

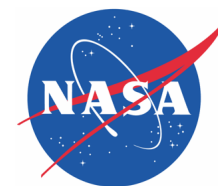
Satellite and Lunar Laser Ranging: fundamental physics and applications

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Feb. 14, 2017



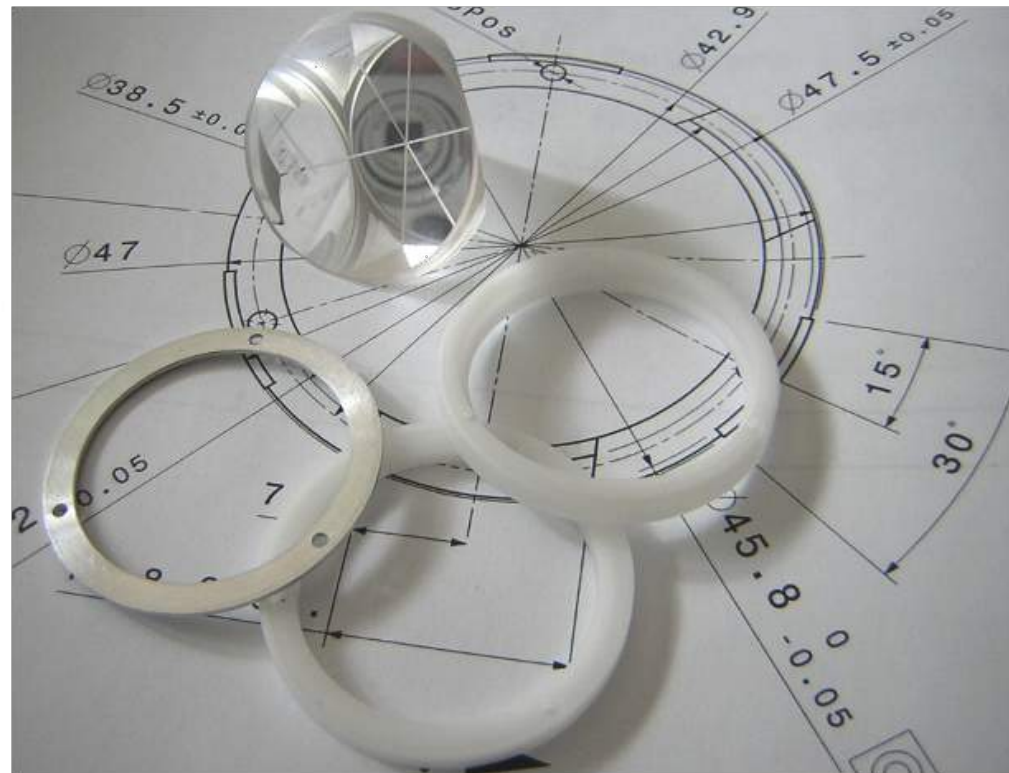
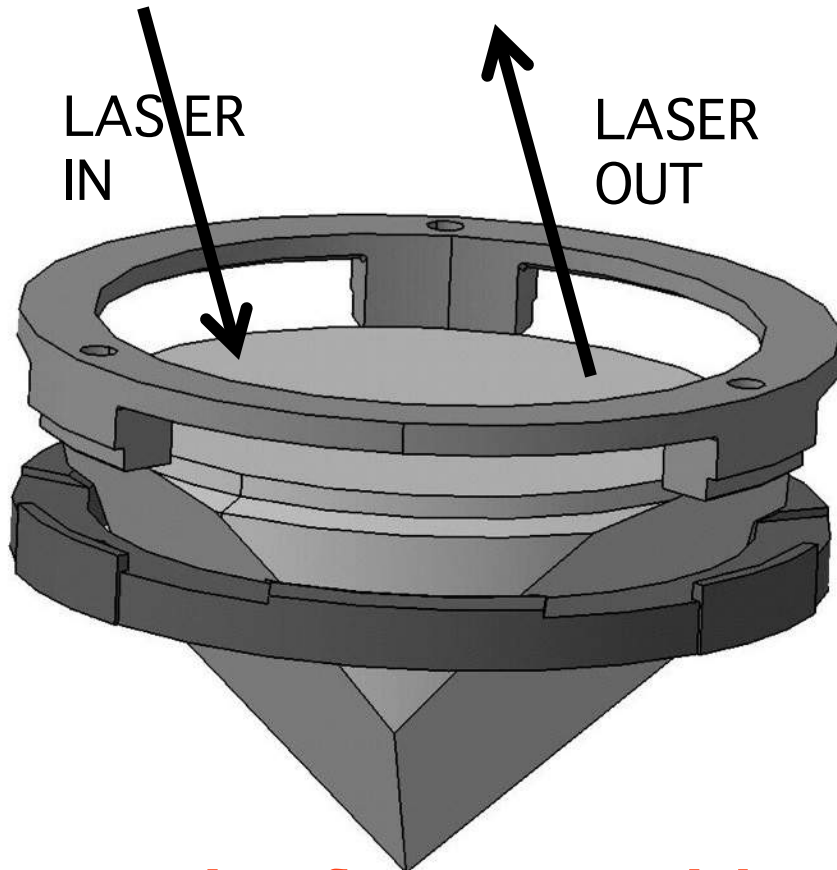
- **Introduction**
 - Missions **Apollo** (USA) / **Luna** (USSR)
 - Satellites **LAGEOS** (USA-Italy)
- **Space Geodesy, Geodynamics**
- **Investigations of General Relativity. And beyond**
- ***Global Navigation Satellite System (GNSS)***
 - GPS, GLONASS, Galileo, IRNSS, ...
- **Next frontier**

“Prisms for ground surveys”, mounted inside spheres

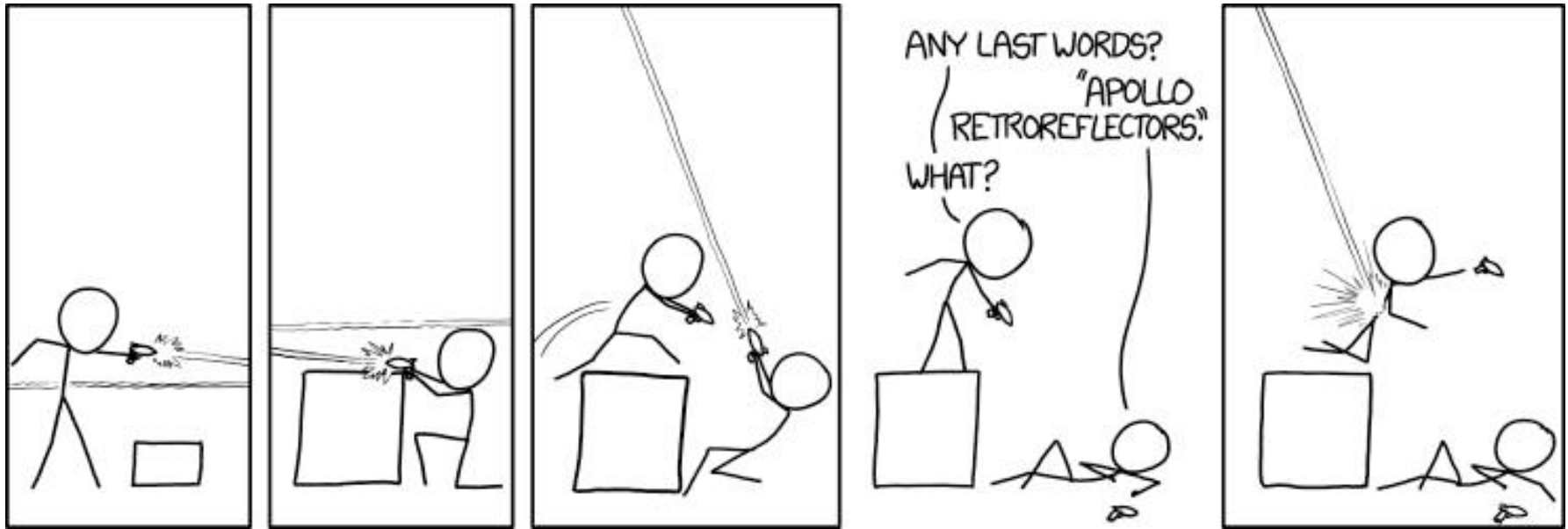


Taylor-Hobson
sphere
for angular
measurements

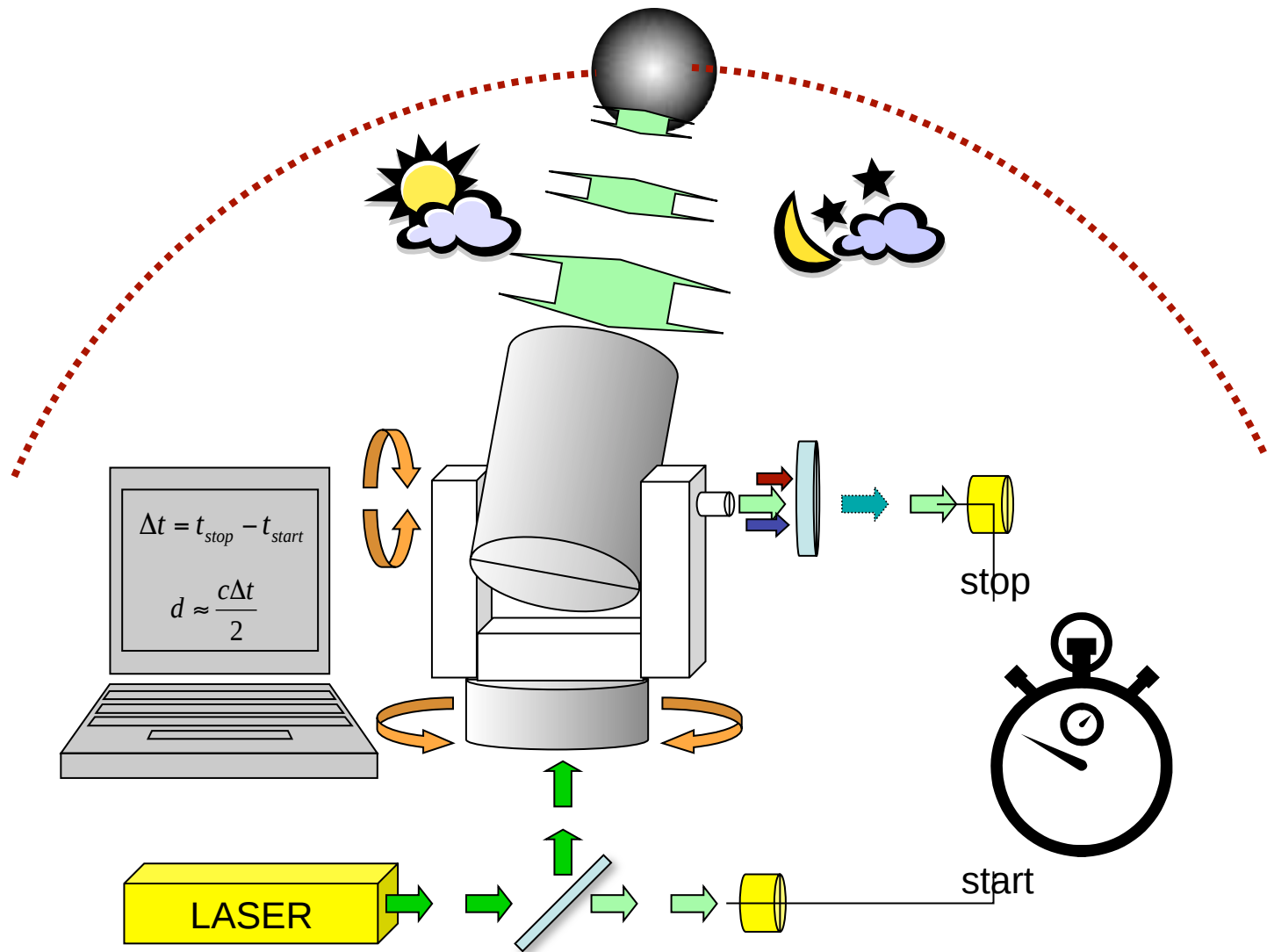
Trick: **cube “corners”**. The triple TIR on back determines an overall retro-reflector IN-OUT



Triple reflection around the corner

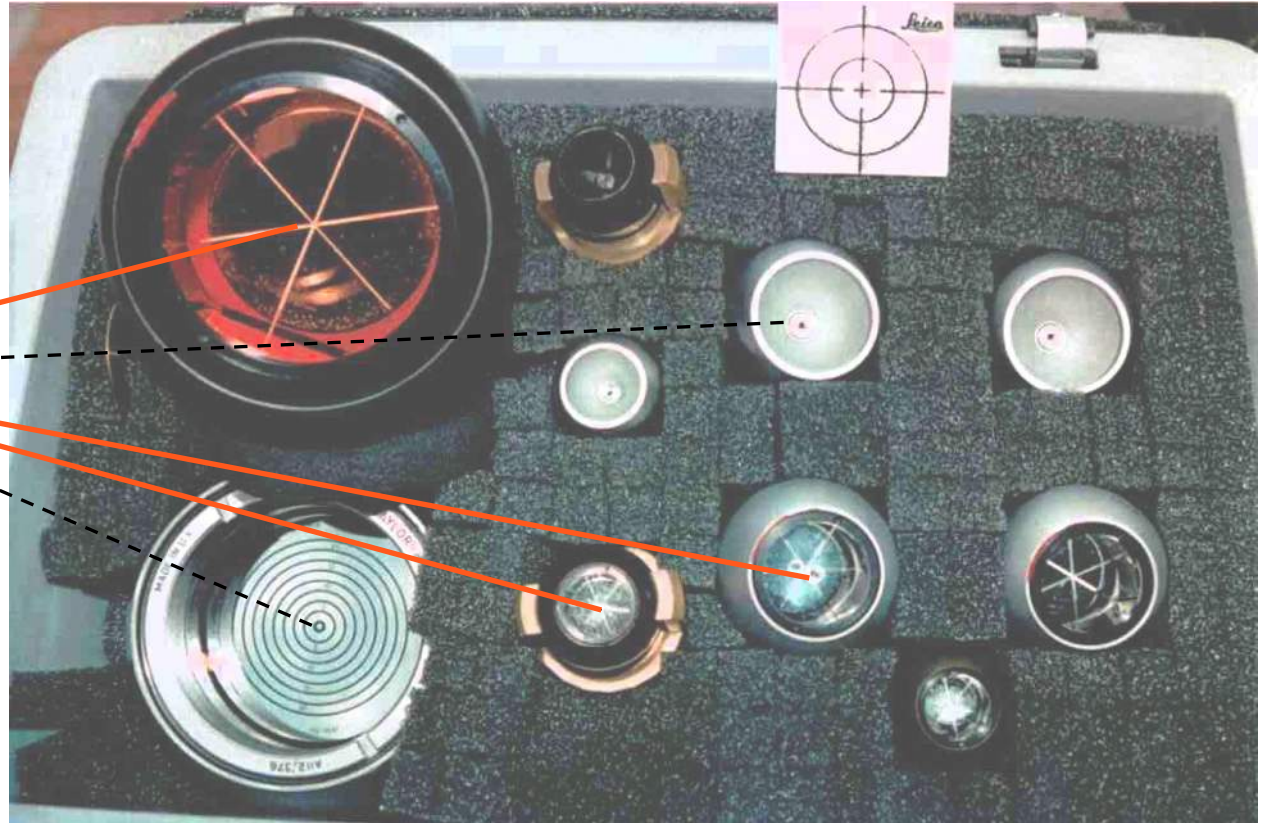


In space: laser pulse, satellite with retroreflectors, receiving telescope/detector, atomic clock, round trip time of flight (ToF)



On ground: theodolite (or laser “tracker”), retroreflectors;
ToF of laser pulses

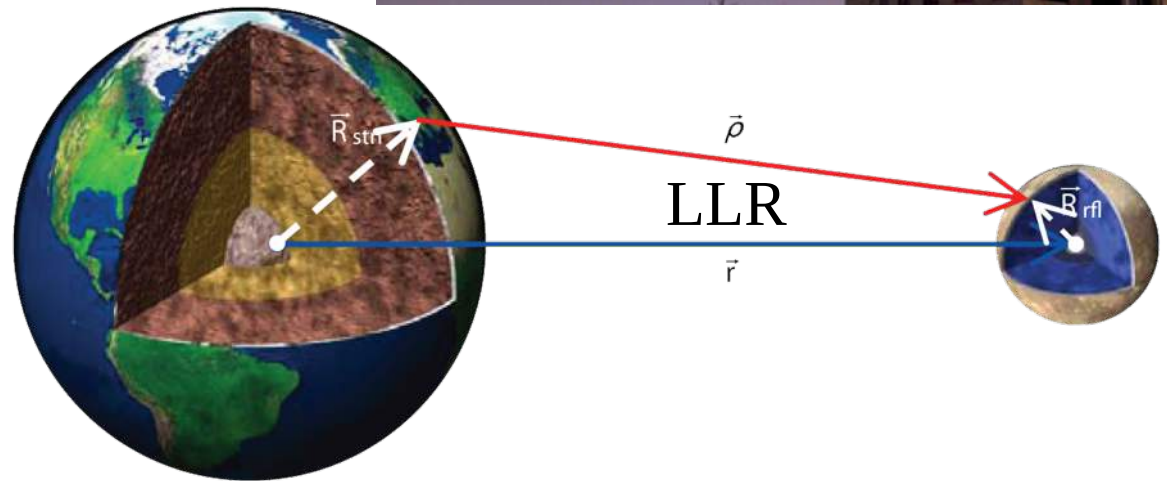
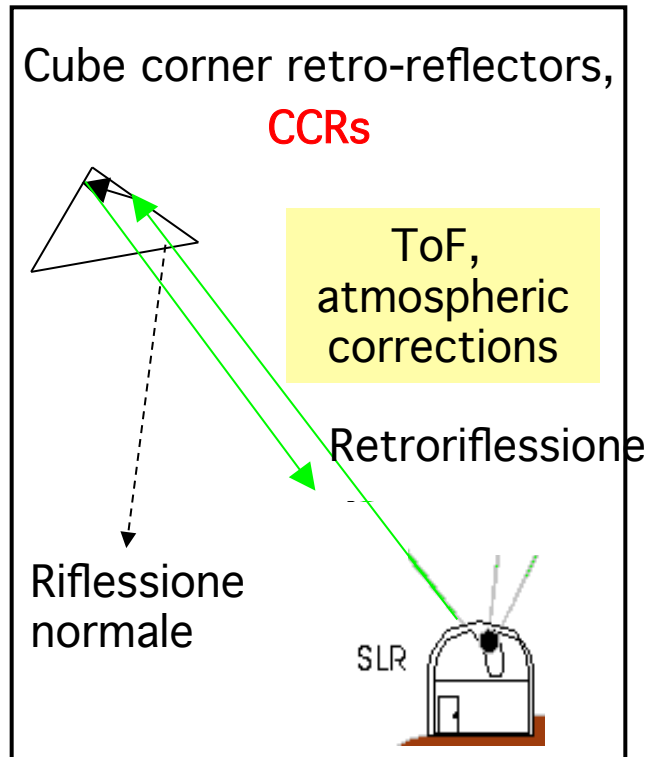
$$d \approx \frac{c\Delta t}{2}$$



Used for positioning metrology of KLOE at DAΦNE,
Here at INFN-LNF

Lunar Laser Ranging (LLR) *Satellite Laser Ranging (SLR)*

ToF measurements

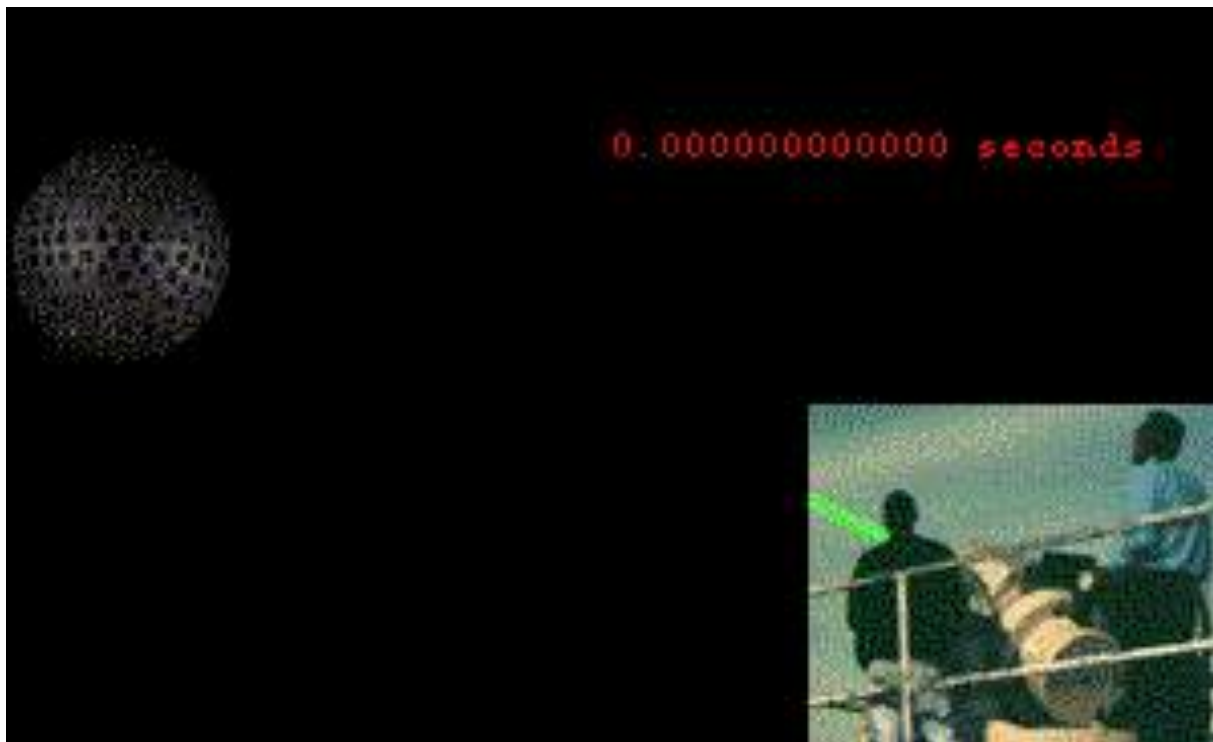


Most accurate space positioning. AND. Absolute (next slides)
(mm to cm) .AND. (100K€ to M€)

Laser interferometry more precise but much more expensive and complex/difficult

- Ground station send laser pulses to satellite
- Telescopio next to laser measures retroreflected pulses
- **LAGEOS satellite spherical satellite, a point-like mass:**
metal => massa, CCRs => laser “traceability”

S
L
R

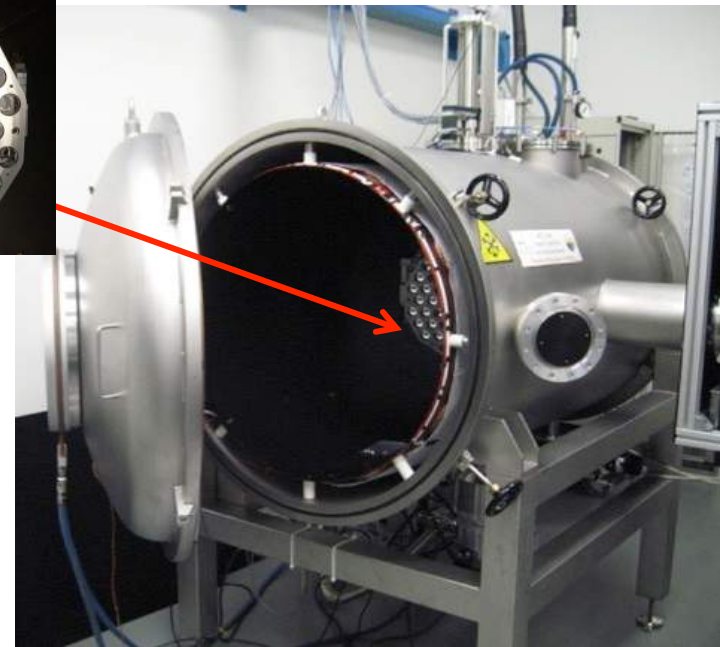
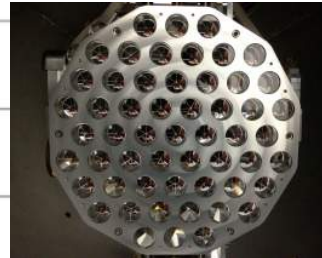
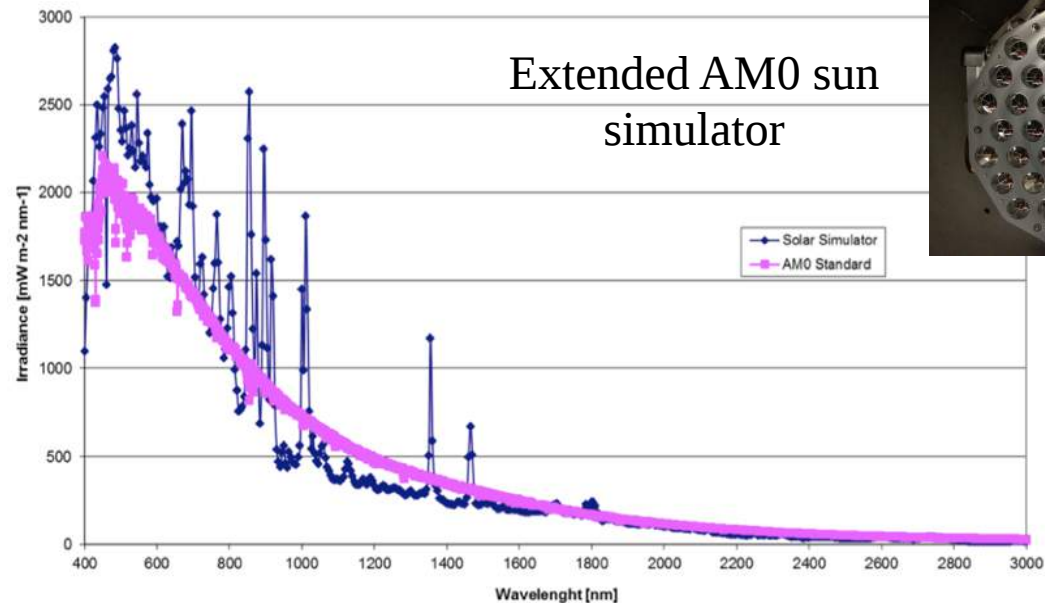




© Dan Long 2014



- **SCF_Lab @ INFN-Frascati**
 - Unique characterization of CCR performance
- **MLRO @ ASI-CGS, Matera Laser Ranging Observatory**
 - SLR/LLR
 - Space Geodesy Center



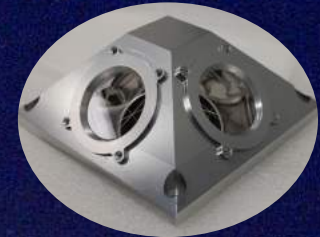
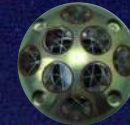
www.lnf.infn.it/esperimenti/etrusco/

SATELLITE/GNSS & LUNAR LASER RANGING (INFN-ASI CCRs)



Matera Laser Ranging
Observatory @ASI – CGS
(1.5 m telescope)

INRRI
 μ -reflector



LEO



Moon

LAGEOS
(model by
NASA)



GNSS

Centro di Geodesia Spaziale *Giuseppe Colombo* Matera, Italy

Un sito la cui posizione è misurata col laser, radio e
navigazione satellitare

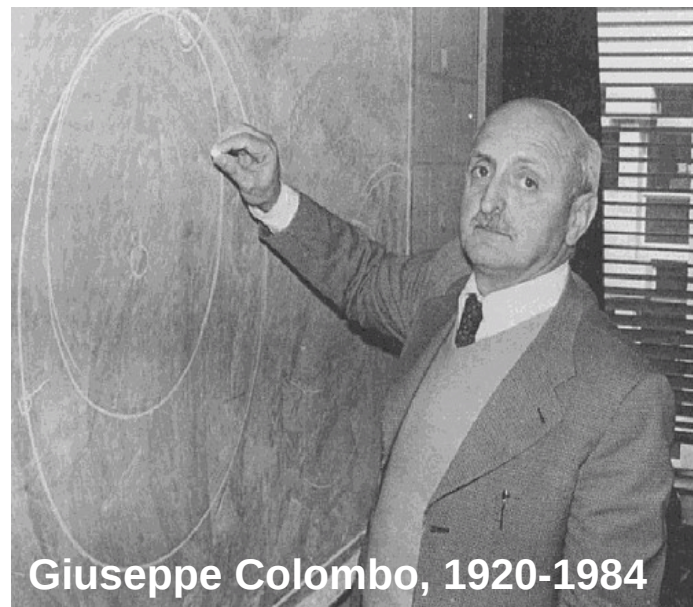
Direttore: Dr. Giuseppe Bianco



MLRO,

Matera Laser Ranging Observatory

Ha fatto il primo LLR il 25 Marzo 2010

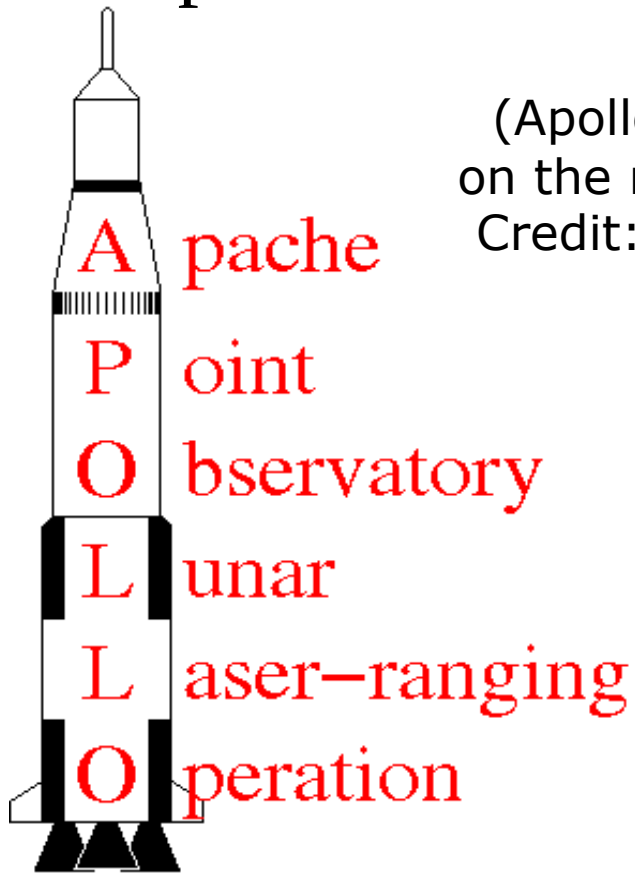


Giuseppe Colombo, 1920-1984

mm precision / orbit segment (normal point)

New Mexico, USA

(Apollo) Laser beams are sent to reflectors on the moon from a telescope in New Mexico.
Credit: Dan Long, Apache Point Observatory



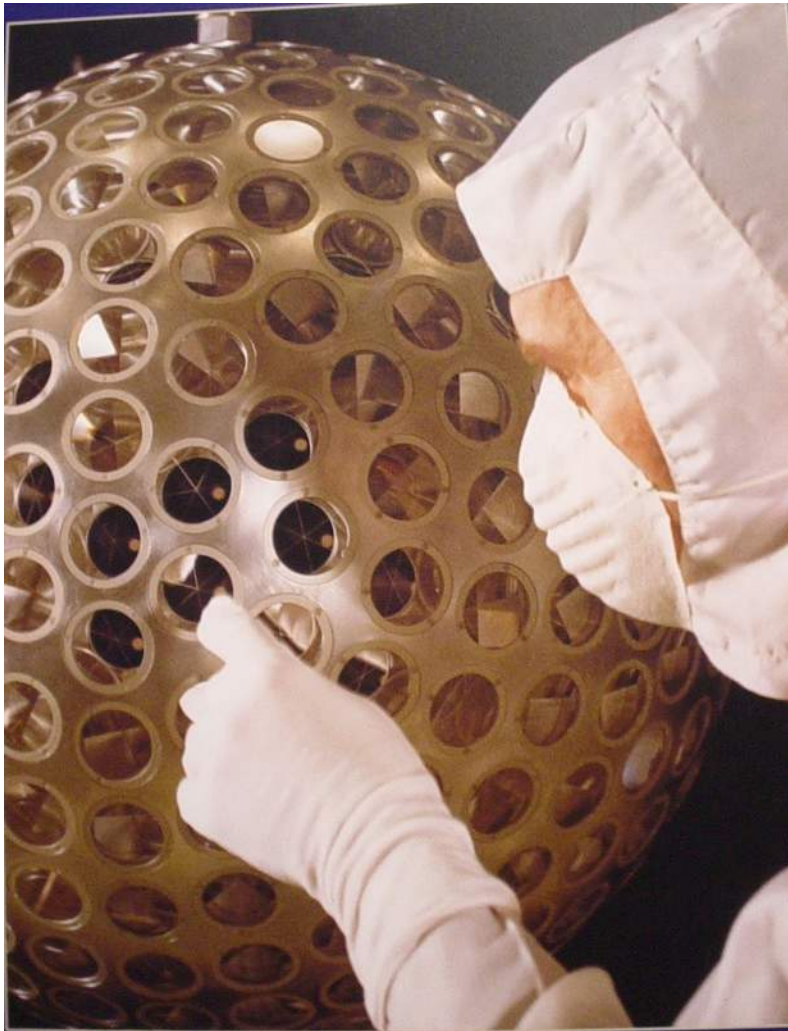
Leader: Tom Murphy,
Univ. of California at San Diego

3 stations in France, OCA-CERGA, Obs. du Calern

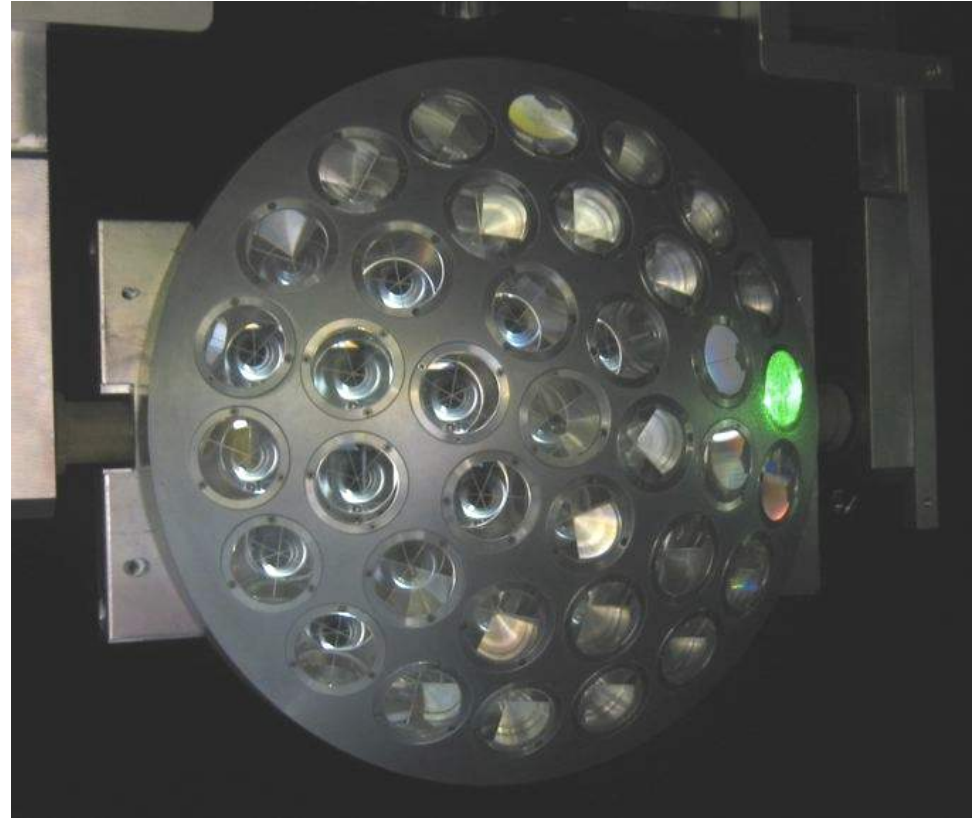


Laser Geodynamics Satellites (LAGEOS)

LAGEOS I (1976; NASA), LAGEOS II (1992; NASA/ASI)

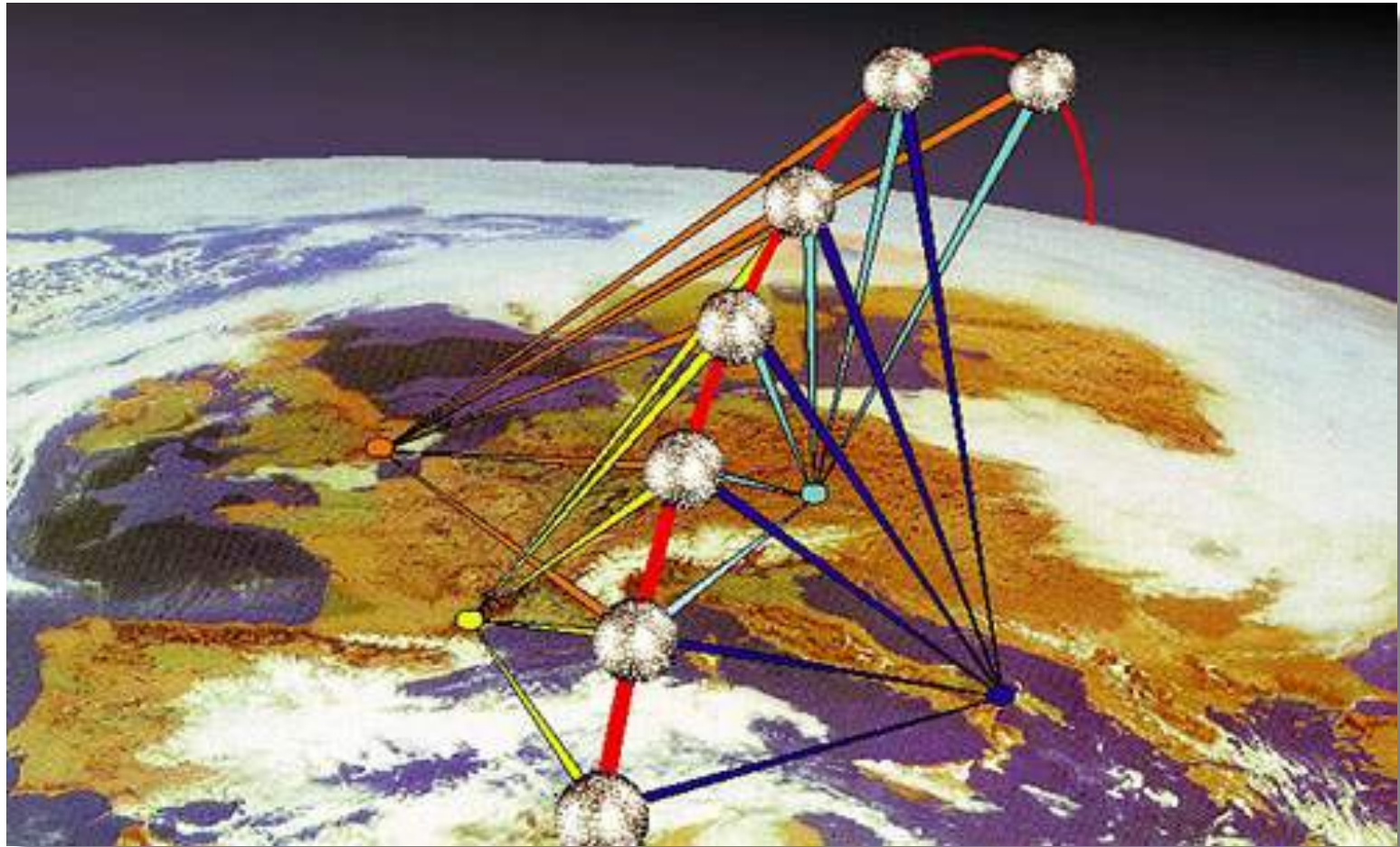


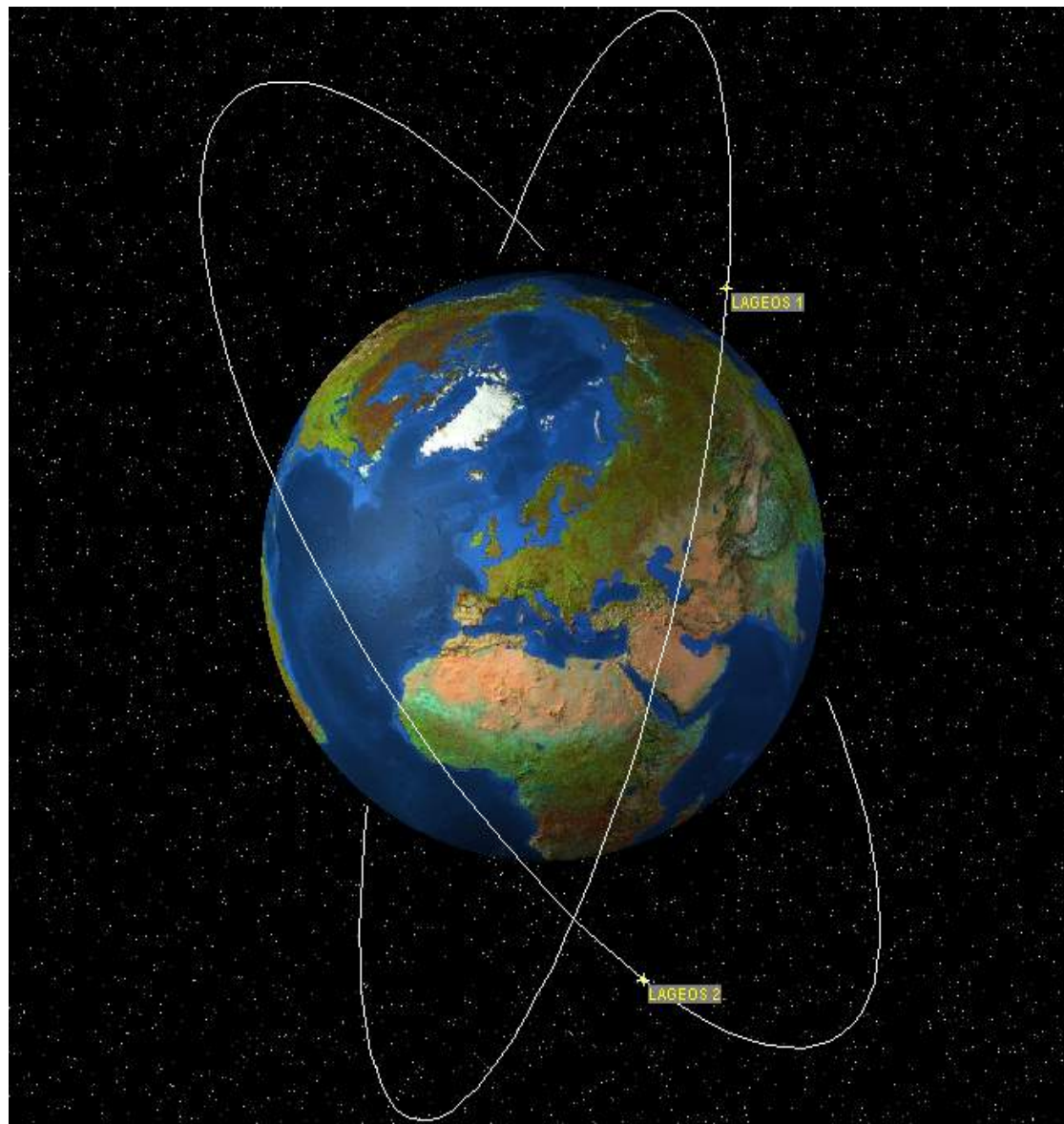
Laser Geodynamic Satellite Experiment (LAGEOS)
LAGEOS satellites reflect laser beams transmitted from ground stations back to sensors on Earth. The first



LAGEOS “Sector”, EM of NASA-GSFC,
at LNF for space tests

LAGEOS tracked by stations at Matera (IT), Herstmonceux (UK), Graz (AT), OCR (FR)

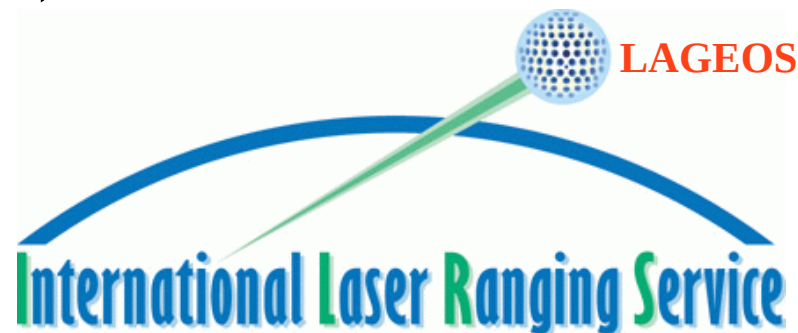




International Terrestrial Reference System (ITRS)

Cartesian frame:

- Origin = Geocenter = centro of mass of the Earth = focii of LAGEOS orbits
- Meter, or scale of length = semi-axis of LAGEOS orbit
- Angles, or orientation of axes = radio-measurement of quasars (“fixed stars”)
- Cartesian coordinates of laser stations
- Distribution of the frame: GPS, Galileo ...





ILRS = SLR/geodesy international Service

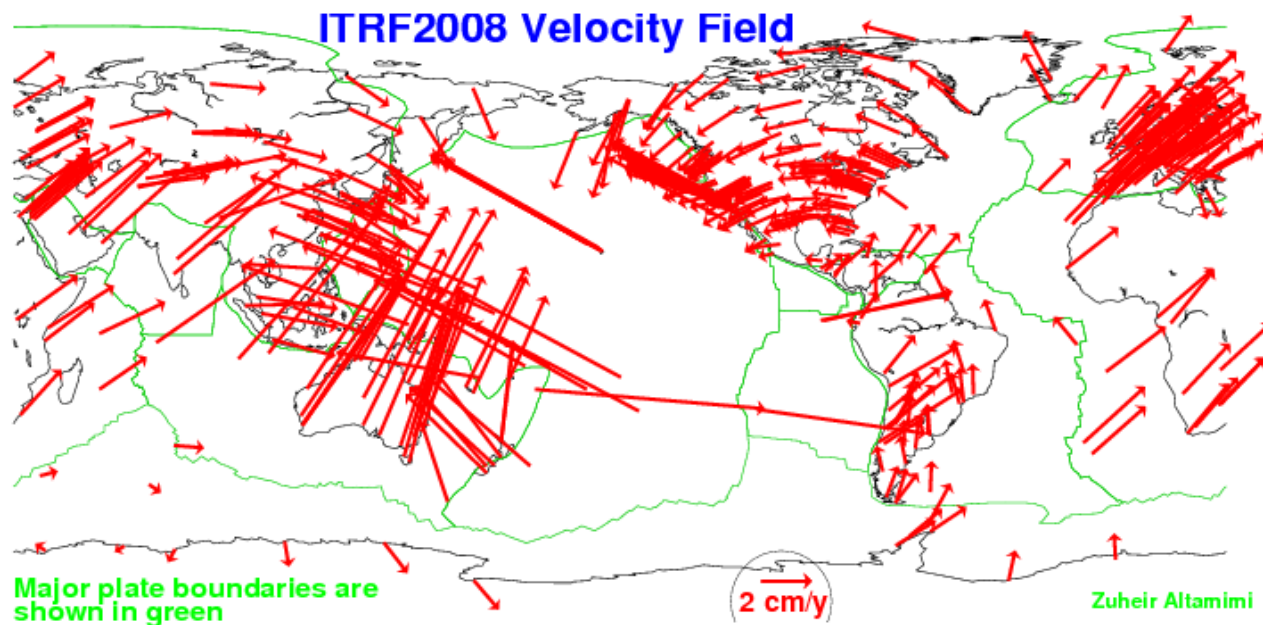
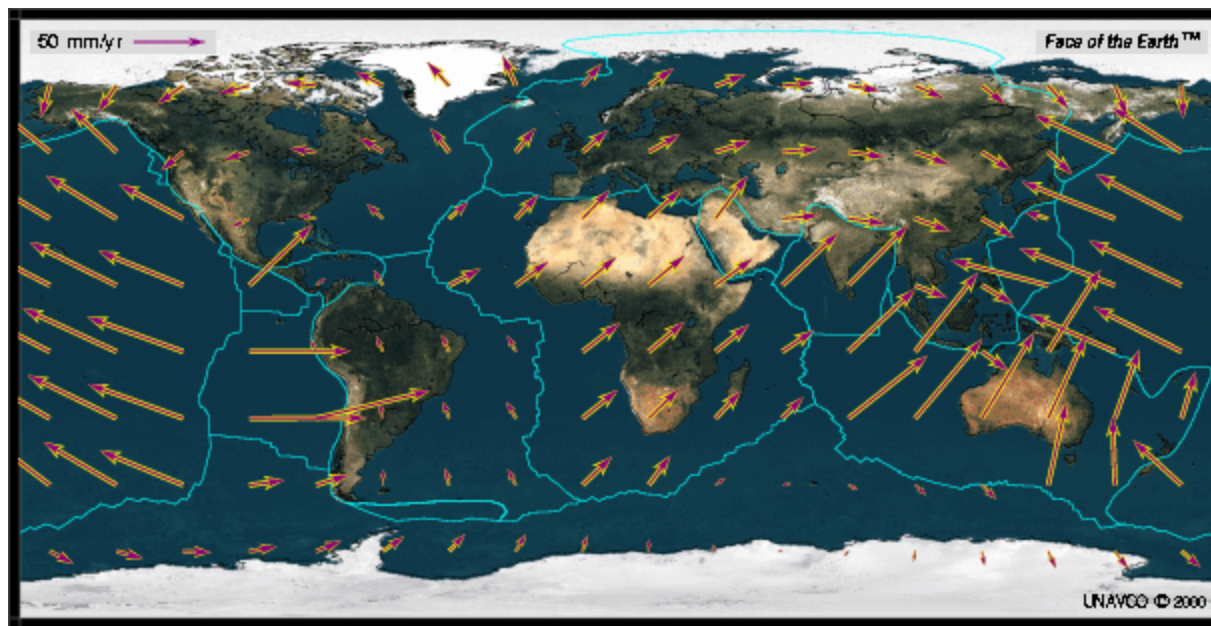


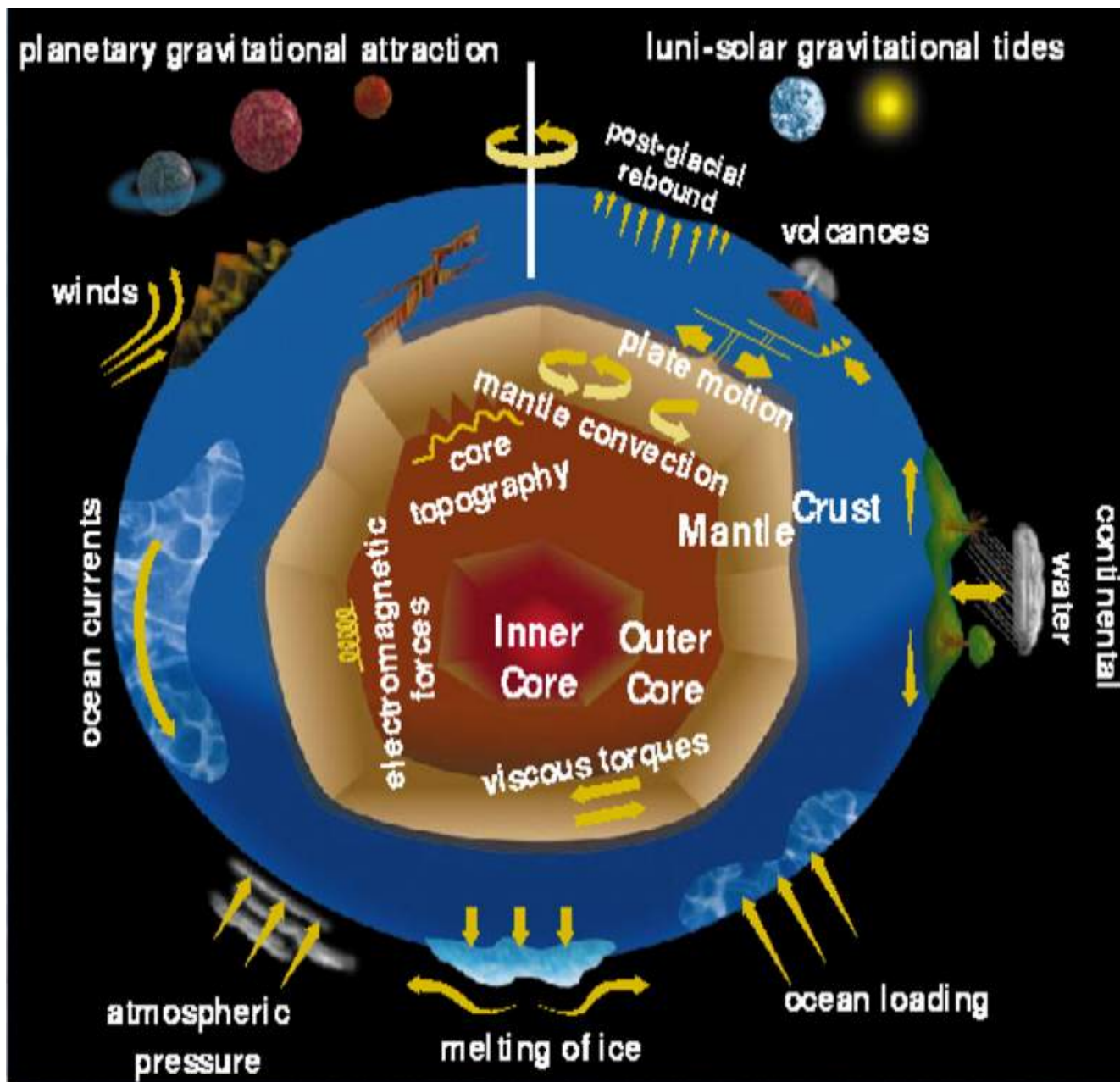
LAGEOS



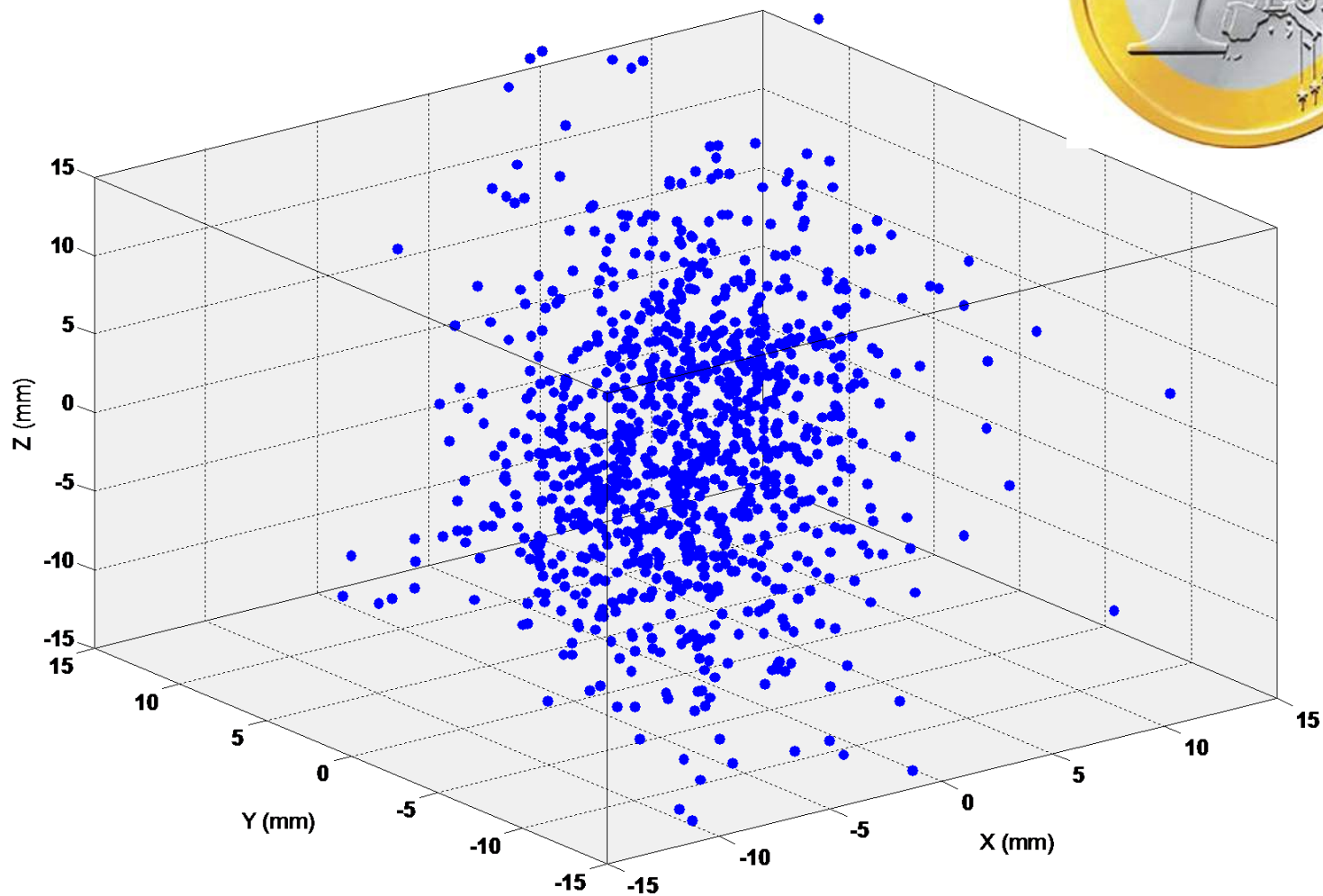
Geodynamics
Plate tectonics

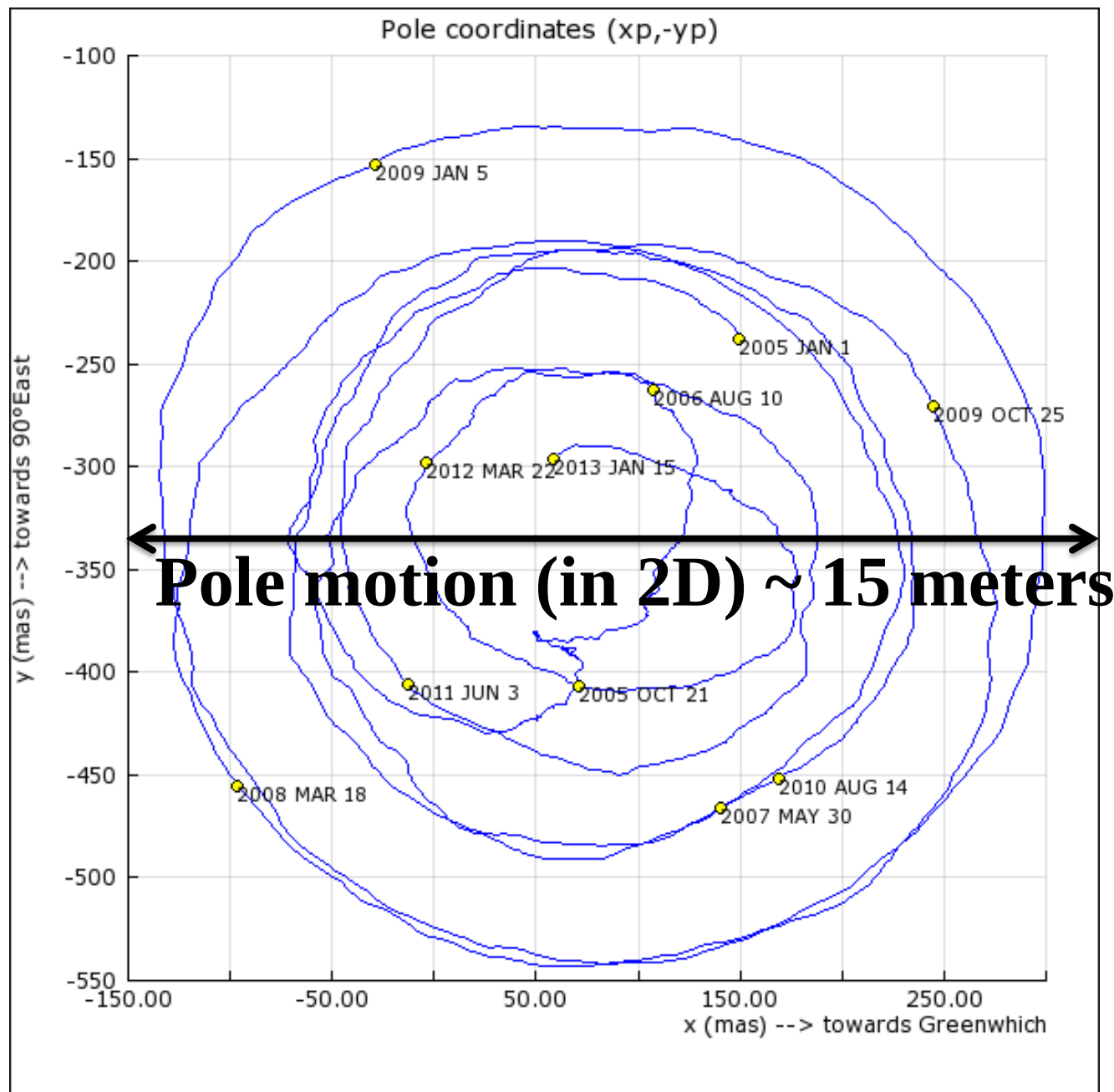




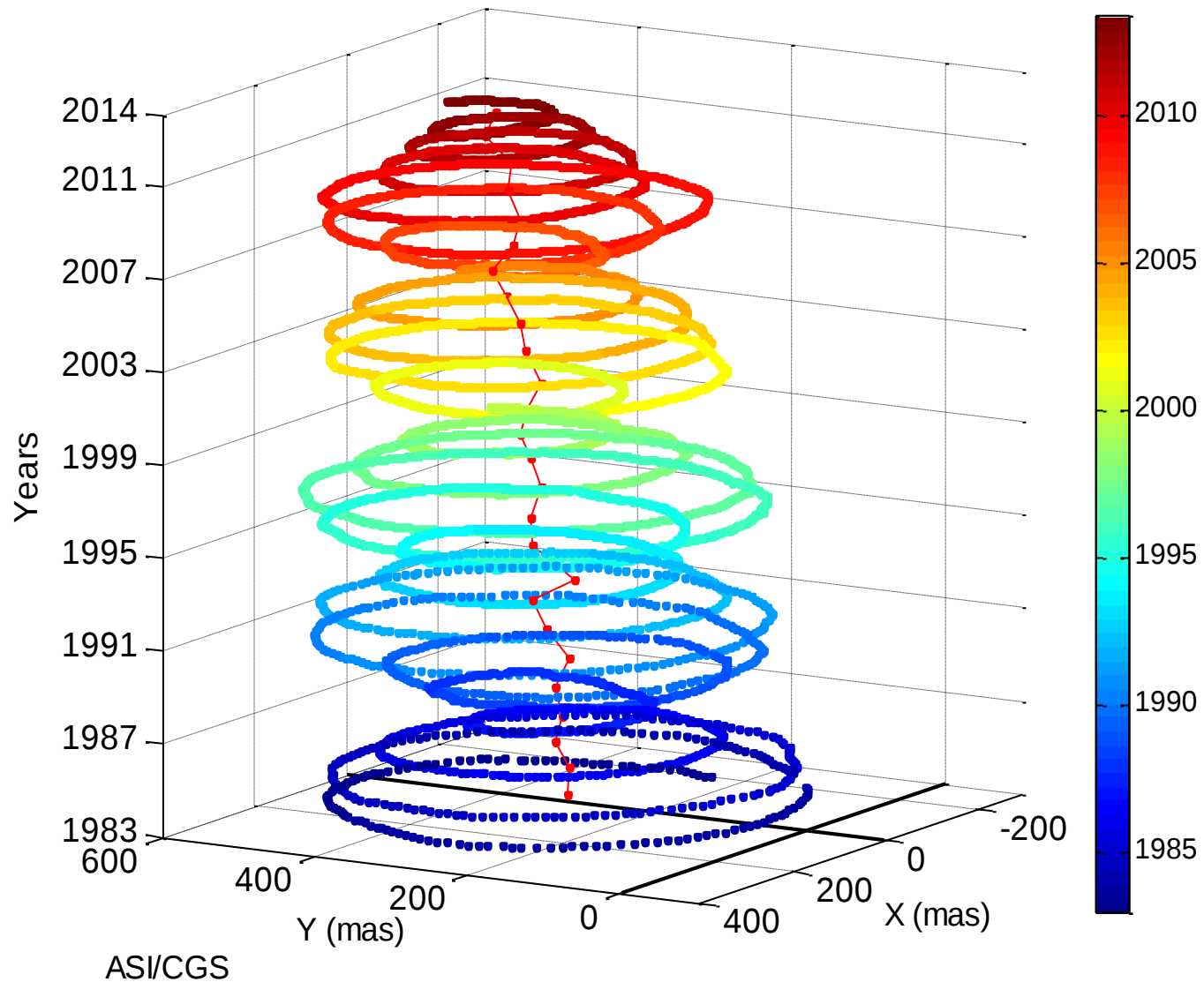


ILRSA 1993-2012 weekly geocenter vs ITRF2008 origin





Pole motion in 3D (tens of meters)

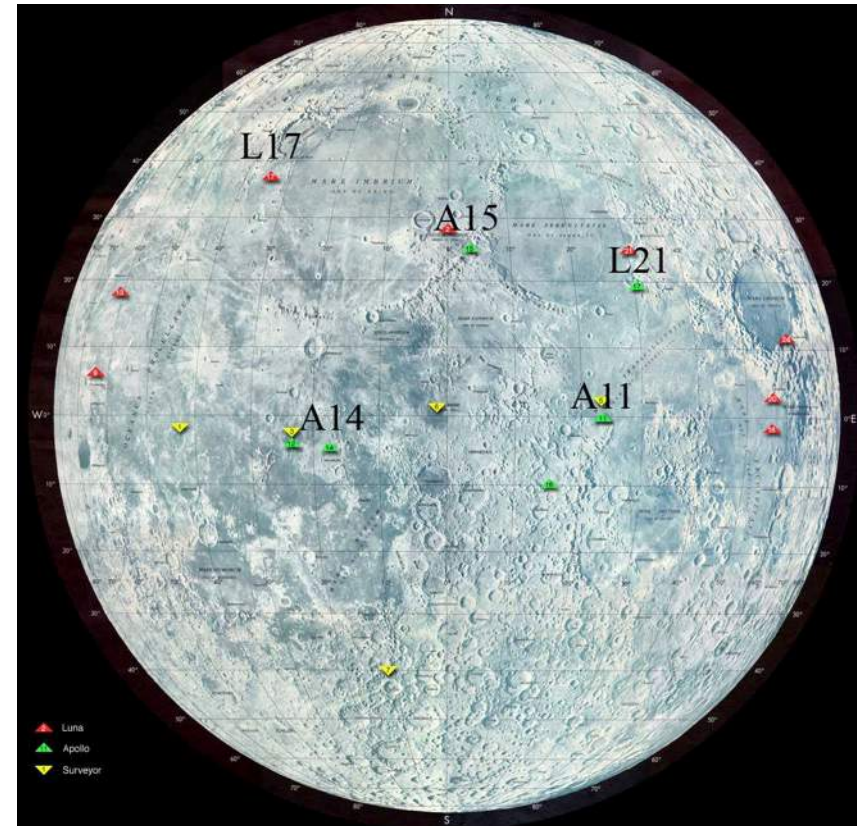




All started with LLR



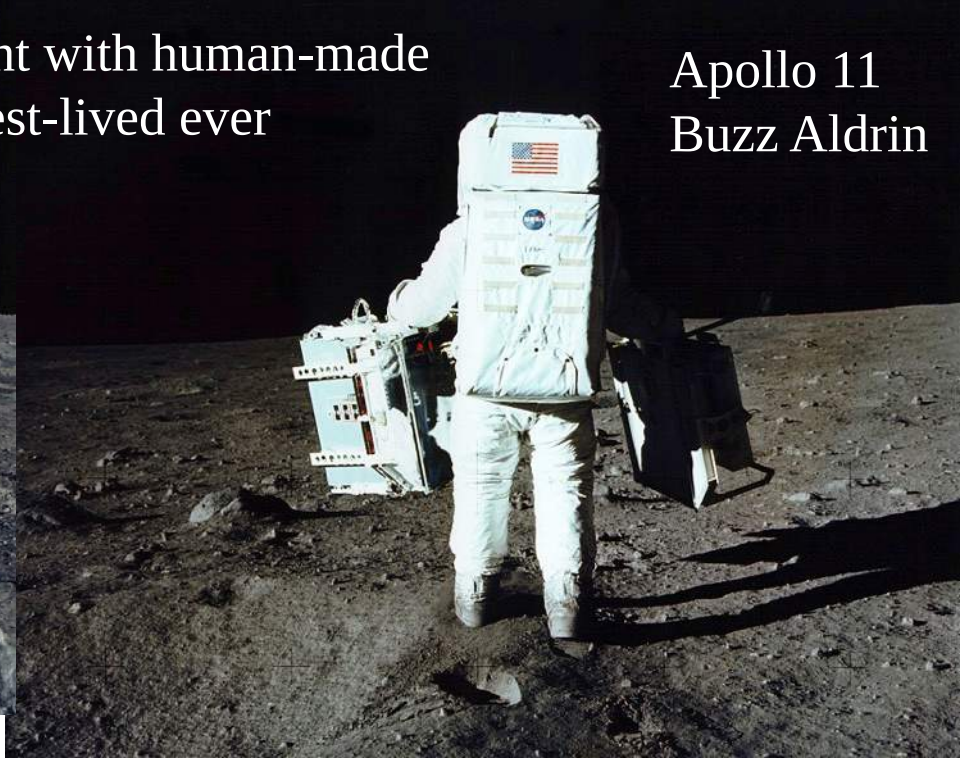
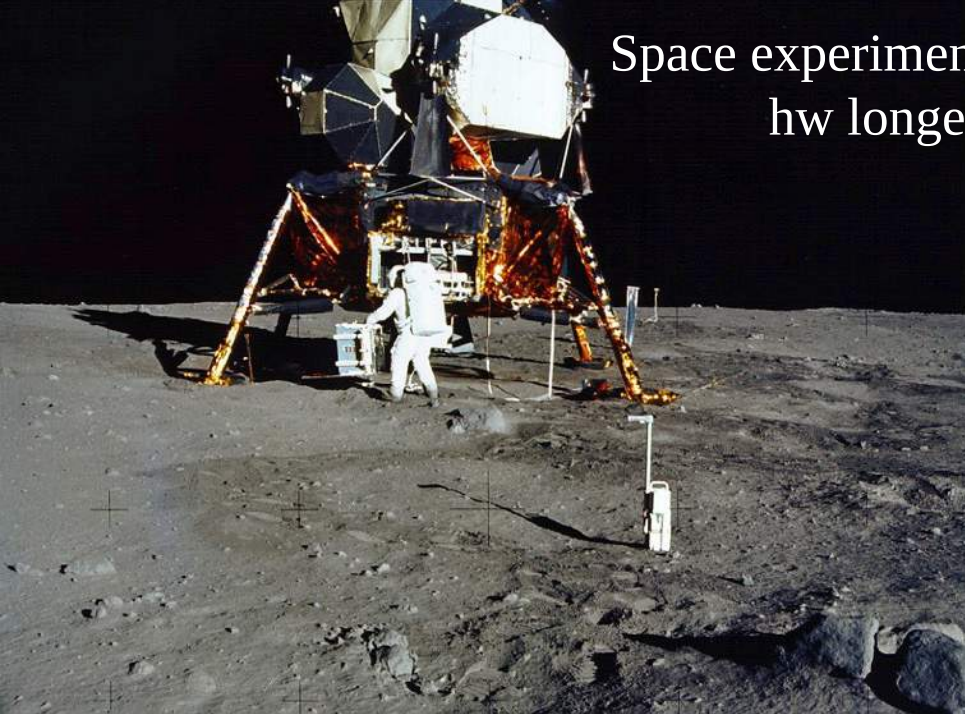
- R. H. Dicke et: precision tracking of satellites using conventional light pulses to test gravity in the 50s
- Laser invented ~1960
- MIT & URSS shoot lasers on lunar surface in the 60s
- CCRs deployed by Apollo missions 11, 14, 15



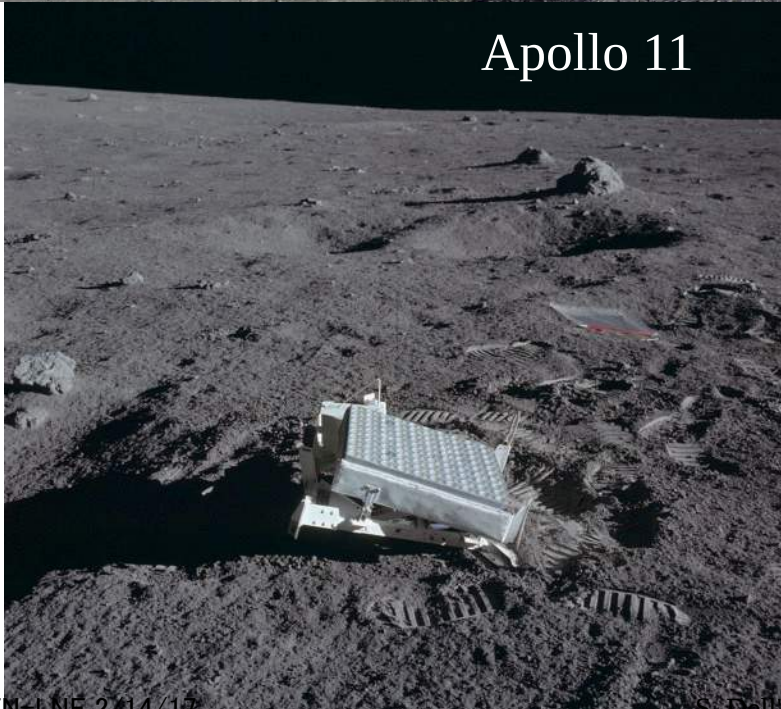
- Apollo: arrays of CCRs of fused silica, circular aperture 3.8 cm each
- Apollo 11 e 14: used 100 CCR
- Apollo 15: 300 CCR

Space experiment with human-made
hw longest-lived ever

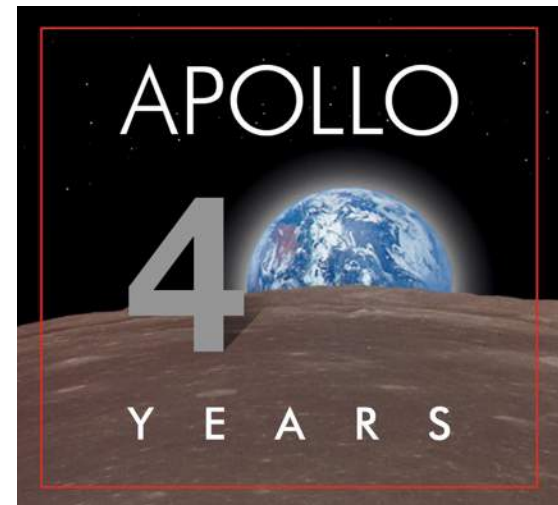
Apollo 11
Buzz Aldrin



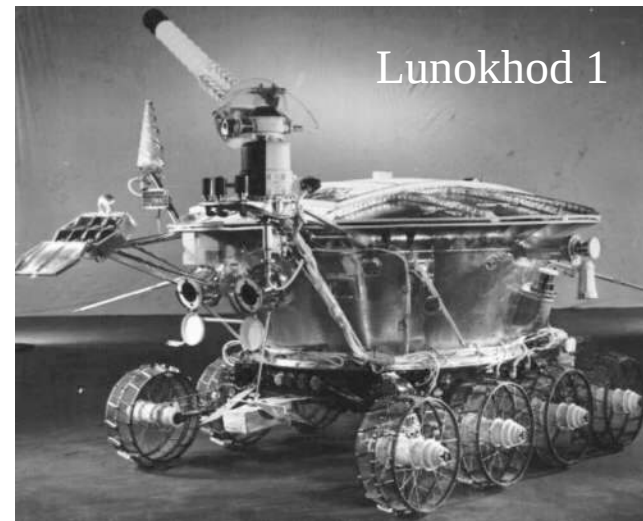
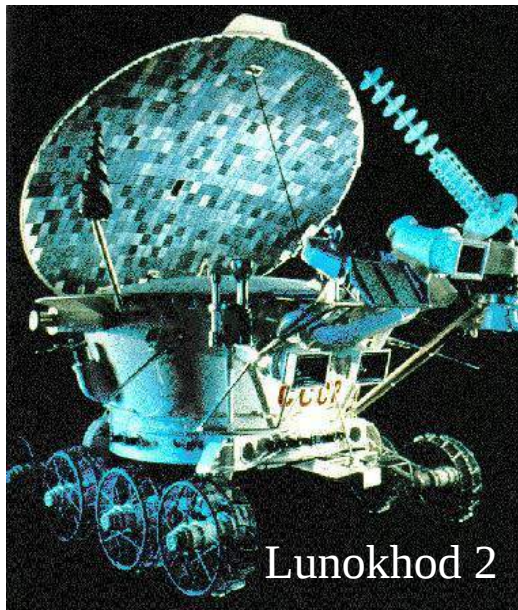
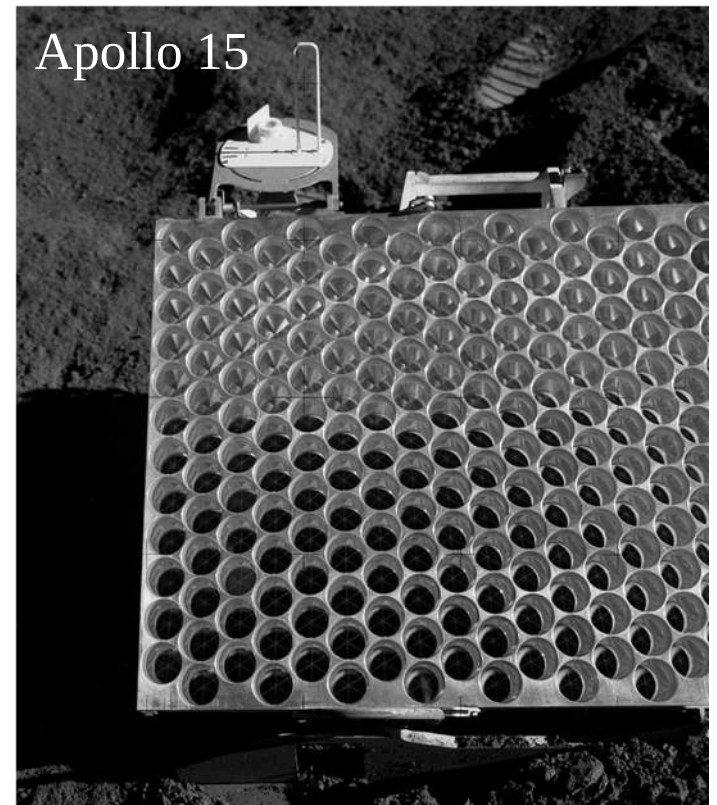
Apollo 11



LLR: only Apollo experiment still giving
data. 2009 logo



1st generation CCRs

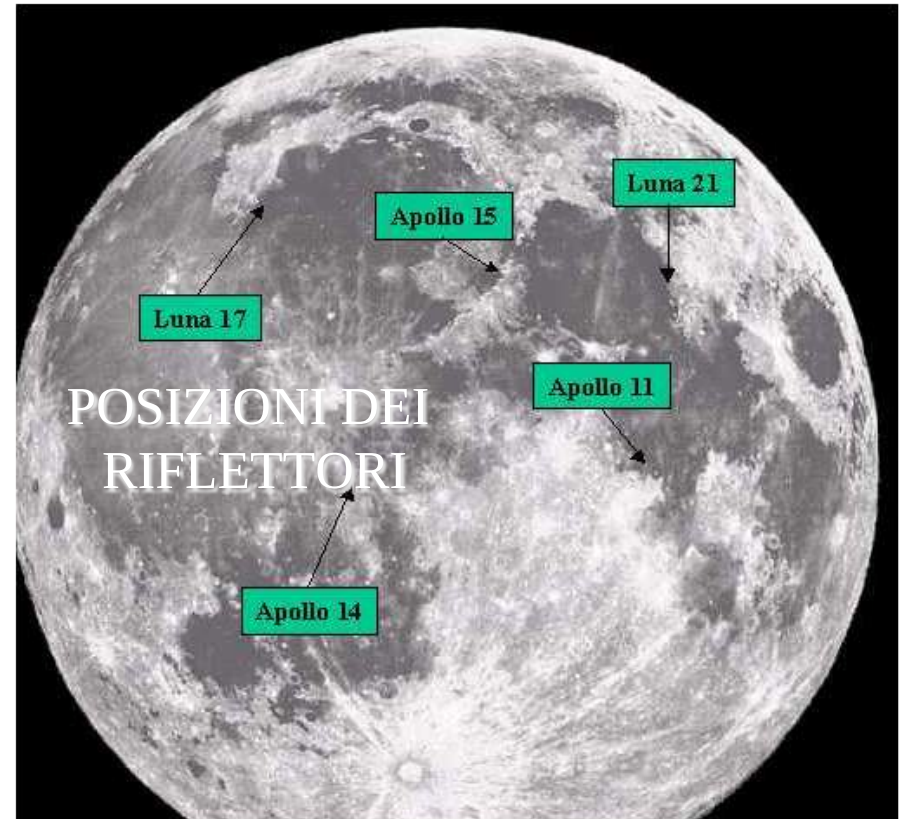


Precision: ~ 2 cm $\sim 5 \times 10^{-11}$ of Earth-Moon distance

GIF: relative sizes, distances, times

ToF ~ 2.6 sec

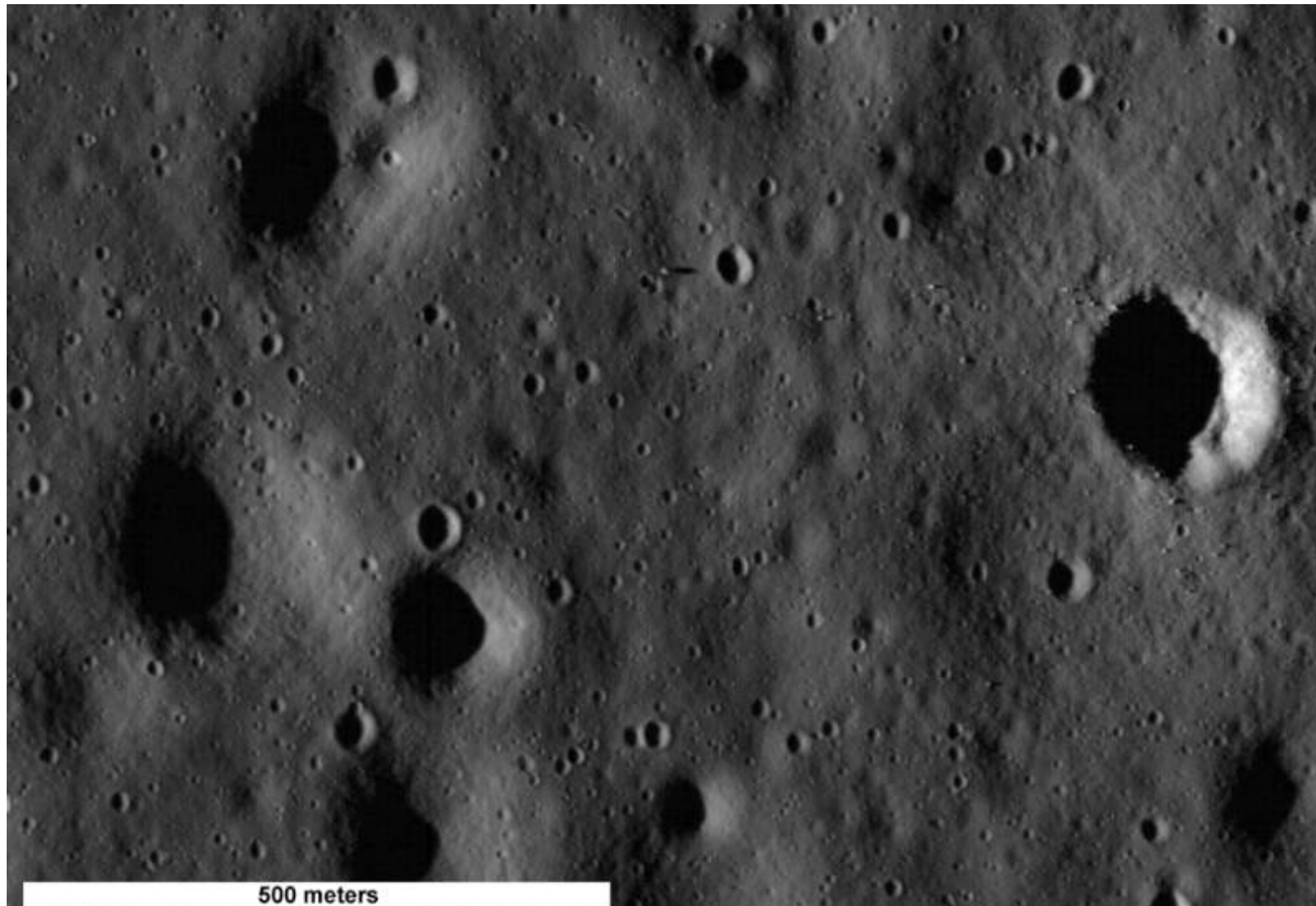
Distance $\sim 384,000$ km



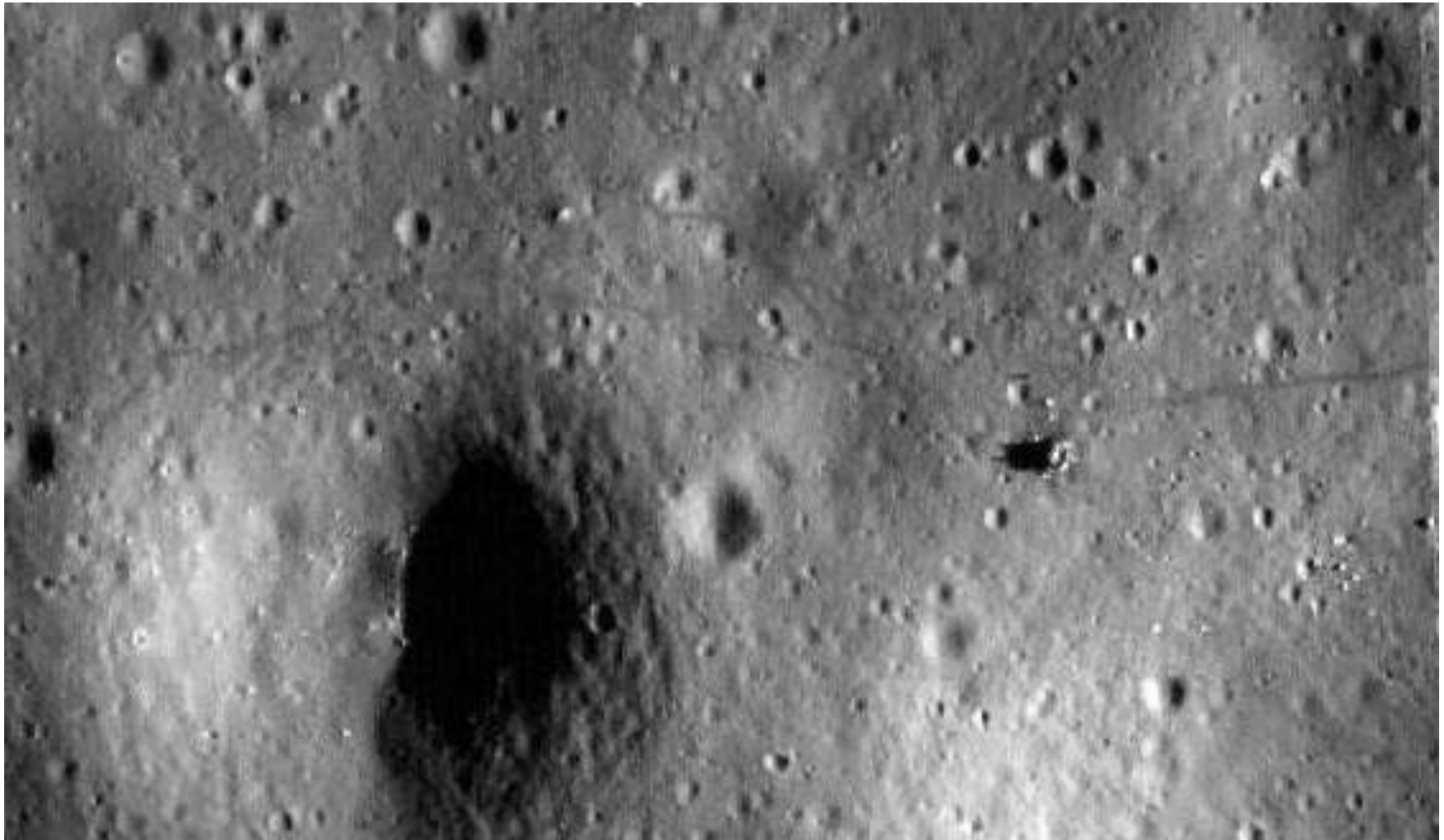
Humanity landed on the Moon?

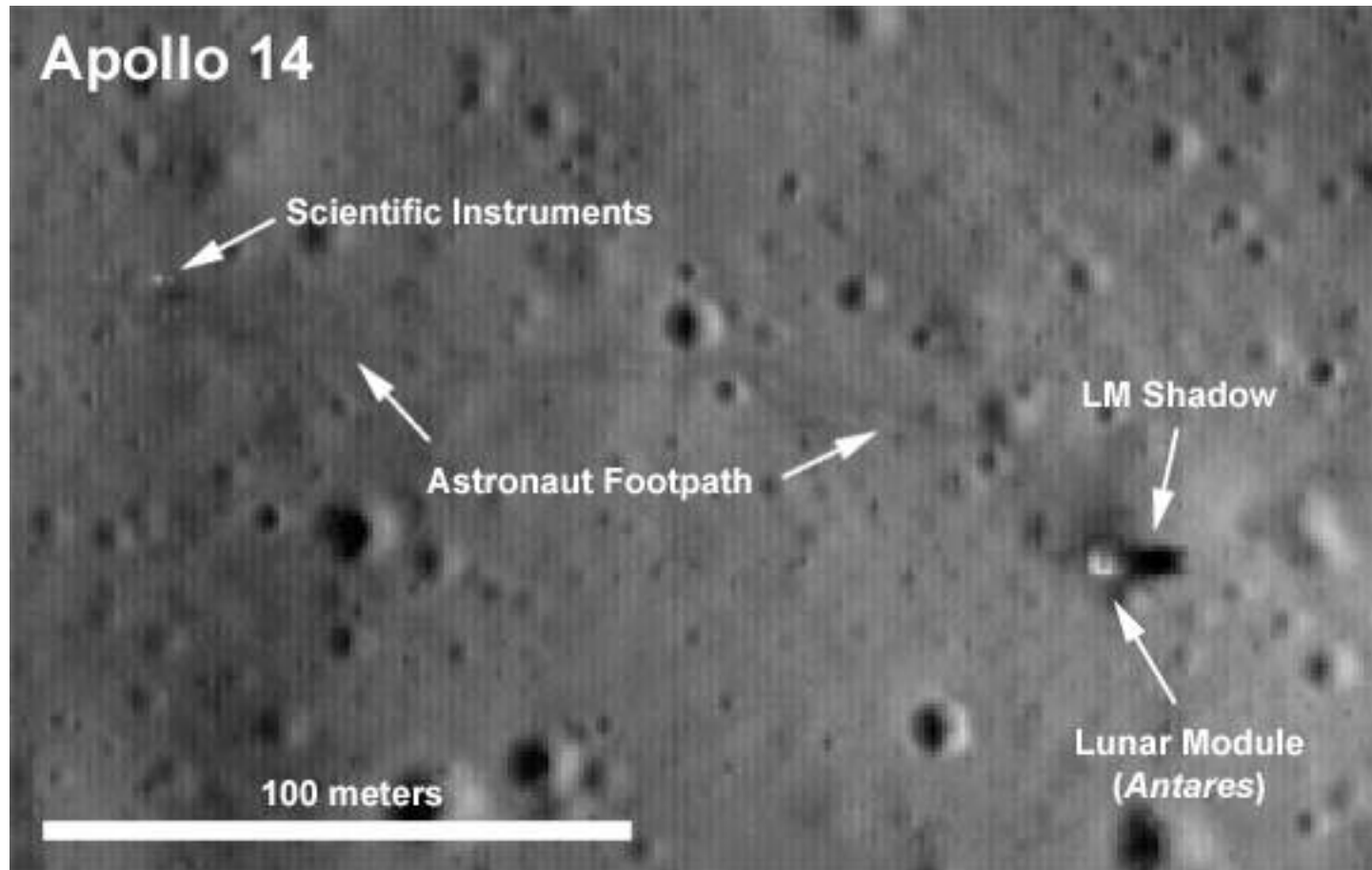
Witnessed since ~48 years by lunar CCRs. Not enough?

Apollo 11 seen by LROC, digital camera of Lunar Reconnaissance Orbiter (NASA)



Apollo 12, Lunar Reconnaissance Orbiter Narrow Angle Camera (LROC NAC)



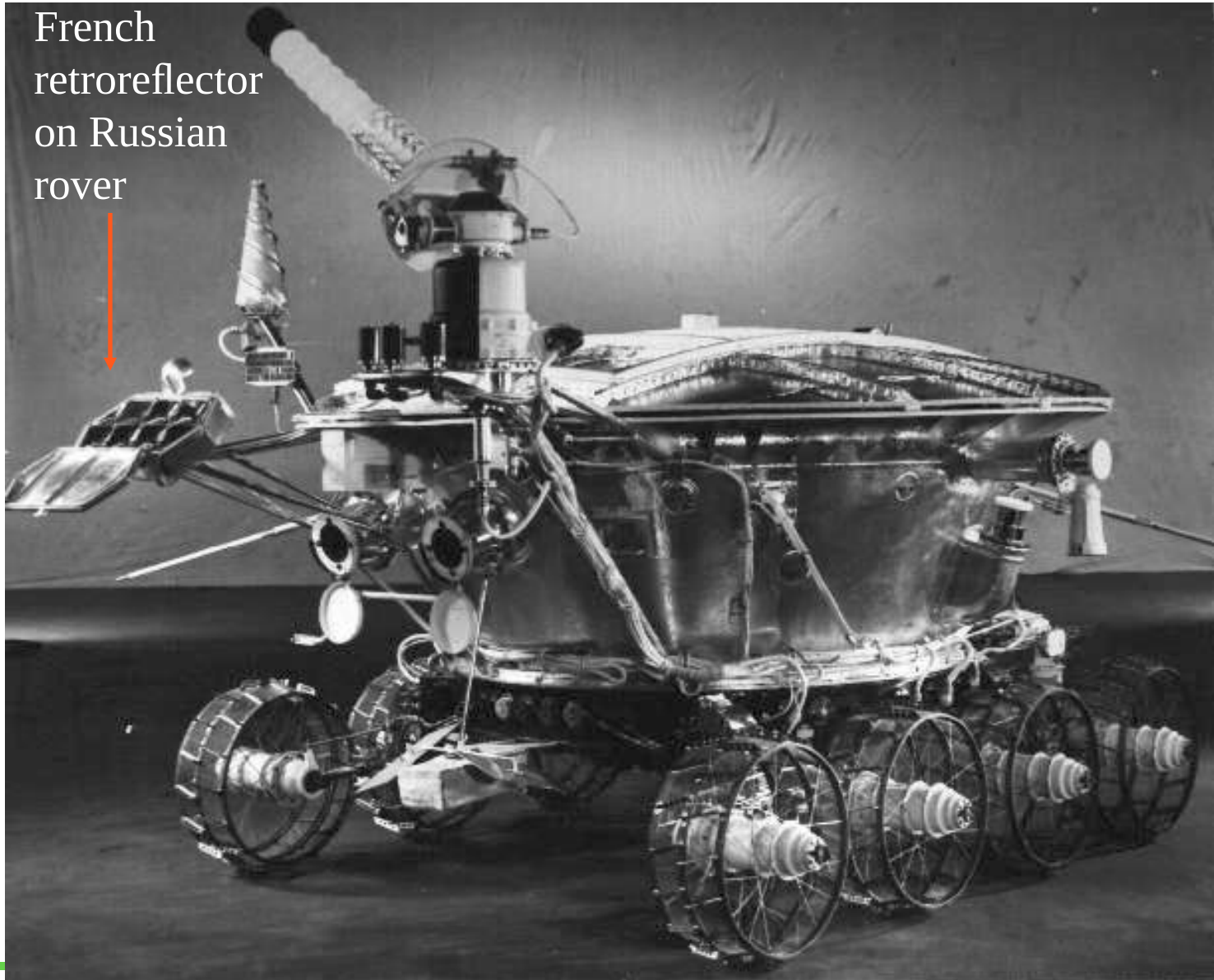


About half million people worked on the Apollo program.

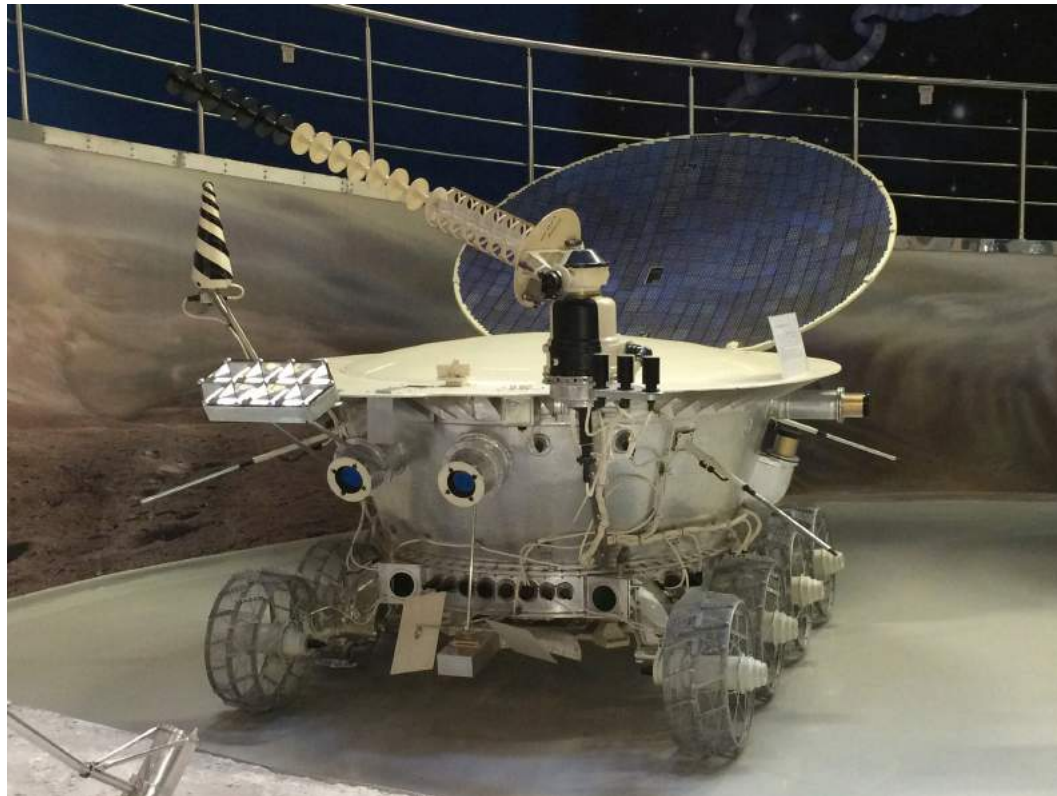
Neil Armstrong: **“it would have been more difficult to fake it than to do it”**

Rover Lunakhod 1, rediscovered by LRO

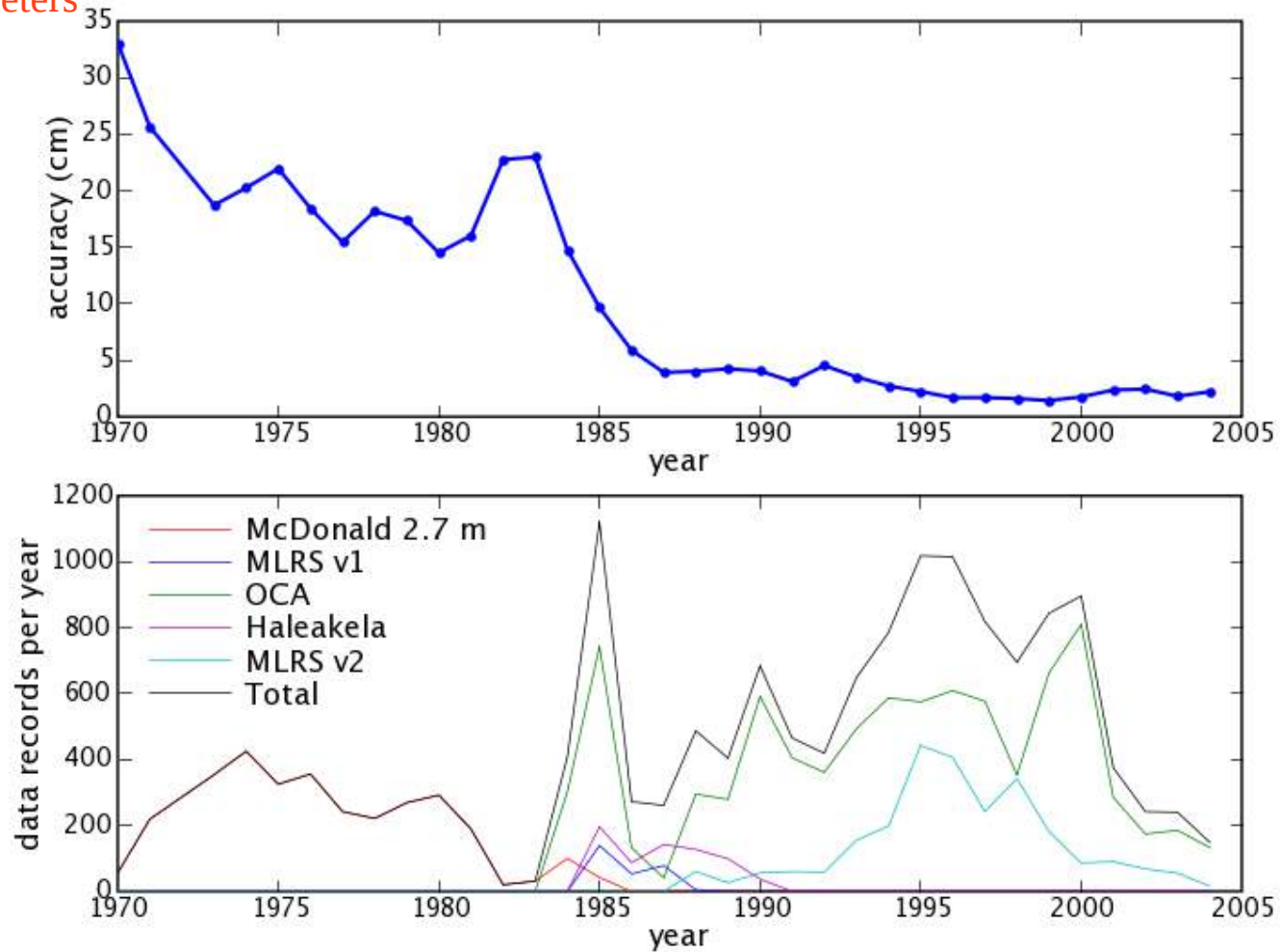
French
retroreflector
on Russian
rover





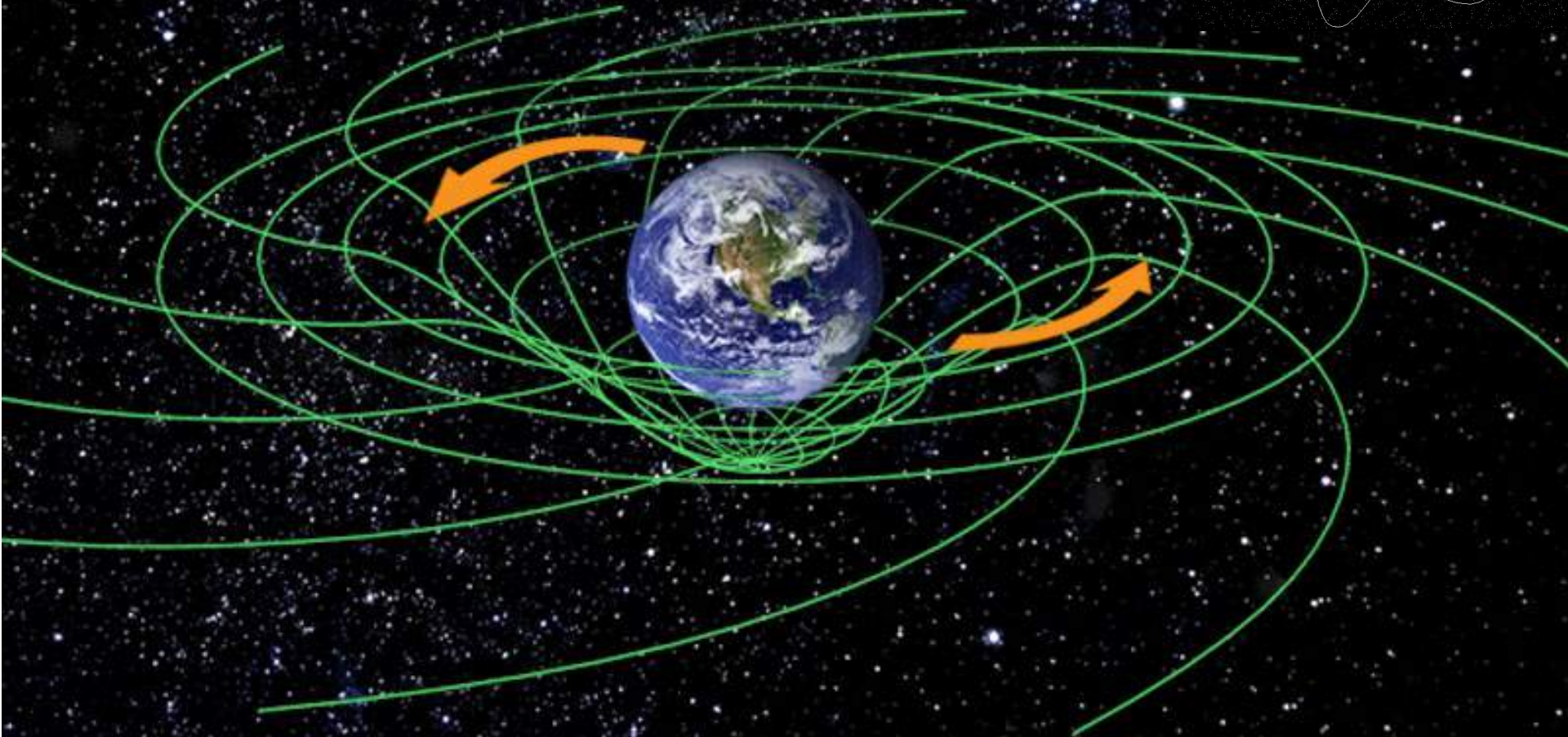
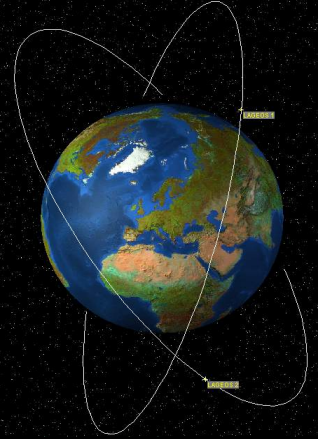


Before:
Hundreds of meters

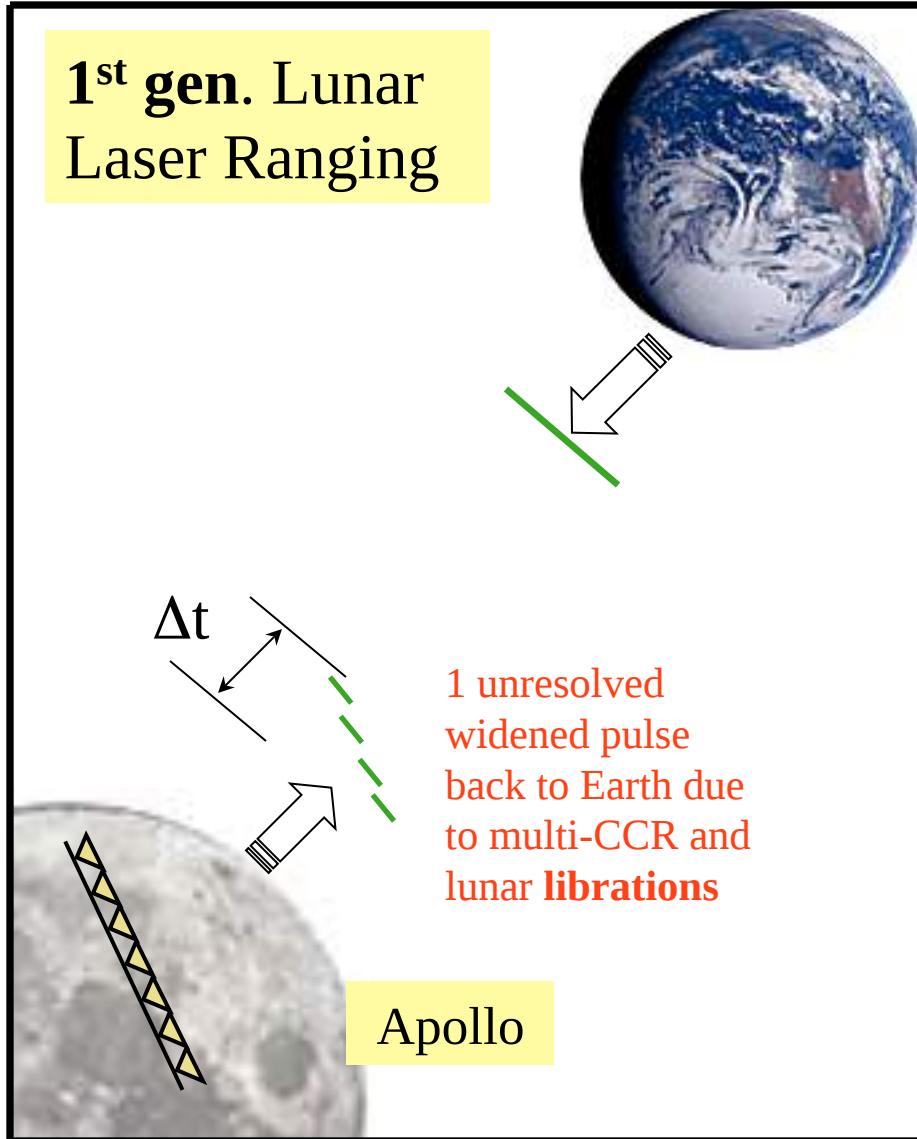


- General relativity (and beyond)
- Technology: SLR/LLR to LAGEOS/Moon
- Examples
 - Precision measurements, but also
 - Tests of violations of
 - Principles (equivalence) and laws ($1/r^2$) of our mother Natura (*de rerum Natura*)
- One example: our LAGEOS brothers
- Three examples: our sistem the Moon

LAGEOS orbits precess around Earth (also) due to the space-time dragging effect related to the mass current given by the Earth spinning around its axis (gravitomagnetic effect, foreseen by general relativity)



1st gen. Lunar Laser Ranging



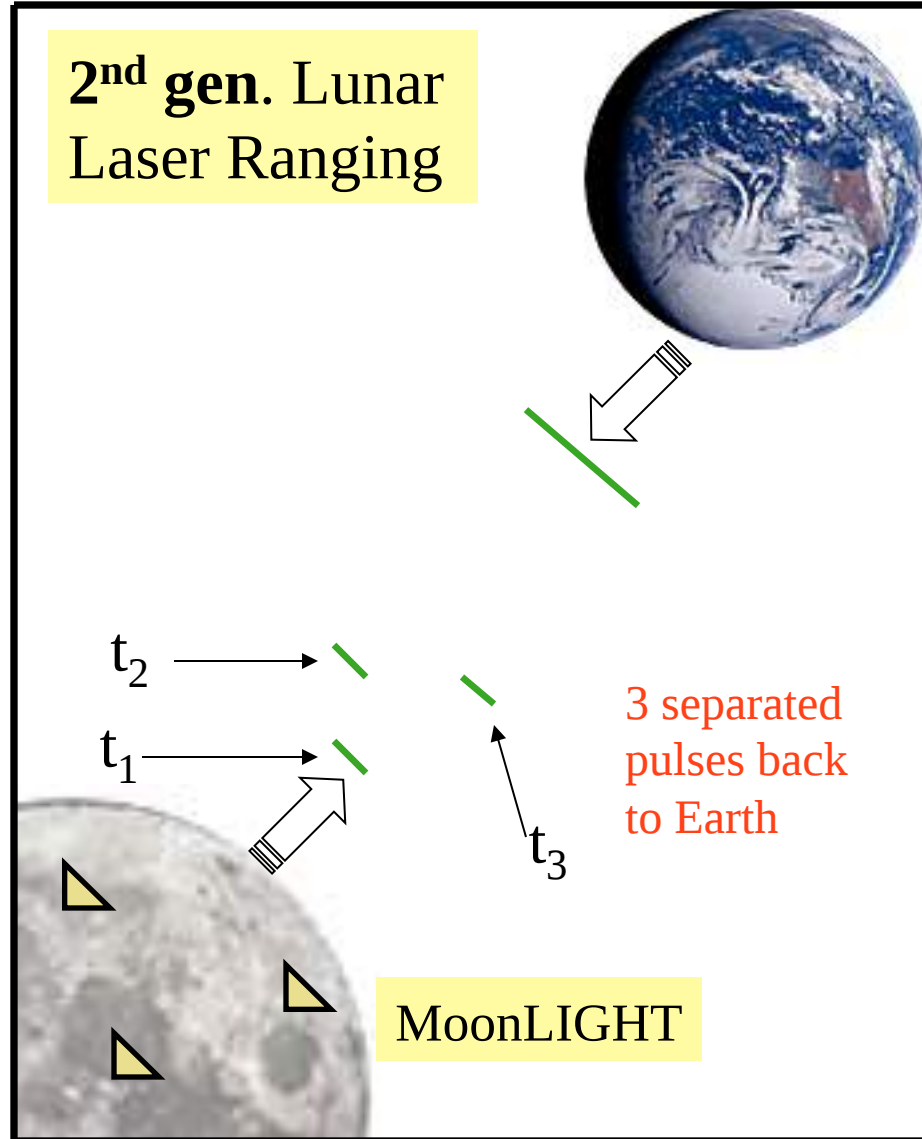
Apollo

1 unresolved widened pulse back to Earth due to multi-CCR and lunar **librations**

Δt



2nd gen. Lunar Laser Ranging



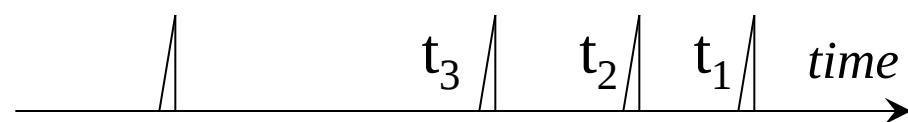
MoonLIGHT

3 separated pulses back to Earth

t_2

t_1

t_3



Pulse to Moon **Pulses to Earth**

Italy-USA next-generation CCR

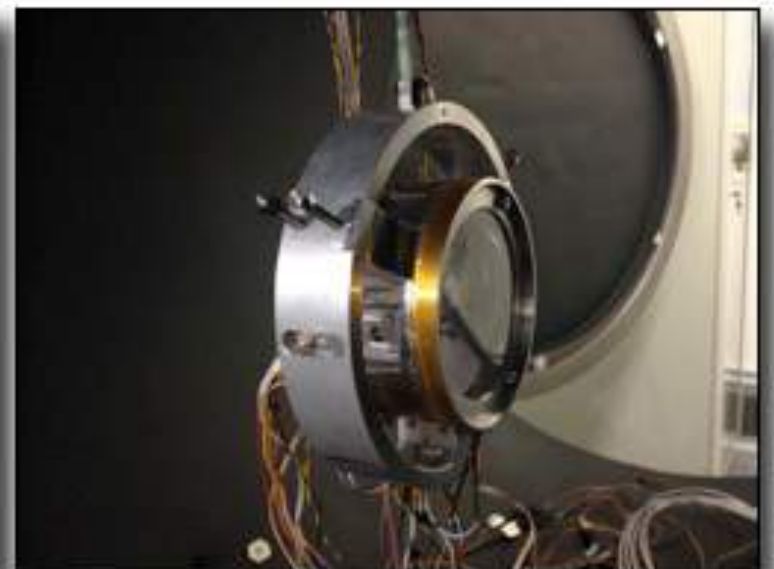
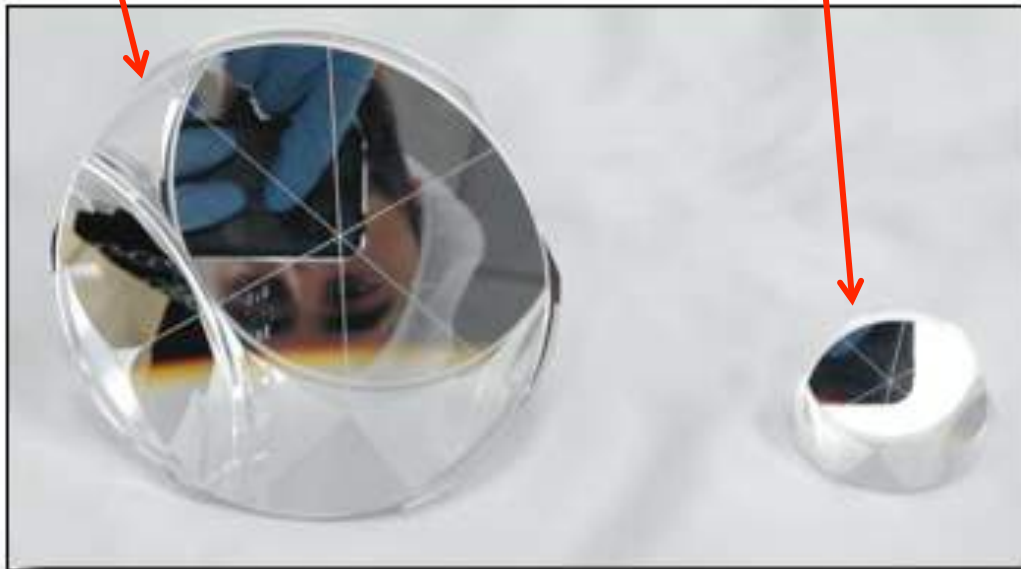


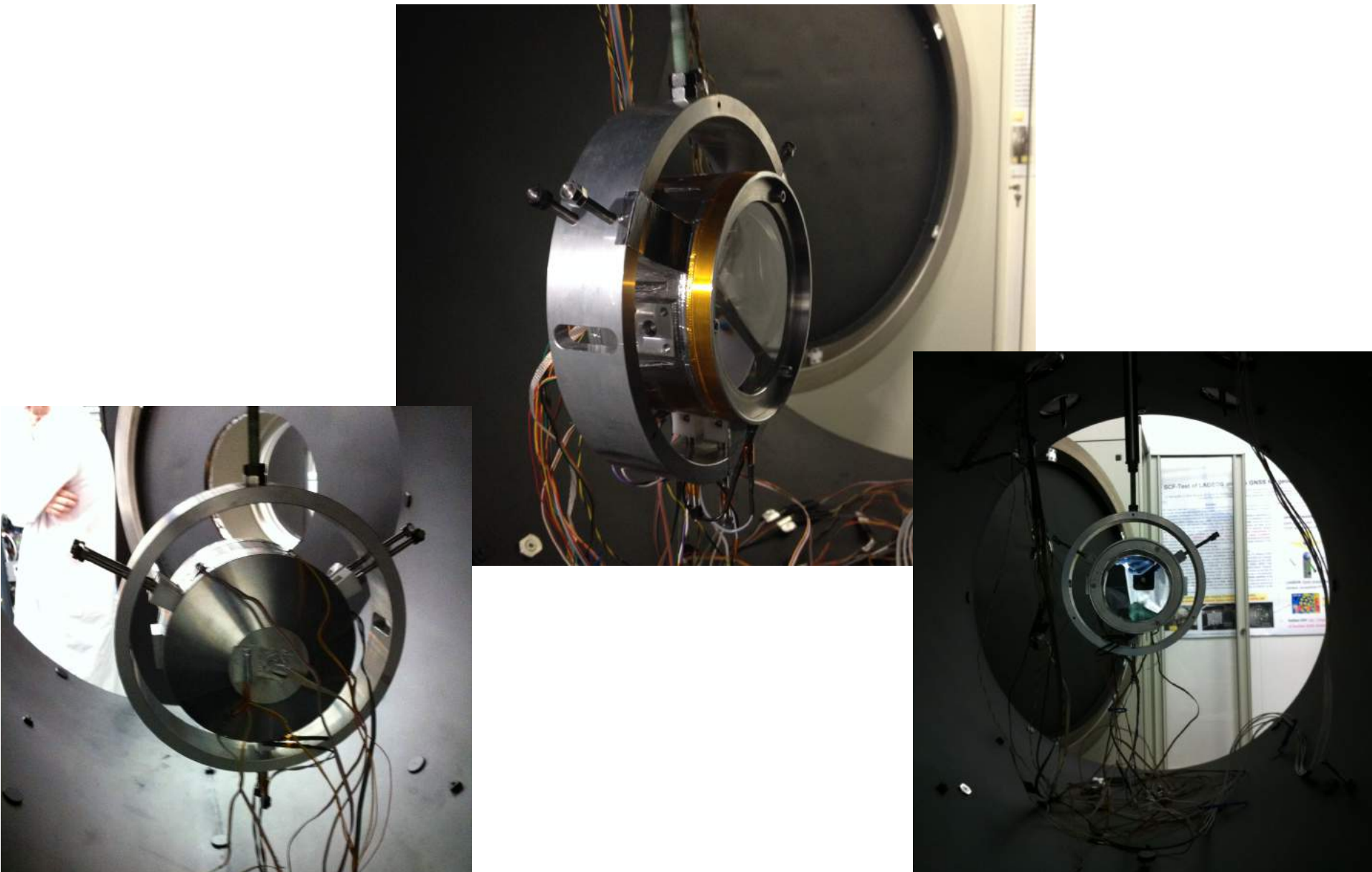
- Collaboration
 - INFN-Frascati, U. Maryland, INFN/Univ. Padova
 - Lunar stations: ASI-MLRO (Italy), APOLLO (US), OCR (Fr.)

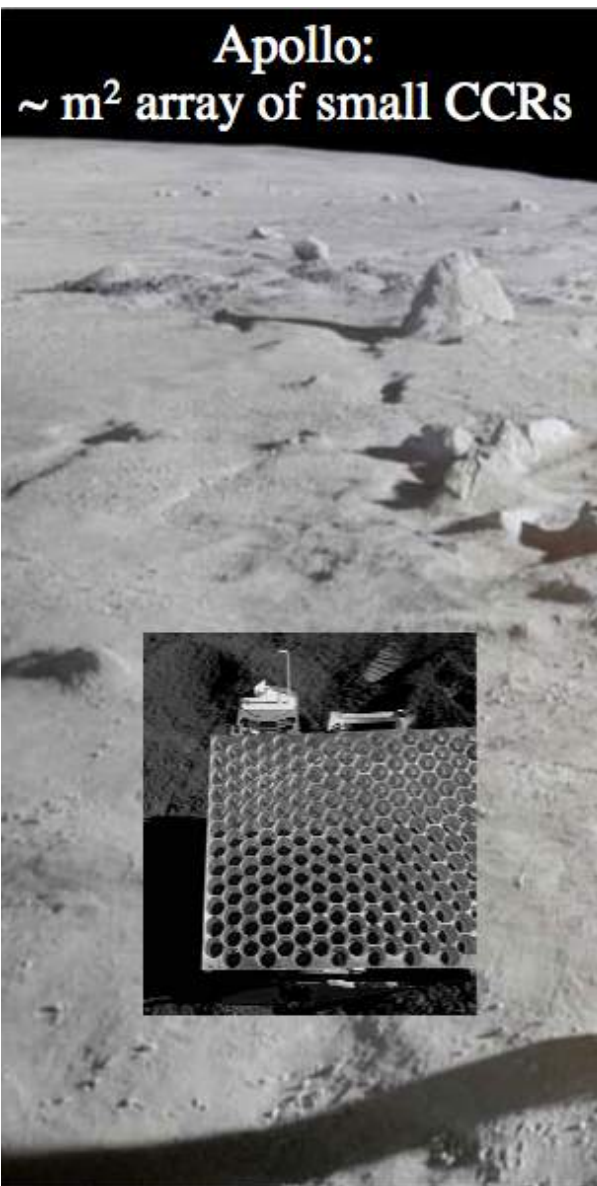
MoonLIGHT (100 mm)

GNSS (33 mm)

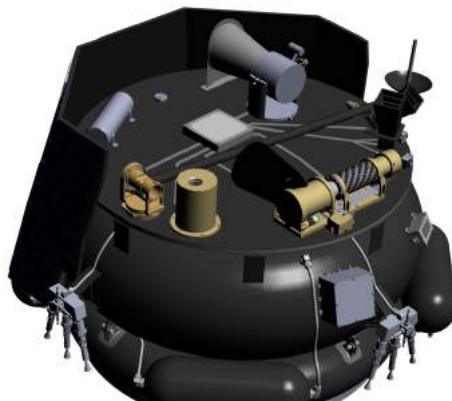
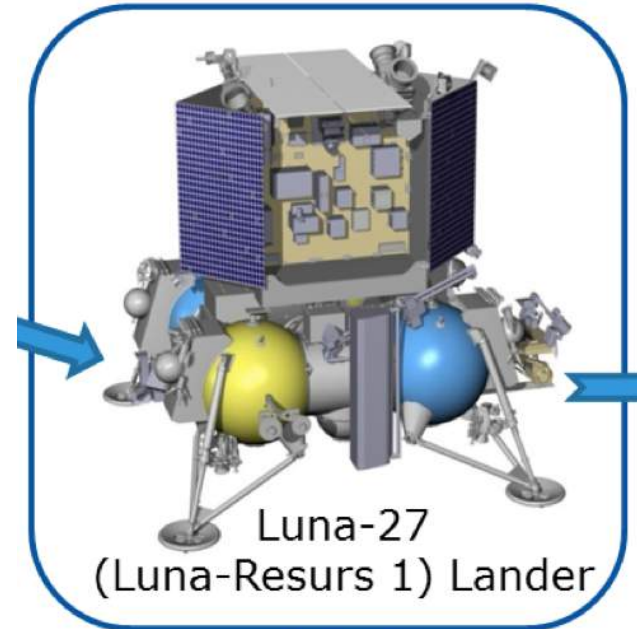
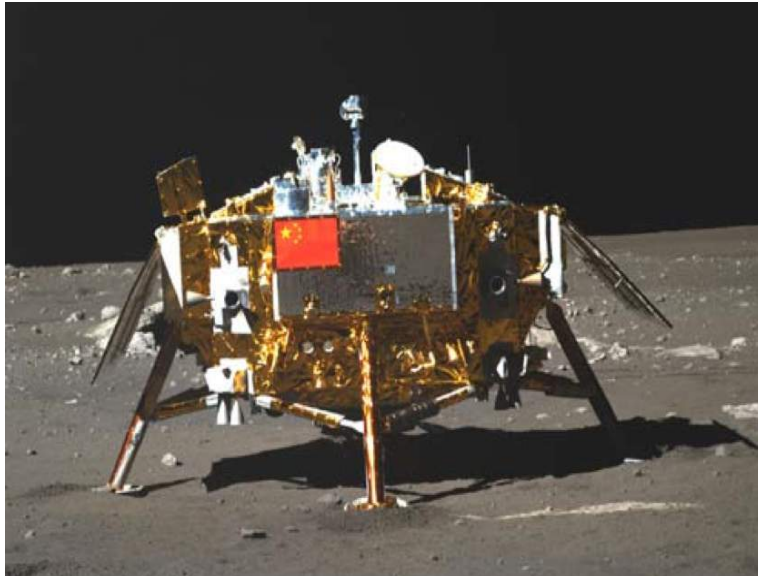
MoonLIGHT package







China, Russia, commercial GLXP, USA ...



Sun-Earth-Moon System

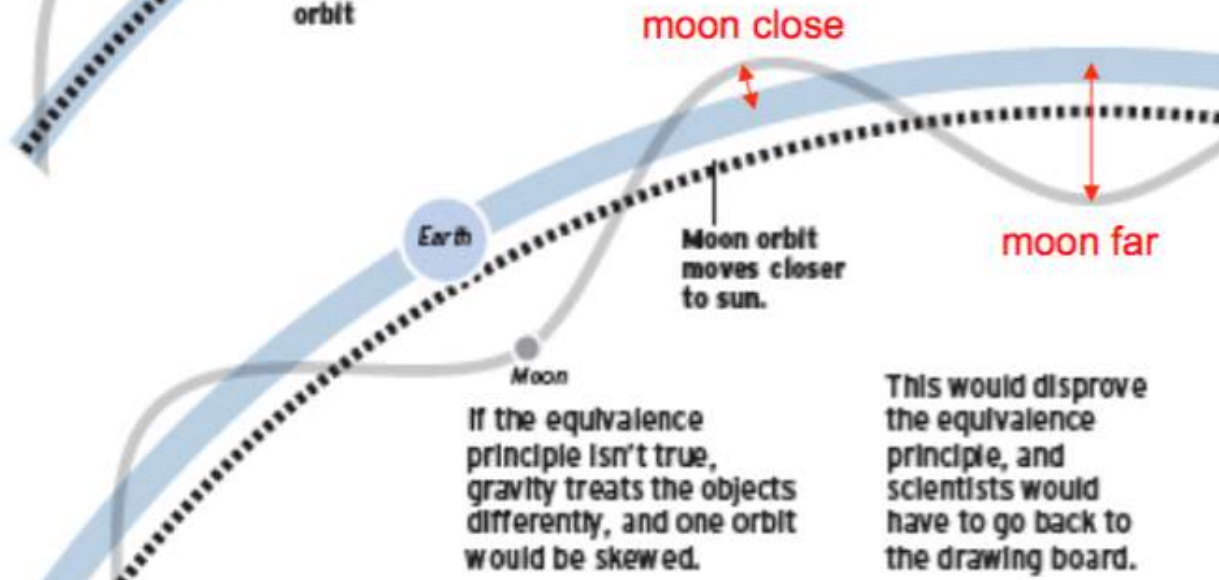
Unique, diversified, very accurate test of Einstein's theory:

- Earth & Moon: two bodies orbiting the Sun
- Equivalence principles: **weak and strong**
 - **Related, respectively to different compositions and quantities of masses of the gravitating bodies**
 - Conceptual experiments of Galileo and/or of the Astronaut
- ‘Geodetica’ precession of the lunar orbit
- Test of force-law **$1/r^2$**

Tests of Equivalence Principles of Galileo-Einstein

WHAT COULD BE FOUND IN THE ORBITS

If the equivalence principle is true, the sun's gravity pulls equally on the Earth and the moon. Therefore Earth's orbit and the moon's average orbit follow the same path.



This would disprove the equivalence principle, and scientists would have to go back to the drawing board.

Weak version (Galileo): hammer and feather have a very different type/quality of mass compositions, but they fall with the same acceleration in the field of the Moon

Strong version (Einstein): the heavy hammer (Earth) and lighter hammer (Moon) have a relatively similar composition (rocky) but in very different quantities. But again they fall with the same acceleration in the field of the Sun – Measured with LLR

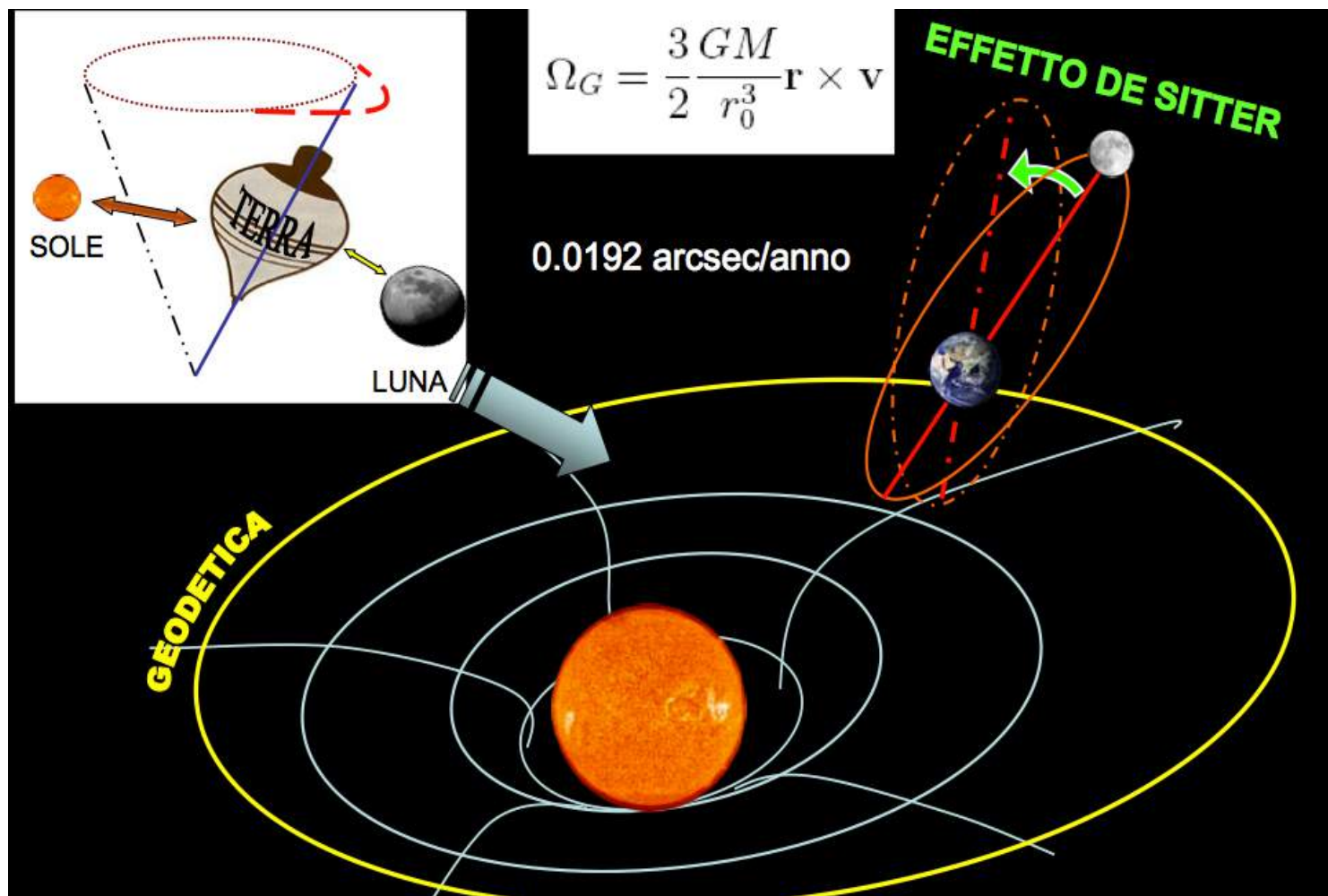






Precession of Moon orbit around the Earth **MLRO**

3-body effect, Sun-Earth-Moon (de Sitter effect) predicted by GR
 $\sim 3 \text{ meters / lunar orbit } (\sim 2''/\text{cy})$

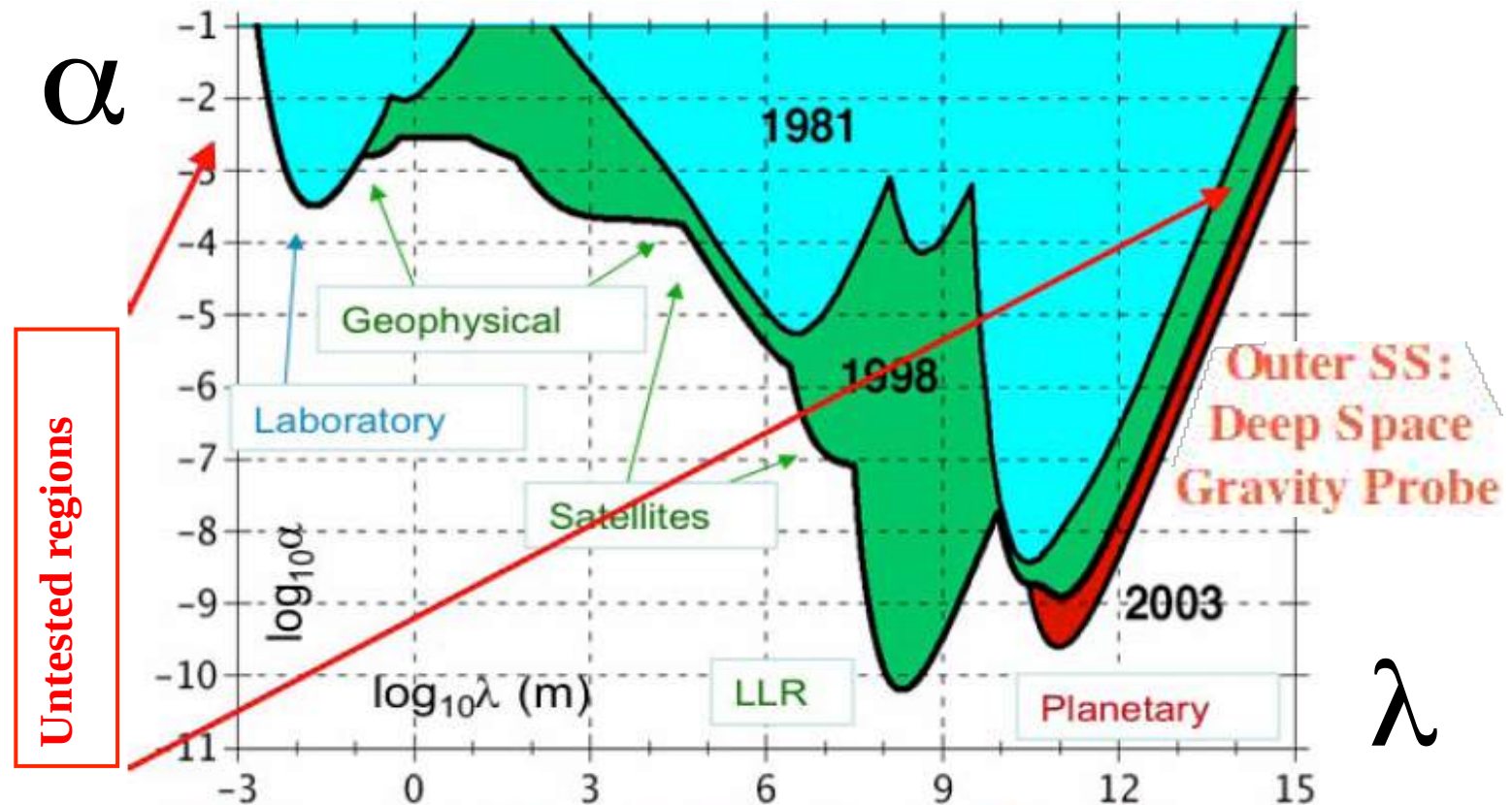


Inverse square law

Yukawa potential?

$$\alpha \propto (\text{Newton potential}) \times e^{-r/\lambda}$$

With our new lunar CCRs we will improve this test x 100

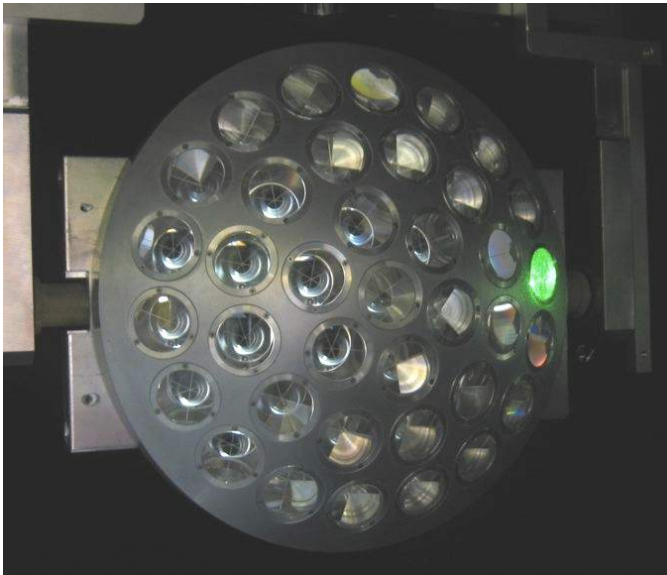


Courtesy : J. Coy, E. Fischbach, R. Hellings, C. Talmadge, and E. M. Standish (2003)

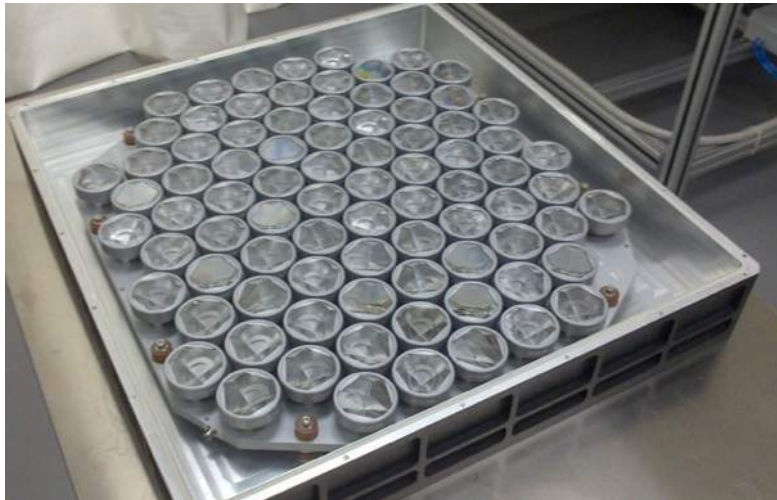
- **Gravitational redshift**
- Atomic clocks of Galileo (Maser e Rubidio) run faster than clock on Earth
- Effect $\sim 25\text{m}$ ($80\text{nsec} \times c$) after 2 min
 $\sim 18\text{Km}$ ($60\mu\text{sec} \times c$) after 1 day (two GPS orbite)

GALILEO, il GNSS Europeo

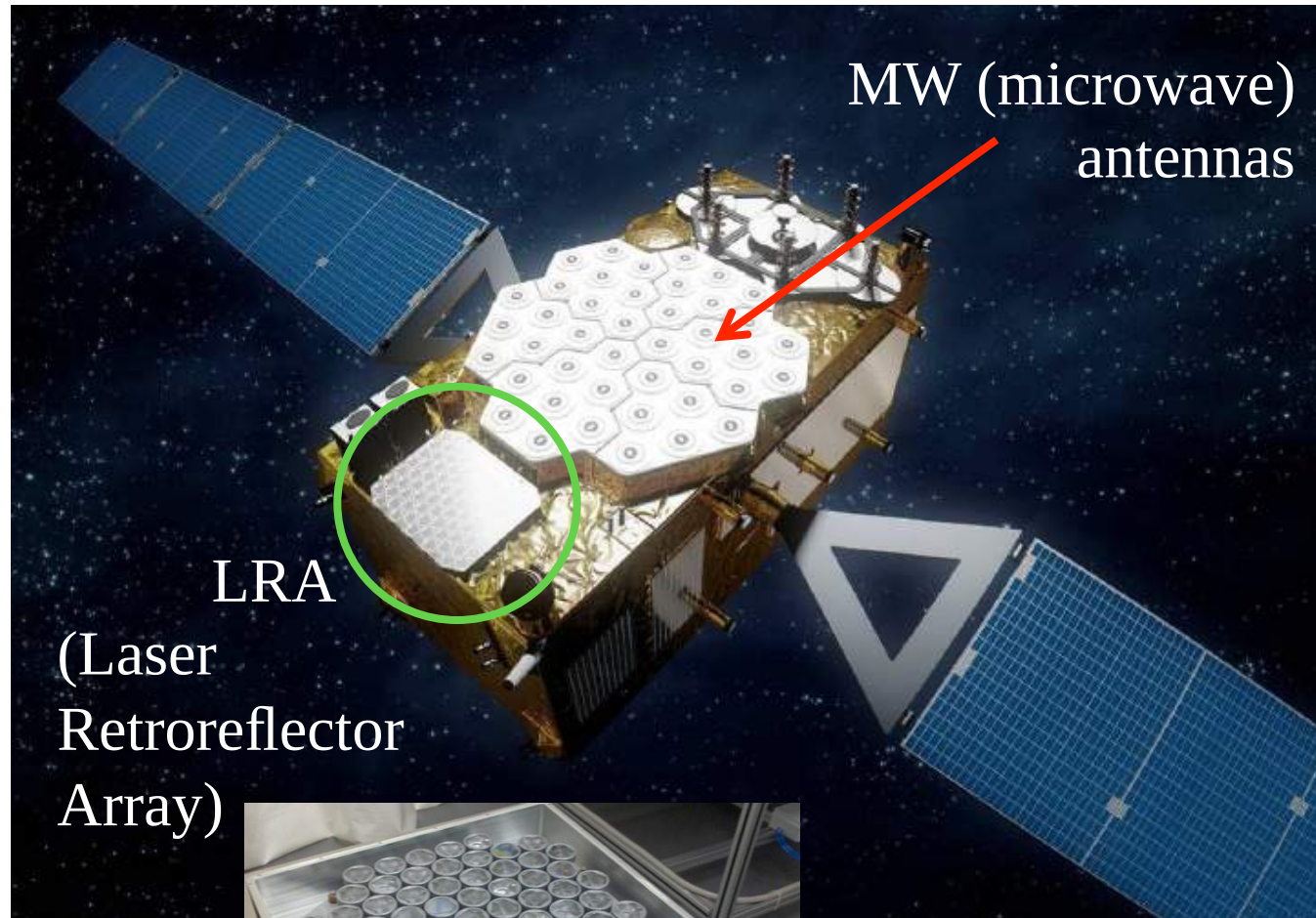




LAGEOS, GPS, Galileo
(clockwise)



Galileo In-Orbit Validation satellite (IOV)



~100 satelliti con retroriflettori laser!



Indian IRNSS: 7+4 regional satellites

SCF-Tested

Galileo europeo:
30 satelliti
Caratterizzati ai LNF



**US GPS:
24 global satellites**

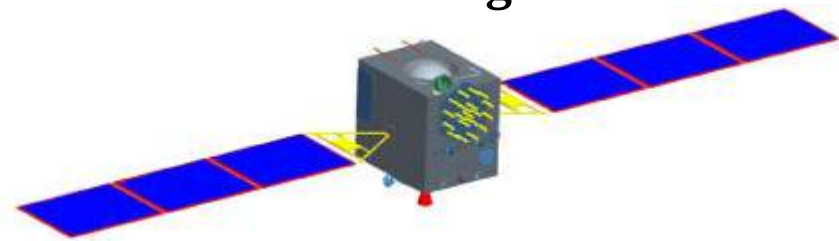


**Japanese QZSS:
3 regional satellites**



Russian GLONASS: 24 global satellites: SCF-Tested

**Chinese Compass/
Beidou: 20 global,
+5 regional satellites**



- Fundamental physics
 - Improve tests from factor 5 to 100
 - Improve by factor 10 the metrology of the Geocenter
 - 1 mm per epoch (10 yers) and stability of 0.1 mm/yr
- New destinations:
 - Mars, far side (not the dark side!!!) of the Moon
 - Asteroids and Comets
 - Icy/rocky moons of Jupiter and Saturn (*all these worlds can be ours, including Europa ...*)

L'inseguimento laser dei satelliti e della Luna (in inglese “Satellite and Lunar Laser Ranging” o SLR/LLR) sono riconosciute come le tecniche più precise e, al tempo stesso, meno costose per realizzare il posizionamento assoluto nel sistema Terra-Luna. I retroriflettori laser portati 40 anni fa sulla superficie lunare dalle missioni Apollo hanno permesso nel tempo di conoscere l'orbita della Luna con la precisione di pochi centimetri, di effettuare molteplici accurati test della Teoria della Relatività Generale di Einstein e importanti misure della struttura interna della Luna. Questo è l'esperimento spaziale costruito dall'uomo più longevo di tutti. Nel 1976 la NASA lanciò il satellite LAGEOS (LAsER GEODynamics Satellite), una sfera metallica massiccia coperta di retroriflettori laser. Nel 1992 NASA e ASI lanciarono LAGEOS II. Le loro orbite sono usate come “metro” spaziale e per definire (in un certo senso “materializzare”) il centro di massa della Terra (altrimenti inaccessibile direttamente). Le missioni Apollo e i due satelliti LAGEOS sono due pietre miliari della fisica fondamentale del sistema Terra-Luna e della geodesia spaziale. Sono anche tra le prime e le più feconde applicazioni scientifiche e tecnologiche del laser, inventato agli inizi degli anni 60. L'acquisizione e l'analisi dei loro dati, che sono gratuiti, continua a ritmo incessante da parte della rete di stazioni di Terra dello ILRS (“International Laser Ranging Service”).

Le due frontiere di questo campo della ricerca spaziale sono l'installazione di retroriflettori laser di seconda generazione sulla superficie della Luna con missioni robotiche e a bordo delle moderne costellazioni navigazione satellitare, come Galileo, il GNSS (Global Navigation Satellite System) dell'Europa. In questi due ambiti è attivo e all'avanguardia il laboratorio di test spaziale SCF_Lab dei LNF, il cui contributo originale alla disciplina dell'inseguimento laser è stata la realizzazione di due apparati sperimentali unici al mondo (SCF e SCF-G) e di una nuova procedura di test spaziale dei retroriflettori laser (SCF-Test). I LNF stanno anche sviluppando nuovi retroriflettori lunari in collaborazione con l'Università del Maryland a College Park (USA), l'istituzione che fu responsabile dei retroriflettori delle missioni Apollo. I nuovi riflettori targati Maryland/Frascati avranno prestazioni migliori fino a un fattore 100 rispetto a quelli Apollo e a quelli installati sui rover Russi Lunokhod.