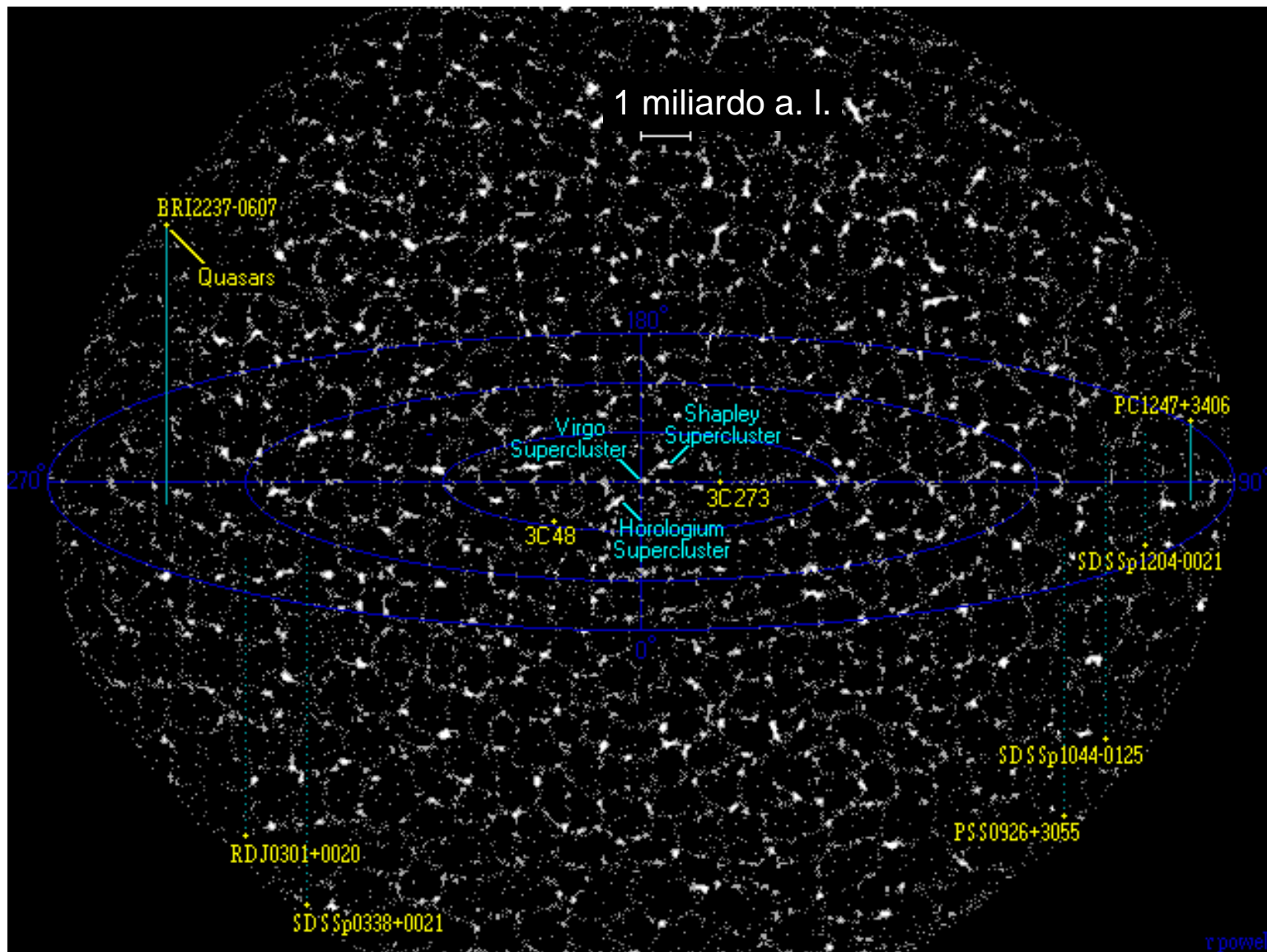


## Teoria delle Stringhe

ulteriore livello  
microscopico: particelle  
non sono puntiformi, ma  
piccoli ( $10^{-33}$  cm) anelli  
oscillanti

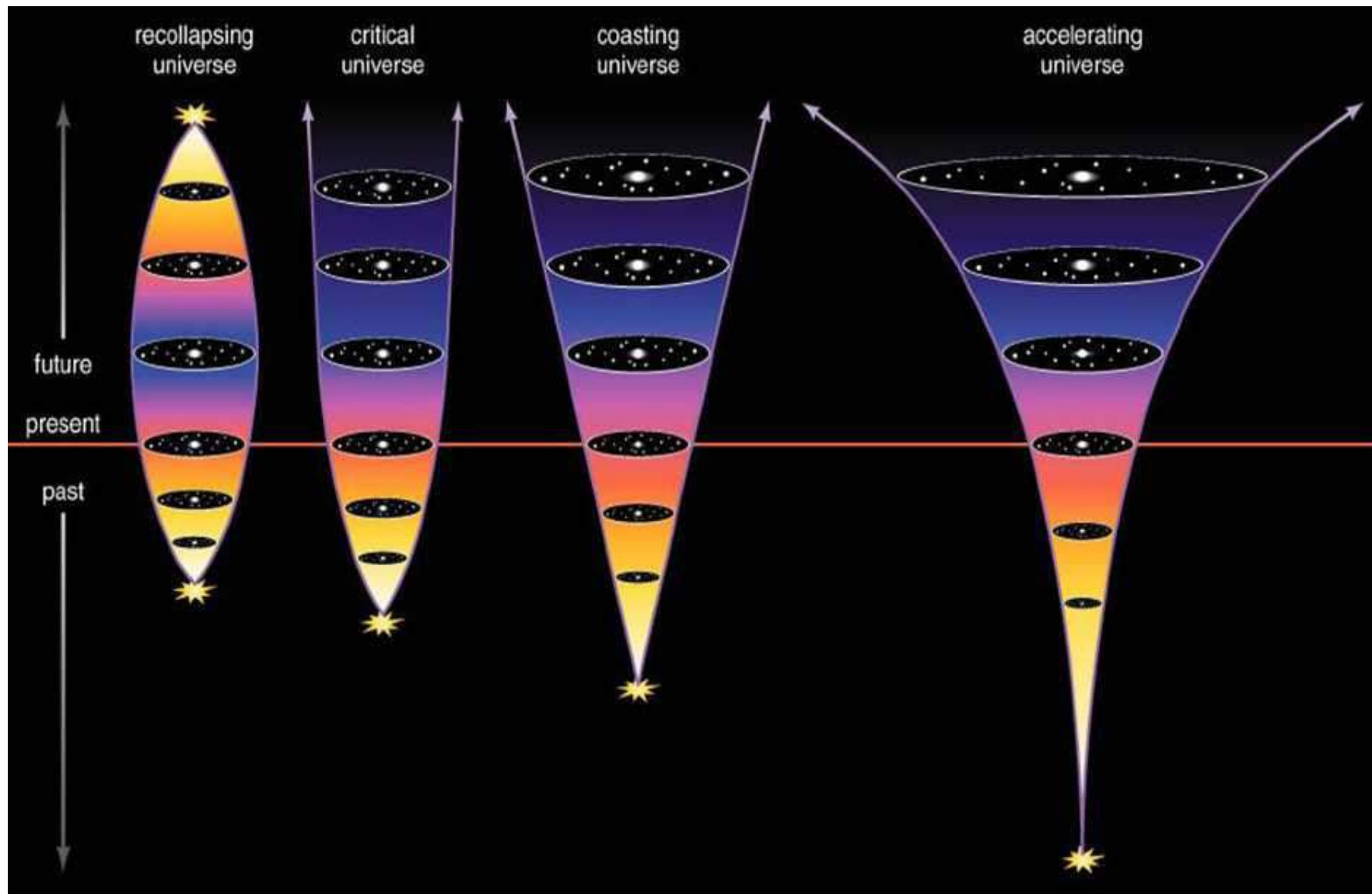
diversi stati di oscillazione  
della stringa →  
particelle diverse





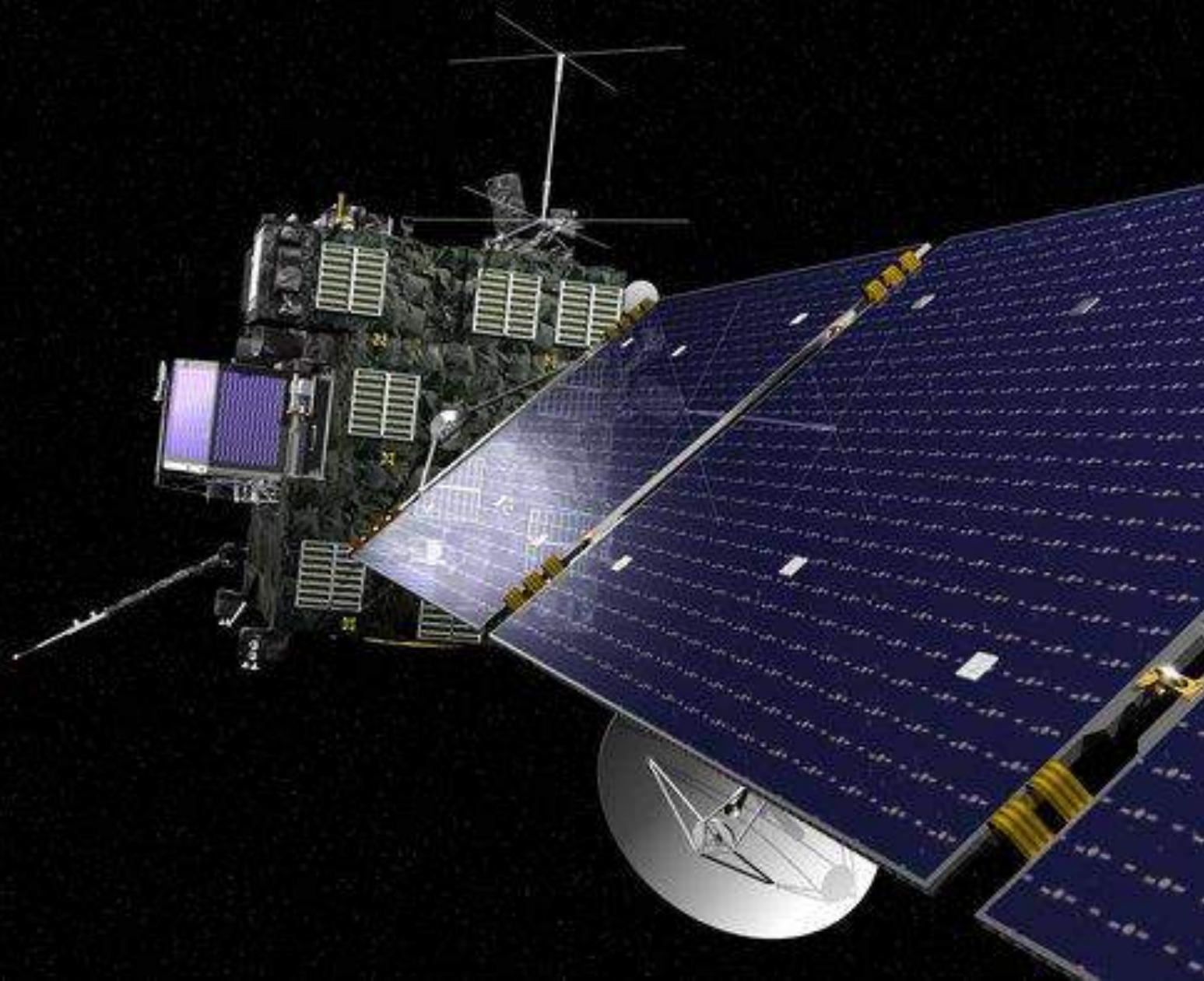


# I modelli Cosmologici











**Stiamo soli nell'Universo?**

# *Feynman – Il valore della scienza*



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Fuori dalla culla  
sulla terra asciutta  
eccolo  
in piedi:  
atomi con la coscienza  
materia con la curiosità.  
In piedi davanti al mare  
meravigliato della propria meraviglia: io  
**un universo di atomi**  
**un atomo nell'universo**

Catalina Oana Curceanu

# Dai buchi neri all'adroterapia

Un viaggio nella fisica moderna



Nata in Transilvania (Bisnov, Romania) Catalina Oana Curceanu è Pitagora Ricercatore dell'Istituto Nazionale di Fisica Nucleare, Laboratori Nazionali di Frascati. Dirige un gruppo di ricercatori che lavorano nel campo della fisica sperimentale adronica e nucleare, conducendo esperimenti sia in Italia sia all'estero, e coordinando vari progetti europei. Ha organizzato varie conferenze internazionali ed è autrice di più di 200 pubblicazioni scientifiche in riviste internazionali. Svolge un'intensa attività di formazione e divulgazione scientifica e scrive per vari giornali e riviste italiane e straniere. Ha la passione di spiegare a tutti quanto sia bello e affascinante il mondo della scienza.