

# *Synchrotron light sources: amazing present and future. Some applications using X-rays*

*Antonella Balerna*

*INFN - Frascati National Laboratory*

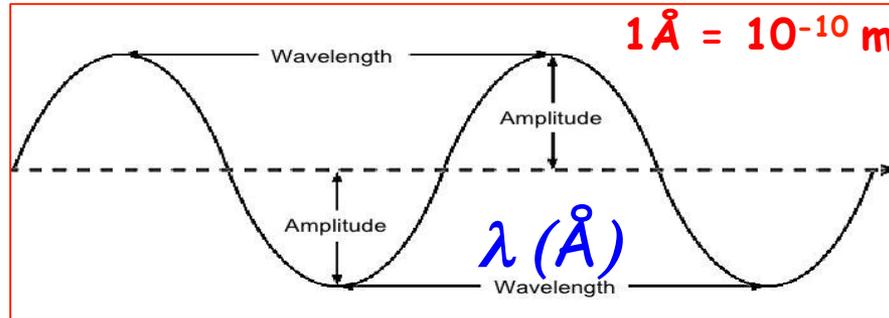


# Outline

- *Units*
- *Light, light sources, brightness and X-rays*
- *Synchrotron light*
- *Main properties*
- *Sources of Synchrotron Radiation*
- *Short history*
- *Present and future*
- *Some applications using X-rays*

# Reminder on some units used

- Unit of Wavelengths ( $\lambda$ ): Angstrom -  $\text{\AA}$



$$1 \text{ nm} = 10^{-9} \text{ m}$$

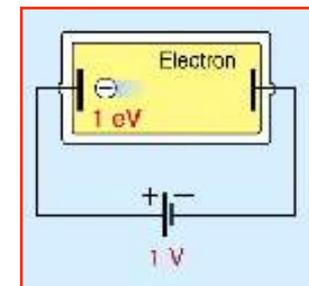
$$1 \text{\AA} = 0.1 \text{ nm}$$

- Unit of Energy: electronvolt - eV

$$1 \text{ eV} = 1.602 \times 10^{-19} \text{ Joules}$$

Energy gained (or lost) by the charge of a **single electron** moving moving across an **electron potential difference of one volt**.

$$\text{GeV} = 10^9 \text{ eV} \quad \text{MeV} = 10^6 \text{ eV}$$



- Unit of time: seconds - s    ps =  $10^{-12} \text{ s}$     fs =  $10^{-15} \text{ s}$



# Light



**Full featured double rainbow in Wrangell-St. Elias National Park, Alaska.**

This file is licensed under the [Creative Commons](https://creativecommons.org/licenses/by-sa/2.5/) Attribution-Share Alike 2.5 Generic license.

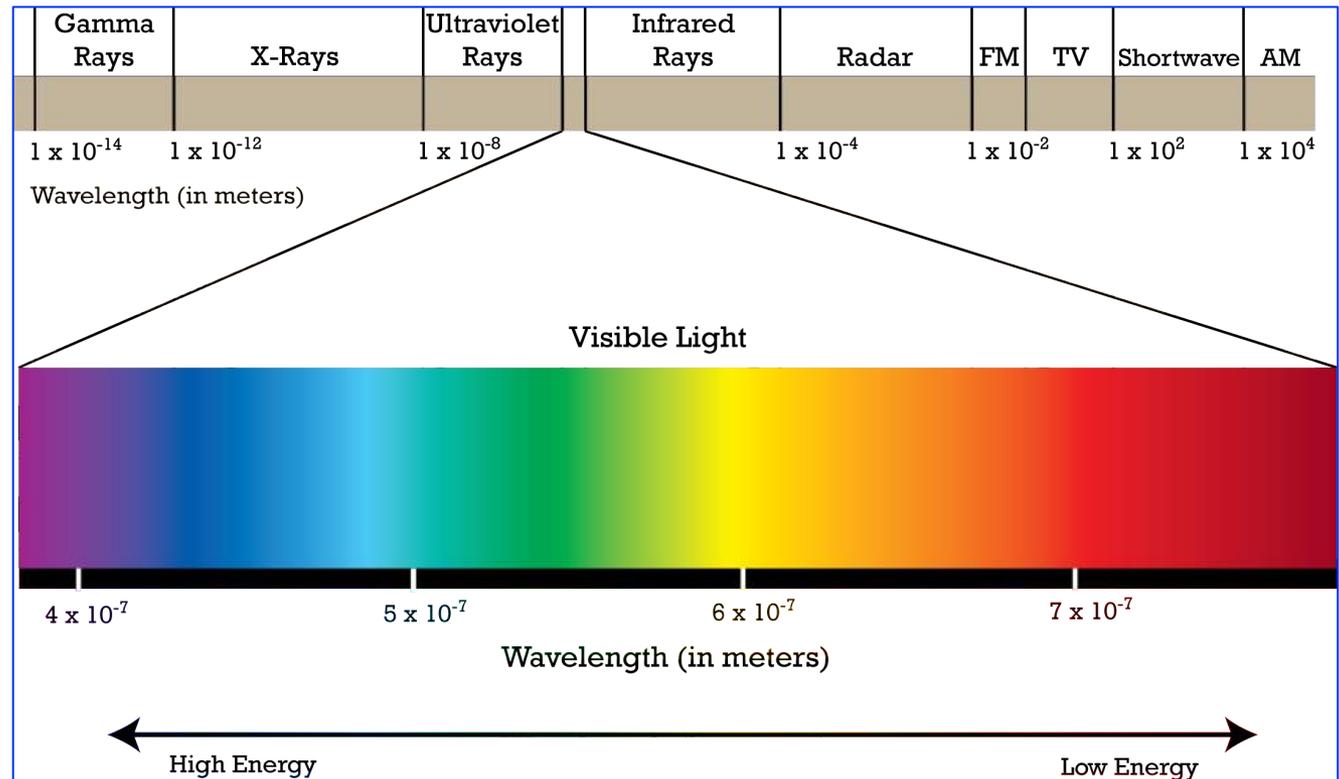
# Visible Light

*Visible light* is only a *tiny slice of the electromagnetic spectrum*. The entire electromagnetic spectrum of light is huge, spanning from *gamma rays on one end to radio waves*.

*Visible Light is the light we can see using our eyes.*

This tiny human spectrum encompasses a very specific range of wavelengths from about:

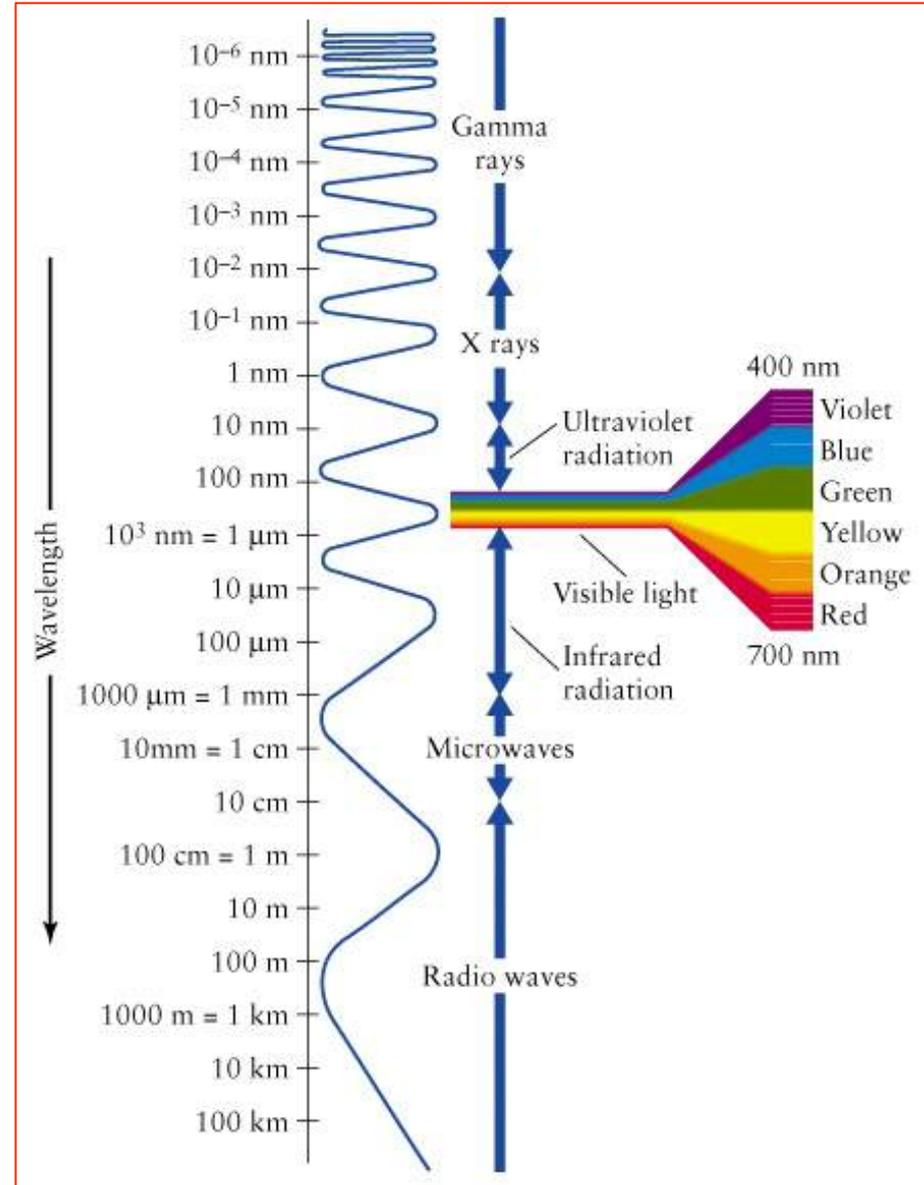
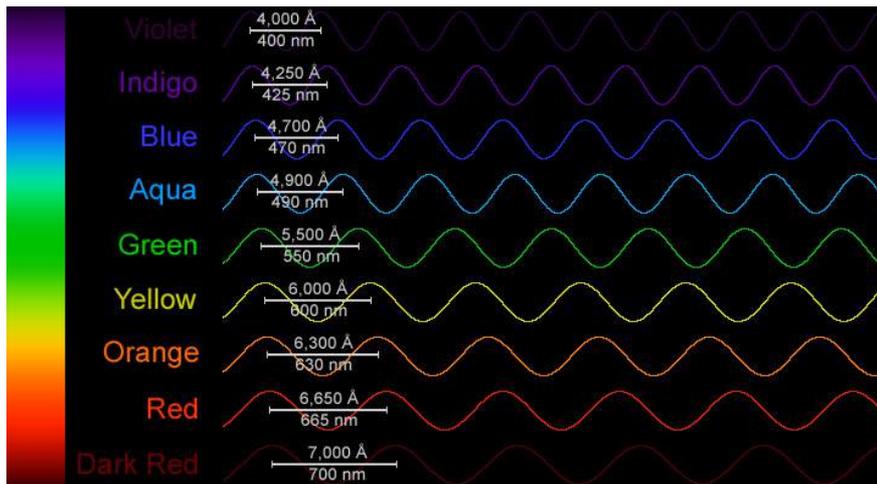
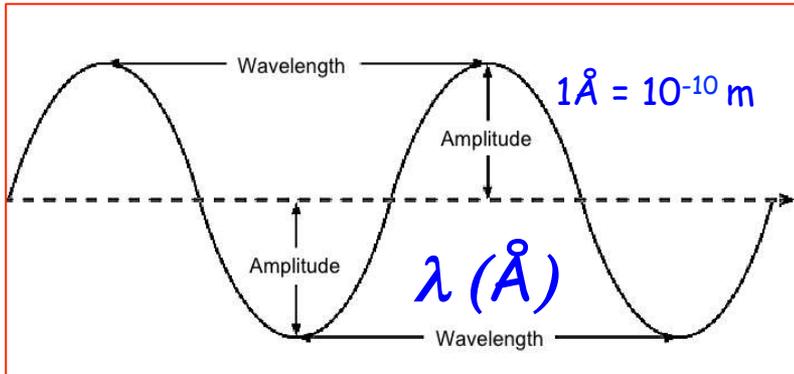
**380 nm to 780 nm.**



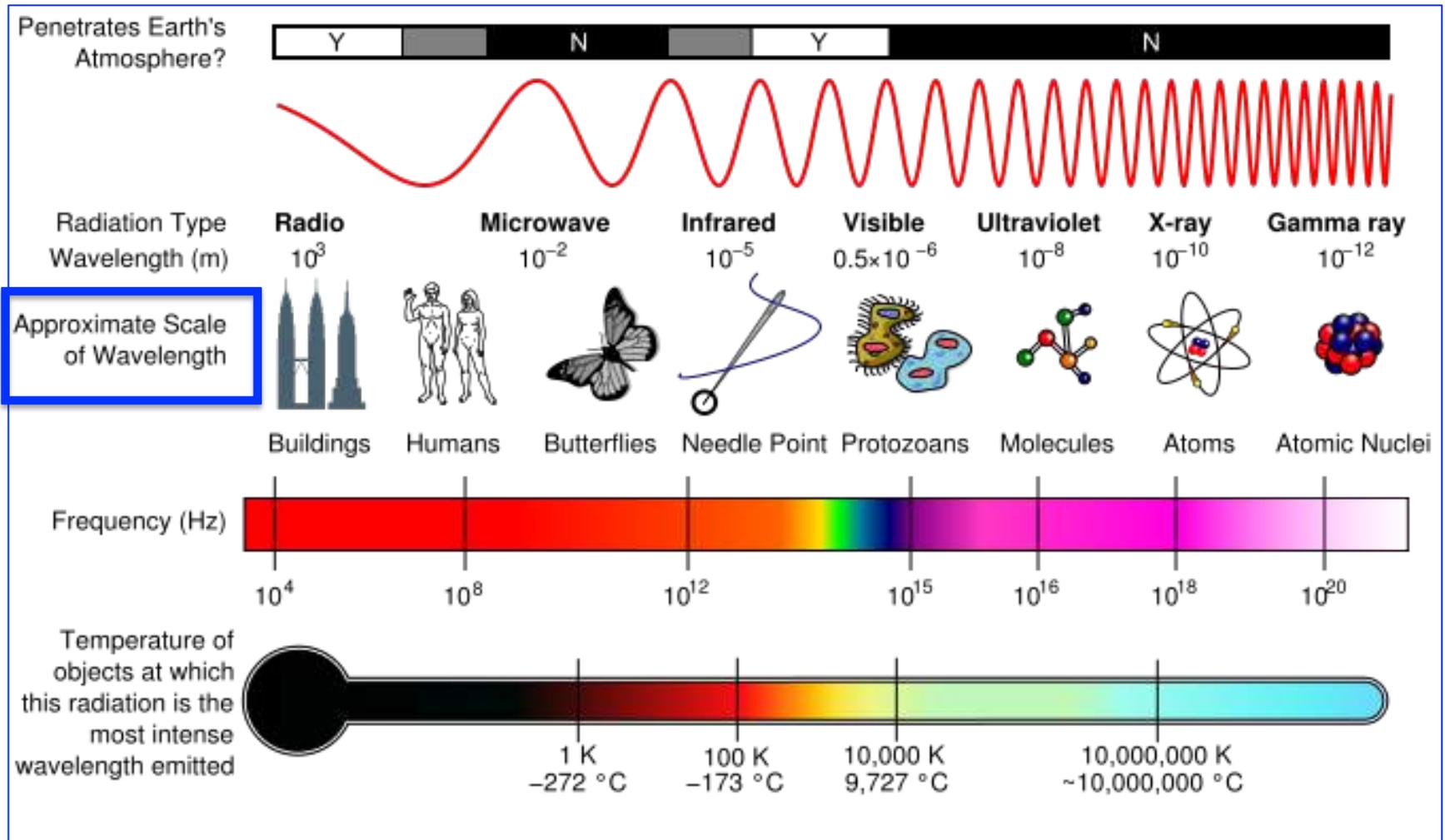
*Physiologically we see these frequencies because the photoreceptors in our retinas are sensitive to them.* When photons of light hit the photoreceptors it creates an electrochemical signal which is the first step in a fascinating process which ultimately results in us seeing colors.

# Light and wavelengths

Light travels as wave of energy. Waves of light have different *wavelengths* (the distance between the top of one wave and the top of the next). *Different colors of visible light have different wavelengths.*



# Electromagnetic Spectrum



The **wavelength** ( $\lambda$ ) and **frequency** ( $\nu$ ) of light are strictly related: the higher the frequency the shorter the wavelength! This is because **all light waves move through vacuum at the same speed** ( $c$  = speed of light) and the equation that relates wavelength and frequency for electromagnetic waves is:  $\lambda\nu = c$   $E = h\nu$

# Light sources

*Fire is not a very useful light source to see small details because its emitted power is spread in all directions!*



*A torchlight is more adequate because due to its small size the emission is concentrated within a narrow angular spread: this a "bright" source!*



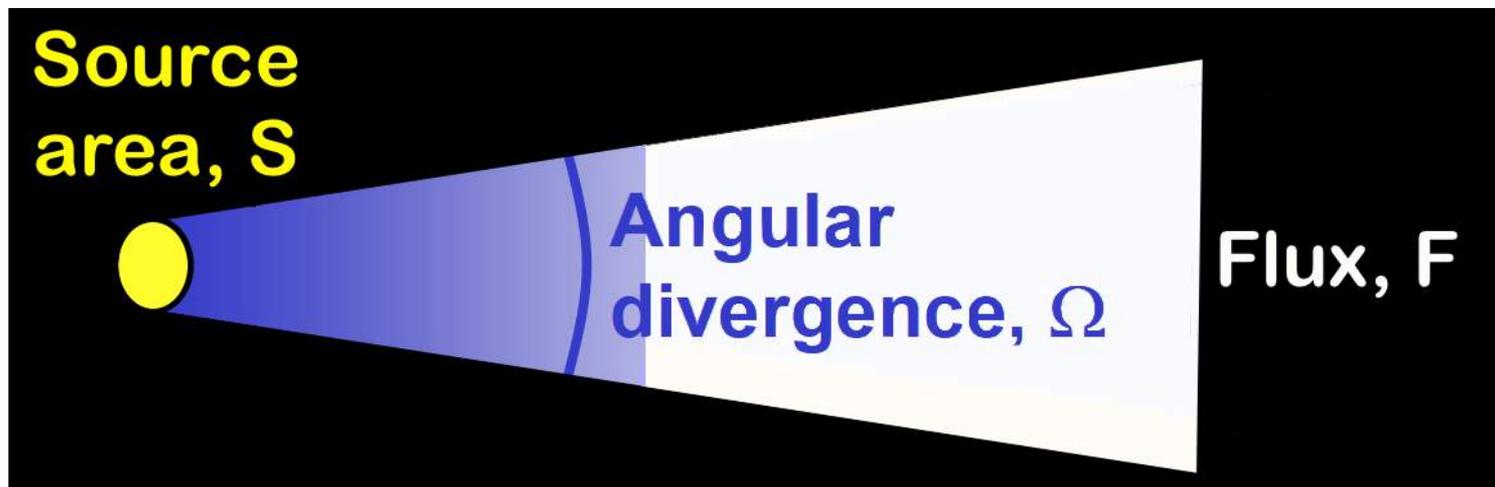
*Synchrotron radiation is a very bright light source that, as will be shown, gives us the chance to study also things that we cannot "see" with our eyes using not visible light but X-rays!*



# Light sources and brightness

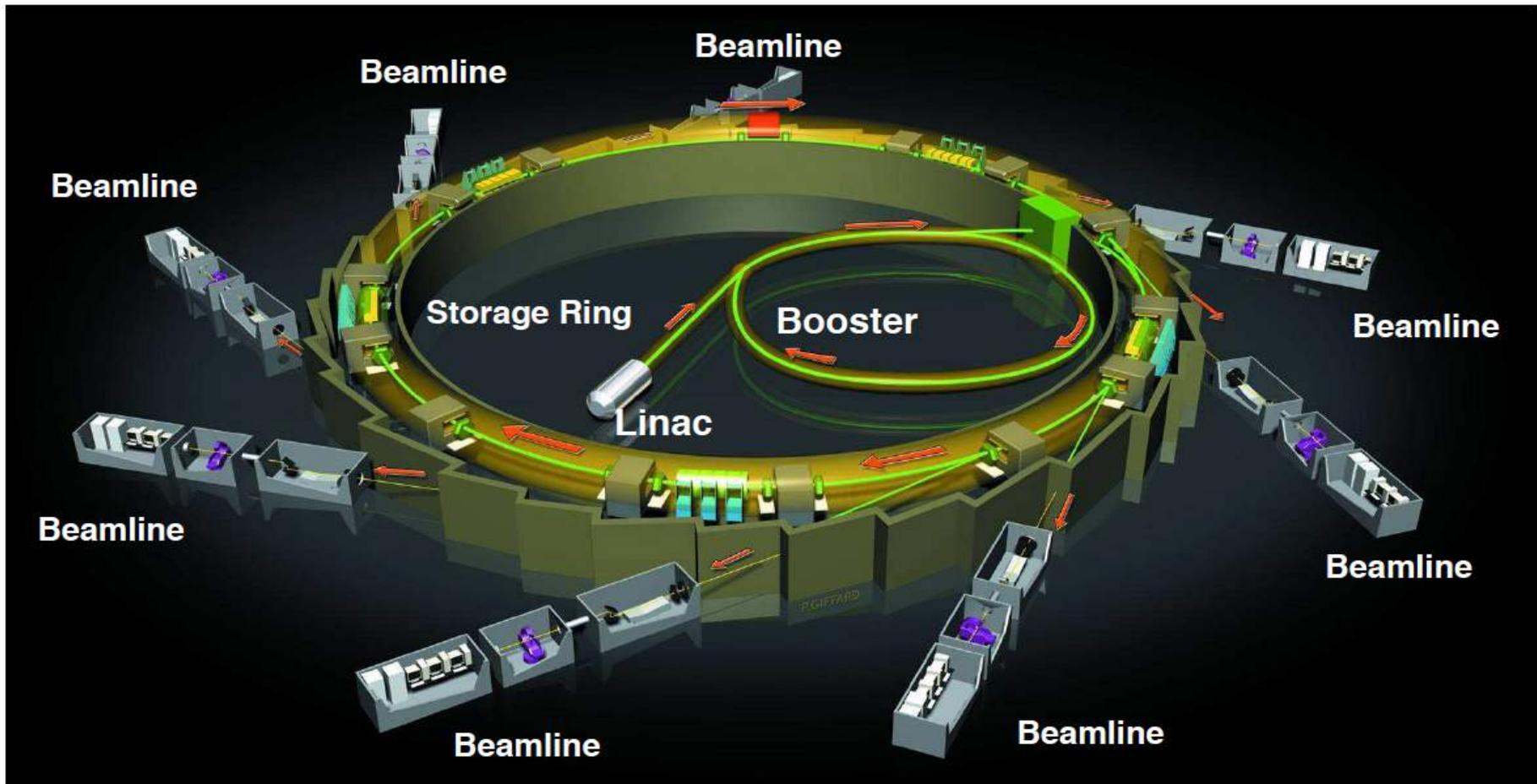
When interested in nm scale details BRIGHTNESS becomes fundamental.

A *bright source* is the one *very effective in illuminating a specific target*.  
If the specific target is small a *bright source* has a *small size with light emission concentrated within a narrow angular spread*.

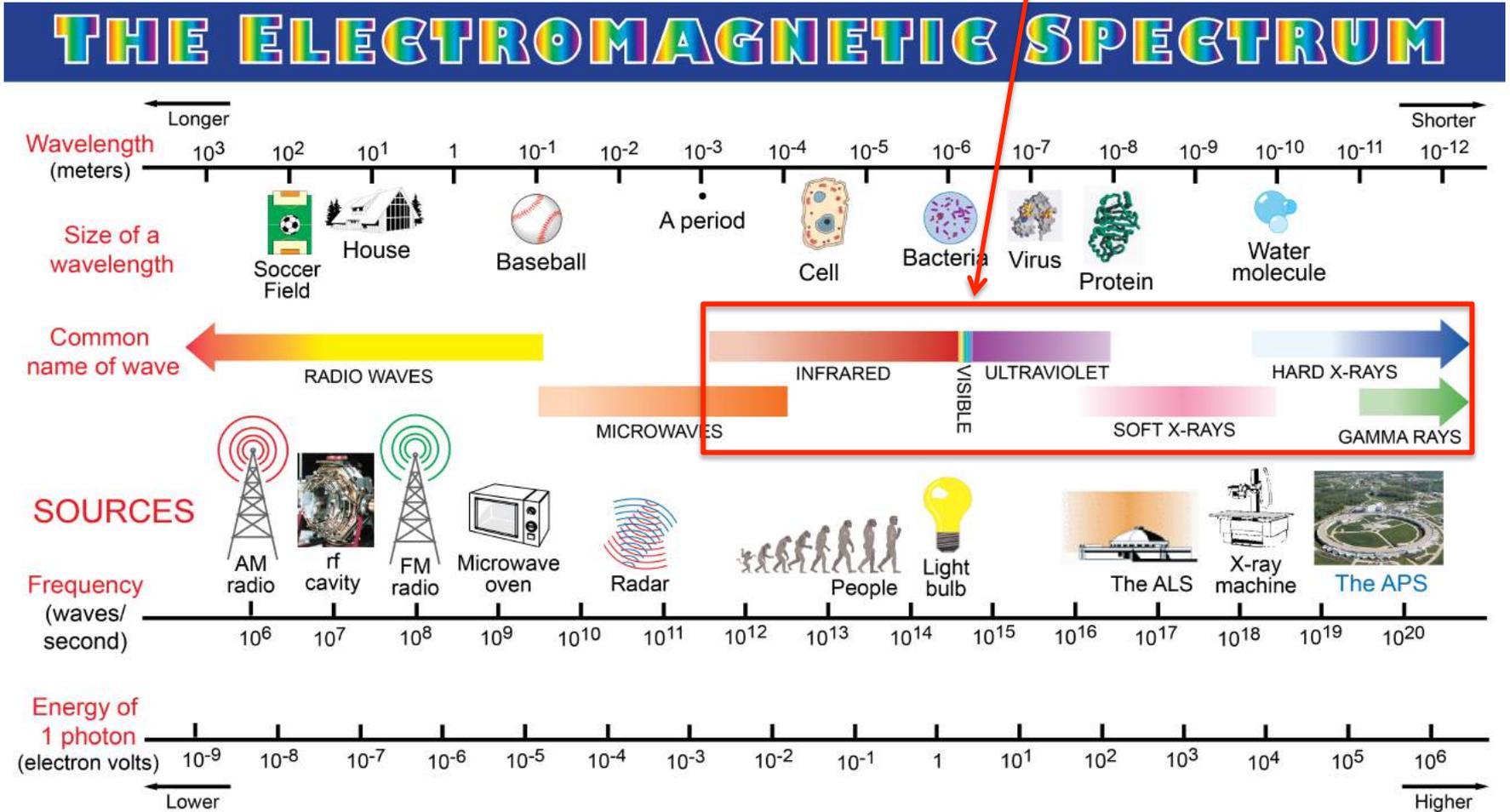


$$\text{Brightness} = \text{constant} \frac{F}{S \times \Omega}$$

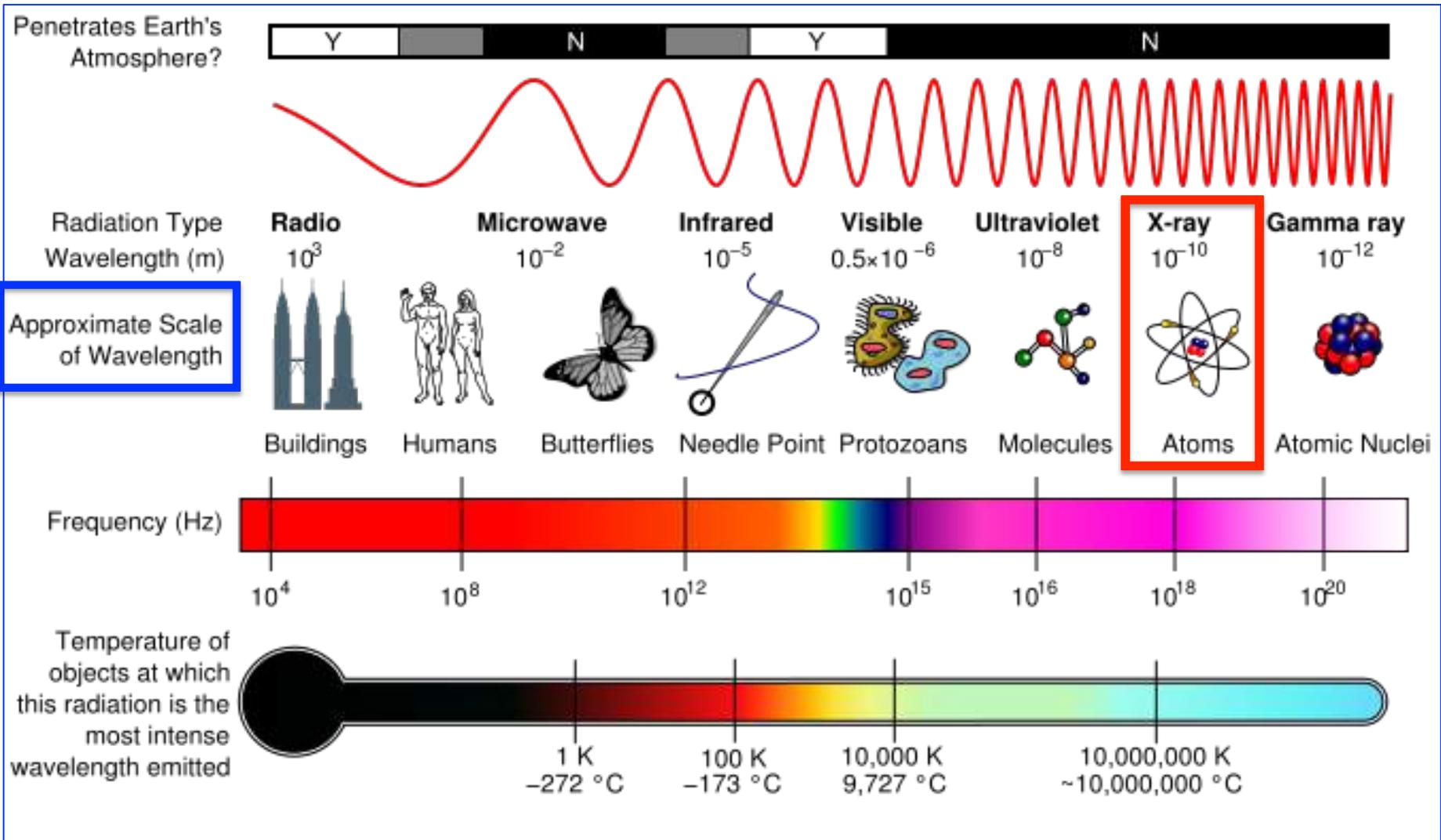
# *Accelerators are bright sources of synchrotron radiation*



# Spectral range covered by Synchrotron Radiation



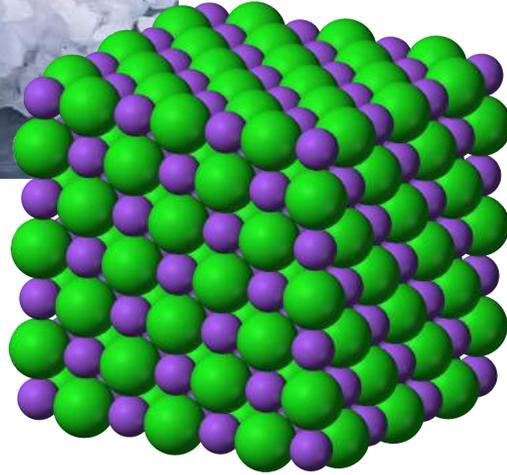
# Electromagnetic Spectrum and X-rays



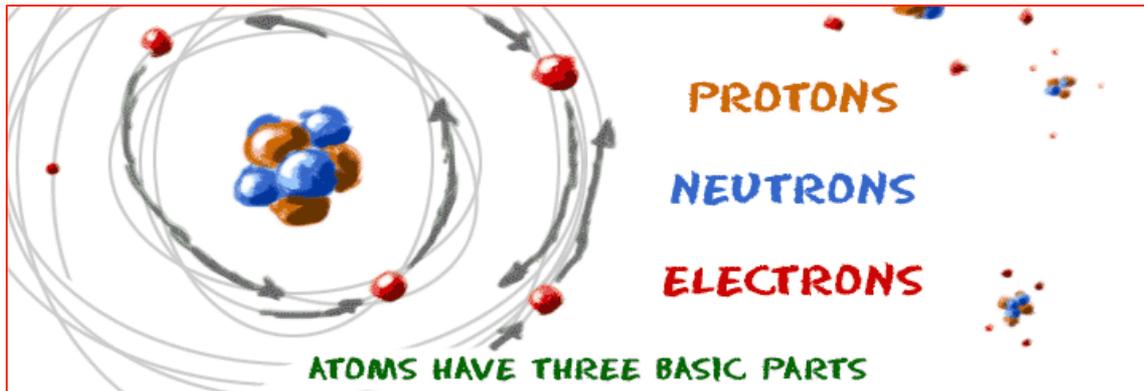
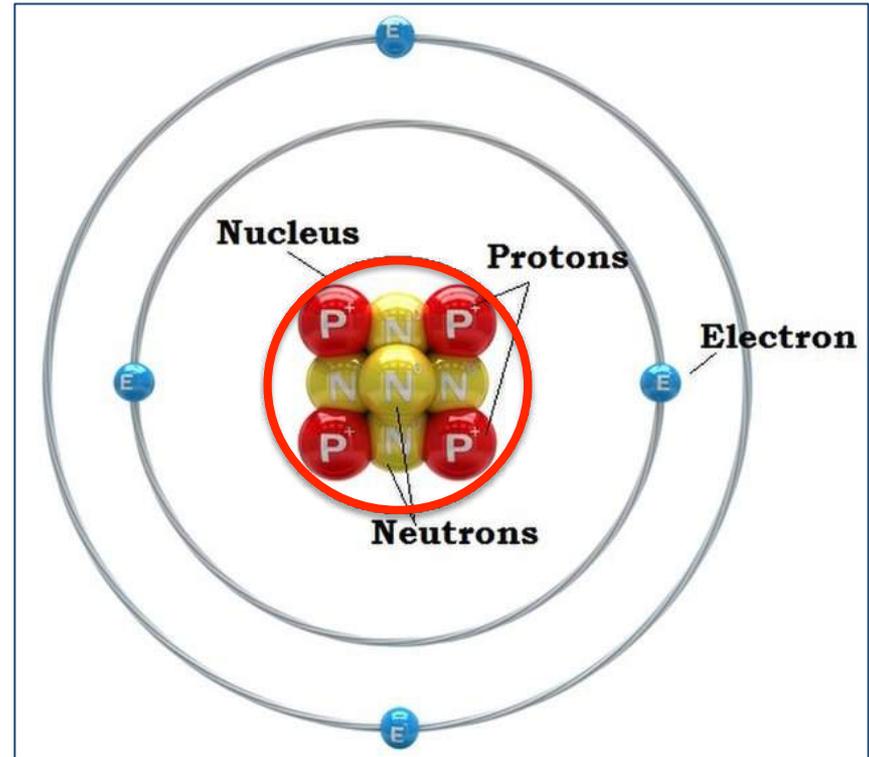
# Atoms



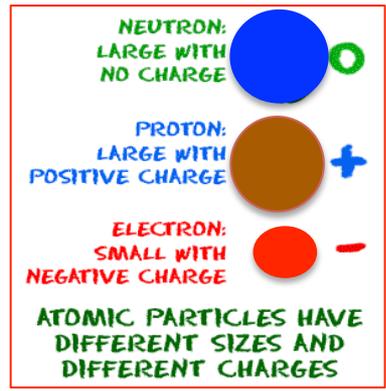
Table salt  
NaCl



Crystal atomic  
structure of  
sodium chloride



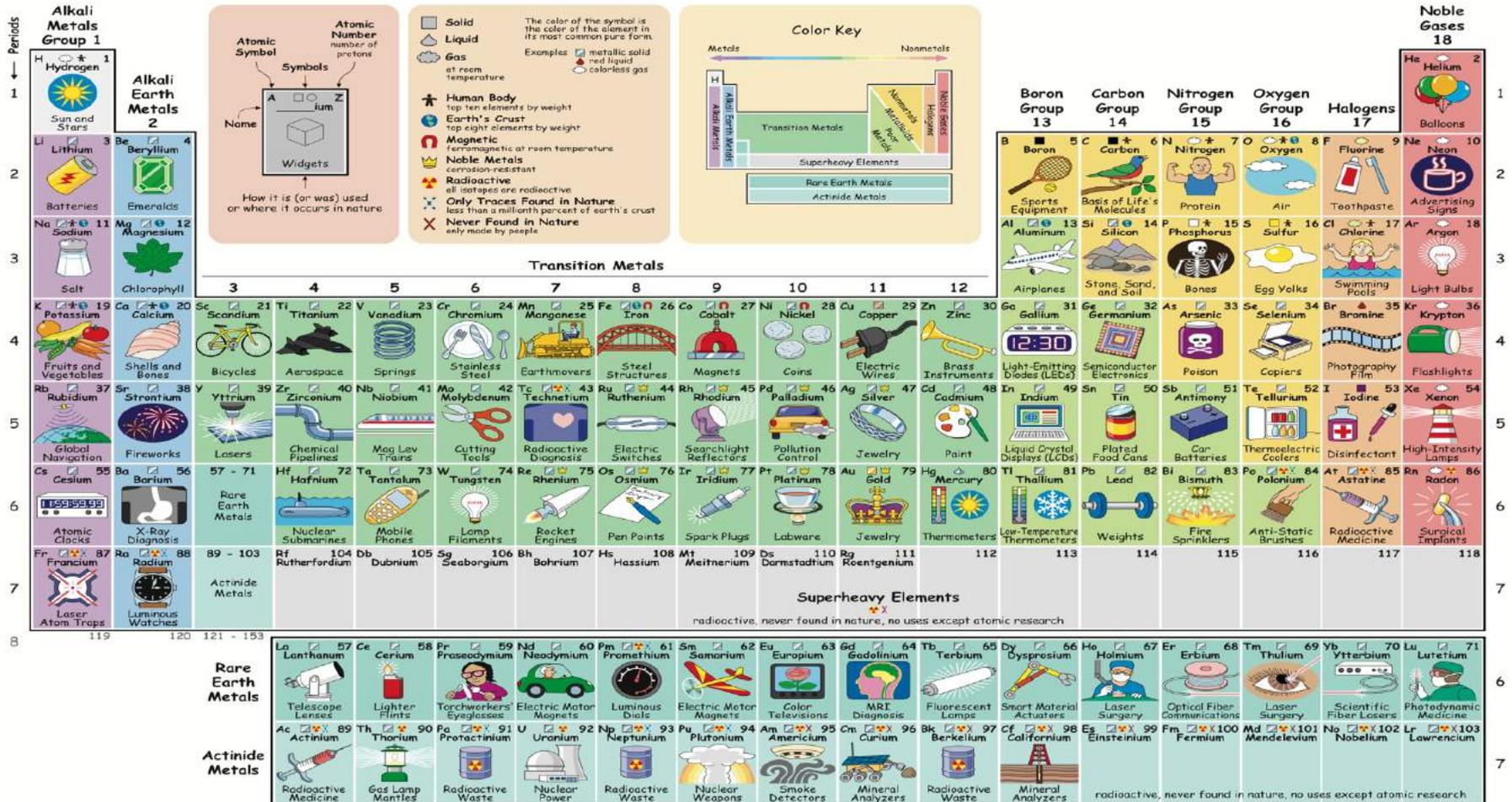
PROTONS  
NEUTRONS  
ELECTRONS



# Atoms

Matter is everything around us! All matter such as solids, liquids, and gases, is composed of atoms. Therefore, atoms are considered to be the basic building block of matter. From the periodic table, it can be seen that there are only about 100 different kinds of atoms. These same 100 atoms form thousands of different substances ranging from the air we breathe to the metal used to support tall buildings.

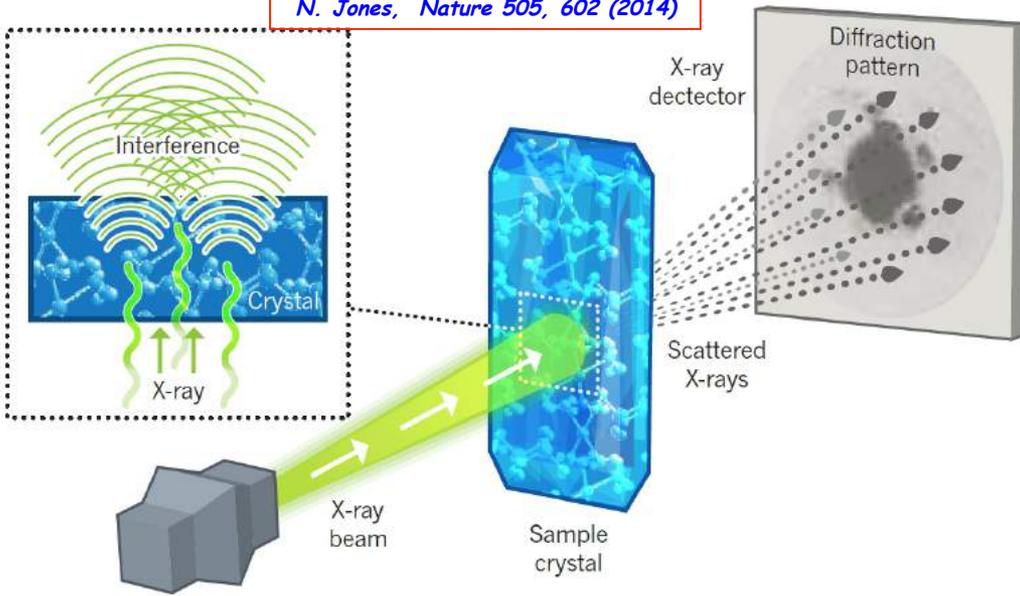
## The Periodic Table of the Elements, in Pictures



Selected by Z = atomic number = number of protons in the nucleus

# X-rays and atoms

N. Jones, Nature 505, 602 (2014)



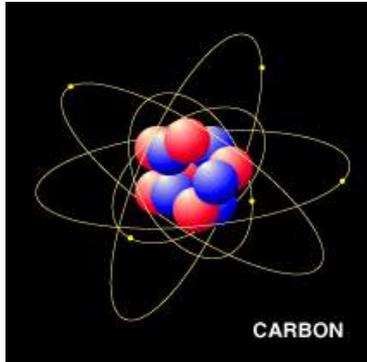
With X-rays we can study atoms because their wavelengths are of the order of  $10^{-10}$  m or 0.1 nm or  $1 \text{ \AA}$   
Matter is composed of **ATOMS!**

Using X-rays we can study the atomic structure of materials. The atomic structure primarily affects the chemical, physical, thermal, electrical, magnetic, and optical properties.



**Why is this important?**

# Atoms and X-rays

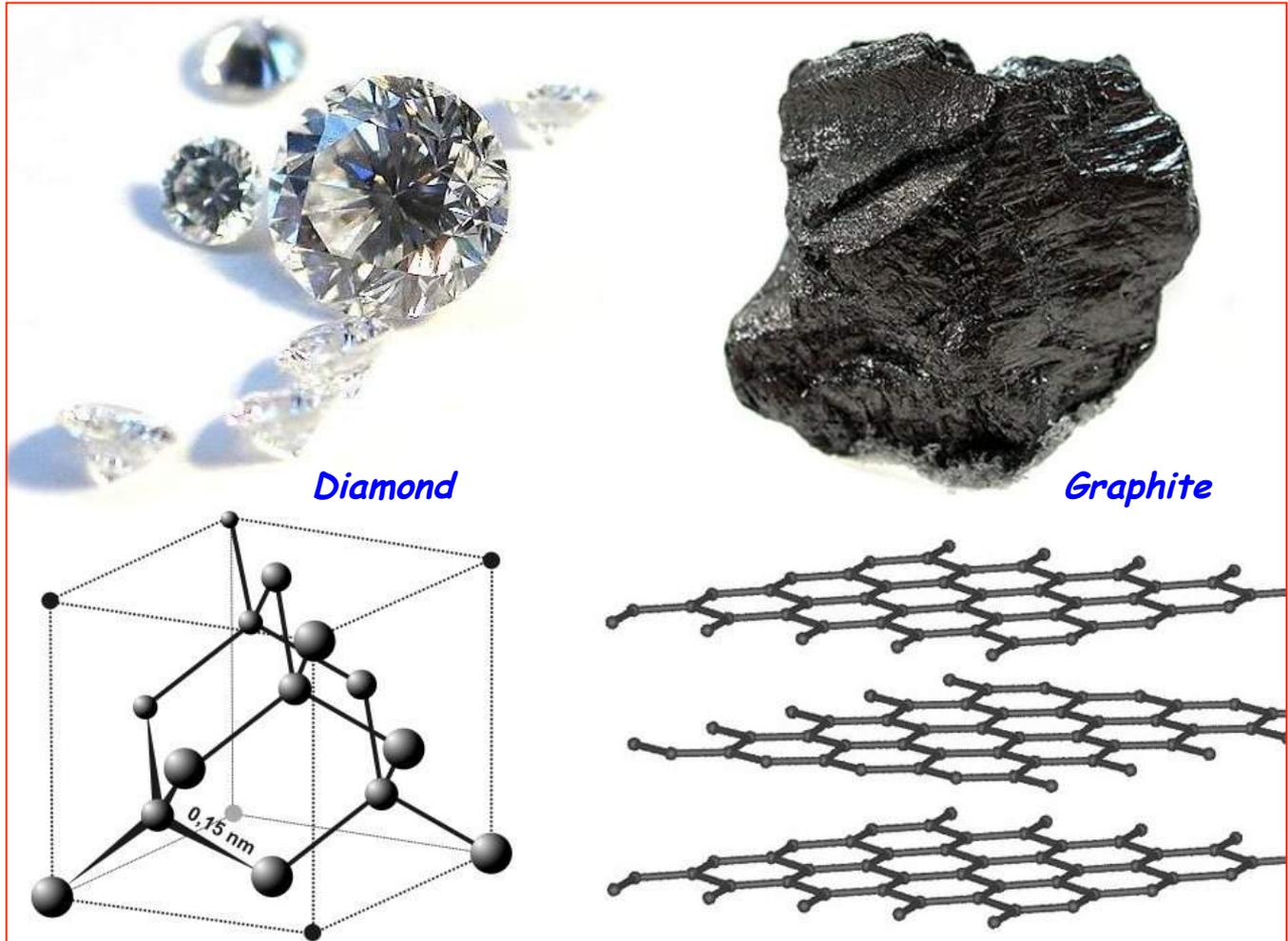


C atomic number  $Z = 6$

*Both diamond and graphite are made entirely out of carbon atoms!*

*The differing properties of graphite and diamond arise from their distinct crystal structures.*

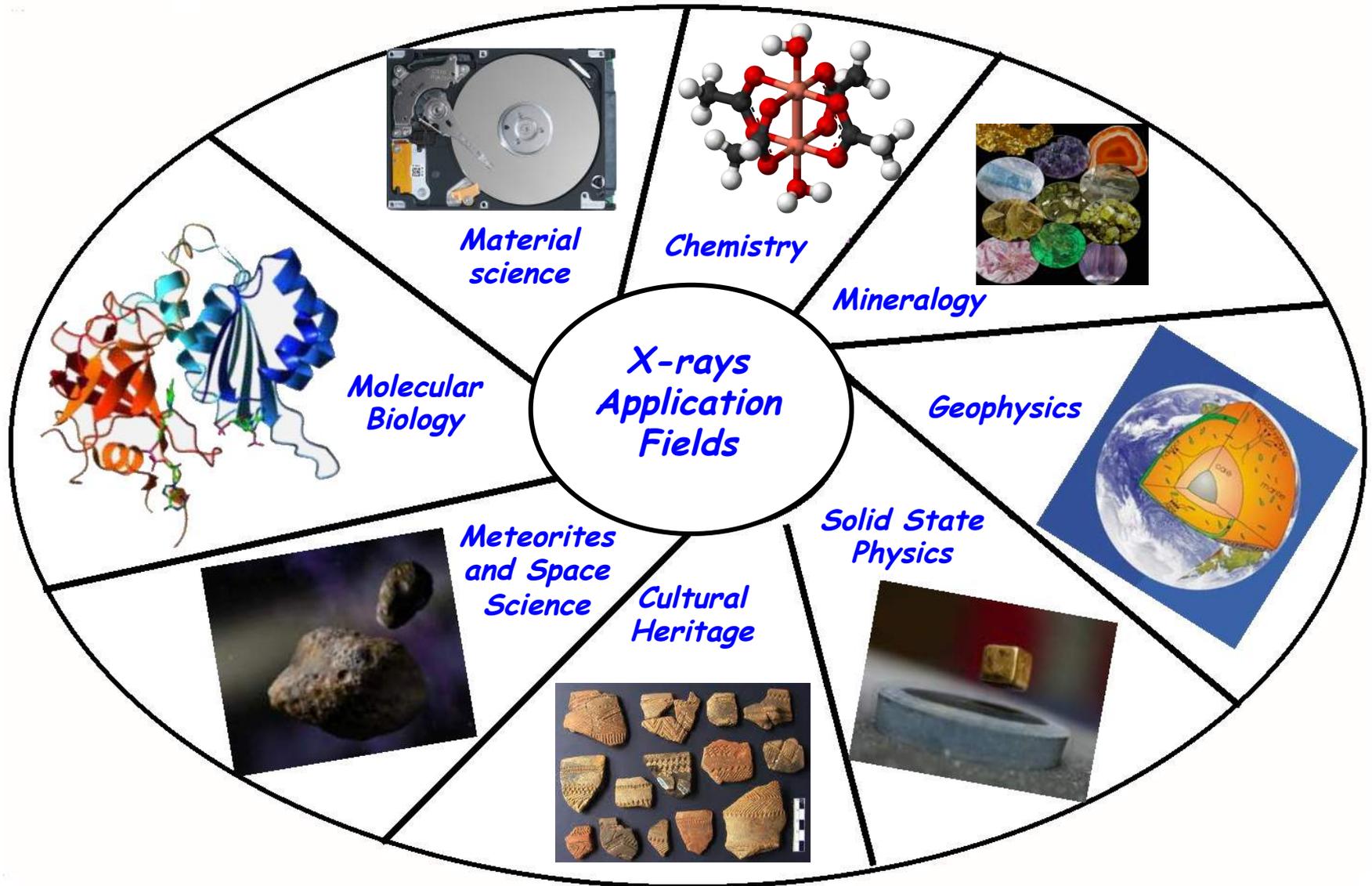
*Graphite is opaque and metallic- to earthy looking, while diamonds are transparent and brilliant.*



In *graphite*, the *individual carbon atoms link up to form sheets of carbon atoms*. Within each sheet every carbon atom is bonded to three adjacent carbon atoms (*covalent bonds*) producing hexagonal rings of carbon atoms. *Weak bonding forces called van der Waals forces hold the sheets together*. Because these *forces are weak, the sheets can easily slide past each other*. The sliding of these *sheets gives graphite its softness for writing and its lubricating properties*.

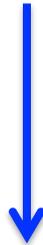
In *diamonds*, each carbon atom is strongly bonded to four adjacent carbon atoms located at the apices of a tetrahedron (a three-sided pyramid). The four valence electrons of each carbon atom participate in the formation of *very strong covalent bonds*. These bonds have the same strength in all directions. *This gives diamonds their great hardness*.

# *X-rays application fields*



*Bright X-ray source?*

*HE Particle accelerators*



*Synchrotron light*

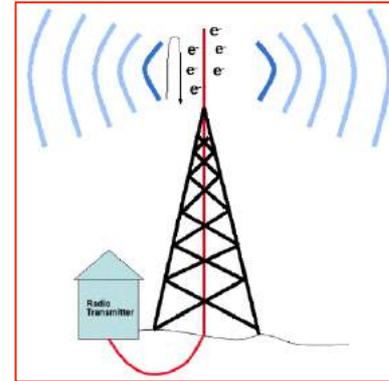
# *Answers to be given*

- *What is synchrotron light?*
- *How is it produced?*
- *History?*
- *Properties?*
- *Sources?*
- *How and why is it used?*
- *Applications?*

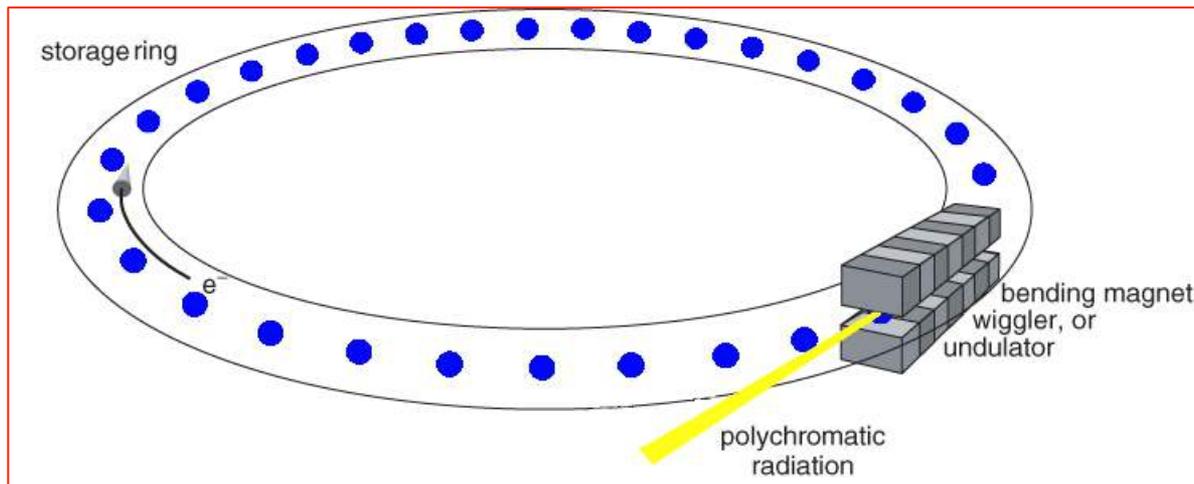
# Synchrotron radiation

Accelerated charged particle, like  $e^+$ ,  $e^-$  and ions, emit electromagnetic radiation.

$$v \ll c \text{ or } \beta = v/c \ll 1$$



When charged particles, moving at relativistic speeds ( $v \approx c$ ), are forced to change the direction of their motion (acceleration), under the effect of magnetic fields, in circular particle accelerators, like synchrotrons, the radiation produced is called **synchrotron radiation**.



$$v \approx c \text{ or } \beta = v/c \approx 1$$

# Synchrotron light is present in nature

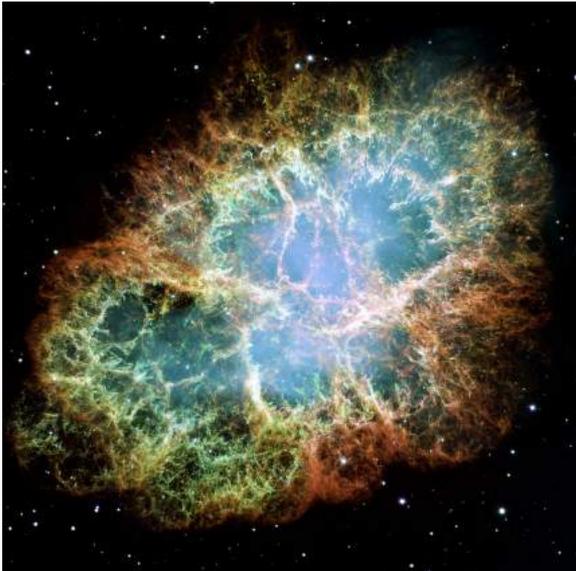
*Synchrotron radiation is a very important emission process in Astrophysics.*

*Crab Nebula: remnant of a supernova explosion seen on earth by Chinese astronomers in 1054, at about 6500 light years from Earth in the constellation Taurus !*

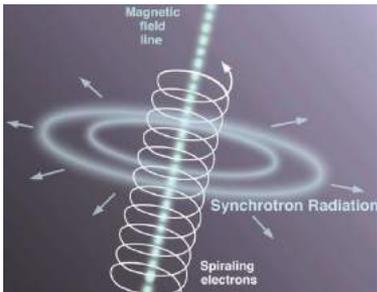
*The heart of the nebula is a rapidly-spinning neutron star, a pulsar, that powers the strongly polarised bluish 'synchrotron' nebula.*

*The Crab pulsar is slowing at the rate of about  $10^{-8}$  sec per day, and the corresponding energy loss agrees well with the energy needed to keep the nebula luminous. Some of this luminosity takes the form of synchrotron radiation, requiring a source of energy for accelerating charged particles.*

*Composite image data from three of NASA's Great Observatories. The Chandra X-ray Observatory image is shown in blue, the Hubble Space Telescope optical image is in red and yellow, and the Spitzer Space Telescope's infrared image is in purple. The X-ray image is smaller than the others because extremely energetic electrons emitting X-rays radiate away their energy more quickly than the lower-energy electrons emitting optical and infrared light. The Crab Nebula is one of the most studied objects in the sky, truly making it a cosmic icon.*

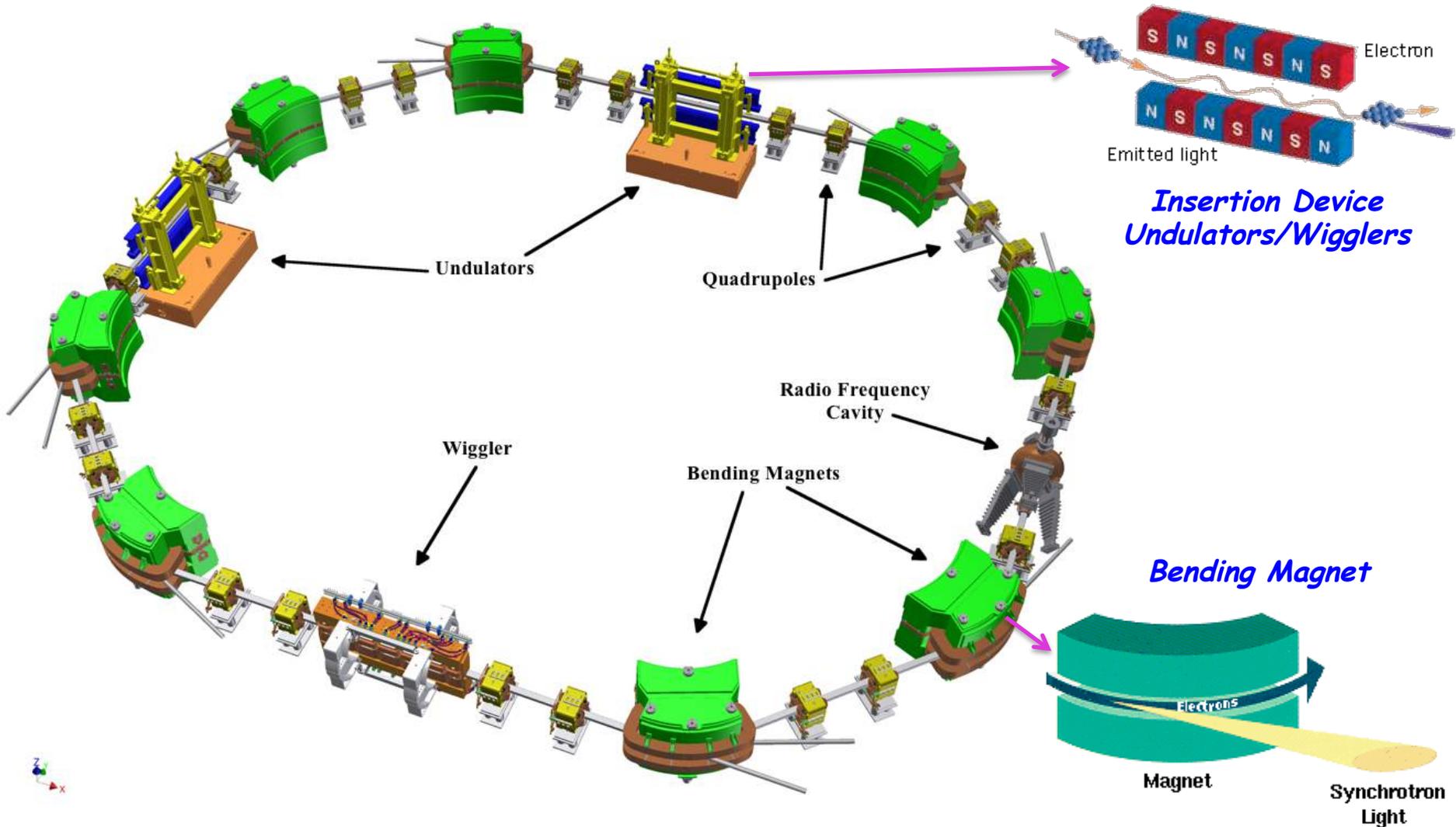


*NASA Hubble Space Telescope image of the Crab Nebula (NASA, ESA and Allison Loll/Jeff Hester (Arizona State University)).*



*NASA's Great Observatories' View of the Crab Nebula X-Ray-blue: NASA/CXC/J. Hester (ASU); Optical-red and yellow: NASA/ESA/J. Hester & A. Loll (ASU); Infrared-purple: NASA/JPL-Caltech/R. Gehrz (Univ. Minn.)*

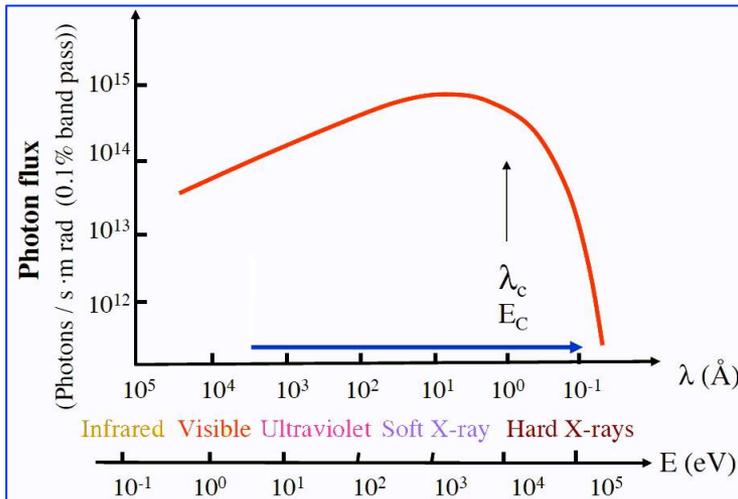
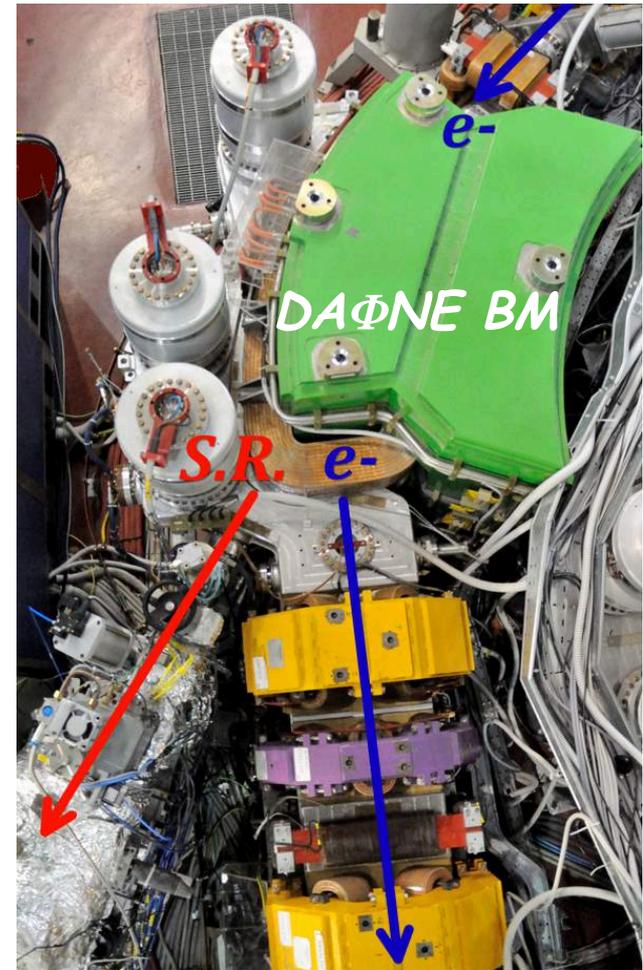
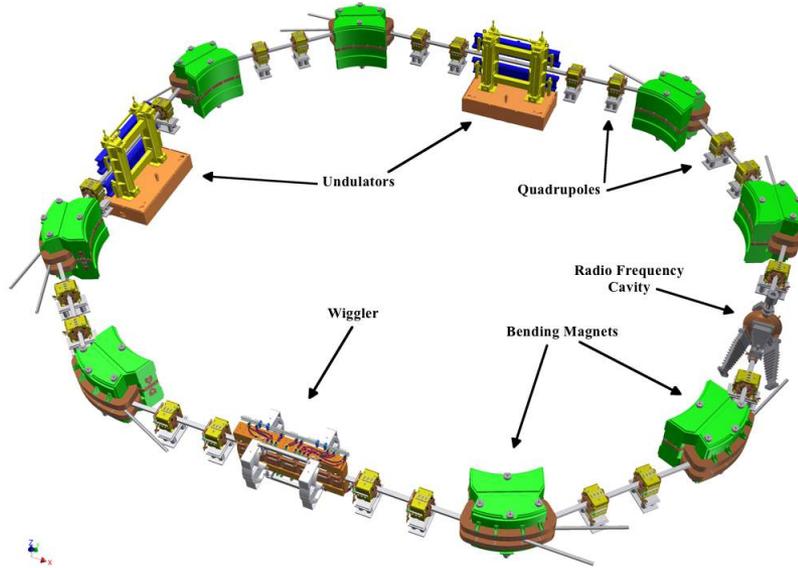
# *Synchrotron light artificially produced by circular particle accelerators*



**Bending magnets and insertion devices are components of the storage ring.**

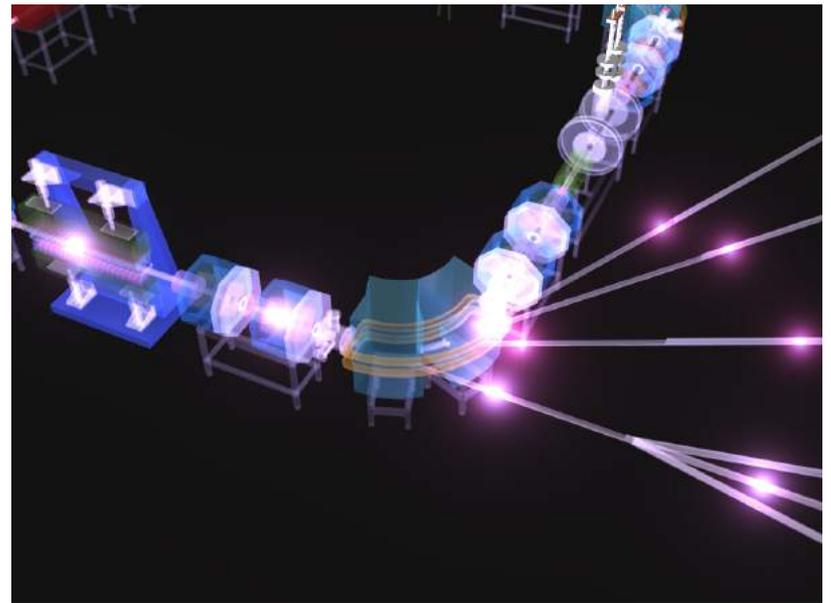
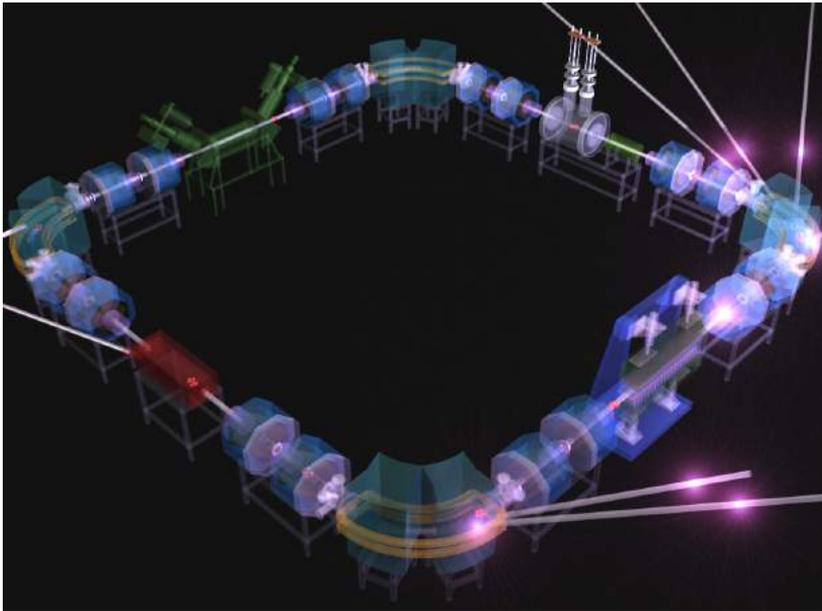
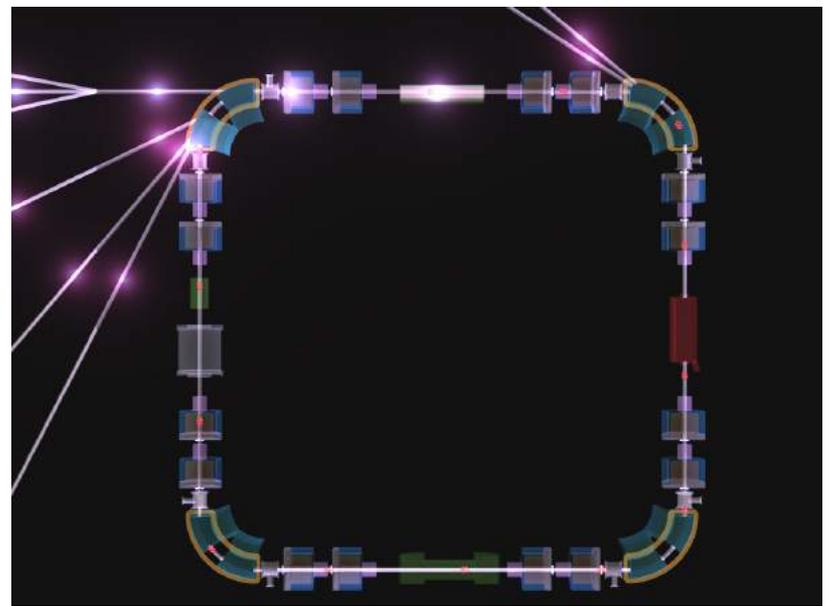
# Bending magnets

Bending magnets are fundamental parts of the storage ring.



$$\epsilon_c [\text{keV}] = 2.218 \frac{E [\text{GeV}]^3}{R [\text{m}]} = 0.665 \cdot E [\text{GeV}]^2 \cdot B [\text{T}]$$

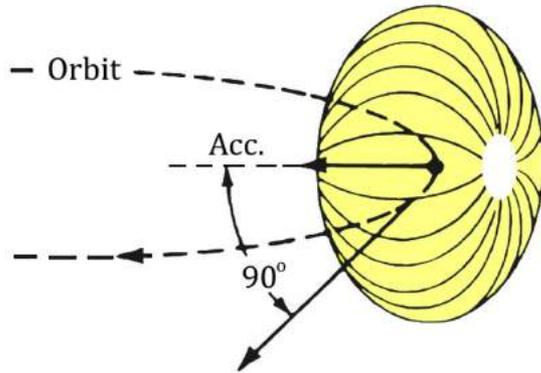
The critical energy,  $\epsilon_c$ , divides the spectrum into two parts of equal radiated power: 50% of the total power is radiated at energies lower than  $\epsilon_c$  and 50% at energies higher than  $\epsilon_c$ .



ASTRID (Aarhus - Denmark) <http://www.isa.au.dk/animations/pictures/pic-index.asp>

[http://www.isa.au.dk/animations/Finalmovie/astrid\\_total\\_v2.mov](http://www.isa.au.dk/animations/Finalmovie/astrid_total_v2.mov)

# Synchrotron radiation: physics



$$\beta \ll 1$$

$$v \ll c \text{ or } \beta = v/c \ll 1$$

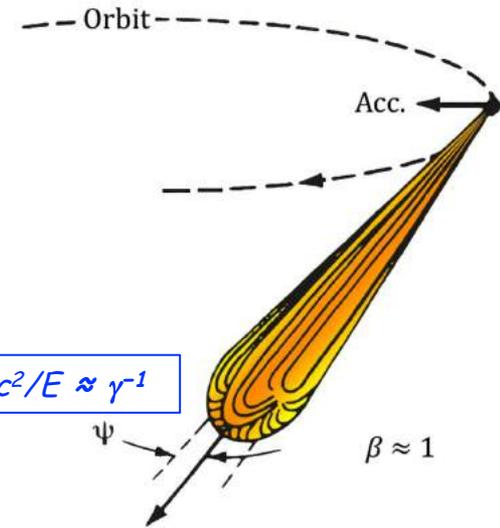
$$P = \frac{2}{3} \frac{e^2 a^2}{c^3} \text{ [W]}$$

$P$  = total emitted power,  $a$  = acceleration

At low electron velocity (non-relativistic) the radiation is emitted in a **non-directional pattern**.

**1897 Lamor:** calculates power radiated by an accelerated charged particle

**1898 Liénard:** extends the theory to relativistic particles in a circular path



$$\psi \approx mc^2/E \approx \gamma^{-1}$$

$$\beta \approx 1$$

$$v \approx c \text{ or } \beta = v/c \approx 1$$

For a relativistic effect, when the speed of the emitting electrons increases to relativistic values ( $v \approx c$ ) the radiation pattern is compressed into a **narrow cone in the direction of motion, resulting into an emission tangential to the particle orbit**.

$$P_{\text{rad}} = \frac{2}{3} \frac{Q^2 c}{R^2} \left[ \frac{E}{mc^2} \right]^4$$

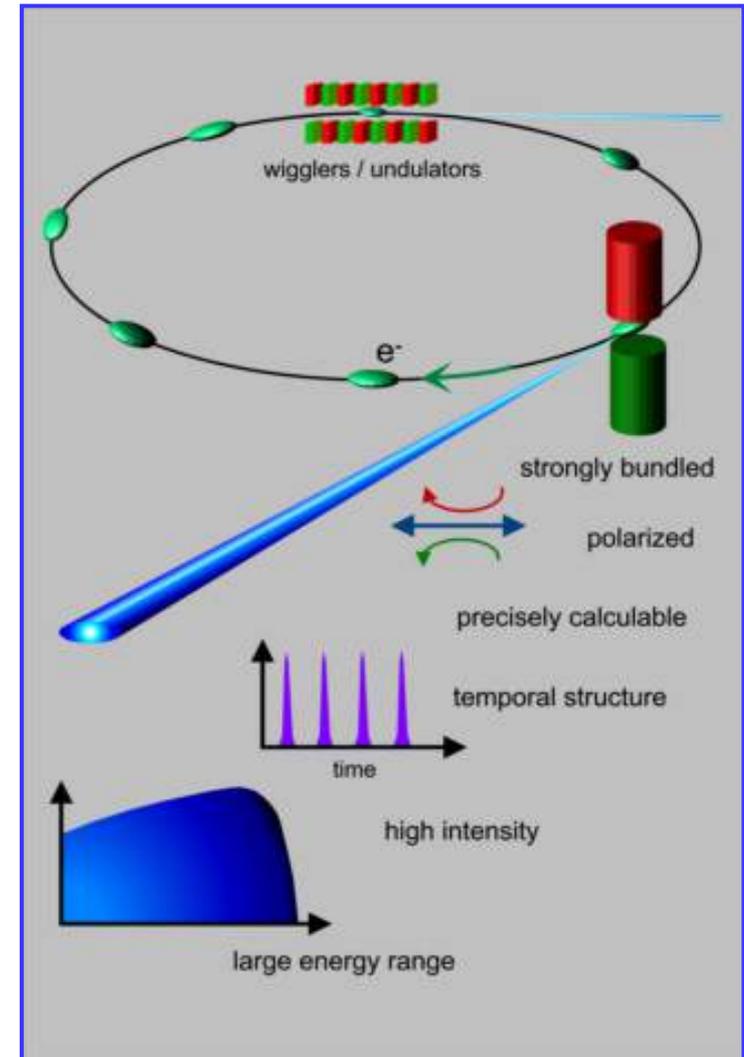
$E$  = particle energy,  $m$  = mass,  $R$  = radius of curvature

**1945 Schwinger:** classical theory of radiation from accelerated relativistic electrons

# Synchrotron Radiation Properties

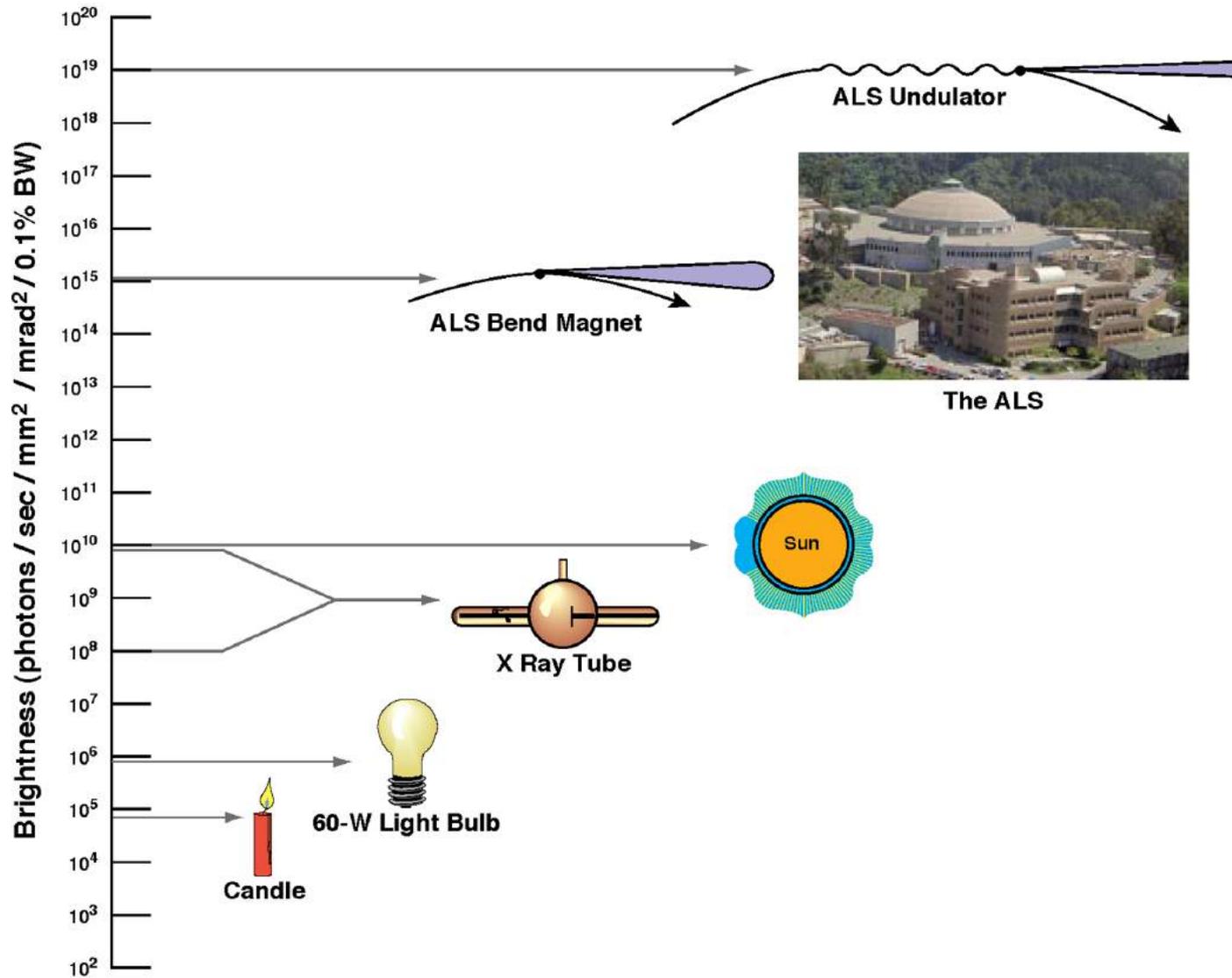
*What makes synchrotron radiation interesting, powerful and unique?*

- **Very high flux and brightness** (with **undulators**) highly collimated photon beam generated by a small divergence and small size source (partial coherence)
- **Broad spectral range (tunability)** which covers from microwaves to hard X-rays: the user can select the wavelength required for experiment- *continuous* (Bending Magnet/Wiggler) - *quasi-monochromatic* (Undulator)
- **Small source size**
- **Collimated beams**
- **High stability** (submicron source stability)
- **Pulsed time structure** - pulsed length down to tens of picoseconds allows the resolution of processes on the same time scale
- **Polarization** (linear, circular, elliptical with Insertion Devices)
- **High vacuum environment**

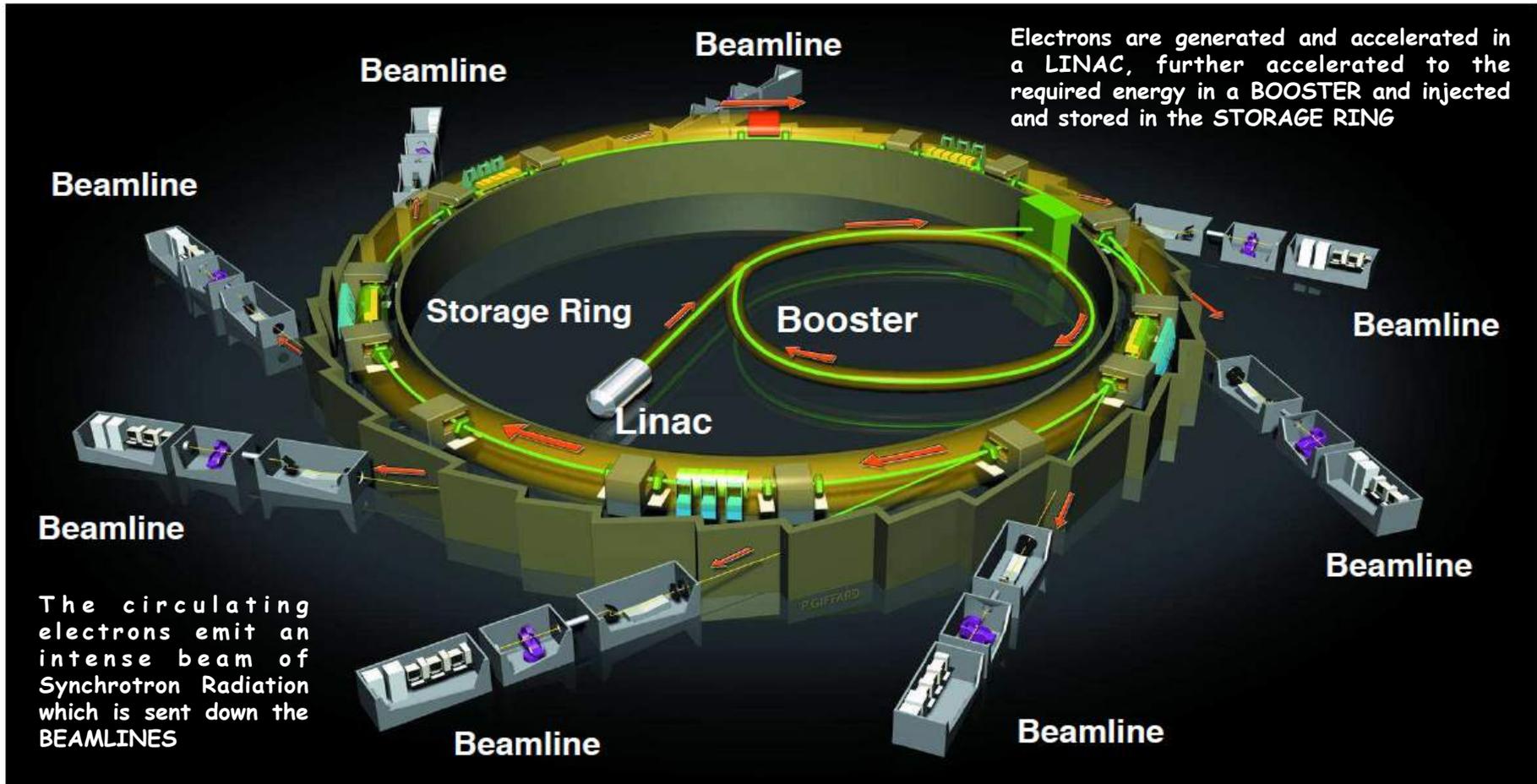


# How Bright Is the Advanced Light Source?

ALS



# Schematic view of a Synchrotron Radiation facility



As a function of the energy range to be used each beamline must be optimized for a particular field of research.

Beamline schematic composition:

- *Front end*
- *Optical hutch*
- *Experimental hutch*
- *Control and computing*

The *front end* isolates the beamline vacuum from the storage ring vacuum; defines the angular acceptance of the synchrotron radiation via an aperture; blocks (beam shutter) when required, the x-ray and Bremsstrahlung radiation during access to the other hutches.

# *Short history*

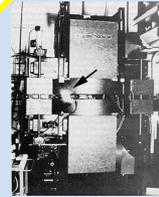
# Synchrotron Light Short History and Name

*Proof of concepts, tests of theories  
1897-1946*



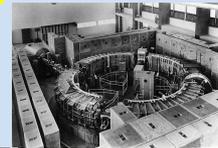
*J. Schwinger Nobel Prize 1965 Classical Relativistic quantum field theory*

*First observation of synchrotron radiation  
1947*



*General Electric Res. Lab. - 70 MeV Electro-Synchrotron ( N.Y. USA)*

*Parasitic use of electro-synchrotrons  
1961*



*Storage rings development  
1960s*



*ADA - B. Touschek - LNF*

*1<sup>st</sup> gen. dedicated ring Tantalus I (USA)  
1968*

*2<sup>nd</sup> gen. dedicated storage ring SRS (UK)  
1981*



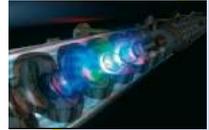
*3<sup>rd</sup> gen. dedicated storage ring ESRF (France)  
1994*



*Brightness increase*

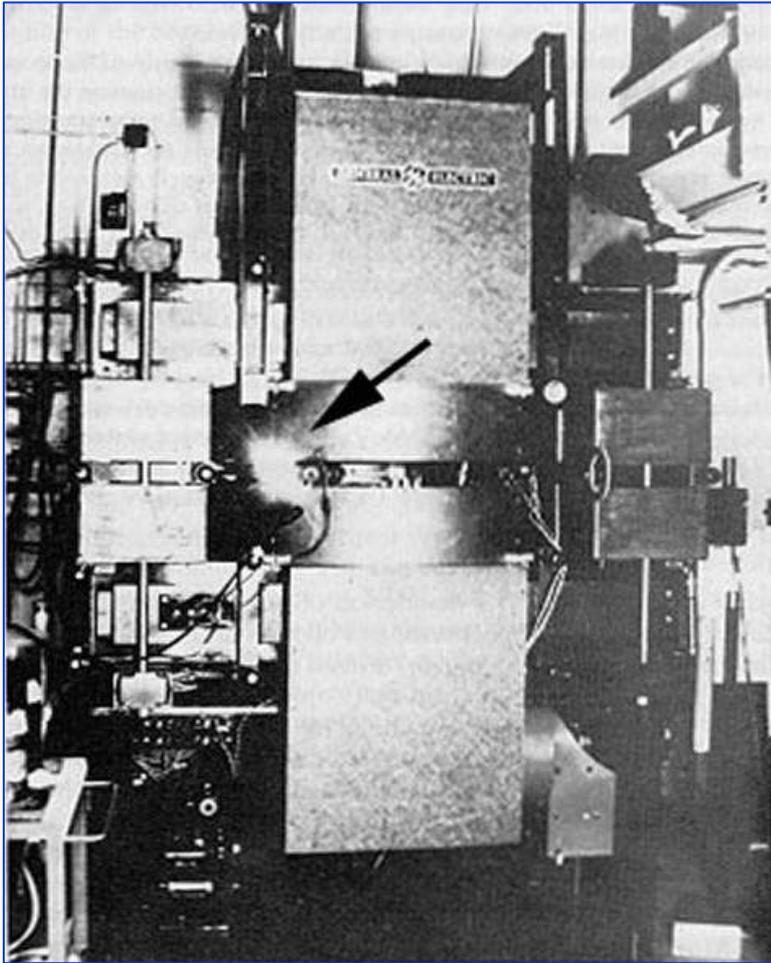
*3<sup>rd</sup>\* gen. ultimate storage rings like MAX IV  
(Sweden) near future*

*4<sup>th</sup> gen. - LINAC based accelerators  
FELs*



# Synchrotron radiation: history

## First generation: parasitic operation and storage rings

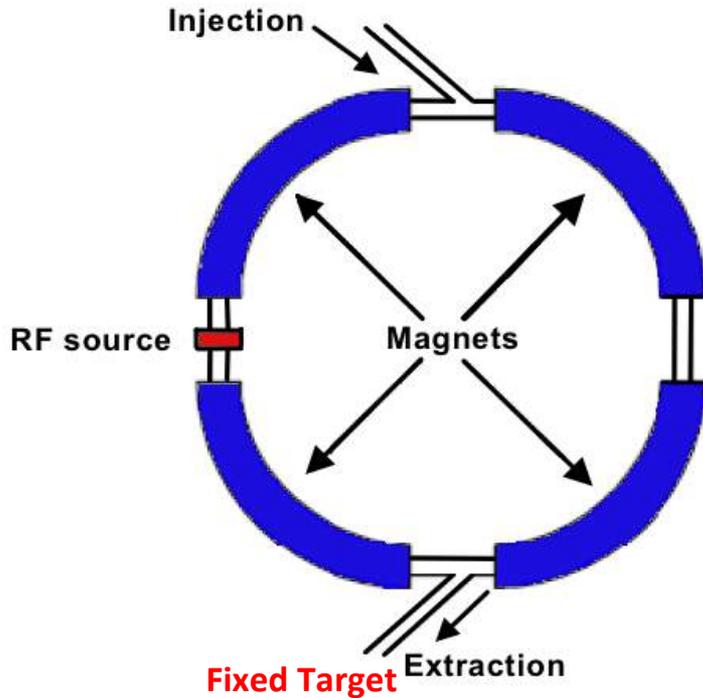


1947 General Electric Res. Lab. - 70 MeV Electron Synchrotron - N.Y. USA

*Starting point: Proof of concepts, tests of theories!*

- In the 50s and 60s machines built for High Energy Physics: synchrotrons (*1947 First 'visual observation of synchrotron radiation*).
- Synchrotron radiation was considered a *nuisance by particle physicists: unwanted but unavoidable loss of energy!*
- 1961 US National Bureau of Standards (now NIST) modified their electron synchrotron : *access to the synchrotron radiation users*.
- Synchrotron radiation scientists became *parasites* of nuclear physics experiments. (*1961 Frascati - CNEN Electro synchrotron - ( 0.4-1.1 GeV)*)
- 1968 *First storage ring dedicated* to synchrotron radiation research: *Tantalus* (University of Wisconsin) only *bending magnets*.

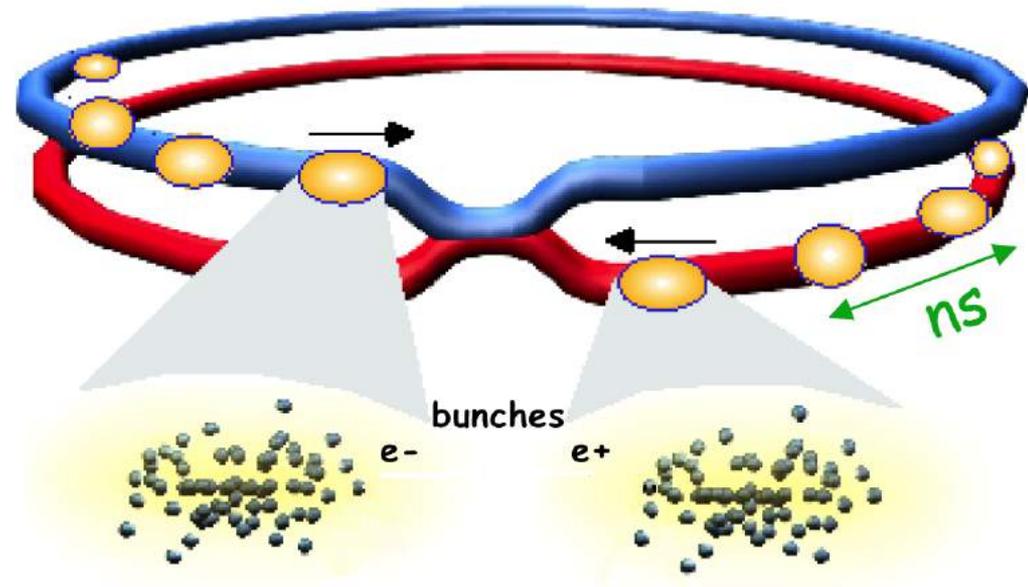
# Synchrotrons and Storage Rings



*Synchrotron*

Particle beam on fixed target

$$E_{CM} = (mE)^{1/2}$$



*Storage rings*

Colliding particle beams

$$E_{CM} = 2E$$

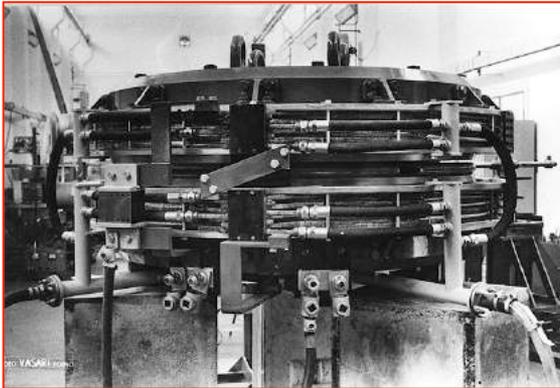
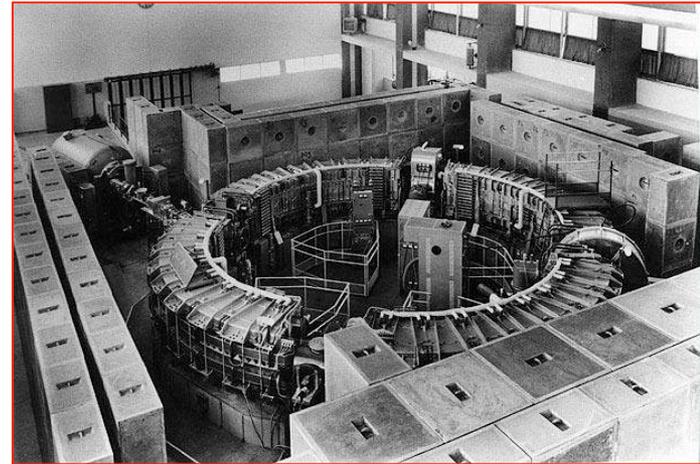
*Colliding beams more efficient*

*E = particle energy  $\gg mc^2$ ;  $E_{CM}$  = centre-of-mass energy*

# Synchrotron radiation: short history

## Frascati: Electro-Synchrotron, ADA and ADONE

Frascati - CNEN (Comitato Nazionale Energia Nucleare)  
Laboratory *Elettro-Sincrotrone* - (0.4-1.1) GeV,  $C = 28$   
m (1959-1975)



LNF *ADA* (Anello Di Accumulazione) - first electron-positron  
storage ring (proposed by B. Touschek) 0.25 GeV,  $C = 5$  m  
(1961-1964)

LNF *ADONE* (big ADA) electron-positron storage ring 1.5  
GeV per beam,  $C = 105$  m  
(1969-1993)



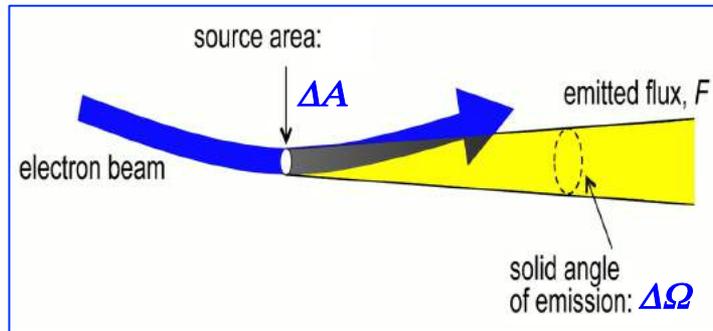
1976-1993 LNF *ADONE* 1.5 GeV parasitic/dedicated use  
for SR experiments after its use for HE experiments.

# Increasing brightness

**Brightness** (flux density in phase space) is an invariant and depends on the *size of the source* ( $\Delta A$ ) (electron beam) and on *the angular divergence of the radiation* ( $\Delta\Omega$ ), given by the convolution of the angular distribution of synchrotron radiation with the angular divergence of the electron beam.

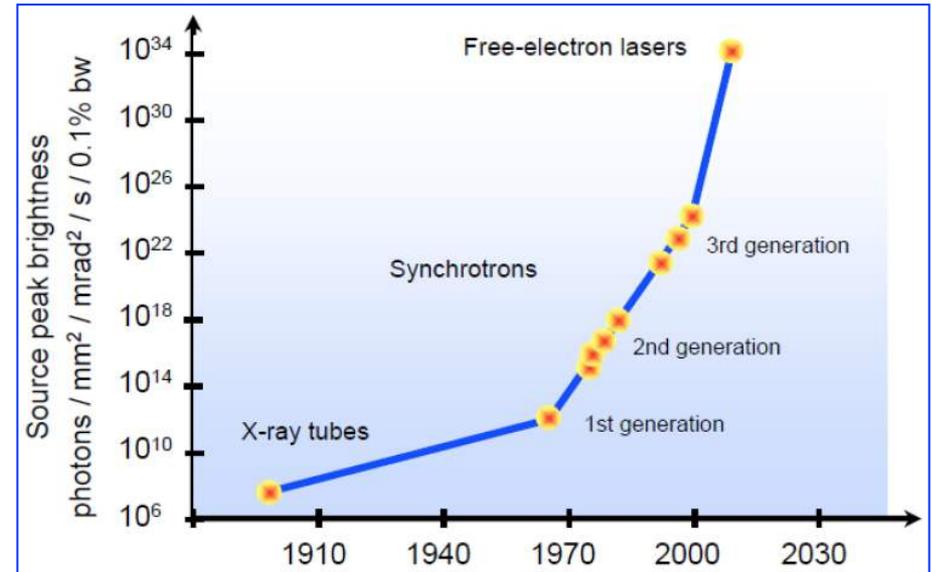
**Brightness more important than flux** (photons/s).

**Brightness = photon flux / [( $\Delta A$ ) ( $\Delta\Omega$ )]**



In a storage ring the *product of the electron beam transverse size* and *angular divergence* is a constant along the ring and is called *emittance* (*vertical* and *horizontal emittance*).

**Brightness is the main figure of merit of synchrotron radiation sources and its huge increase, was obtained designing low emittance machines, minimizing the source size and the beam divergence.**



**Increase of a factor 1000 every 10 years!!!**

$$\text{Spectral Brightness} = \frac{\text{photons}}{\text{second} \cdot \text{mrad}^2 \cdot \text{mm}^2 \cdot 0.1\% \text{BW}}$$

# *Synchrotron radiation: short history*

## *Third generation: optimized sources*

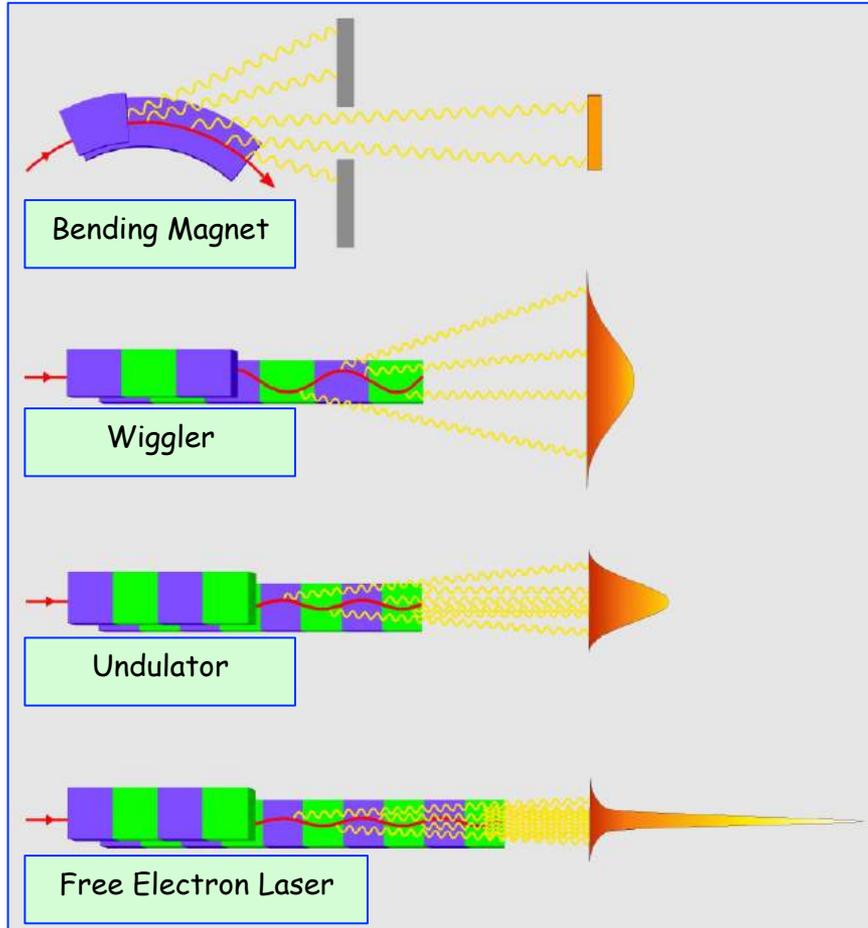
*Synchrotron light is now a unique tool for science!*



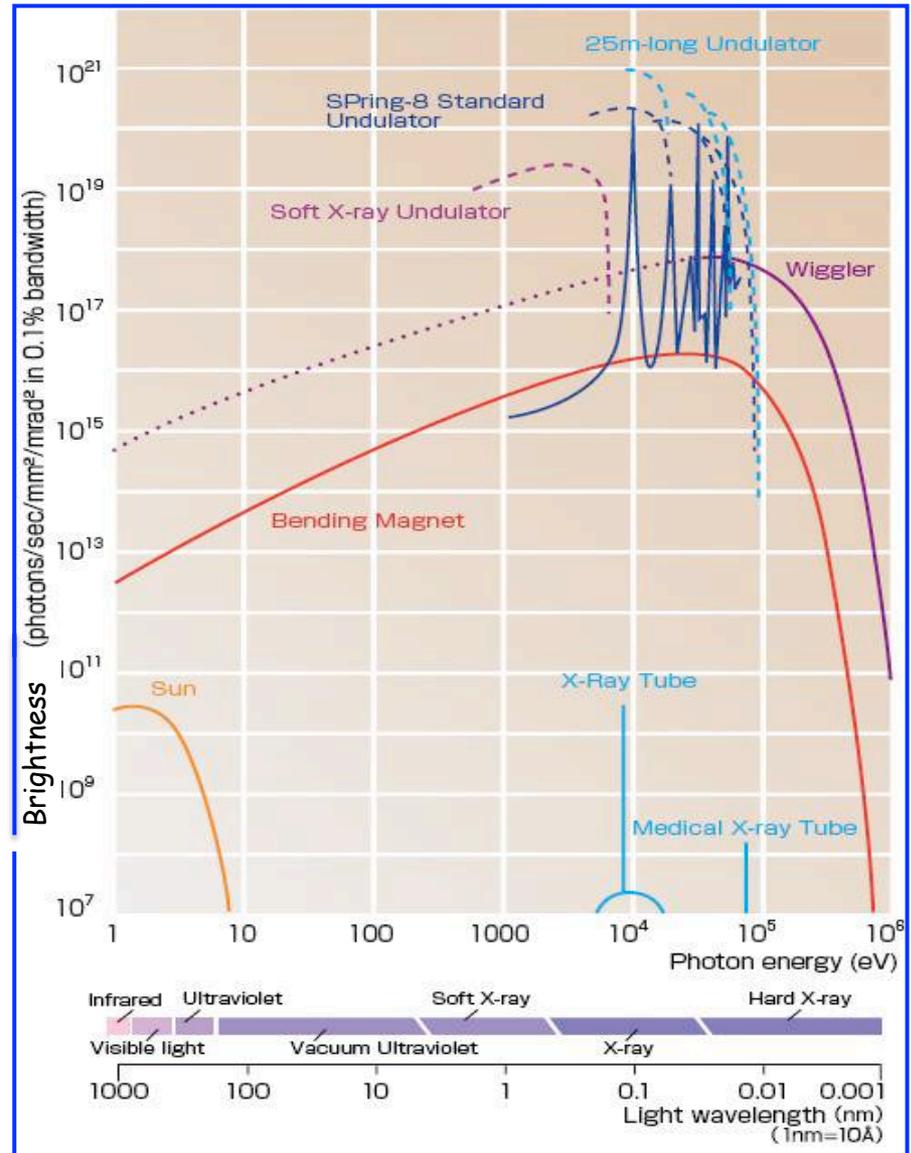
*ESRF, Grenoble - France 6 GeV,  $C = 844$  m  
opened to users in 1994*

- Sources designed specifically for high brightness or low emittance.*
- Emphasis on research with insertion devices like undulators!*
- High-energy machines able to generate hard x-rays*
- Larger facilities to support rapidly growing user community, many beamlines high number of users.*

# Comparing the achievable brightness



Courtesy DESY

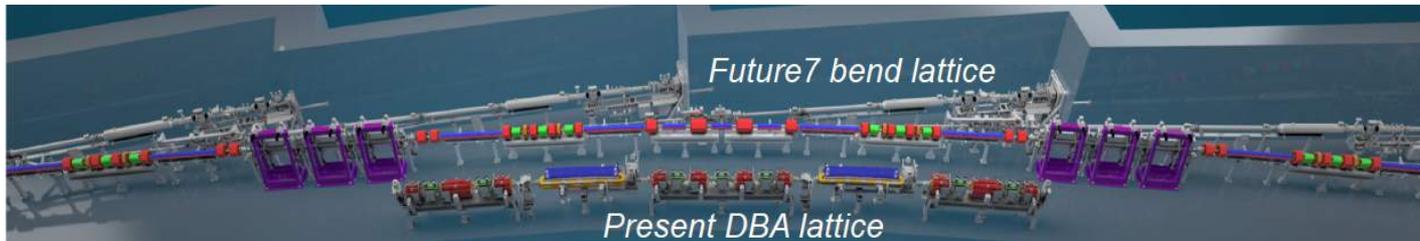


Courtesy SPring-8

# Synchrotron radiation: history

## Present and future: Ultimate Storage Rings

Brightness and transverse coherence increase in the X-ray range with implementation of **low emittance lattices** (multi-bend achromat schemes).



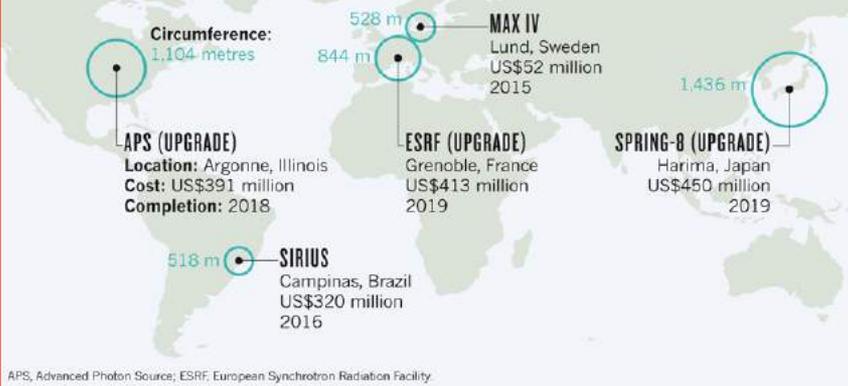
$$\epsilon_{x_0, min} \propto \frac{E^2}{N_B^3}$$

*J. Jacob, Status of the ESRF operation & upgrade, 2013*

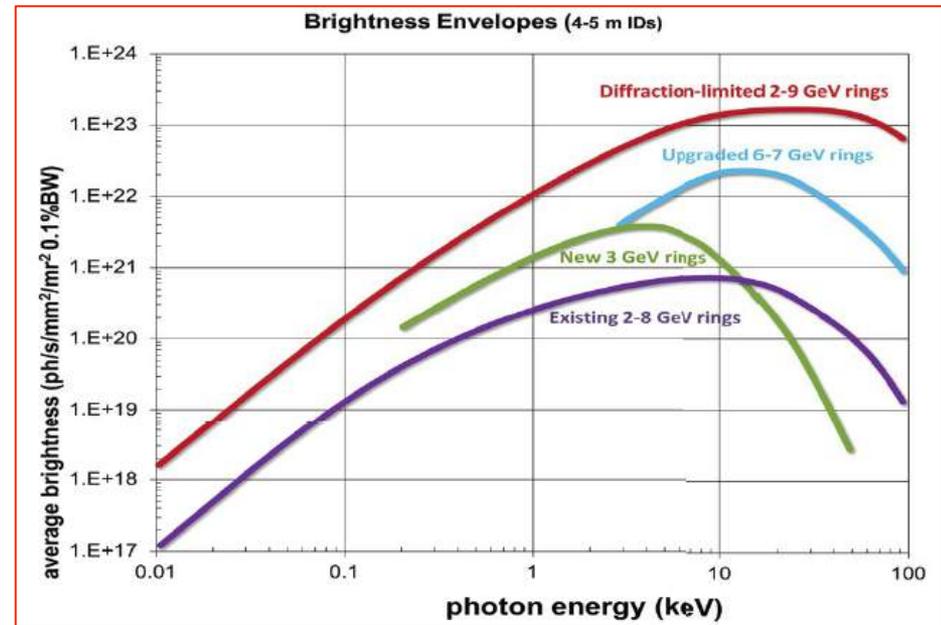
ESRF  $\epsilon_x = 4$  nm;  $\epsilon_x = 5$  pm in the future  $\epsilon_x = 0.16$  nm

### FOCUSED BEAMS

Five synchrotron facilities are developing special magnets so that they can become ultimate storage rings.



*E.S. Reich, Ultimate upgrade for US synchrotron, Nature, 2013*



*H. Owen - Univ. of Manchester (UK)*

# *3rd Generation SR Light Sources*



*ESRF - France*



*DIAMOND - UK*



*ALBA - Spain*

## *Under construction - Ultimate SR facilities*



**Max-IV**

*Lund - Sweden*



*Sirius - Brazil*



**SSRF**

*Shanghai - China*

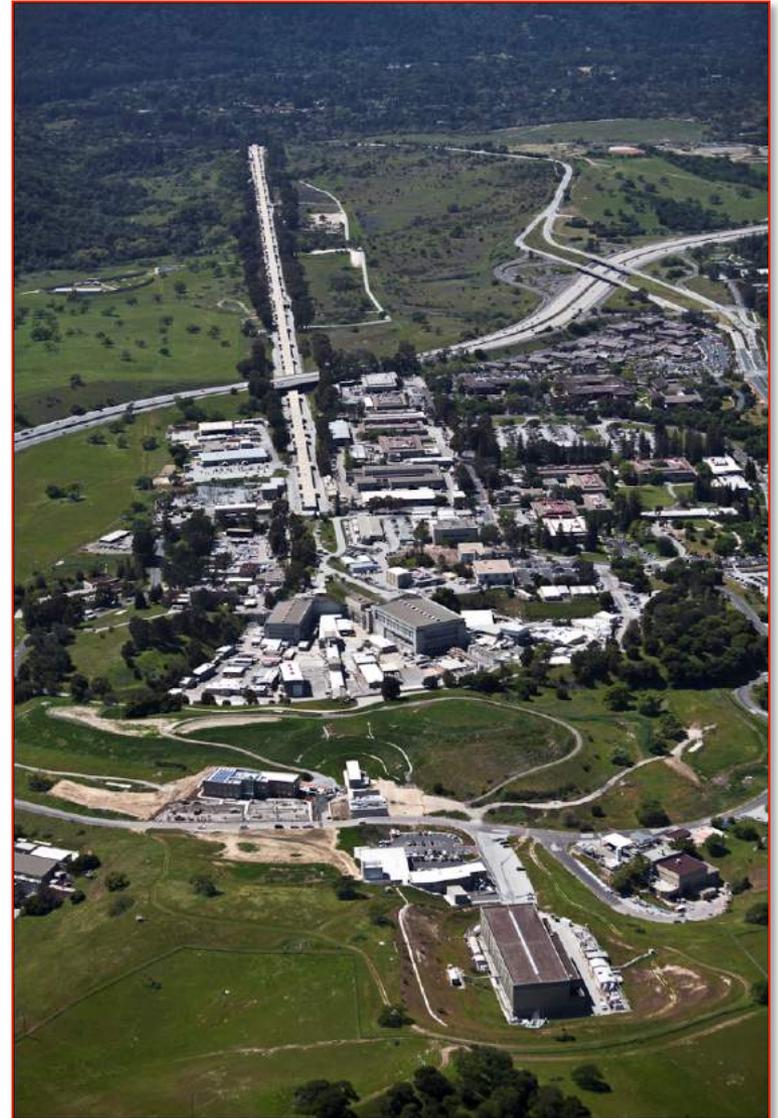
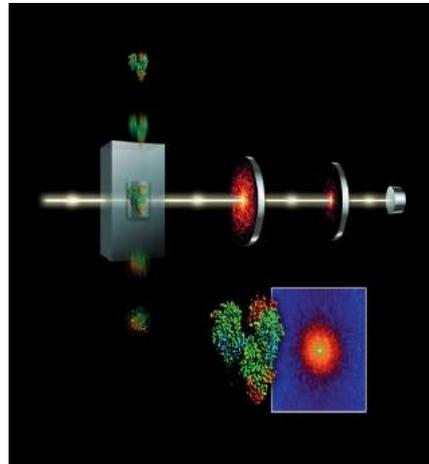
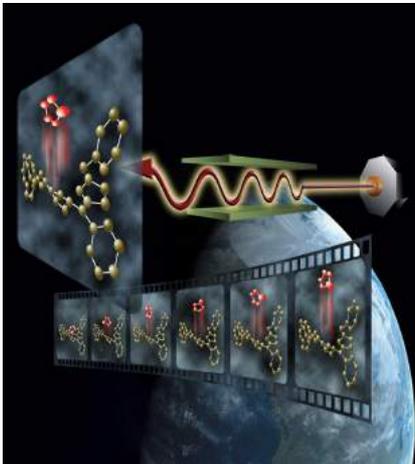
# Fourth generation: X-ray free electron lasers XFELs

*Electrons in a FEL are not bound to atoms or molecules. The "free" electrons traverse a series of alternating magnets, called "undulators," and radiate light at wavelengths depending on electrons energy, undulator period and magnetic field.*

# *Synchrotron radiation: history*

## *Fourth generation: LINAC based sources and Free Electron Lasers*

- Extremely bright and coherent sources*
- Ultrafast pulses*
- Already working in IR to UV and X-ray (LCLS April 2009) ranges*
- European XFEL started activity in 2017*
- Filming chemical reactions as they occur*
- Protein crystallography no longer needed -image molecules directly*



# XFELs present and future

LCLS-I, II 2009, 2019  
14.5 GeV, 120 Hz NC



SACLA 2011  
8.5 GeV, 60 Hz NC



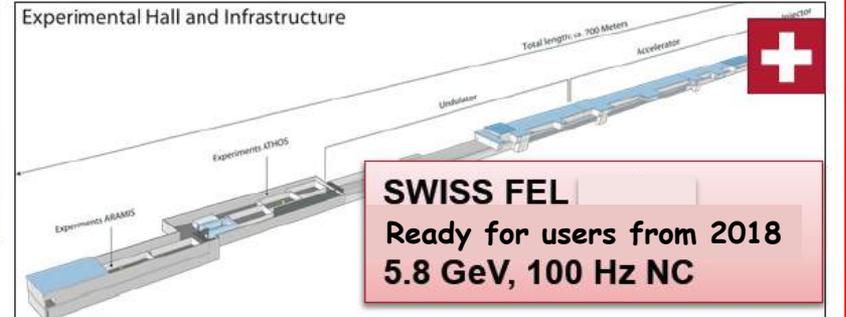
PAL XFEL  
Operational from 2017  
10 GeV, 100 Hz NC



XFEL Operational from 2017  
17.5 GeV, 2700 x 10 Hz SC



Experimental Hall and Infrastructure



SWISS FEL  
Ready for users from 2018  
5.8 GeV, 100 Hz NC

NC: normal conducting acceleration, SC: super conducting acceleration

***FUTURE***

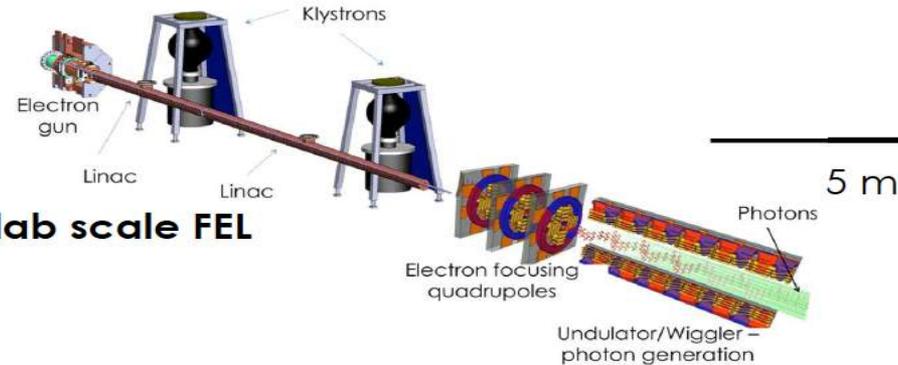
***Fifth Generation Light Sources***

***... re-invent XFELs to fit in campus laboratory.***

The X-ray FEL is enormous. Challenge is to fit this...



Into this...



**The university lab scale FEL**

*J. Rosenzweig, UCLA - Fifth Generation Light Sources X-ray FELs Based on New Accelerator and Undulators - 2014*

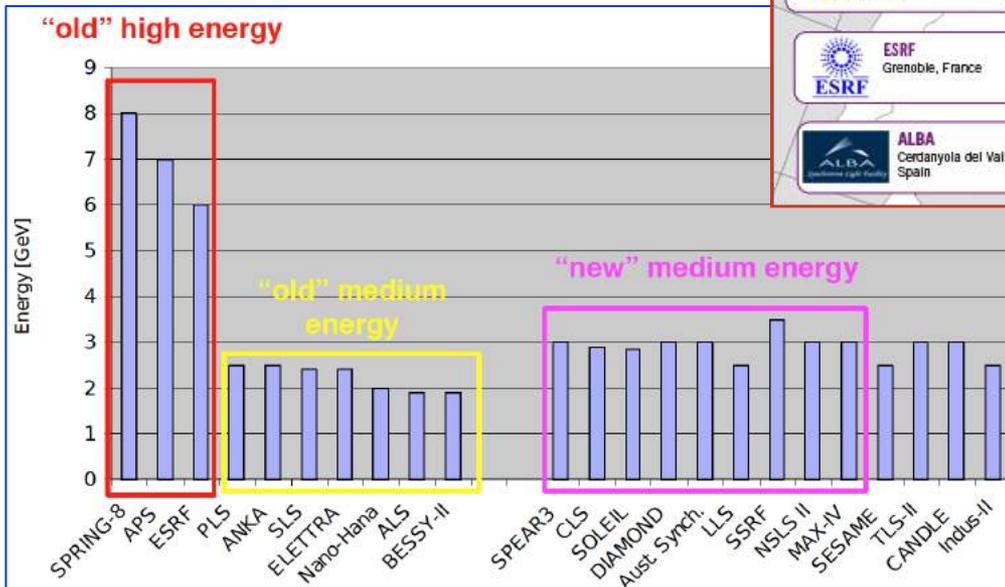
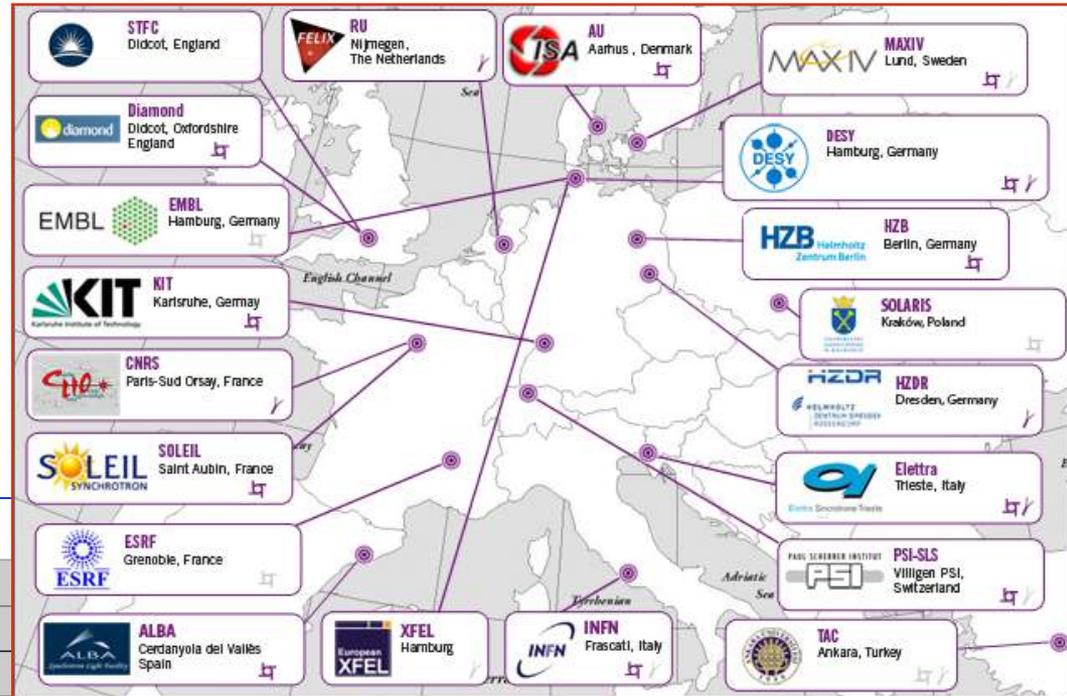
*Ongoing research on laser/plasma/wakefield accelerators, high frequency, high repetition rate linacs and electron beam injectors can lead in the future to very compact, university scale, X-ray FELs.*

*C. Pellegrini, UCLA - 5th Generation light sources - 2011*

*SR facilities*

# Synchrotron radiation facilities

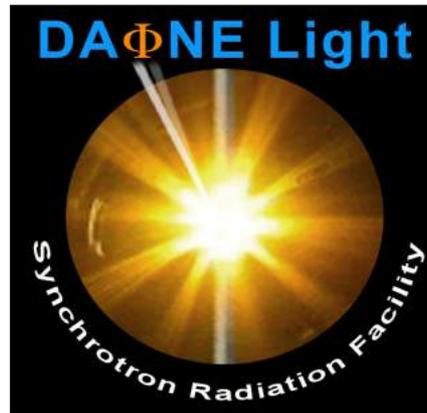
18 in America  
25 in Asia  
25 in Europe  
1 in Oceania  
including facilities under design and  
FELS



EUROPE

Info on European Synchrotron Radiation Facilities: [www.wayforlight.eu](http://www.wayforlight.eu)  
About 67 operational Synchrotron Radiation Facilities Around the World information on: [www.lightsources.org](http://www.lightsources.org)

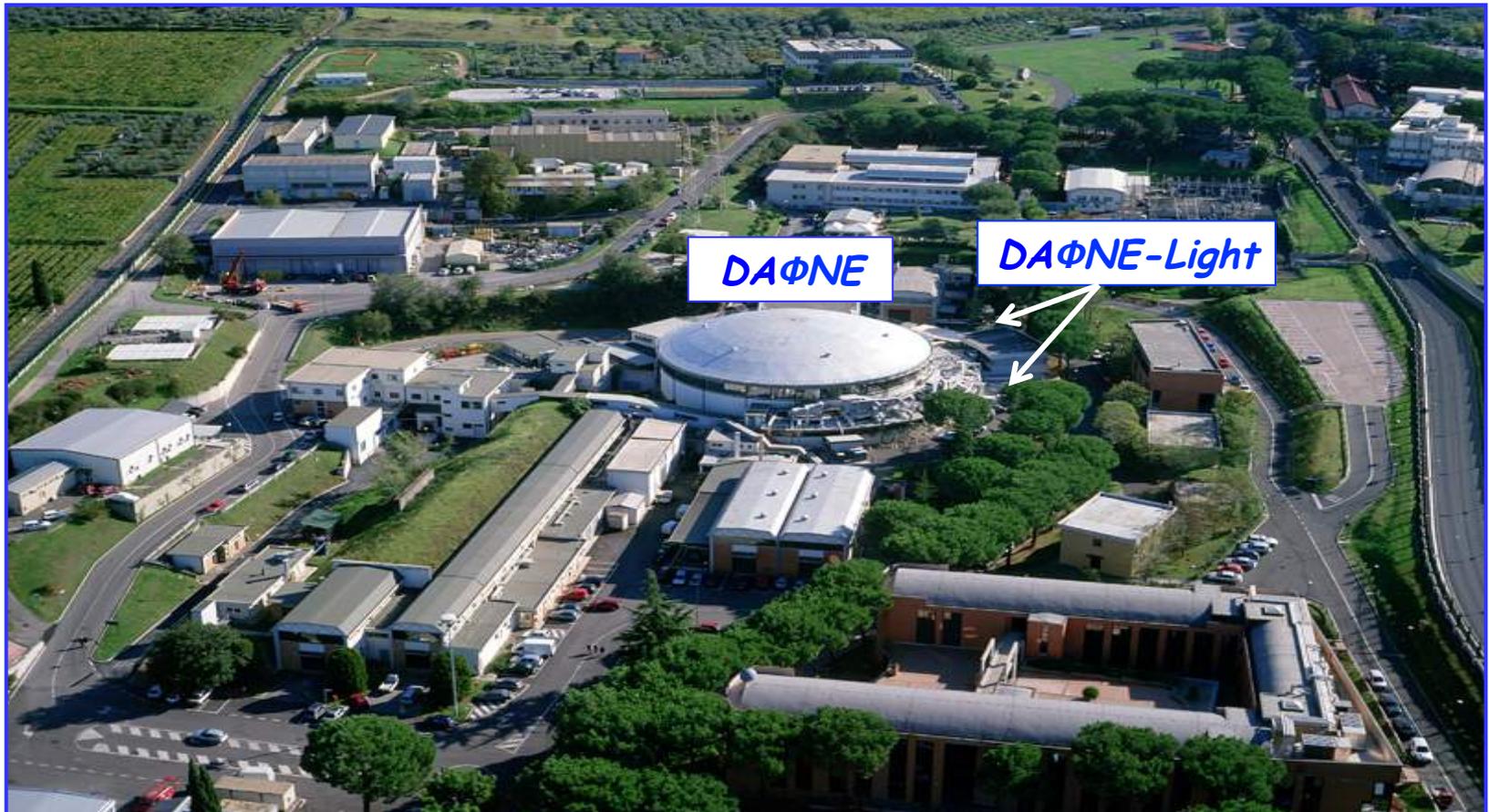
# *Synchrotron radiation @ INFN-Frascati National Laboratory*



# *DAΦNE-Light*



## *INFN-LNF Synchrotron Radiation Facility*



# Beamlines @ DAΦNE

1) *SINBAD - IR beamline (1.24 meV - 1.24 eV)*

2) *DXR1- Soft x-ray beamline (900-3000 eV)*

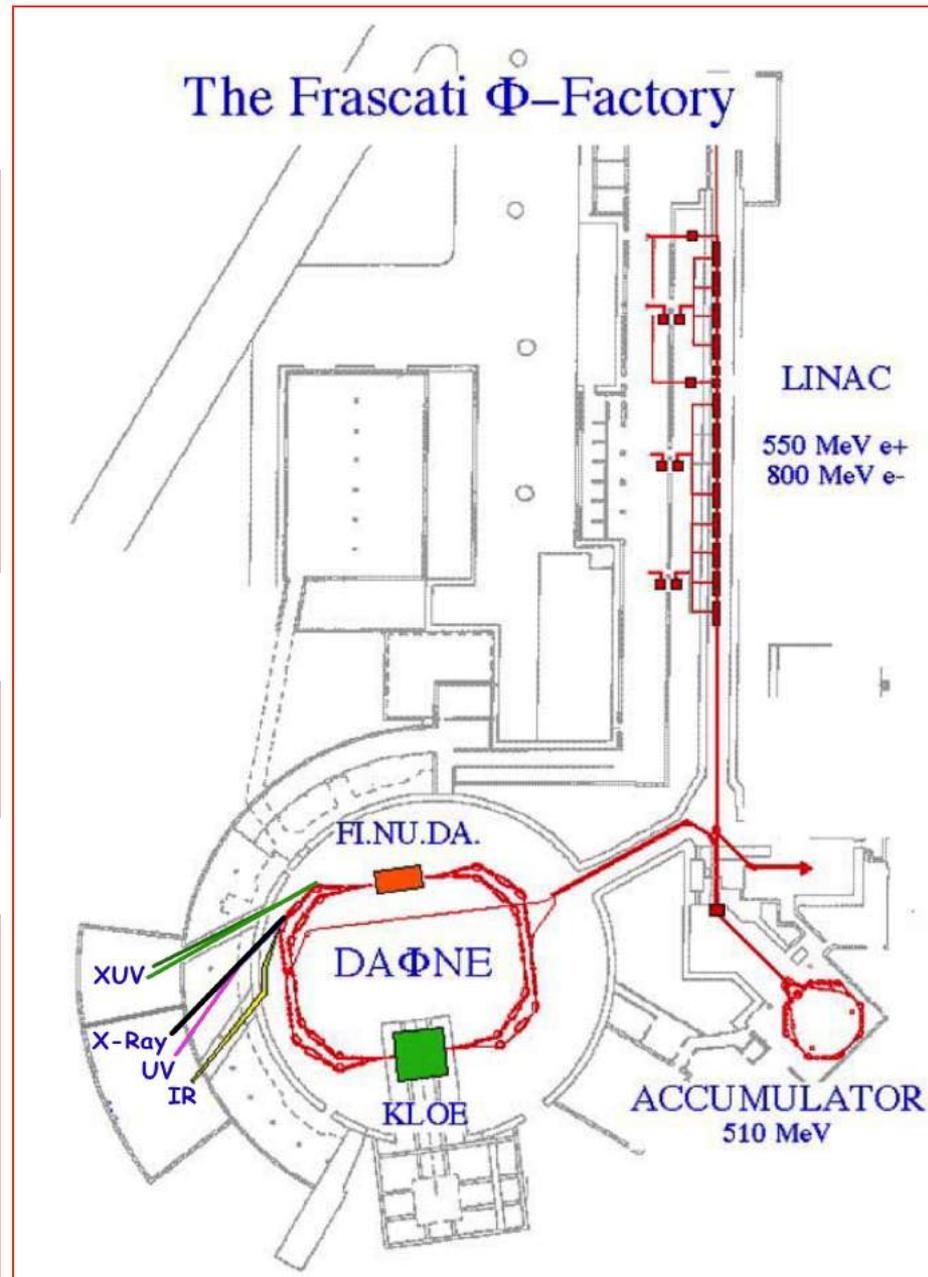
*Open to Italian and EU users*

3) *DXR2 - UV-VIS beamline (2-10eV) new setup.*

## *XUV beamlines*

4) *Low Energy Beamline (35-200 eV) under commissioning;*

5) *High Energy Beamline (60-1000eV) under commissioning.*

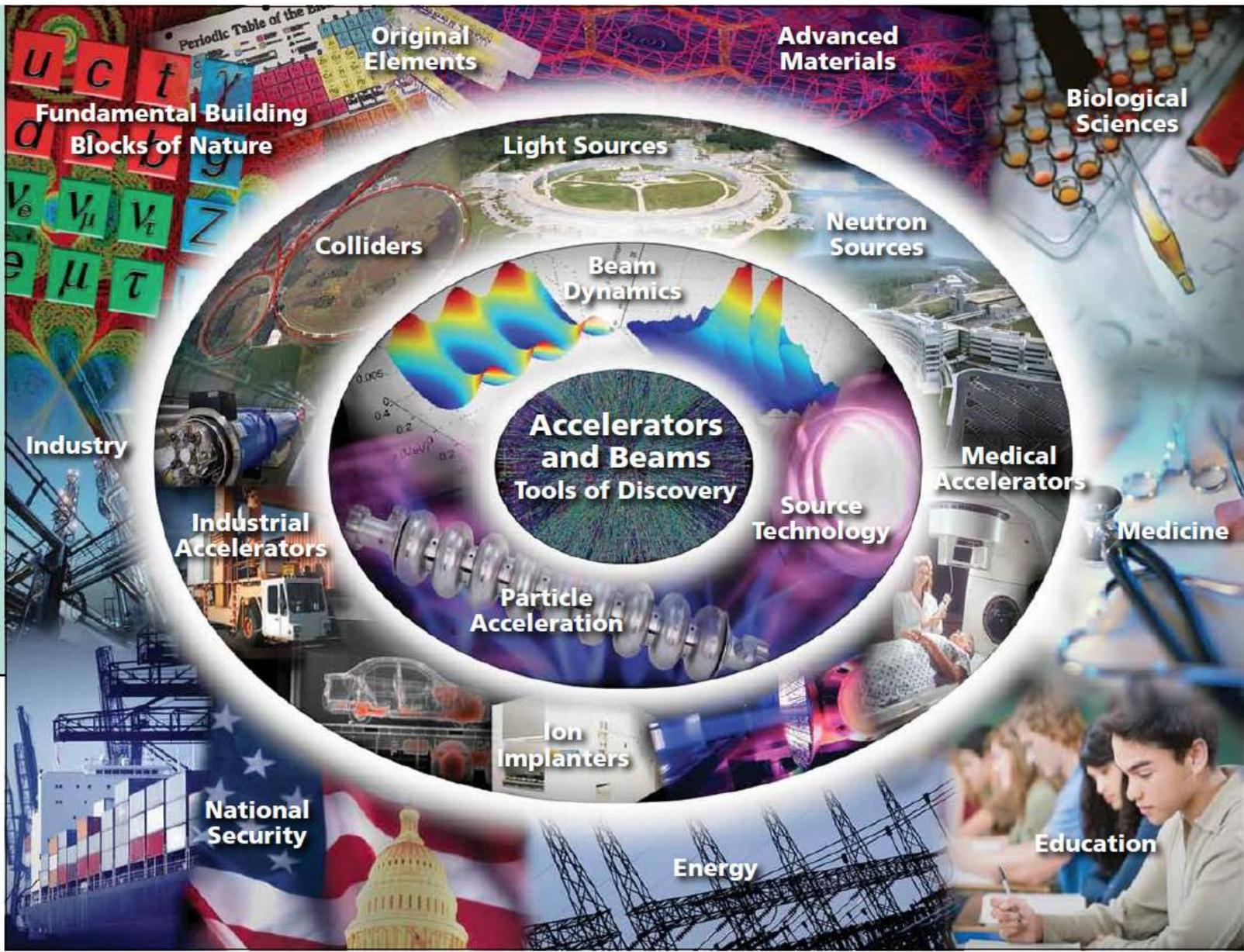


# *Available techniques*

- *FTIR spectroscopy, IR microscopy and IR imaging*
- *UV-Vis absorption spectroscopy*
- *Photochemistry: UV irradiation and FTIR micro-spectroscopy and imaging.*
- *Soft x-ray spectroscopy: XANES (X-ray Absorption Near Edge Structure) light elements from Na to S*
- *SEY (secondary electron yield) and XPS (X-ray photoelectron spectroscopy) - by electron and photon bombardment*

*From accelerators to applications*

COURTESY OAK RIDGE NATIONAL LABORATORY.



# *X-Ray Interaction with Matter*

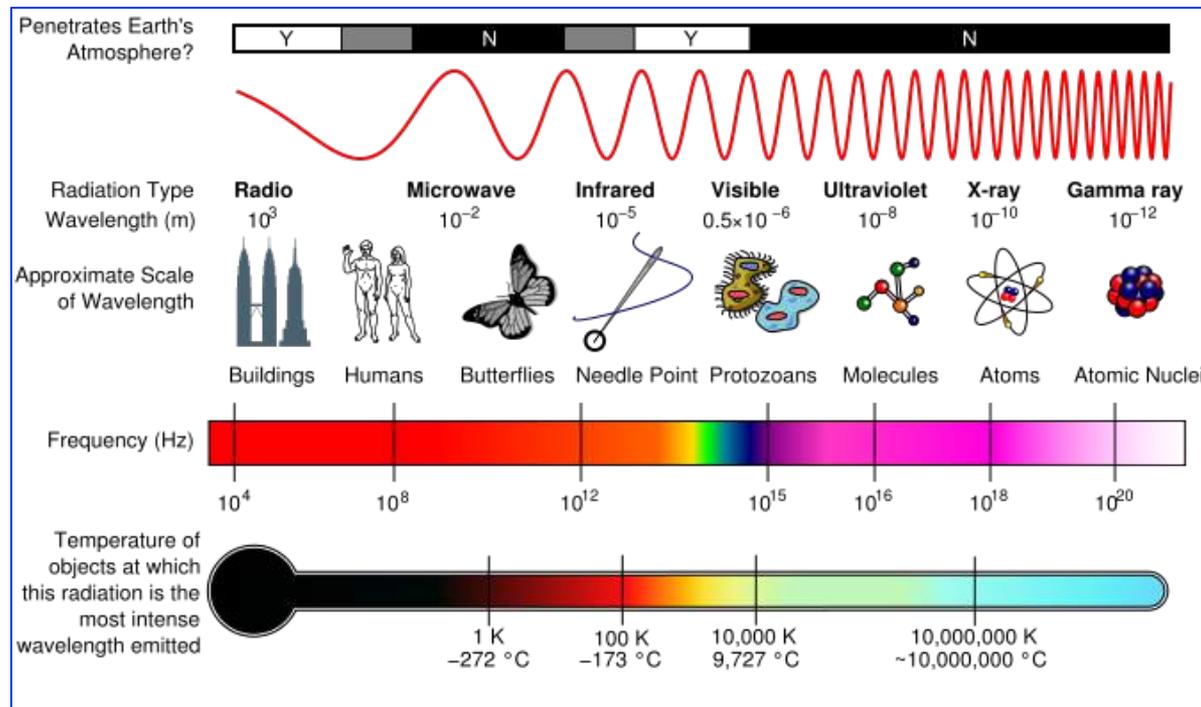
Light waves across the electromagnetic spectrum behave in similar ways.

When light waves encounter objects, they are either transmitted, reflected, absorbed, refracted, polarized, diffracted, or scattered depending on the composition of the object and the wavelength of the light.

# Interaction of radiation with matter

The different kind of radiation (IR, VIS, UV, X) have very **different interaction with matter.**

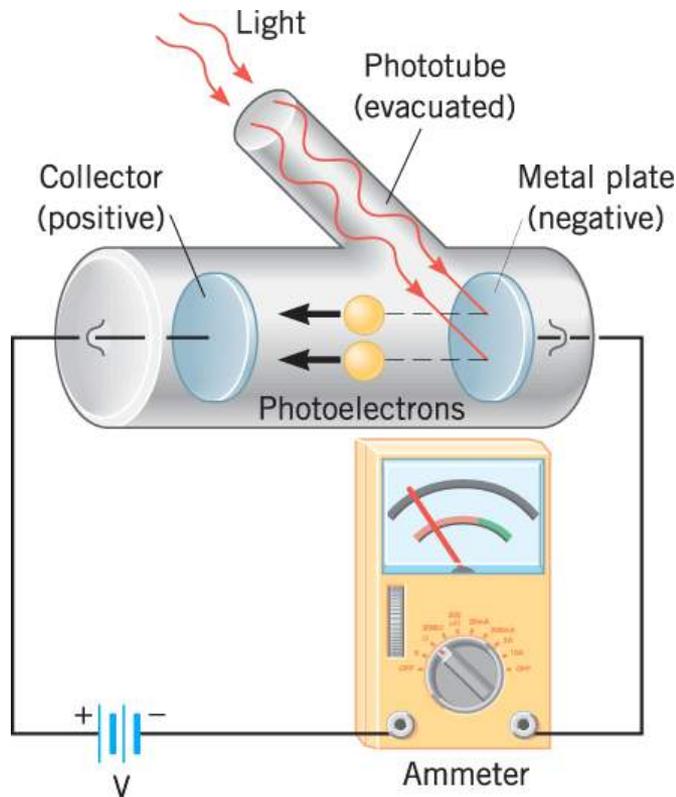
The human body is quite transparent to low frequency radio waves. You can listen to your radio at home since the radio waves pass freely through the walls of your house.



Moving upward through microwaves and infrared to visible light, UV and X-rays radiation is absorbed more and more strongly.

# Photoelectric effect

Electromagnetic radiation can be used to **knock out electrons, freeing them from the surface of a solid**. This process is called the **photoelectric effect** (or **photoelectric emission** or **photoemission**), a material that can exhibit this phenomena is said to be **photoemissive**, and the ejected electrons are called **photoelectrons**; but there is nothing that would distinguish them from other electrons. All electrons are identical to one another in mass, charge, spin, and magnetic moment.



The photoelectric effect does not occur when the frequency of the incident light is less than the **threshold frequency**. **Different materials have different threshold frequencies.**

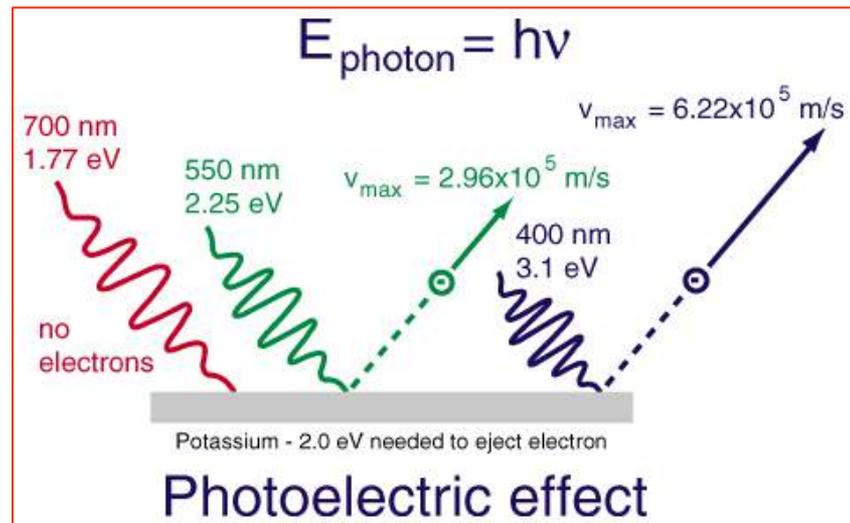
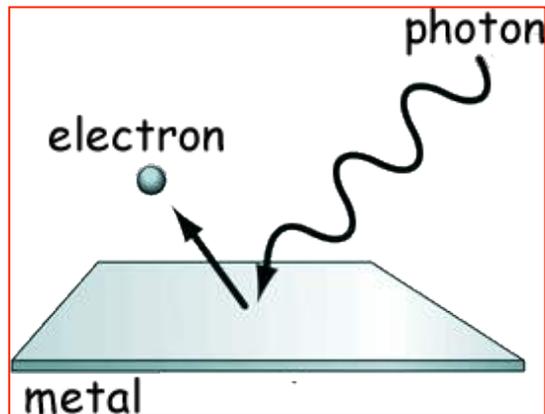
# Photoelectric effect

## Classical physics cannot explain why...

- no photoelectrons are emitted when the incident light has a frequency below the threshold,
- the maximum kinetic energy of the photoelectrons increases with the frequency of the incident light.

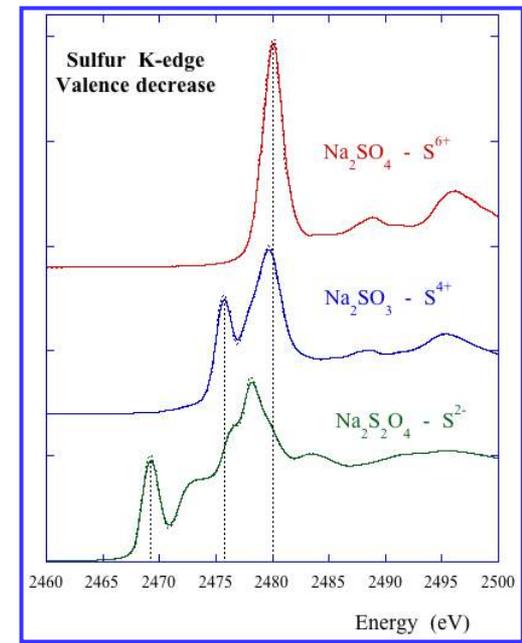
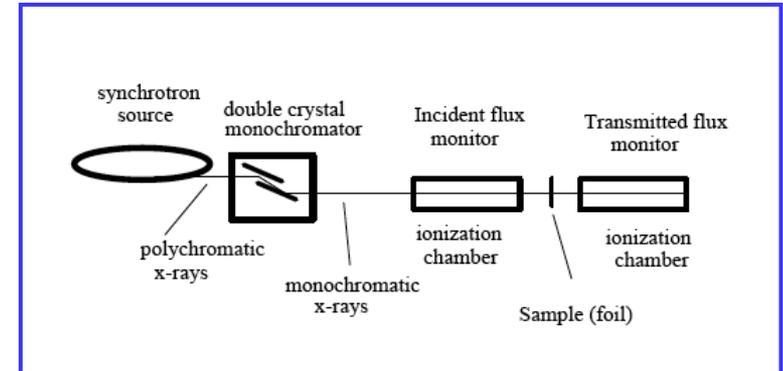
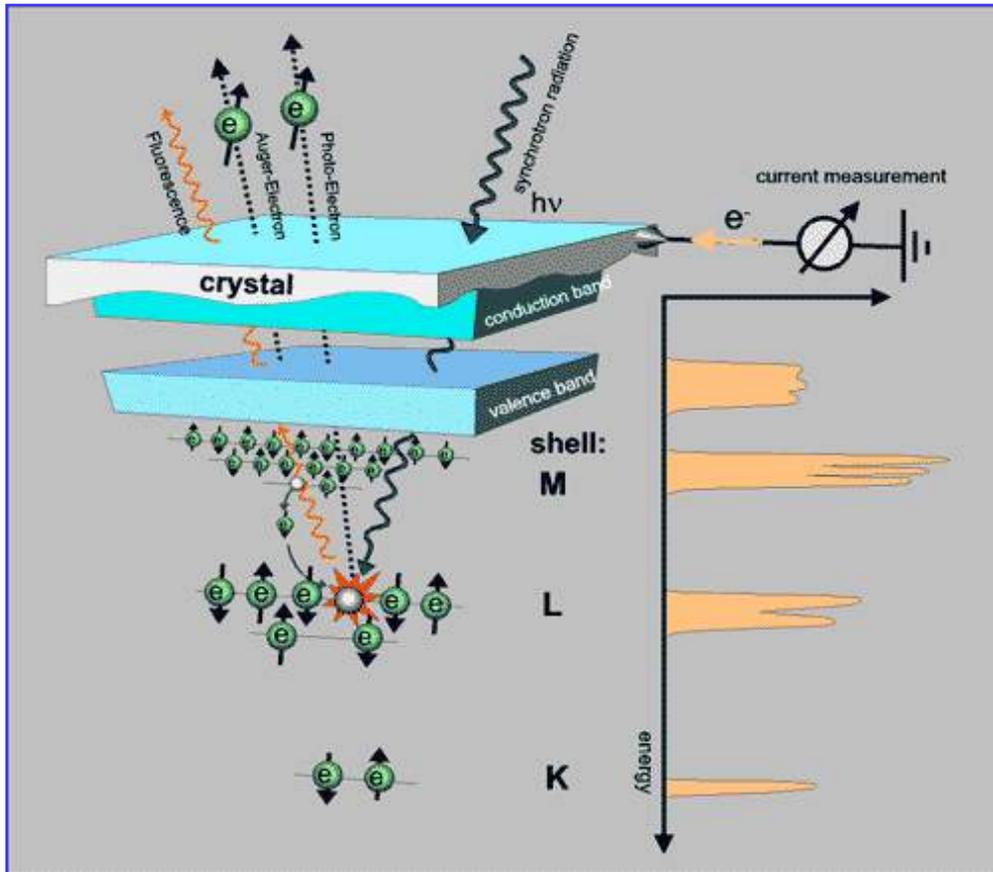
## Modern physics states that...

- electromagnetic radiation is composed of discrete entities called photons
- the energy of a photon is proportional to its frequency
- the work function of a material is the energy needed per photon to extract an electron from its surface



In 1905, *Albert Einstein realized that light was behaving as if it was composed of tiny particles* (initially called quanta and later called photons) and that *the energy of each particle was proportional to the frequency of the electromagnetic radiation* (**Nobel Prize in Physics in 1921**).

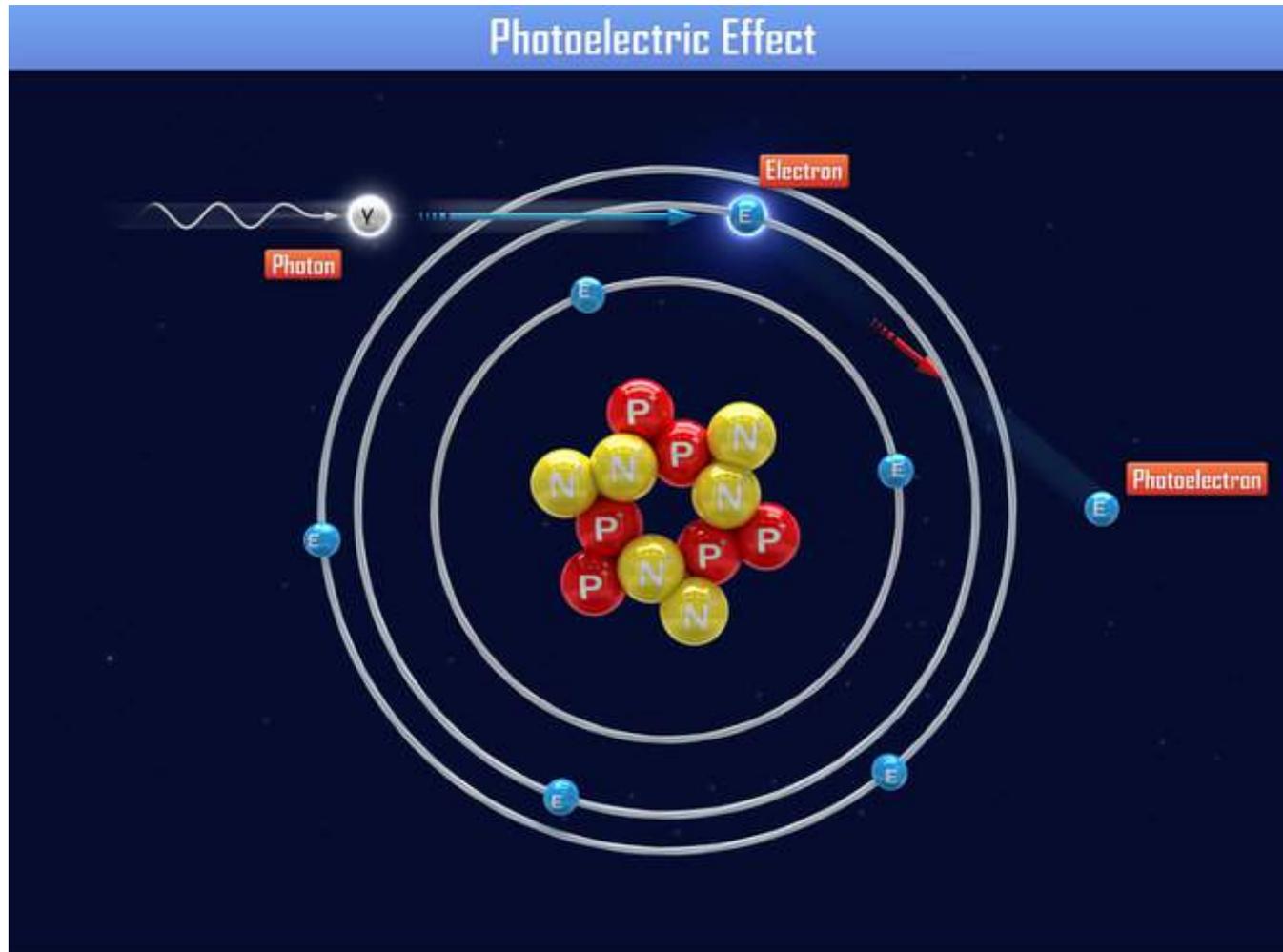
# Photoelectric effect and X-rays



X-ray absorption spectra at the Sulfur K edge - XANES

# *X-ray Absorption and Fluorescence Decay*

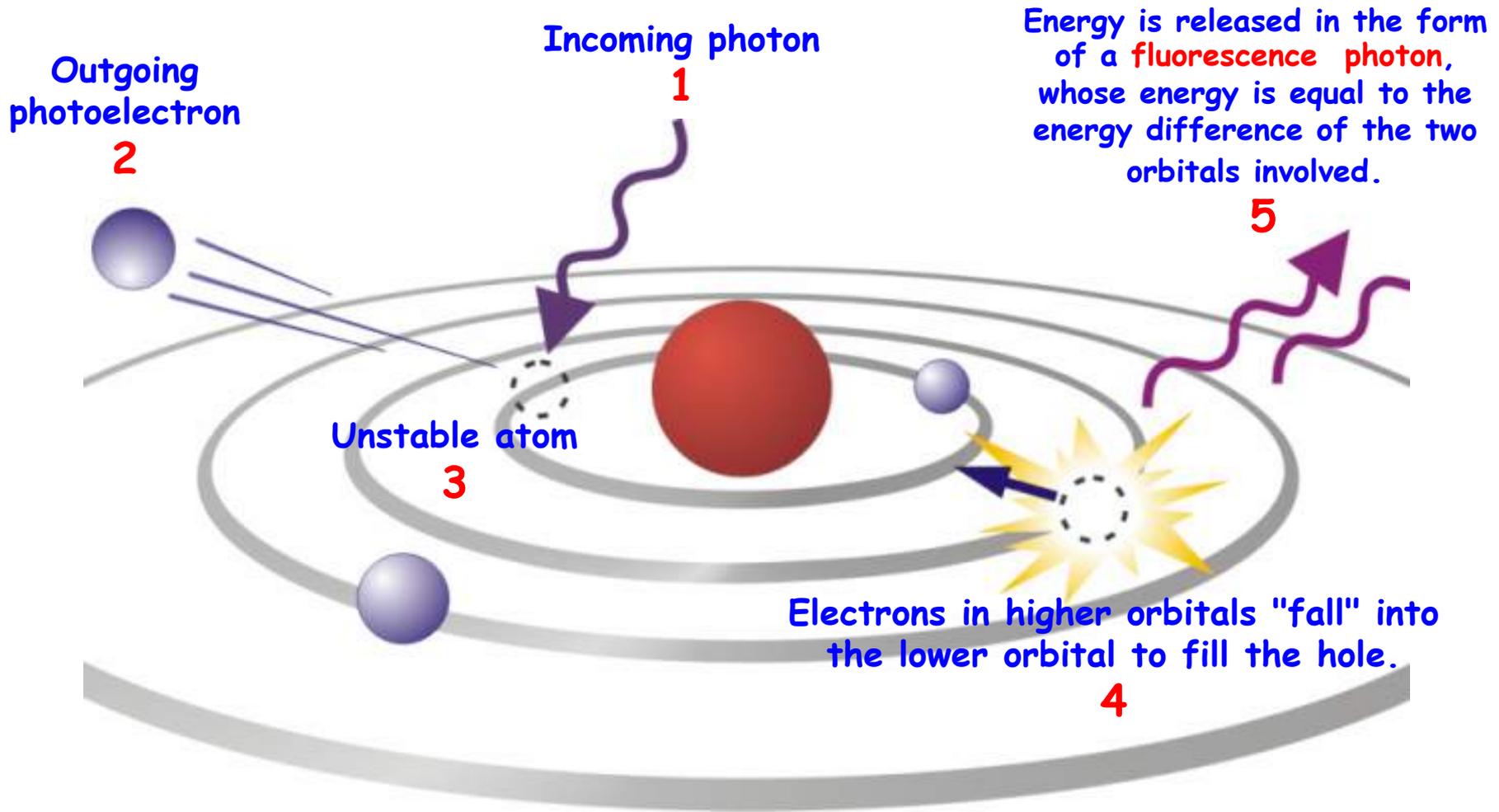
# Photoelectric effect



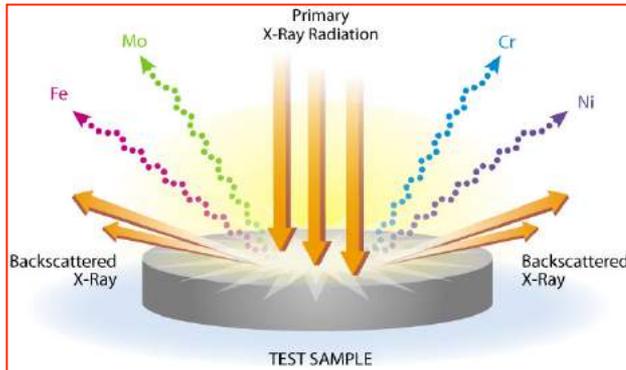
A photon (Y) strikes an electron (E), it knocks it loose and creates a photoelectron.

Credit: general-fmv

# Absorption and fluorescence decay



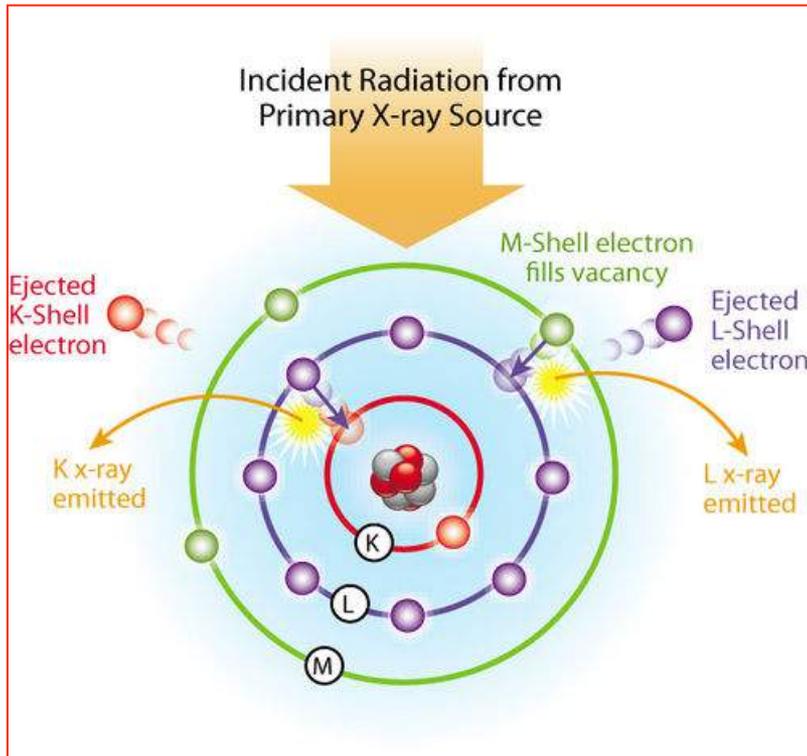
# Decay Process: X-ray Fluorescence



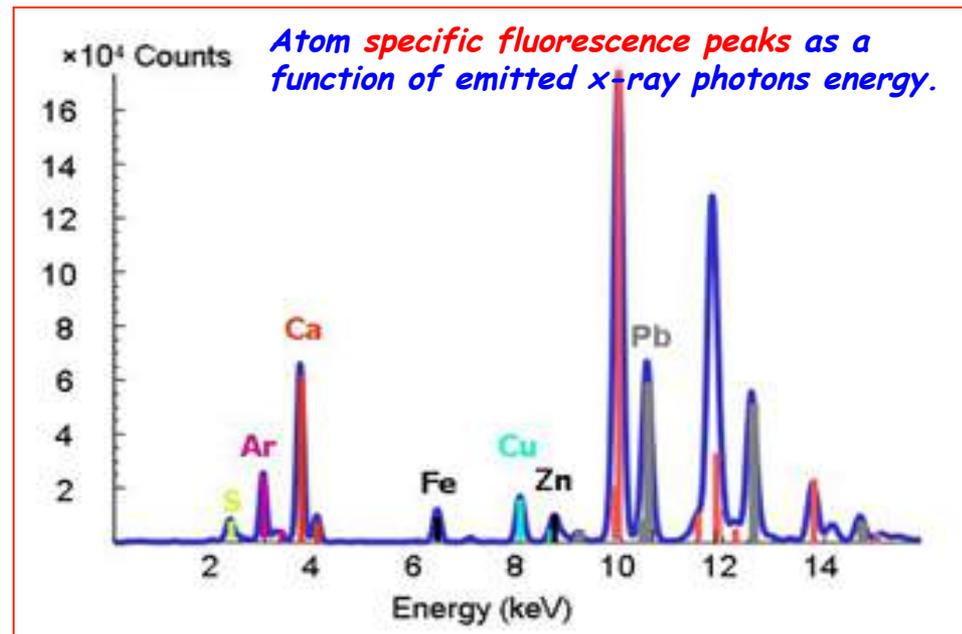
X-ray fluorescence spectrometry (XRF) can be used to accurately **measure the atomic composition of a material.**

In an atom, the electrons orbit around the nucleus in characteristic patterns referred to as "shells."

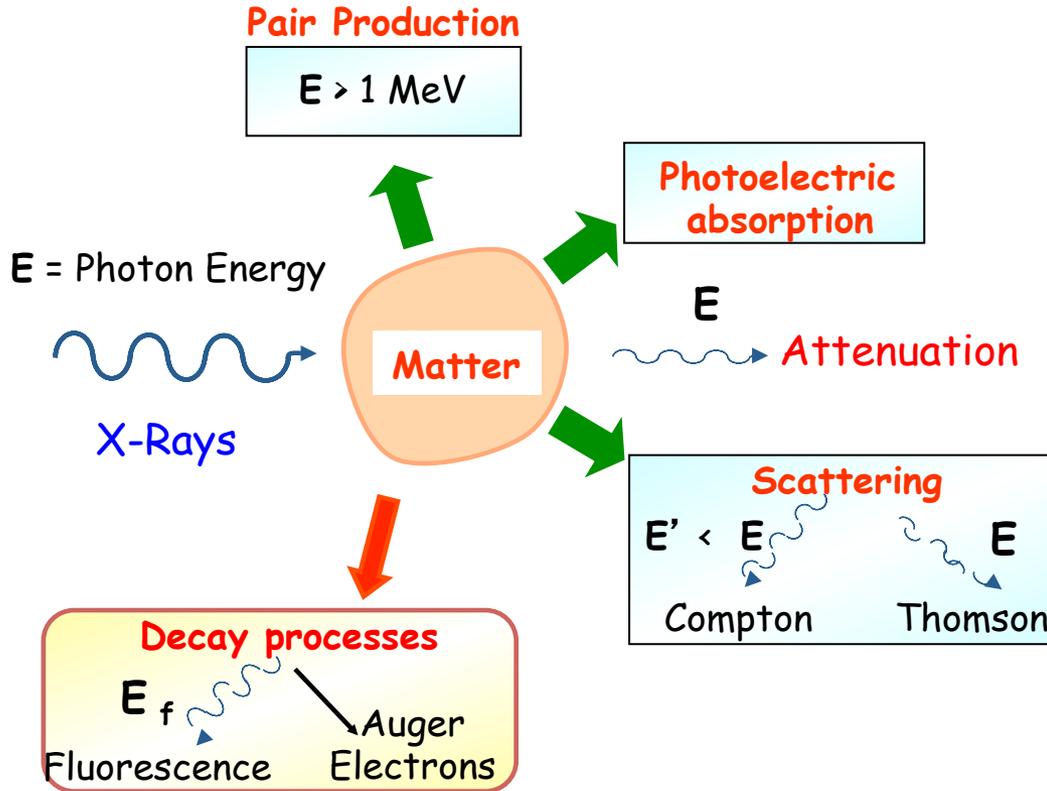
X-rays have enough energy to knock electrons out of a shell. If an electron is extracted from an inner shell, an electron from an outer shell will move to replace it. When the electron moves from the outer shell to the inner shell, it **releases energy in the form of a photon (light)**. The energy of the "fluorescent" photon released is distinct for each atomic element creating a measurable "fingerprint" for that element. XRF produces a **non-destructive chemical analyses of any kind of sample.**



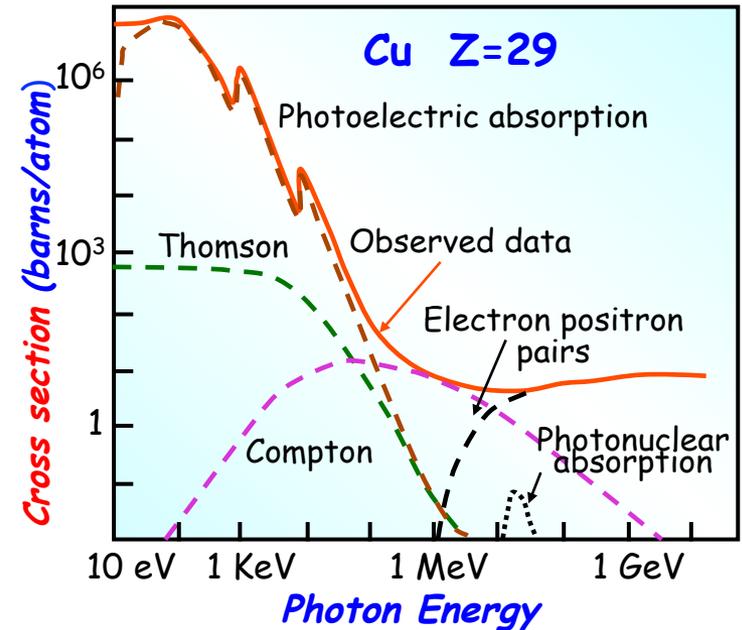
$$\text{Moseley's Law : } E \approx Z^2$$



# Interaction of X-rays with matter

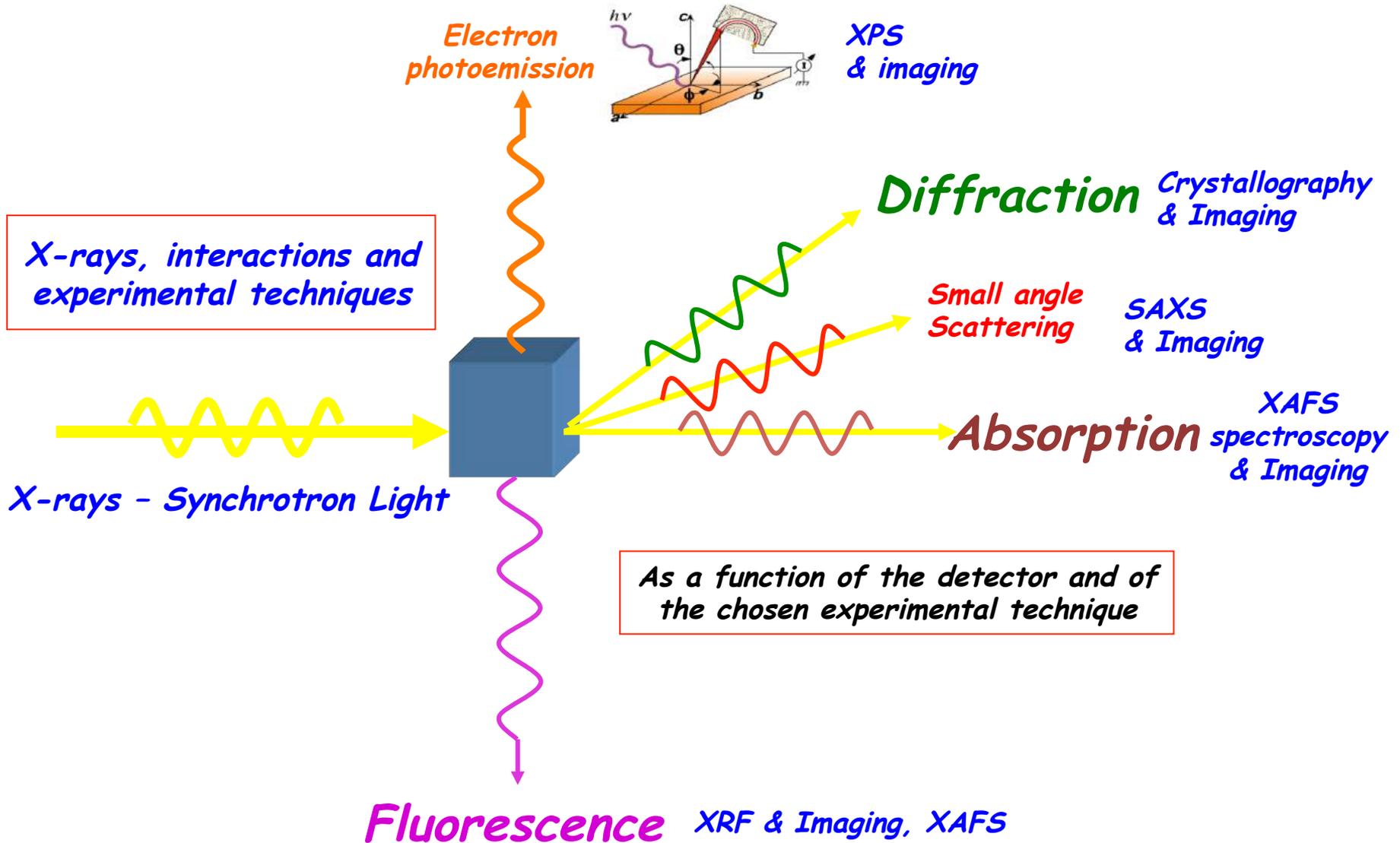


There are different types of interaction of X-rays with matter but taking into account the energy range of interest the ones that will be taken into account are **absorption** and **elastic** or **Thomson scattering**.



The **PROBABILITY** for any given reaction to occur is in **PROPORTION** to its **CROSS SECTION**.

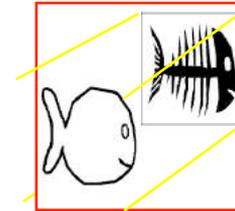
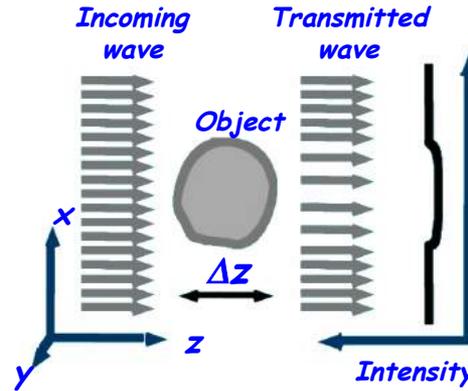
# X-rays interactions with matter and experimental techniques



# Some X-ray techniques

## Imaging

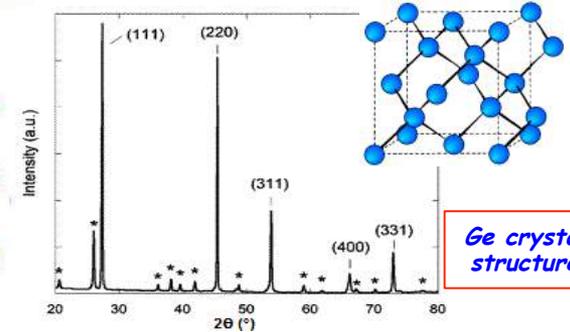
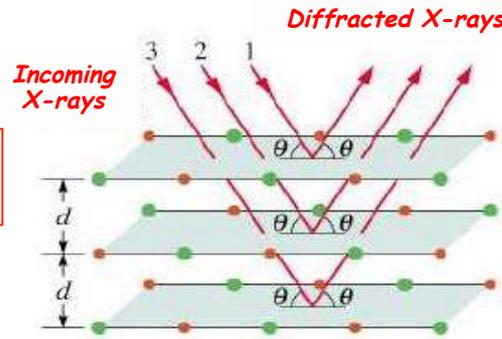
Conventional radiology relies on X ray absorption



## Diffraction

Elastic scattering : Thomson (elastic) if  $E < E_{binding}$

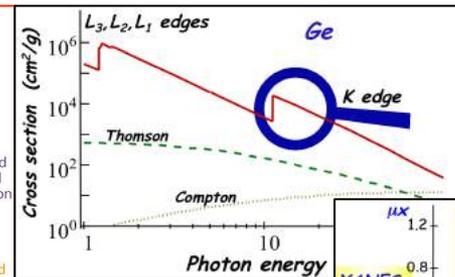
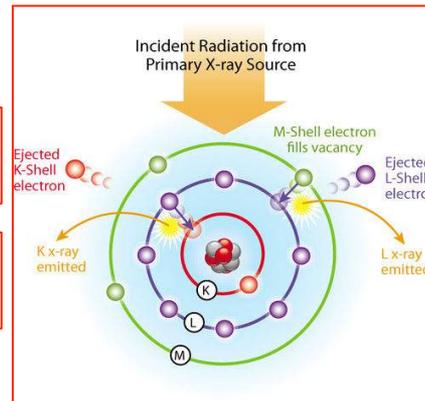
$$n\lambda = 2d \sin\theta \quad \text{Bragg's Law}$$



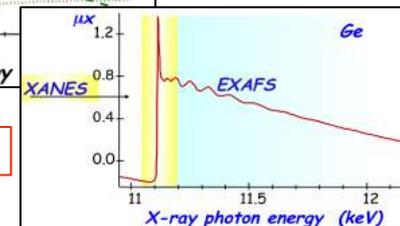
## Absorption Spectroscopy

X ray Absorption Fine Structure (XAFS)

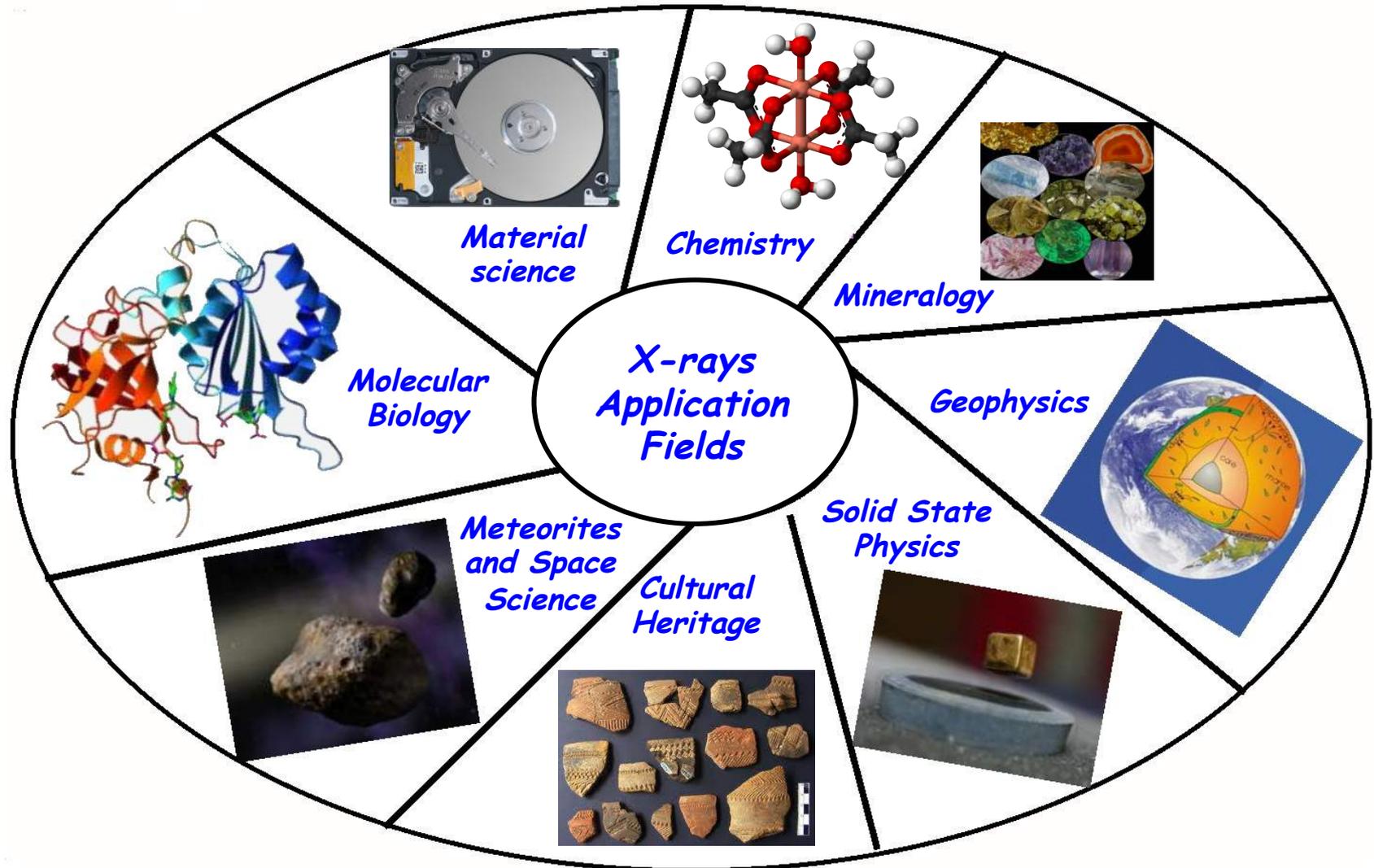
Transmission and fluorescence mode



XAFS = XANES + EXAFS



# *Synchrotron radiation applications using X-rays*



# *Some applications using X-rays (synchrotron light)*

*X-rays and Cultural Heritage*

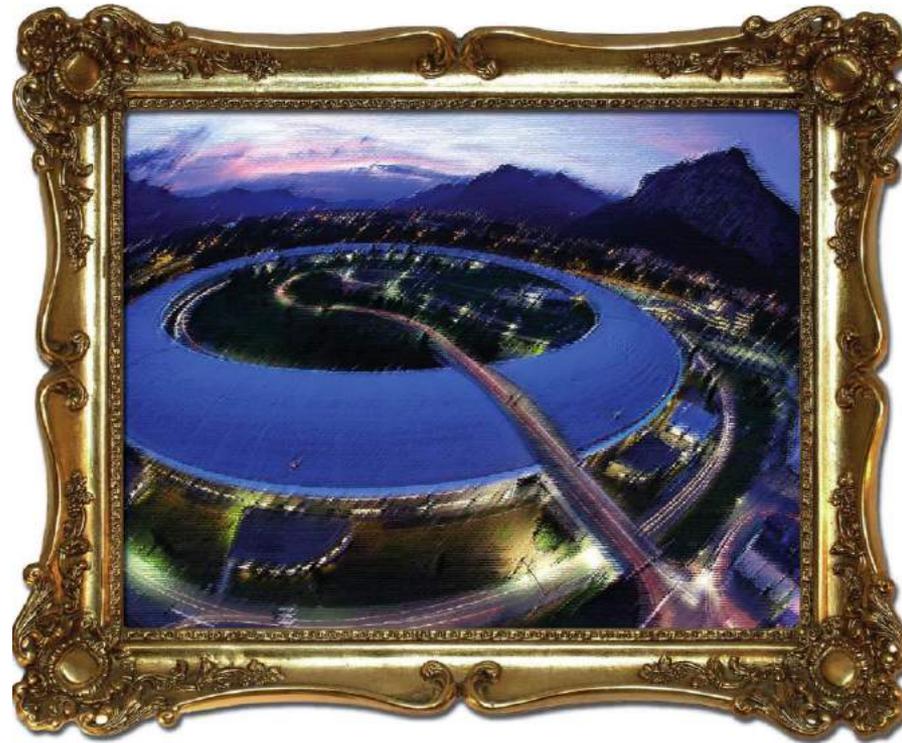
*X-rays and Paleontology*

*X-rays and Food*

*X-rays and Biology*

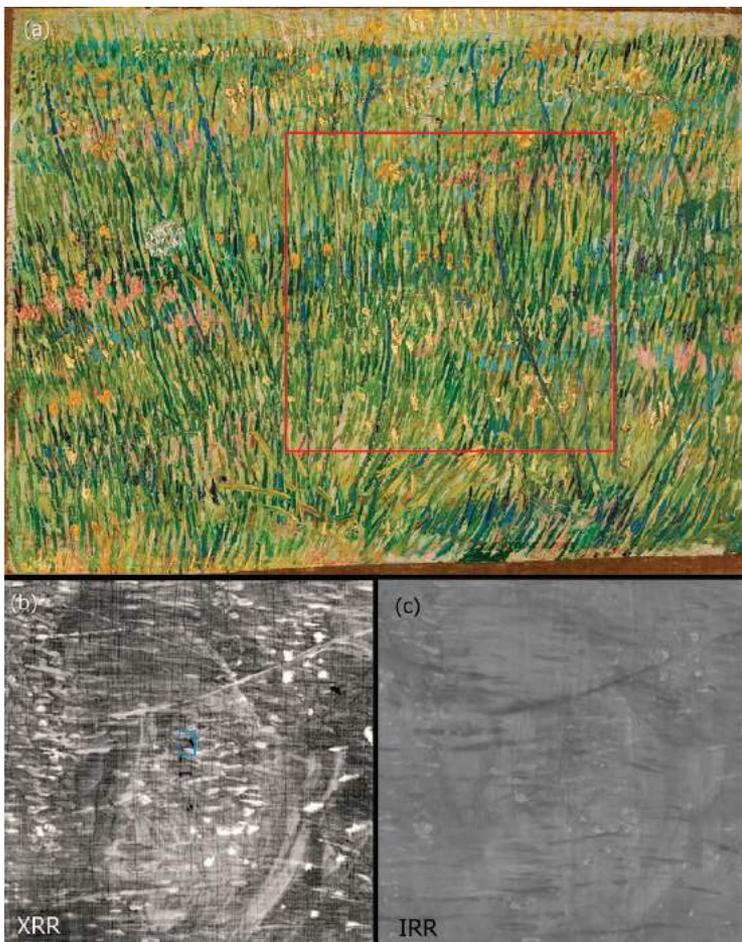
*X-rays and Extreme Conditions*

# *Applications in the field of cultural heritage*



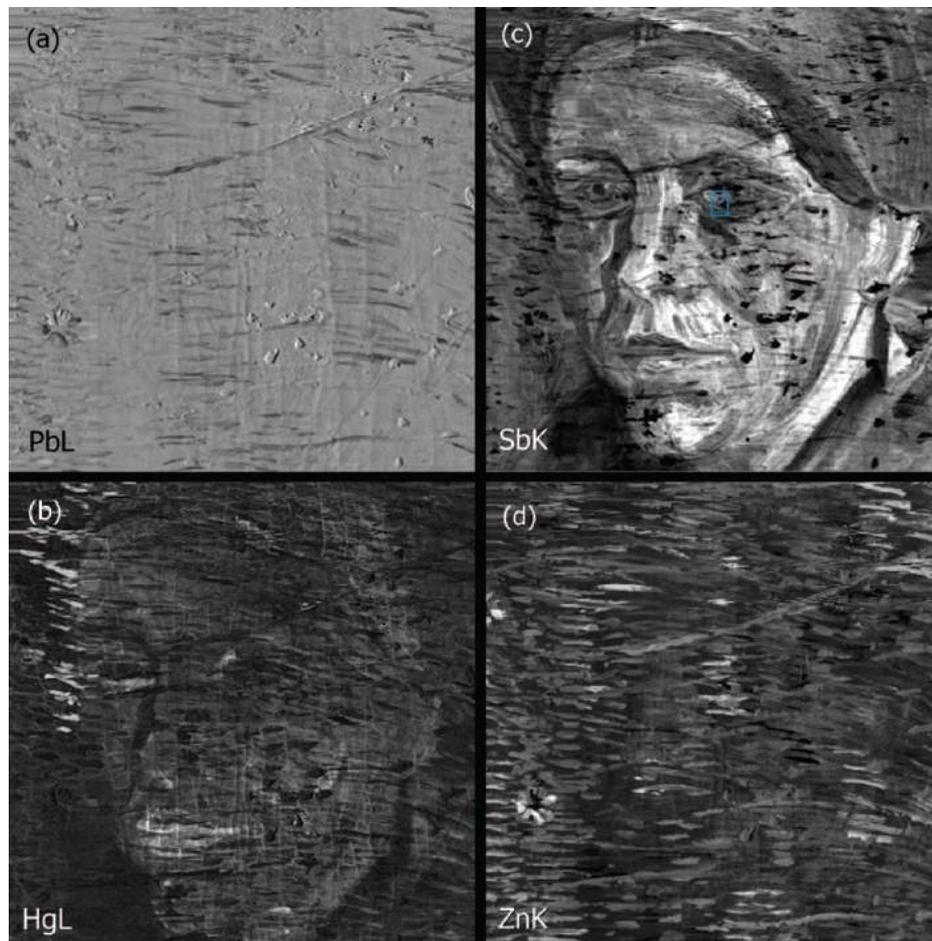
# Visualization of a Lost Painting of van Gogh using XRF

Vincent van Gogh, *Patch of Grass*, Paris 1887, Kroller-Muller Museum, Otterlo, The Netherlands, (KM 105.264; F583/JH1263).



Conventional X-Ray Radiation (XRR) transmission radiography

Infrared reflectography (IRR)



Synchrotron Radiation - XRF (black, low intensity and white, high intensity). Hg L shows the distribution of vermillion, Sb K of Naples yellow and Zn K of zinc white.

# Visualization of a Lost Painting of van Gogh



a) *Tritonal color reconstruction* of Sb (yellowish white) and Hg (red) (b) Detail from Vincent van Gogh, *Head of a Woman*, Nuenen 1884-85, Kro Iler-Muller Museum, Otterlo (KM 105.591:F154/JH608). (c) Detail from Vincent van Gogh, *Head of a Woman*, Nuenen 1884-85, Van Gogh Museum, Amsterdam (F156/JH569).

Vincent van Gogh (1853-1890), is best known for his vivid colors and his short but highly productive career. His productivity is even higher than generally realized, as many of his known paintings cover a previous composition. Van Gogh would often reuse the canvas of an abandoned painting and paint a new or modified composition on top. These hidden paintings offer a unique and intimate insight into the genesis of his works.

# Blackening of Pompeian Cinnabar paintings studied using X-rays



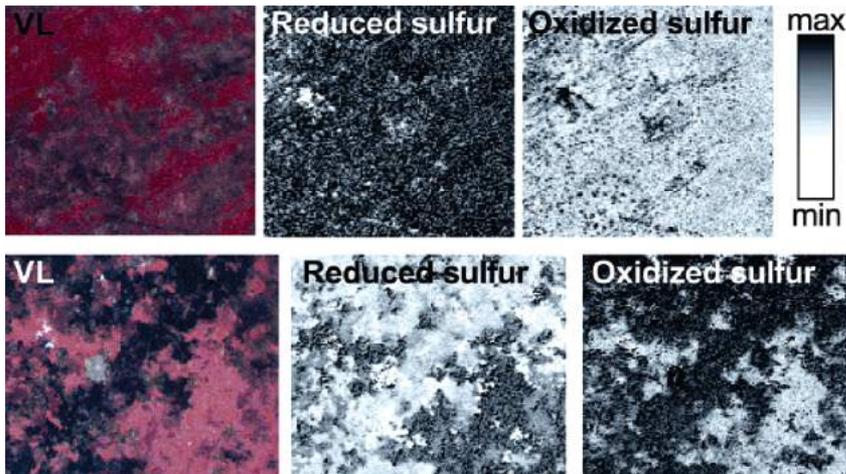
A wall painted red in the remains of Pompei: turning black

## Painting alterations

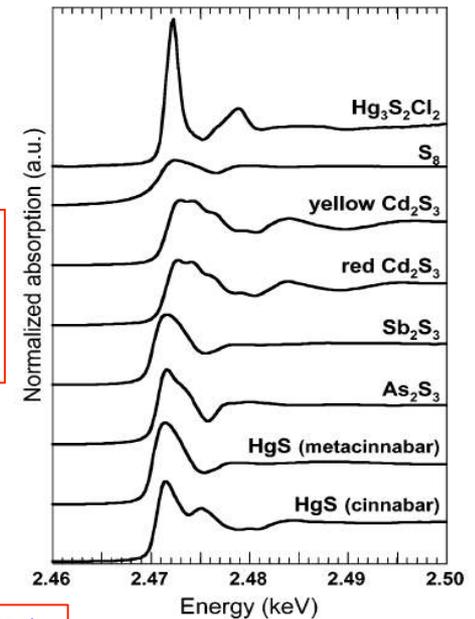
Scientists have *wondered for many years why the red in Pompeii's walls*, a dye that is made from cinnabar (HgS), turns black.

The ESRF scientists found that the *chemical composition in the affected samples* was different from that of cinnabar, which indicated that *some important chemical reactions had taken place*.

On the one hand, *cinnabar had reacted with chlorine*, which led to the formation of grey chlorine mercury compounds. The chlorine came from the sea and possibly punic wax (the wax that was used in the frescoes). *Reduced and oxidized sulfur distributions reveal that the sulfated black coating consists of a 5- $\mu$ m-thick layer covering intact cinnabar*.



X-ray absorption spectra of Sulfur-XANES reflecting the sulfur oxidation-state diversity.



# Archimedes Palimpsest

The Archimedes Palimpsest is a 1000-year-old manuscript on Parchment that contains seven treatises of the Greek mathematician, Archimedes.

Eight hundred years ago, the writings were erased and overwritten (parchments were very precious) becoming a Christian prayer book.

In 1998, the manuscript was purchased by a private collector (2.200.000 \$). Over the last ten years, a team of researchers has devoted a significant effort to recover the erased writing of Archimedes.



The book has suffered very severely from mold



<http://archimedespalimpsest.org/>

# Archimedes Palimpsest

A synchrotron X-ray beam at the SSRL facility (Stanford USA) illuminated an erased work, written in **iron gall ink on parchment** and **obscured by gold paint**.



<https://news.stanford.edu/news/2005/may25/archimedes-052505.html>

# *X-rays and Paleontology*

# Imaging and paleontology

*Amber has always been a rich source of fossil evidence. X-rays now make it possible for paleontologists to study opaque amber, previously inaccessible using classical microscopy techniques. Scientists from the University of Rennes (France) and the ESRF found 356 animal inclusions, dating from 100 million years ago, in two kilograms of opaque amber from mid-Cretaceous sites of Charentes (France).*



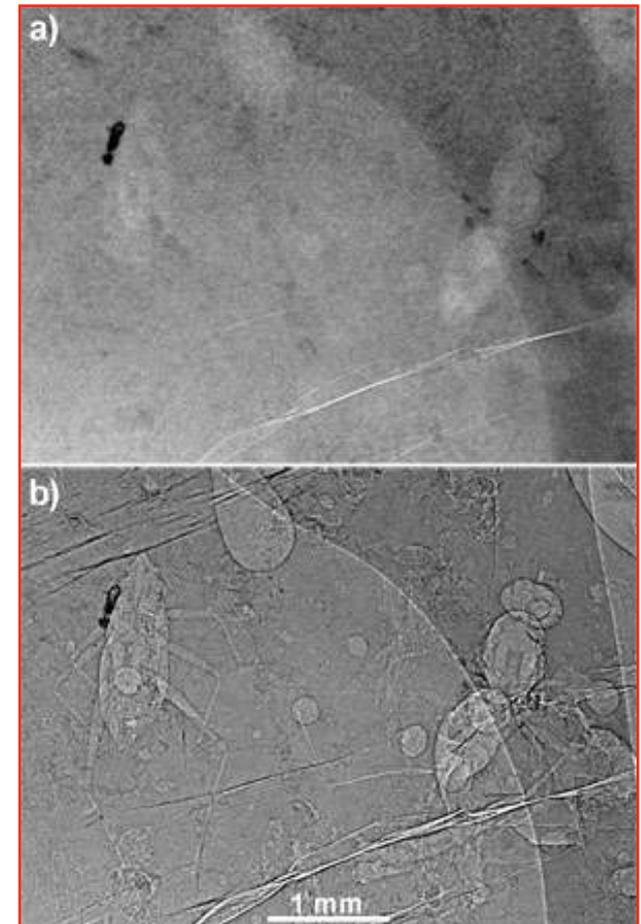
*Pieces of opaque amber-  
Image credits: V. Girard/D. Néraudeau,  
UMR CNRS 6118 Rennes*

*Radiography of an  
amber block with  
inclusions viewed in  
absorption mode.*

*Propagation phase  
contrast X-ray  
synchrotron imaging  
(990 mm prop. distance)*

*M. Lak, D. Néraudeau, A. Nel, P. Cloetens, V. Perrichot and P. Tafforeau,  
Phase Contrast X-ray Synchrotron Imaging: Opening Access to Fossil Inclusions in  
Opaque Amber, *Microscopy and Microanalysis*, (2008), 14:251-259*

## Fossil inclusions in opaque amber



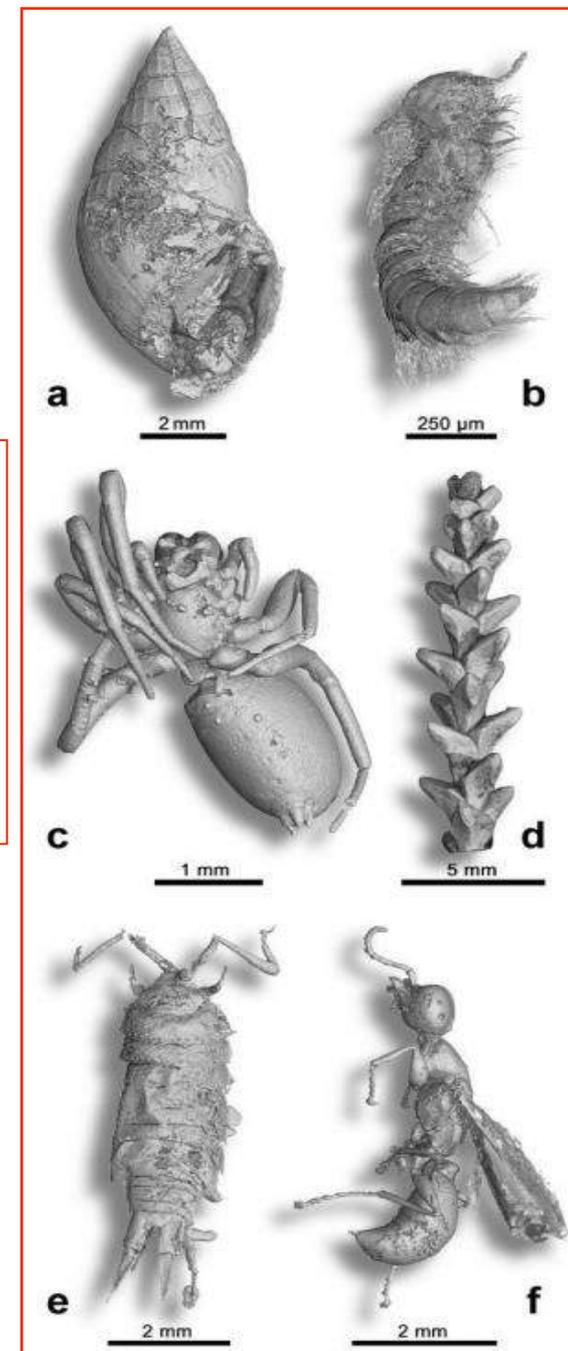
# Imaging and paleontology

Synchrotron X-ray micro tomography was used to determine the 3D reconstruction and allowed the paleontologists to study the organisms in detail and to describe them.



*Cretaceous beetle*

Examples of virtual 3D extraction of organisms embedded in opaque amber:  
a) Gastropod Ellobiidae; b) Myriapod Polyxenidae; c) Arachnid; d) Conifer branch (*Glenrosa*); e) Isopod crustacean *Ligia*; f) Insect hymenopteran *Falciformicidae*.



*M. Lak, D. Neraudeau, A. Nel, P. Cloetens, V. Perrichot and P. Tafforeau, Phase Contrast X-ray Synchrotron Imaging: Opening Access to Fossil Inclusions in Opaque Amber, Microscopy and Microanalysis, (2008)*

## MILLION YEARS AGO

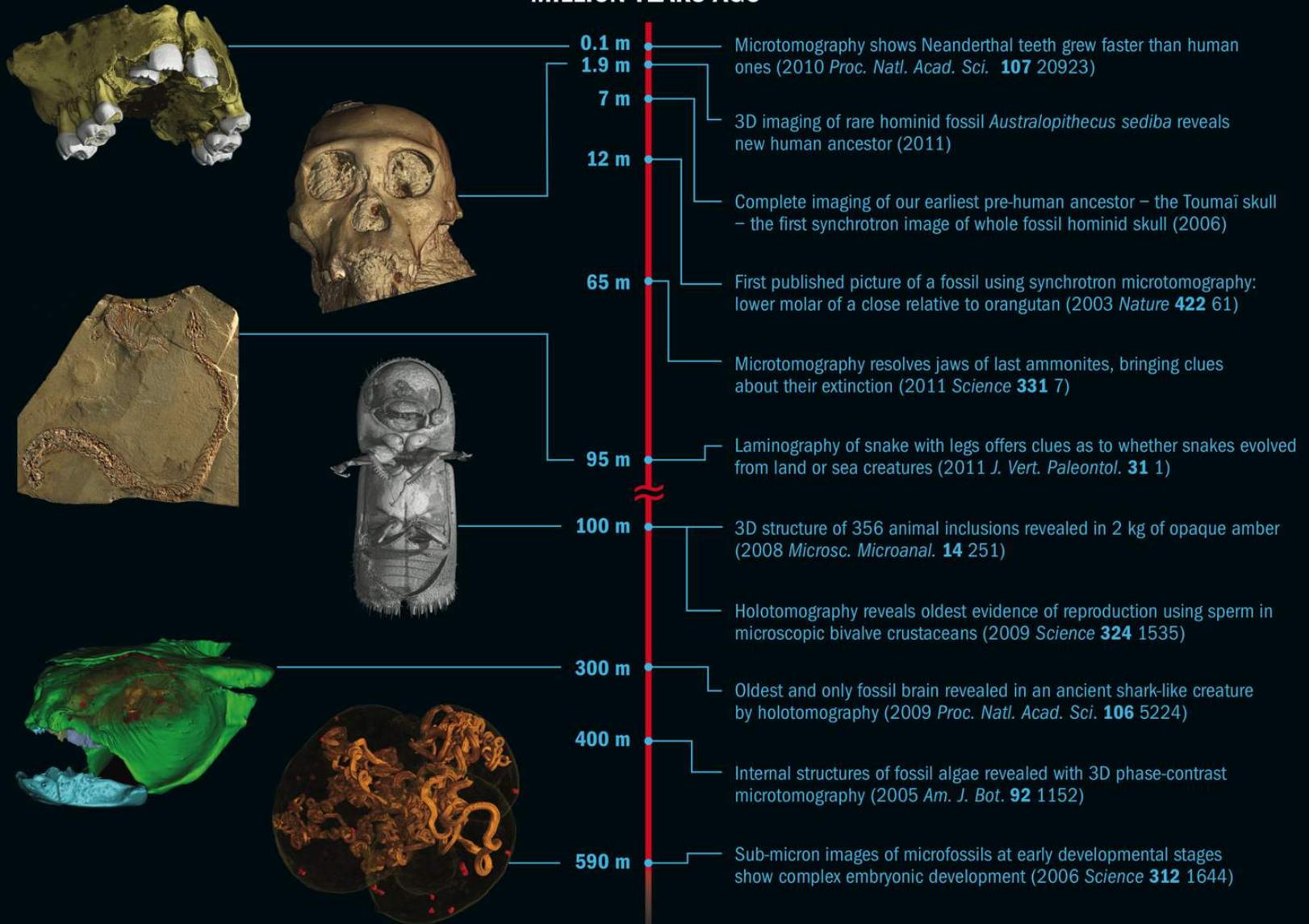


Image credits (top to bottom): Van Gogh Museum, Amsterdam; Mary Rose Trust; Royal Museum of Fine Arts, Antwerp; American Chemical Society; C Reyes-Valerio; D Bagault C2RMF; P Tafforeau, ESRF; P Tafforeau, ESRF; A Houssaye; P Tafforeau, ESRF; A Pradel/CNRS; P Tafforeau, ESRF.

# *X-rays and food*

# Structure of Chocolate

Most chocolate eaters will have had the surprising experience that a newly opened bar of chocolate (pure or milk) has a greyish-white layer instead of the familiar chocolate color.



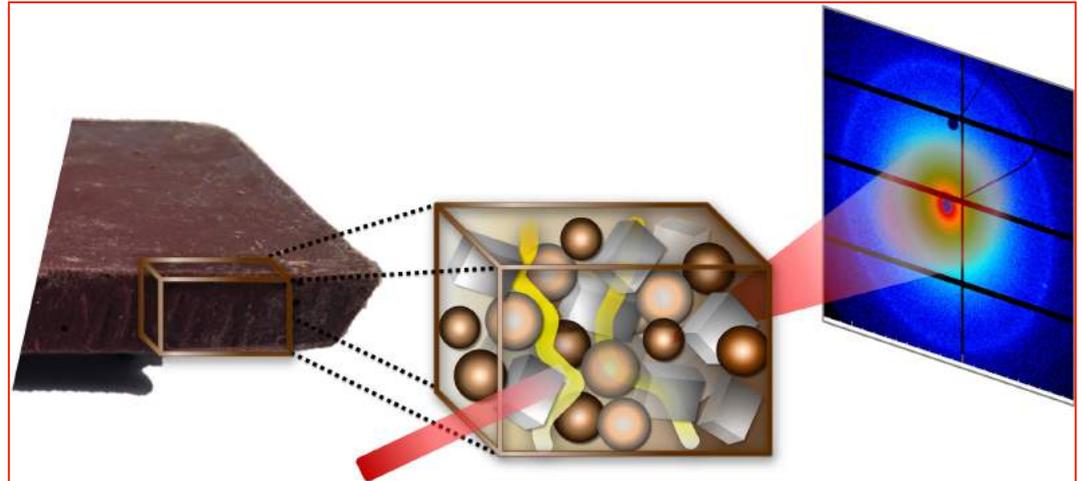
A basic chocolate recipe consists of roughly one-third cocoa butter, a fat that crystallizes easily. How the butter crystallizes determines the quality of the chocolate.

<http://www.esrf.eu/UsersAndScience/Publications/Highlights/2004/SCM/SCM8>

Even though fat bloom does not actually constitute any deterioration in the quality of the product, the visual alteration associated with it can lead to a large number of consumer complaints.

Cocoa butter crystallizes in six different crystal forms. The amount of fluid also depends on the form of the crystals.

**Manufacturers can also limit fat bloom by controlling crystallization.**

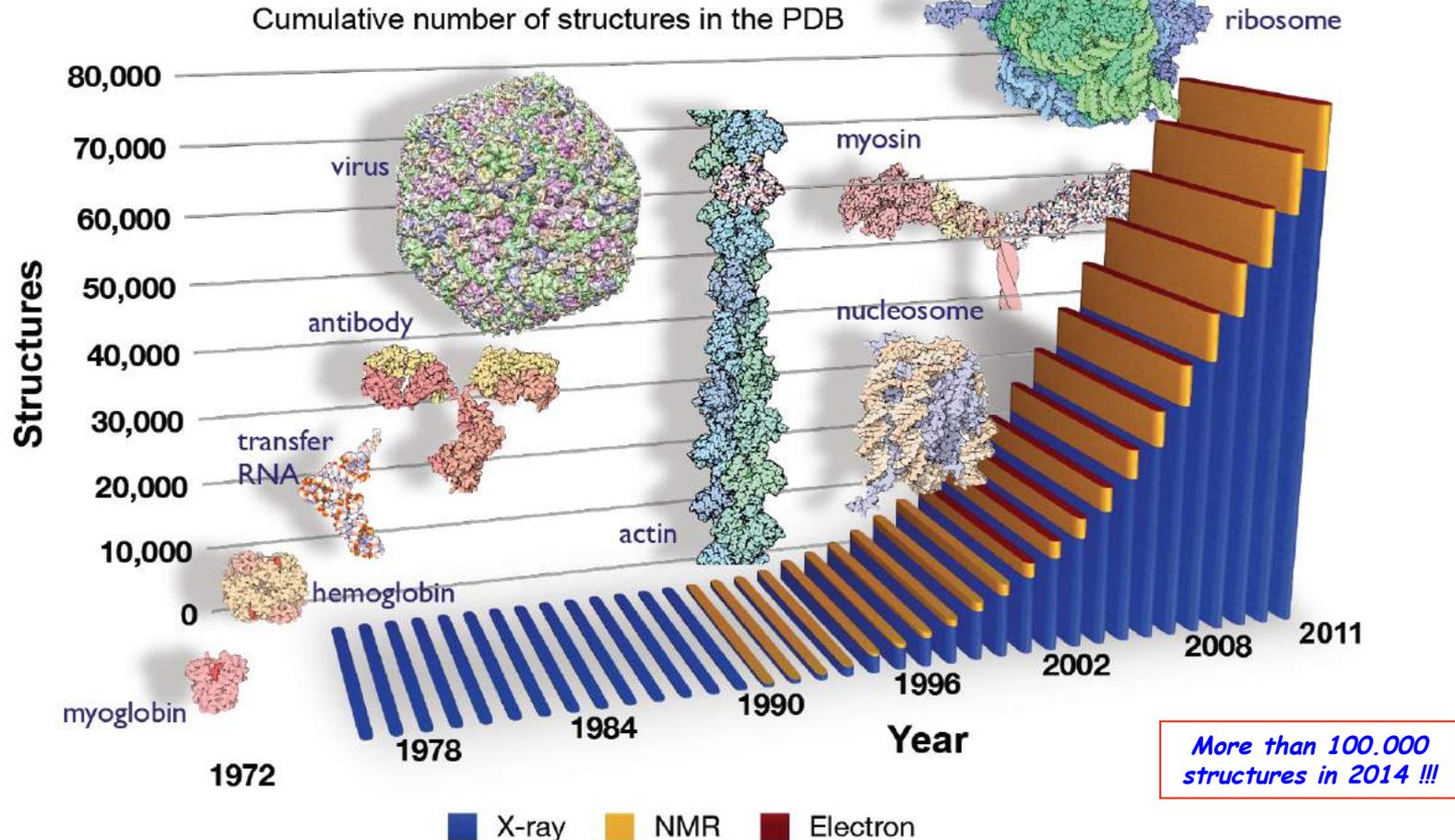


*Svenja K. Reinke et al., Tracking Structural Changes in Lipid-based Multicomponent Food Materials due to Oil Migration by Microfocus Small-Angle X-ray Scattering, ACS Appl. Mater. Interfaces 2015, 7, 9929*

# *Bio-crystallography*

# Bio-crystallography

The number of protein structures solved is now increasing linearly

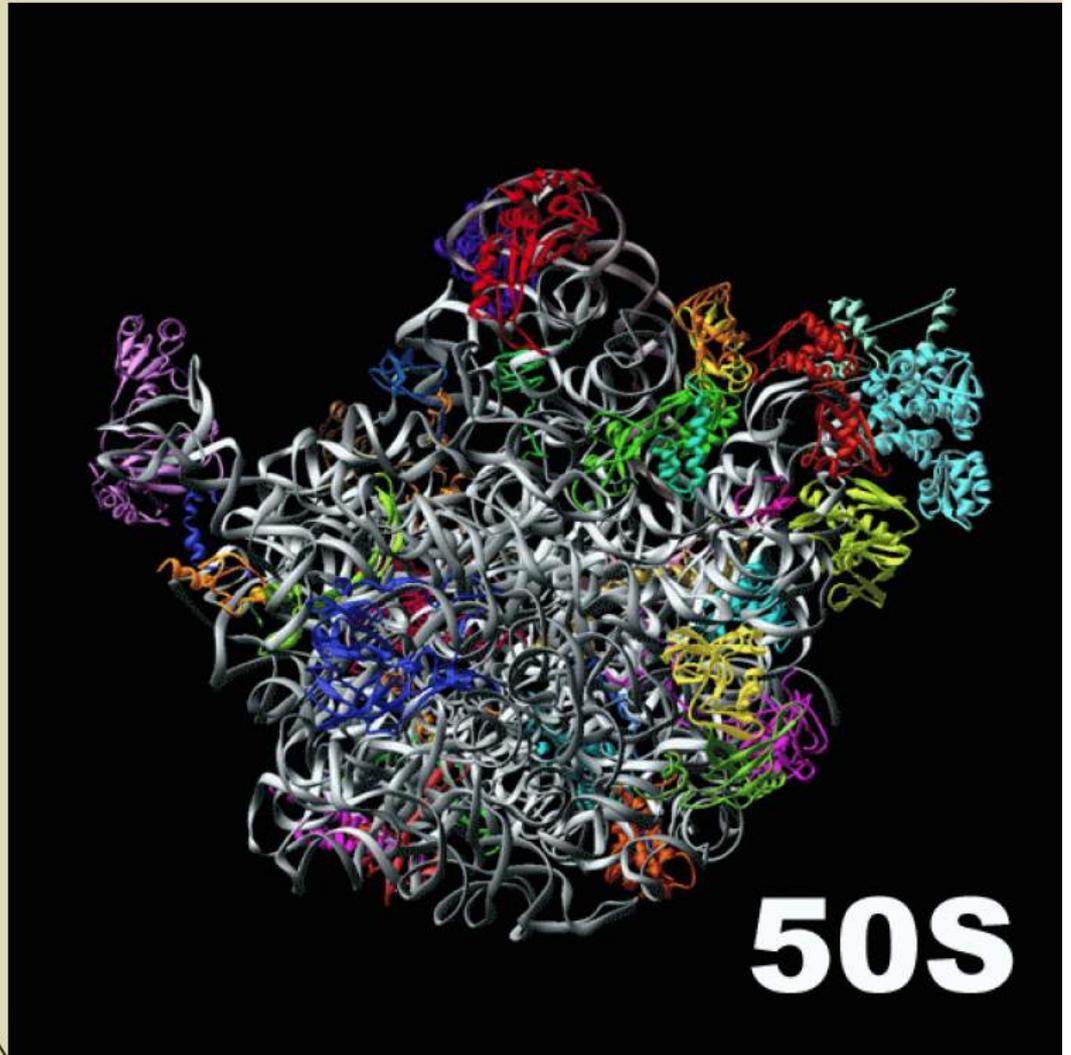




Myoglobin  
16 kDa

**50 YEARS**

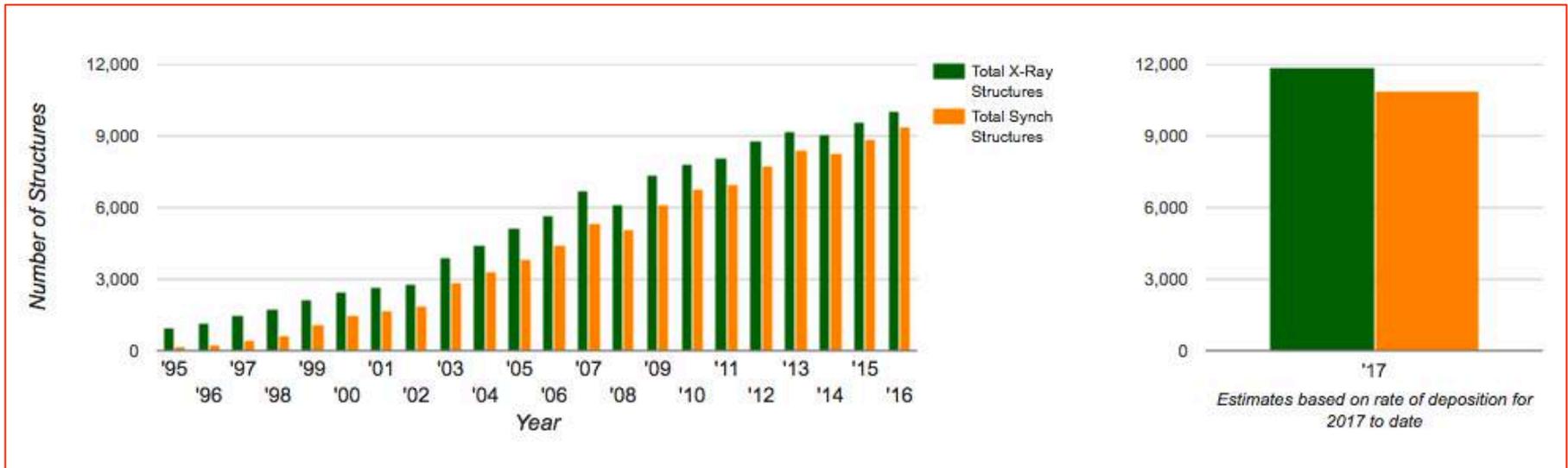
Peso molecolare:  
 $1\text{kDa} = 1000\text{ g mol}^{-1}$



**50S**

Ribosome 2.5 MDa

# *X-ray bio-crystallography and synchrotron radiation*



2016 <http://biosync.sbkb.org/index.jsp>

# Very recent Noble Prizes in Chemistry



2009 "for studies of the structure and function of the ribosome"

*Biocrystallography vs. Structural Biology*



Photo: MRC Laboratory of  
Molecular Biology

**Venkatraman  
Ramakrishnan**



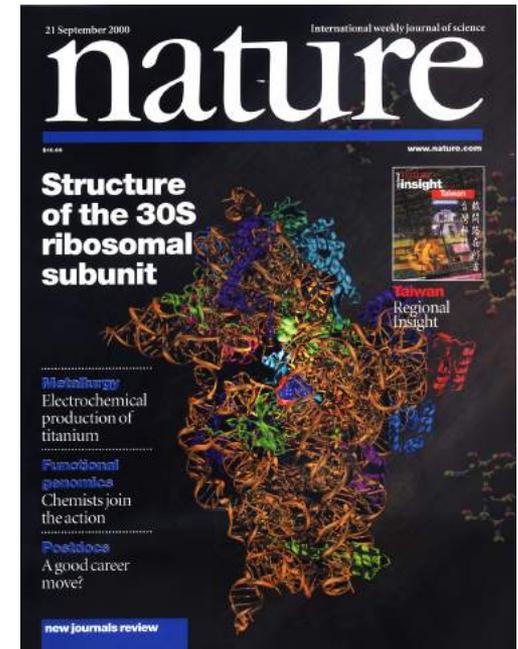
Credits: Michael  
Marsland/Yale University

**Thomas A. Steitz**



Credits: Micheline  
Pelletier/Corbis

**Ada E. Yonath**



*Using Synchrotron radiation Research*

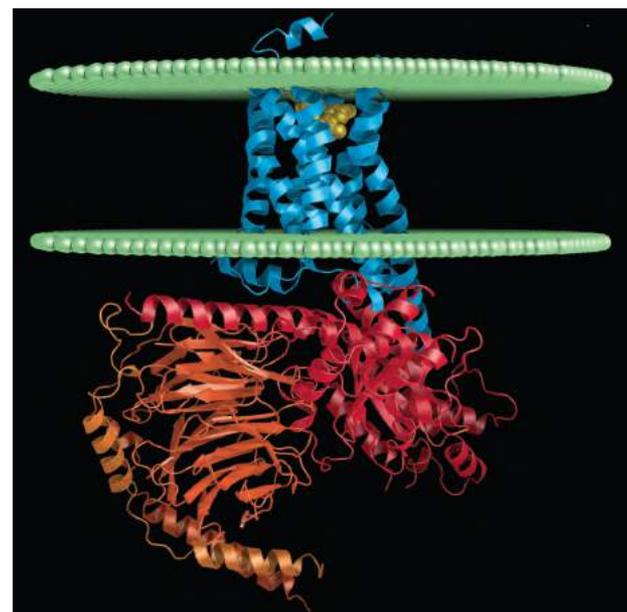
An understanding of the ribosome's innermost workings is important for a scientific understanding of life. This knowledge can be put to a practical and immediate use; many of today's antibiotics cure various diseases by blocking the function of bacterial ribosomes. Without functional ribosomes, bacteria cannot survive. This is why ribosomes are such an important target for new and more efficient antibiotics.

# Very recent Noble Prizes in Chemistry



2012 "for studies of G-protein-coupled receptors"

*Biocrystallography vs. Structural Biology*



*Using Synchrotron radiation Research*

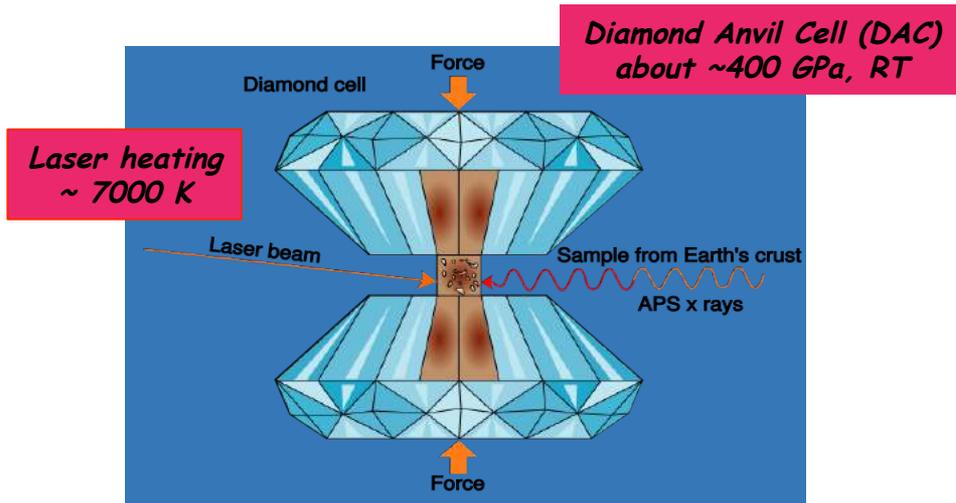
*G-Protein Coupled Receptor (blue) sits within lipid bilayer (green) to respond to hormone (yellow)-  
Image by Wayne Decatur - <http://www.hhmi.org/bulletin/winter2013/features/index.html>*

G protein coupled receptors (GPCRs) represent *the largest family of membrane proteins* (about 800 different proteins) *controlling body functions, drug transit across membranes* and representing the richest source of targets for the pharmaceutical industry.

*X-rays and extreme  
conditions*

# X-rays and extreme conditions (P, T): new opportunities

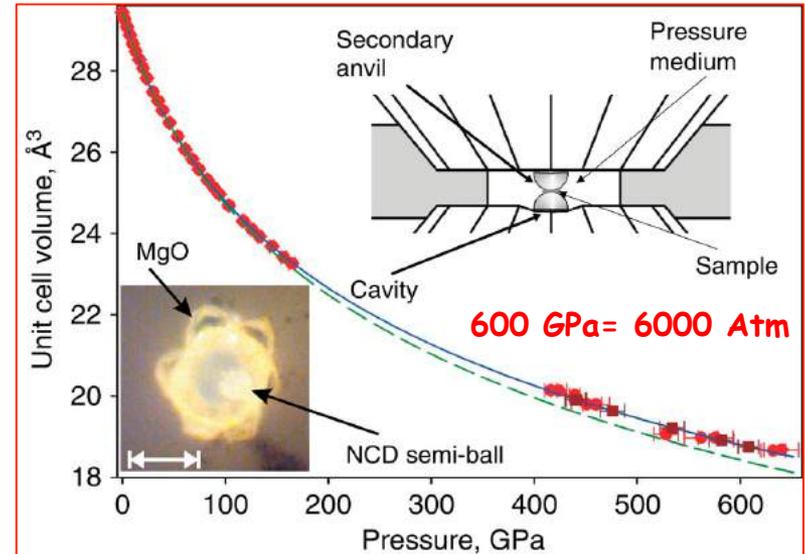
P. W. Bridgman (Nobel Prize in Physics in 1946 - discoveries made in the field of high pressure physics)- "*Compression offers a route to breaking down the electronic structure of the atoms themselves and to the possibility of creating entirely different bulk properties*".



**100GPa=1Mbar**

Application fields:

- Earth and Planetary Science
- Condensed Matter Physics
- Chemistry and Materials Science
- Biology and Soft Matter

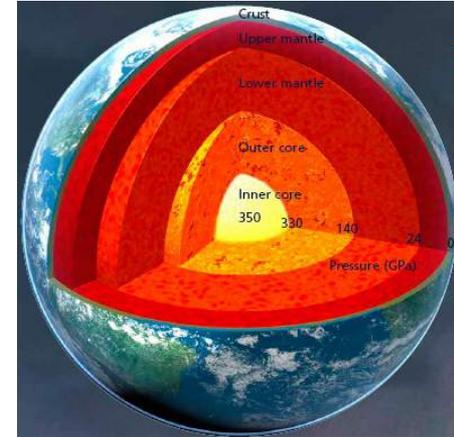
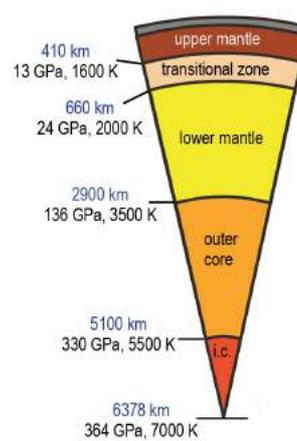
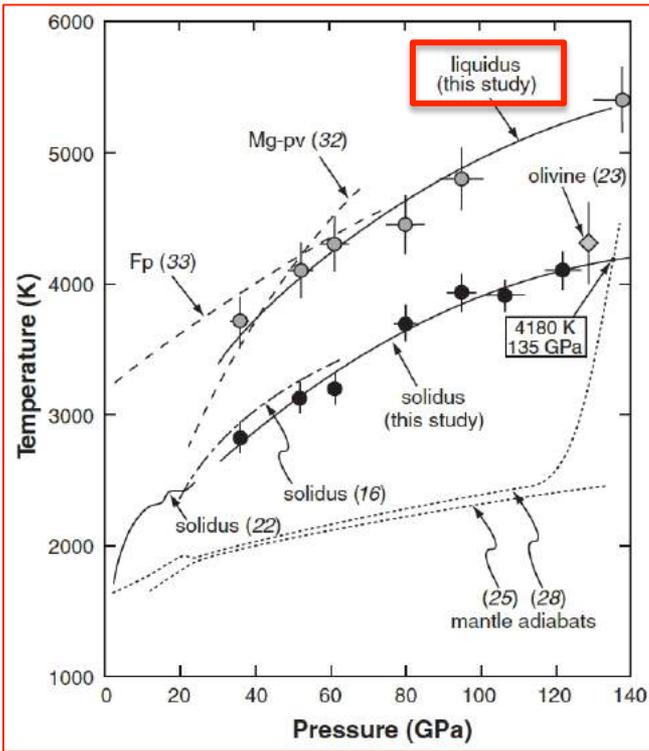


*New system to increase Pressure*

*L. Dubrovinsky et al. Implementation of micro-ball nano-diamond anvils for high-pressure studies above 6 Mbar - Nature Comm. (2012)*

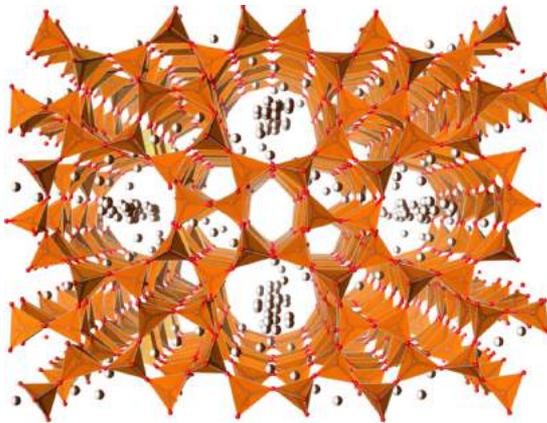
- Novel transformations: solids, liquids, glasses
- Structures: unexpected complexity
- Molecules break down, but new ones form
- Novel electronic and magnetic phenomena
- New chemical reactions: low to high pressure
- New recoverable materials

# Extreme conditions (P, T)



**Melting of Peridotite (Olivine and iron-magnesium silicates)**  
**First direct evidence (ESRF ID27) that the layer located at the bottom of the Earth's mantle (2900 km depth) contains partially molten minerals.** This result supports the existence of a deep magmatic ocean. (P: 36 -140 GPa T:2500-5000 K)

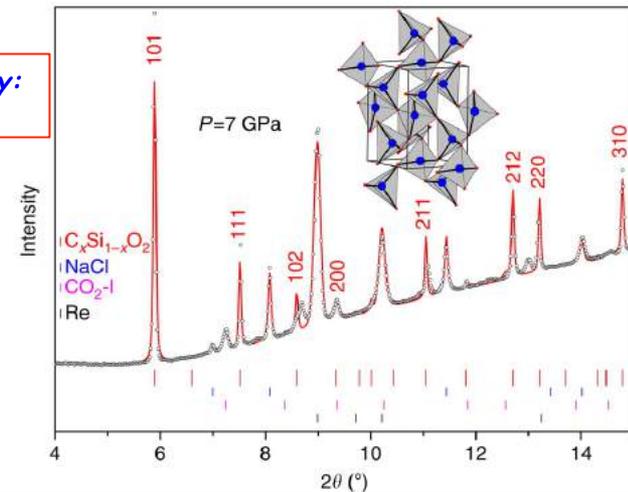
*G. Fiquet, et al. Melting of Peridotite to 140 GPa Science (2010)*



**New materials: conducting polymers**

*D. Scelta et al. High Pressure Polymerization in a Confined Space: Conjugated Chain/Zeolite Nanocomposites. Chem. Mater. (2014)*

**New oxide chemistry:**  
 $C_{0.6(1)}Si_{0.4(1)}O_2$



*M. Santoro et al. Carbon enters silica forming a cristobalite-type  $CO_2-SiO_2$  solid solution, Nature Comm. (2014)*

# *Conclusions*

- Synchrotron radiation has surely revolutionized X-ray applications.
- Most of the SR facilities in the world have beamlines dedicated to different X-ray applications.
- Synchrotron radiation X-ray applications still have a very bright future.

*Thank you for your attention*

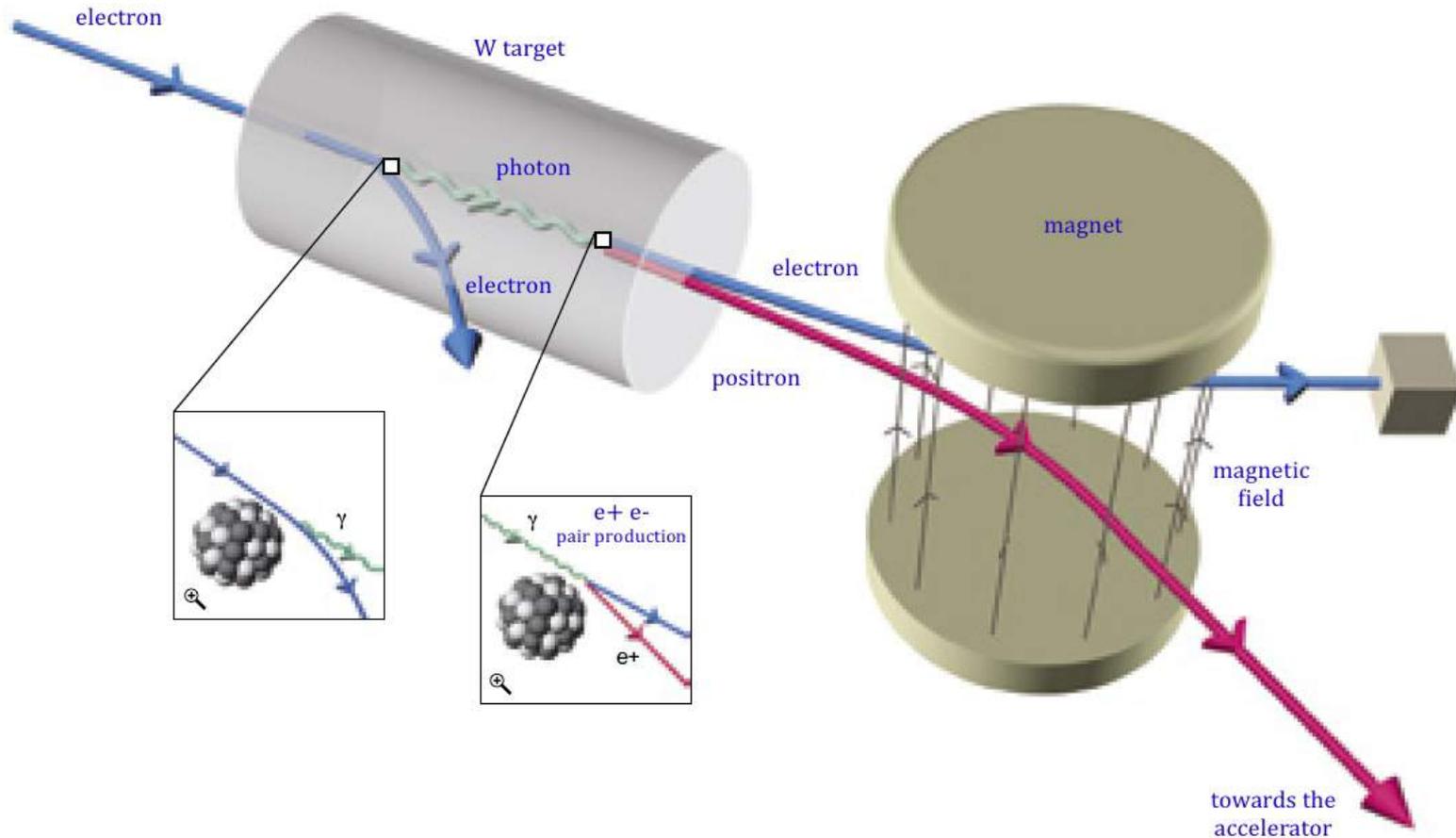


*Supplementary material- f.y.k.*

# Reminder

- **Speed of light**  $c = 2.99792458 \times 10^8 \text{ m/s}$
- **Electron charge**  $e = 1.6021 \times 10^{-19} \text{ Coulombs}$
- **Electron volts**  $1 \text{ eV} = 1.6021 \times 10^{-19} \text{ Joule}$
- **Energy and rest mass**  $1 \text{ eV}/c^2 = 1.78 \times 10^{-36} \text{ kg}$ 
  - Electron**  $m_0 = 511.0 \text{ keV}/c^2 = 9.109 \times 10^{-31} \text{ kg}$
  - Proton**  $m_0 = 938.3 \text{ MeV}/c^2 = 1.673 \times 10^{-27} \text{ kg}$
- **Relativistic energy, E**  $E = mc^2 = m_0 \gamma c^2$
- **Lorentz factor,  $\gamma$**   $\gamma = 1/[(1-v^2/c^2)^{1/2}] = 1/[(1-\beta^2)^{1/2}]$   
 $\beta = v/c$
- **Relativistic momentum, p**  $p = mv = m_0 \gamma \beta c$
- **E-p relationship**  $E^2/c^2 = p^2 + m_0^2 c^2$   
*for ultra-relativistic particles*  $\beta \approx 1, E = pc$
- **Kinetic energy**  $T = E - m_0 c^2 = m_0 c^2 (\gamma - 1)$

# Anti-matter positron production



*M. Calvetti, Antiparticelle accelerate, Asimmetrie 7, 16-21 (2008)*

# X-rays discovery

## EINE NEUE ART VON STRAHLEN.

VON  
DR. W. RÖNTGEN,  
O. O. PROFESSOR AN DEN K. UNIVERSITÄT WÜRZBURG.

WÜRZBURG.  
VERLAG UND DRUCK DER STAHEL'SCHEN K. Hof- UND UNIVERSITÄTS-  
BICH- UND KUNSTHANDLUNG.  
Kode 1895.

60 S.

*On a New Kind of Rays*

While **Wilhelm Roentgen** was working on the effects of **cathode rays** during **1895**, he discovered X-rays. His experiments involved the passing of electric current through gases at extremely low pressure. On **November 8, 1895** he observed that certain rays were emitted during the passing of the current through discharge tube. His experiment that involved working in a totally dark room with a well covered discharge tube resulted in the **emission of rays which illuminated a barium platinocyanide screen. The screen became fluorescent even though it was placed two meters away from discharge tube.**



**Wilhelm Conrad Roentgen**

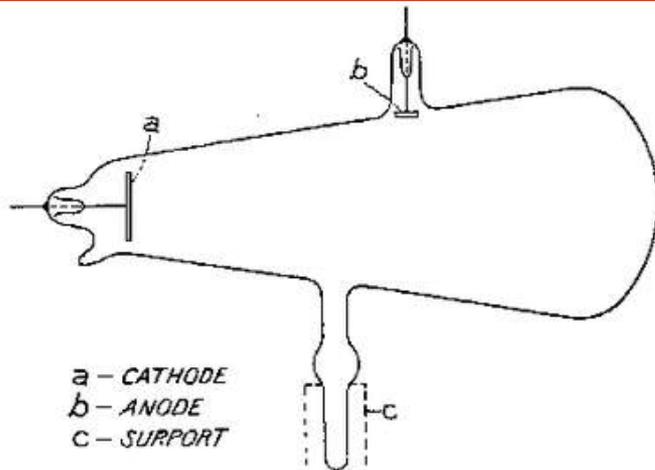


Fig. 1. Earliest type of roentgen tube.

**Gas tube:** electrons are freed from a cold cathode by positive ion bombardment, thus necessitating a certain gas pressure.

He continued his experiments using **photographic plates** and generated the very **first "roentgenogram"** by developing the image of **his wife's hand** and analyzed the variable transparency as showed by her bones, flesh and her wedding ring.



*X-ray tubes*

# X-rays: conventional sources

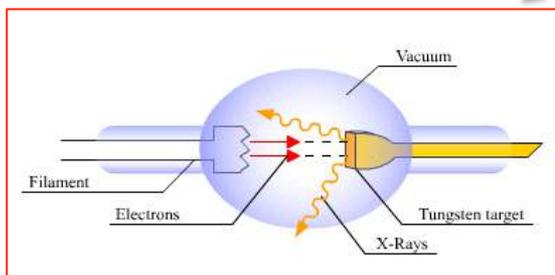
From gas tubes (cold cathode) to high vacuum tubes (hot cathode)



*Crookes tube*

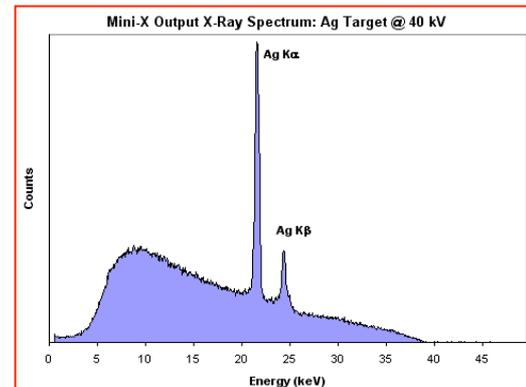
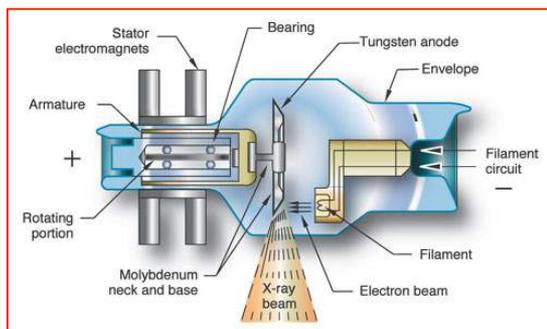


*Coolidge tube*

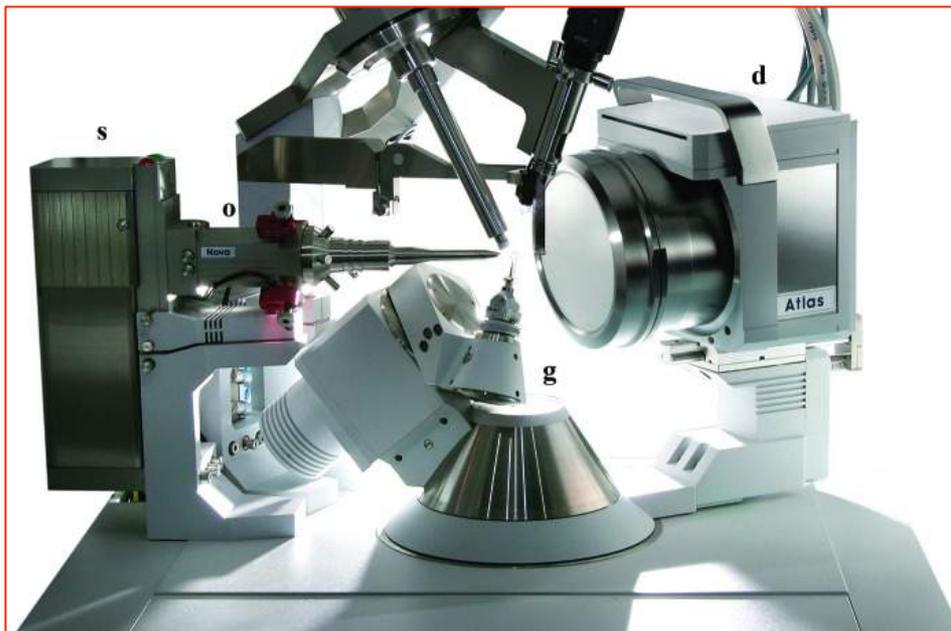


The *Coolidge tube (1913)*, also called *hot cathode tube*, is the most widely used. Electrons are produced by thermionic effect from a tungsten filament heated by an electric current. The filament is the cathode of the tube. The high voltage potential is between the cathode and the anode, the electrons are accelerated, and hit the anode.

The *rotating anode tube* is an improvement of the *Coolidge tube* anode surface (water cooled) is always moving, so heat is spread over a much larger surface area giving a 10-fold increase in the operating power.



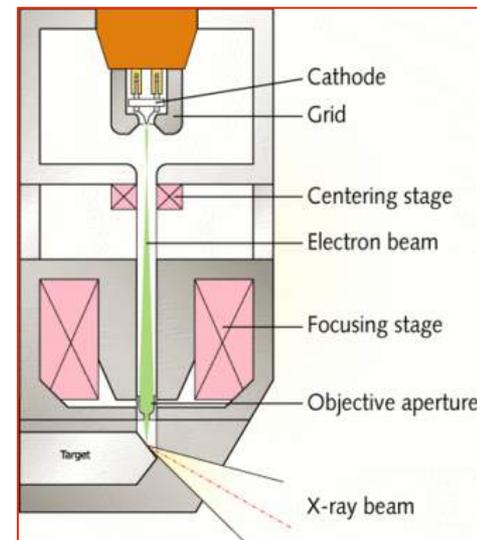
# X-rays: conventional sources



## EVOLUTION

A compact X-ray in-house laboratory system consisting of a **microfocus sealed-tube source** (s), focusing multi-layer optics (o), a four-circle goniometer (g) and a CCD detector (d).

Pictures of **Agilent Technologies**.

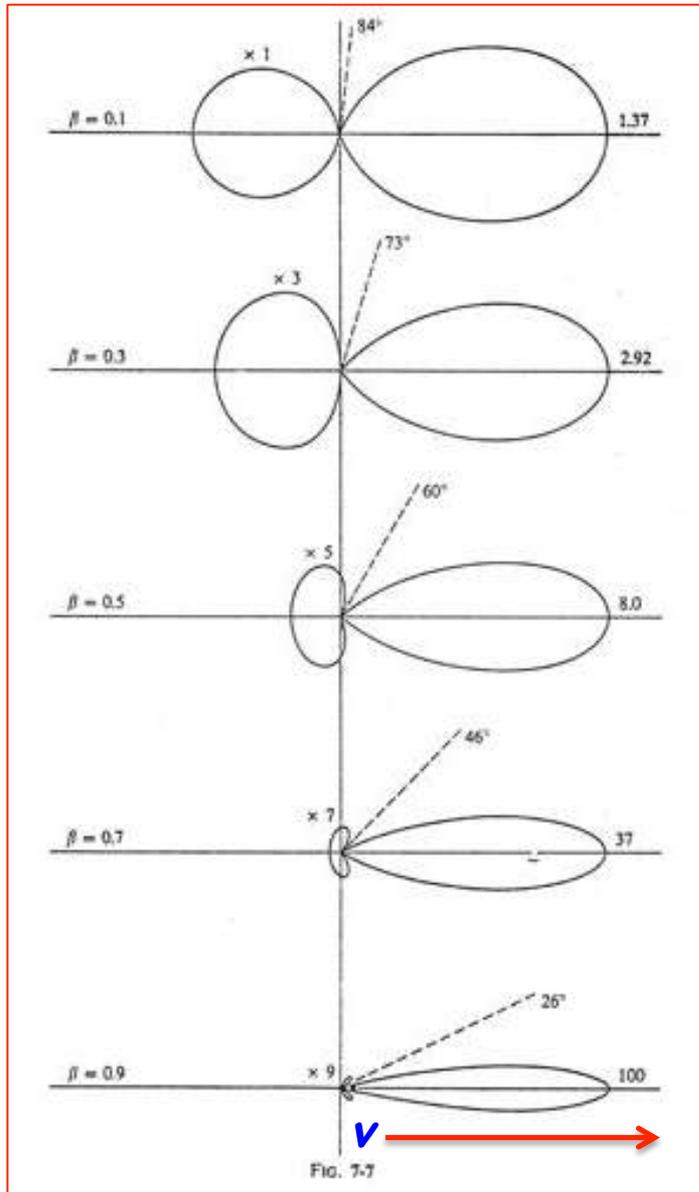


Approximate X-ray beam brilliance for the main types of in-house sources with optics

System	Power (W)	Actual spot on anode ( $\mu\text{m}$ )	Apparent spot on anode ( $\mu\text{m}$ )	Brilliance ( $\text{photons s}^{-1} \text{mm}^{-2} \text{mrad}^{-1}$ )
Standard sealed tube	2000	10000 $\times$ 1000	1000 $\times$ 1000	0.1 $\times 10^9$
Standard rotating-anode generator	3000	3000 $\times$ 300	300 $\times$ 300	0.6 $\times 10^9$
Microfocus sealed tube	50	150 $\times$ 30	30 $\times$ 30	2.0 $\times 10^9$
Microfocus rotating-anode generator	1200	700 $\times$ 70	70 $\times$ 70	6.0 $\times 10^9$
State-of-the-art microfocus rotating-anode generator	2500	800 $\times$ 80	80 $\times$ 80	12 $\times 10^9$
Excillum JXS-D1-200	200	20 $\times$ 20	20 $\times$ 20	26 $\times 10^9$

*Synchrotron radiation*

# Synchrotron radiation: physics



$$v \ll c \text{ or } \beta = v/c \ll 1$$

As  $\beta$  approaches 1:

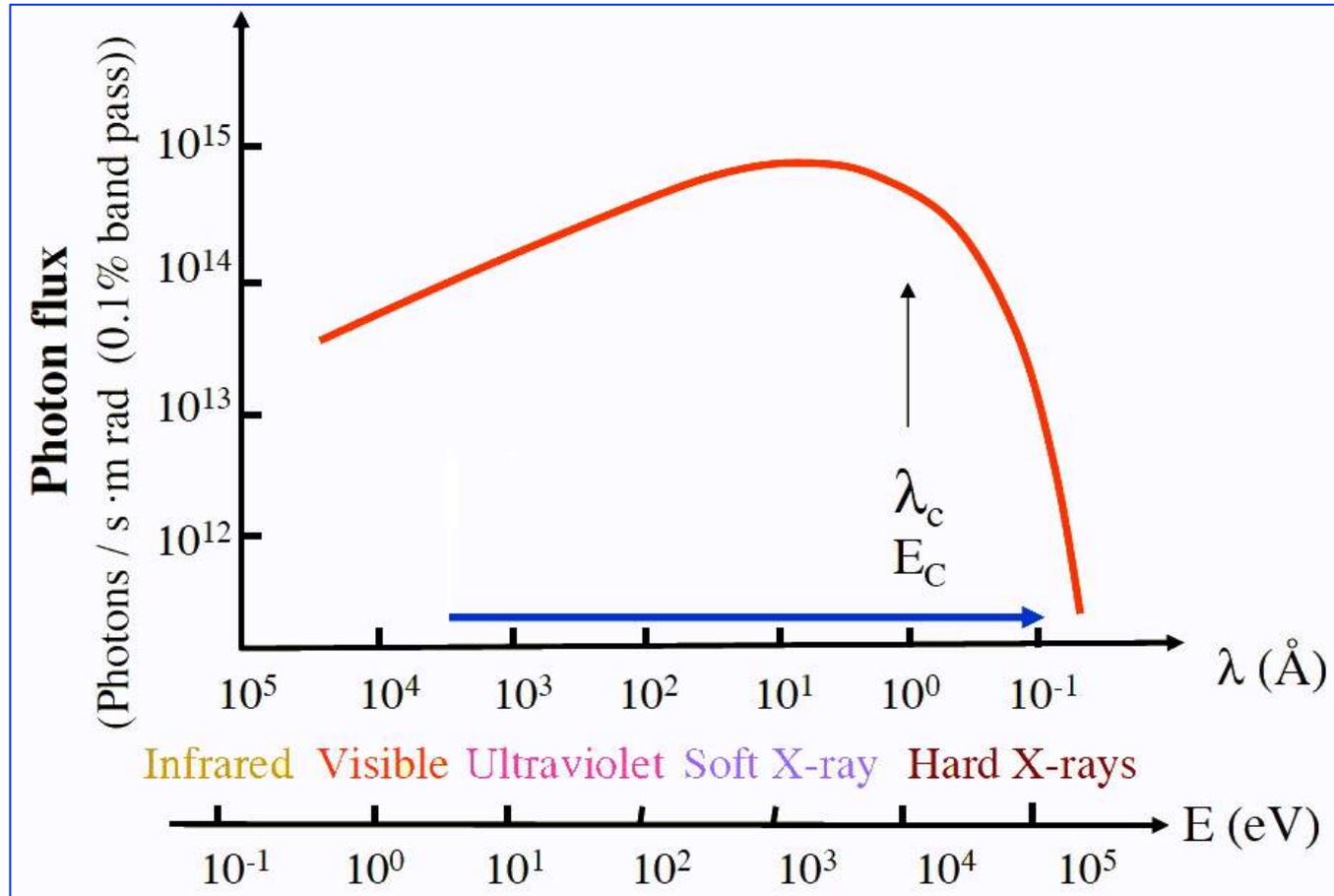
- 1) The shape of the radiation pattern changes: it is more in the forward direction!
- 2) the node at  $\theta' = 90^\circ$  in the frame of the radiating particle transforms to:

$$\tan \theta_{lab} = \frac{\sin \theta'}{\gamma(\cos \theta' + \beta)} = \frac{1}{\gamma\beta} \approx \frac{1}{\gamma}$$

$$\gamma^{-1} \approx mc^2/E$$

$$v \approx c \text{ or } \beta = v/c \approx 1$$

# Spectral distribution: universal synchrotron radiation function



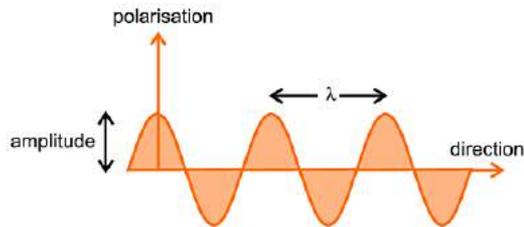
$E_c$  and  $\lambda_c$  respectively critical energy and critical wavelength

$$E_c [\text{keV}] = 2.218 \frac{E [\text{GeV}]^3}{\rho [m]} = 0.665 \cdot E [\text{GeV}]^2 \cdot B [\text{T}]$$

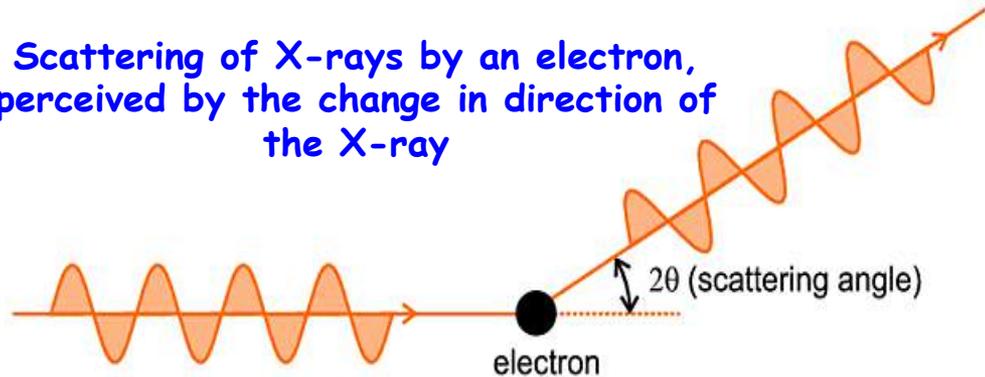
# Elastic Scattering and Diffraction

# Elastic or coherent Scattering

Incoming X-rays

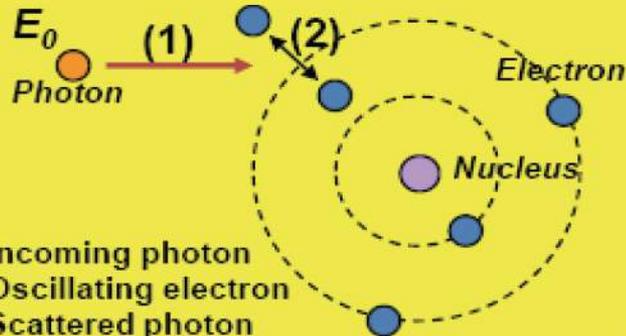


Scattering of X-rays by an electron, perceived by the change in direction of the X-ray



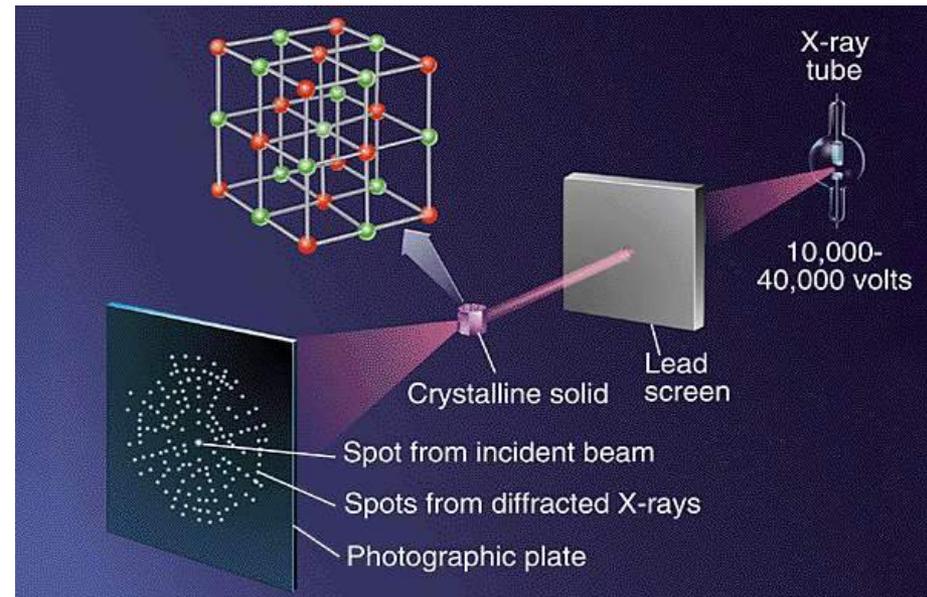
Coherent scattering

$$E_0 < E_{\text{binding}}$$



- (1) Incoming photon
  - (2) Oscillating electron
  - (3) Scattered photon
- No loss of energy.

Scattering of X-rays by the electrons of the atoms.

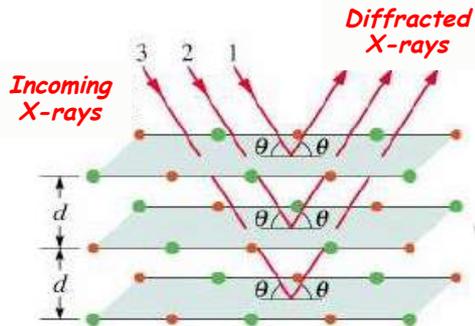


Scattering and diffraction of X-rays by the electrons of the atoms of a crystalline solid.

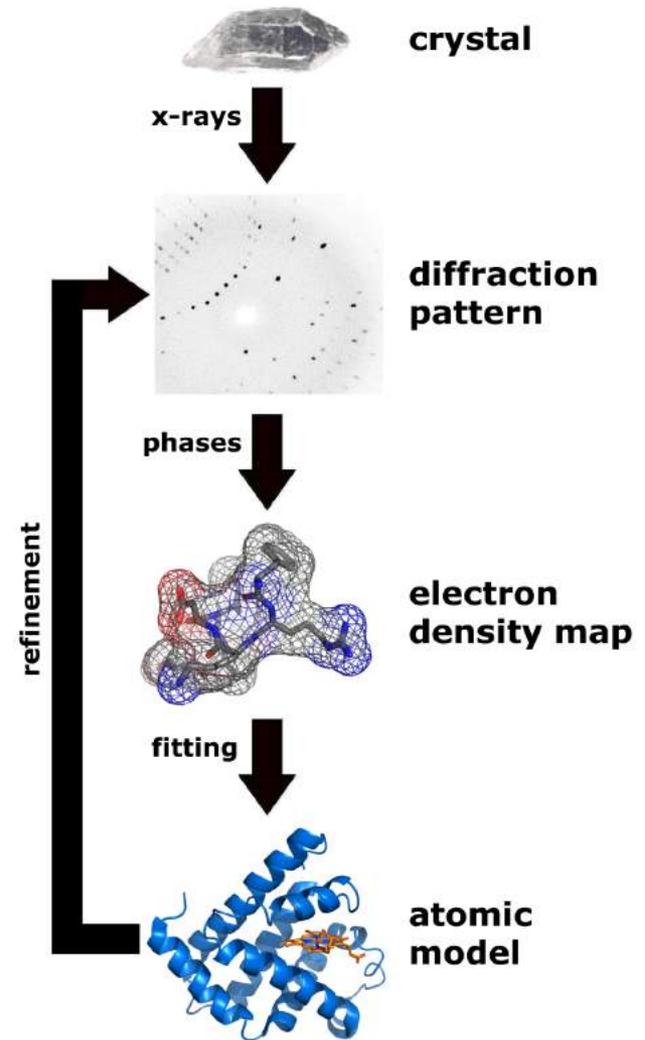
# Elastic scattering and diffraction

*X-ray diffraction is an important tool used to identify phases by comparison with data from known structures, quantify changes in the cell parameters, orientation, crystallite size and other structural parameters. It is also used to determine the (crystallographic) structure (i.e. cell parameters, space group and atomic coordinates) of novel or unknown crystalline materials.*

*The interference pattern of X-rays scattered by crystals (XRD or X Ray Diffraction pattern) can be used study the atomic structure of interest. Bragg's law explains the relation between:  $d$ , the distance between atomic layers in a crystal,  $\lambda$  is the wavelength of the incident X-ray beam and  $\theta$  the angle of incidence at which the faces of crystals appear to reflect X-ray beams.*

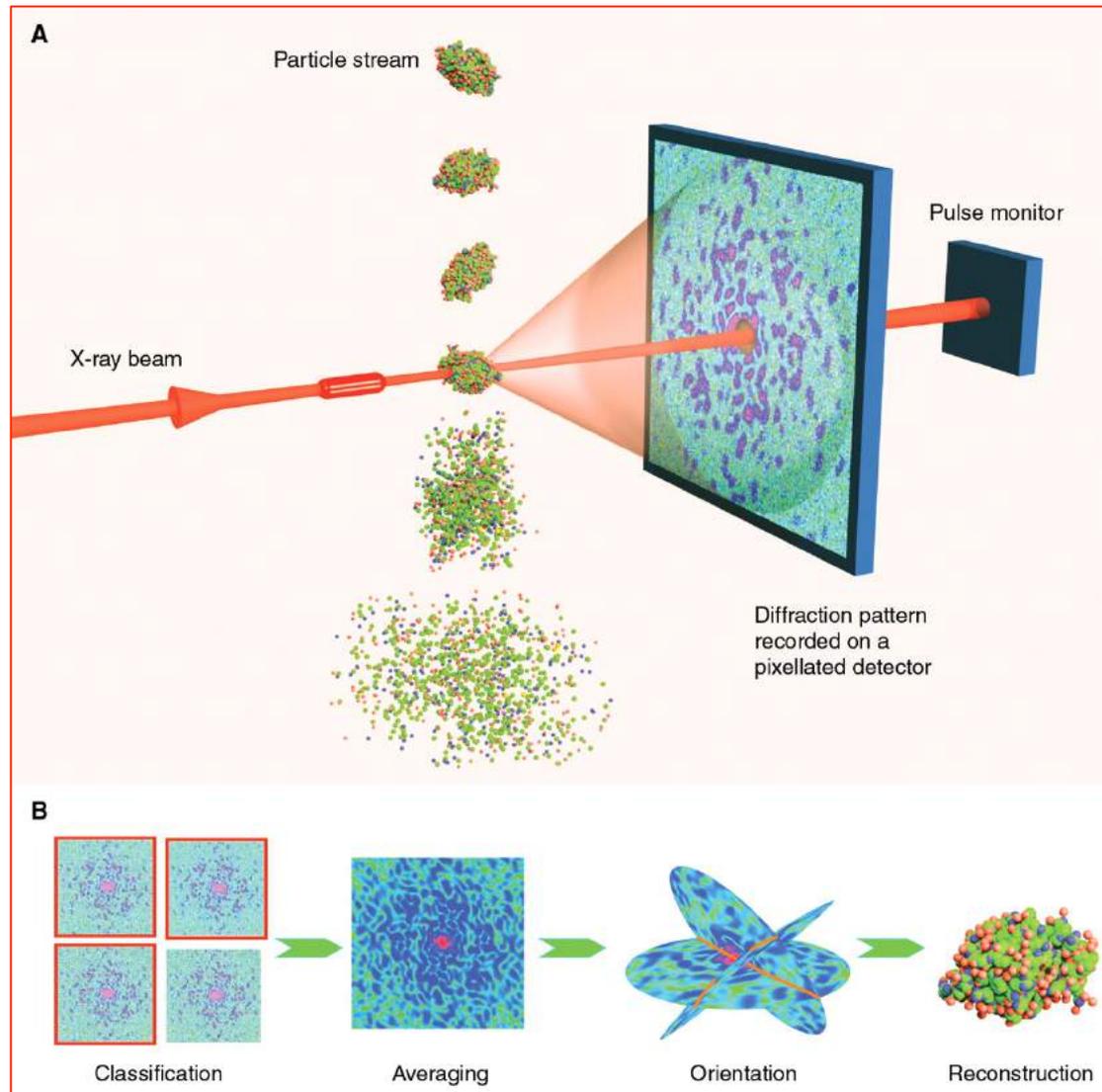


$$\text{Bragg's Law: } n\lambda = 2d \sin\theta$$



Structure determination by X-ray crystallography - Thomas Splettstoesser ([www.scistyle.com](http://www.scistyle.com))  
<https://commons.wikimedia.org/w/index.php?curid=1248574>

# Scattering and imaging



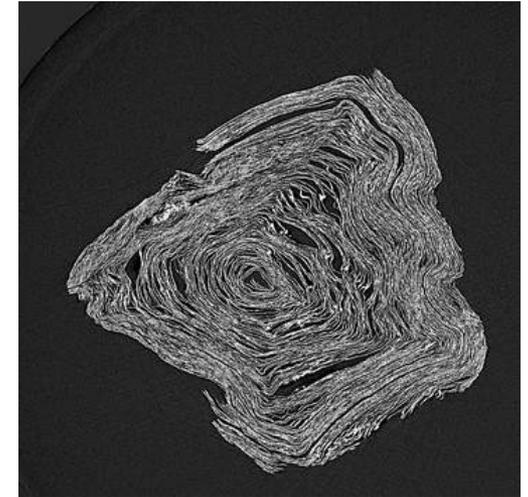
# Revealing letters in rolled Herculaneum papyri using X-ray phase-contrast tomography



Close up photograph of *Herculaneum Papyrus* scroll PHerc.Paris.4. The photographed zone is 5cm. (Credit: E. Brun)

Hundreds of papyrus rolls, buried by the eruption of Mount Vesuvius in 79 AD and belonging to the only library passed on from Antiquity, were discovered 260 years ago at Herculaneum.

These carbonized papyri are extremely fragile and are inevitably damaged or destroyed in the process of trying to open them to read their contents.



*A section of papyrus.* Letter sequences are found in a fragment of a hidden layer. (Credit: CNRS-IRHT UPR 841 / ESRF / CNR-IMM Unité de Naples)

V. Mocella et al., Nature Communications - DOI: 10.1038/ncomms6895- January 2015

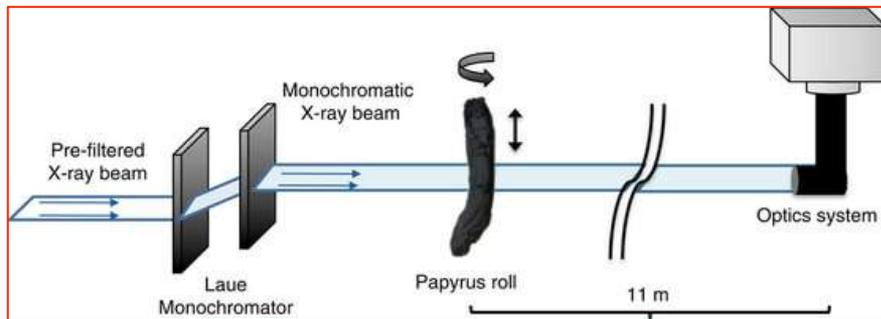
<https://www.youtube.com/watch?v=d3aWBgNYOCU>

# Revealing letters in rolled Herculaneum papyri.

In recent years, new imaging techniques have been developed to read the texts without unwrapping the rolls. *Until now, specialists have been unable to view the carbon-based ink of these papyri, even when they could penetrate the different layers of their spiral structure.*

For the first time *X-ray phase-contrast tomography* (beamline ID17 of the ESRF, Grenoble, France) can *reveal various letters hidden inside the precious papyri without unrolling them.*

This attempt opens up new opportunities to read many Herculaneum papyri, which are still rolled up, thus enhancing our knowledge of ancient Greek literature and philosophy.



The *papyrus alphabet from PHerc.Paris.4. as revealed by the XPCT experiment*, primarily from the *innermost region of the papyrus*, where the individual coils are more distinguishable, are reported on lines 1 and 2; on *line 3 the infrared images of the same letters from the unrolled (opened) papyrus PHerc. 1471*, which was used as a reference for the writing style of the scroll PHerc.Paris. 4.

