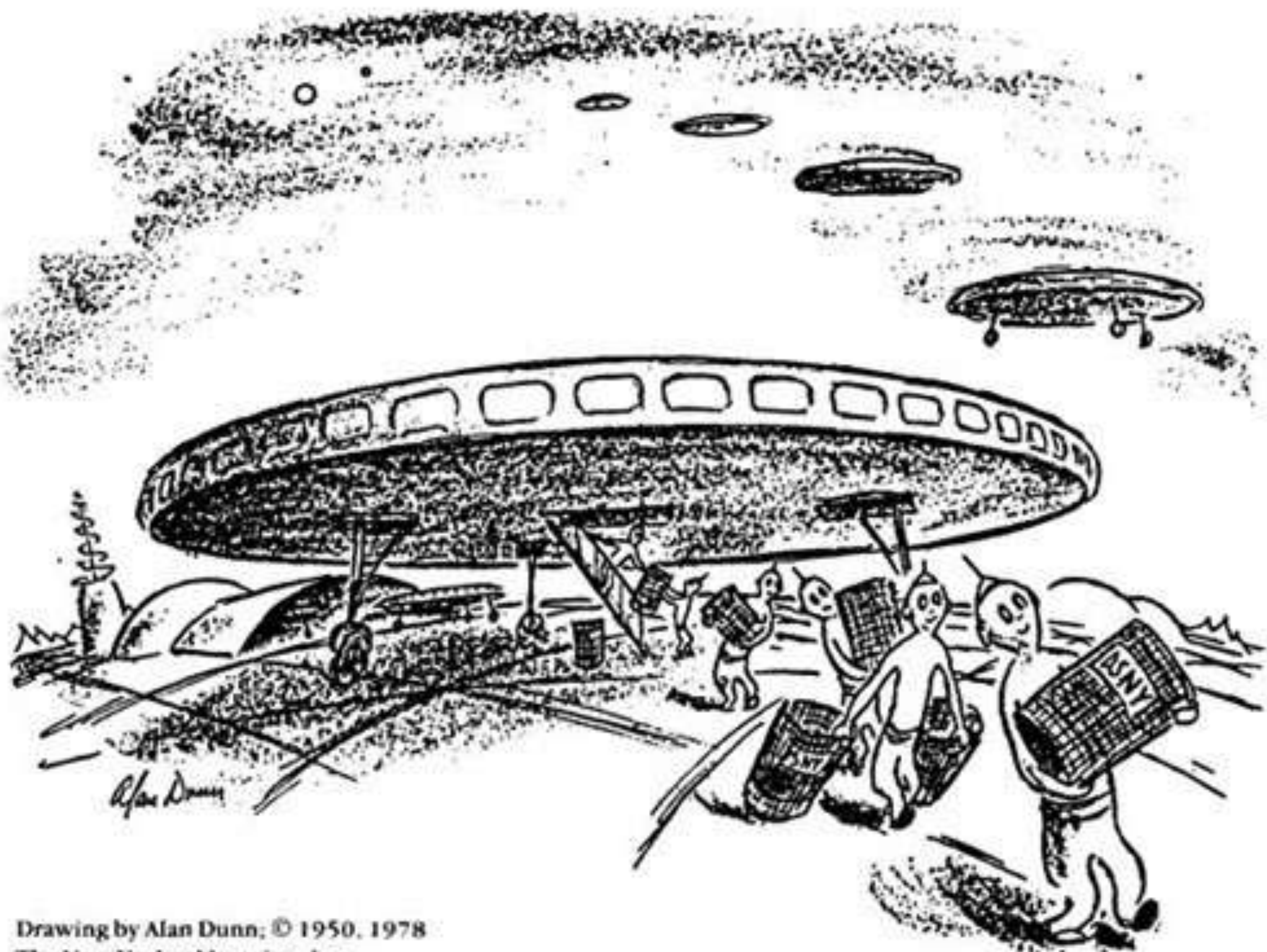


The background of the slide is a composite image. On the right side, there is a large, detailed view of the Earth from space, showing the Western Hemisphere with North and South America visible. The Earth's surface is covered in clouds and landmasses, with a blue ocean. In the upper left corner, there is a bright, glowing sun, and a few small, distant stars are visible in the blackness of space.

Where is everybody? Searching for life in the Universe

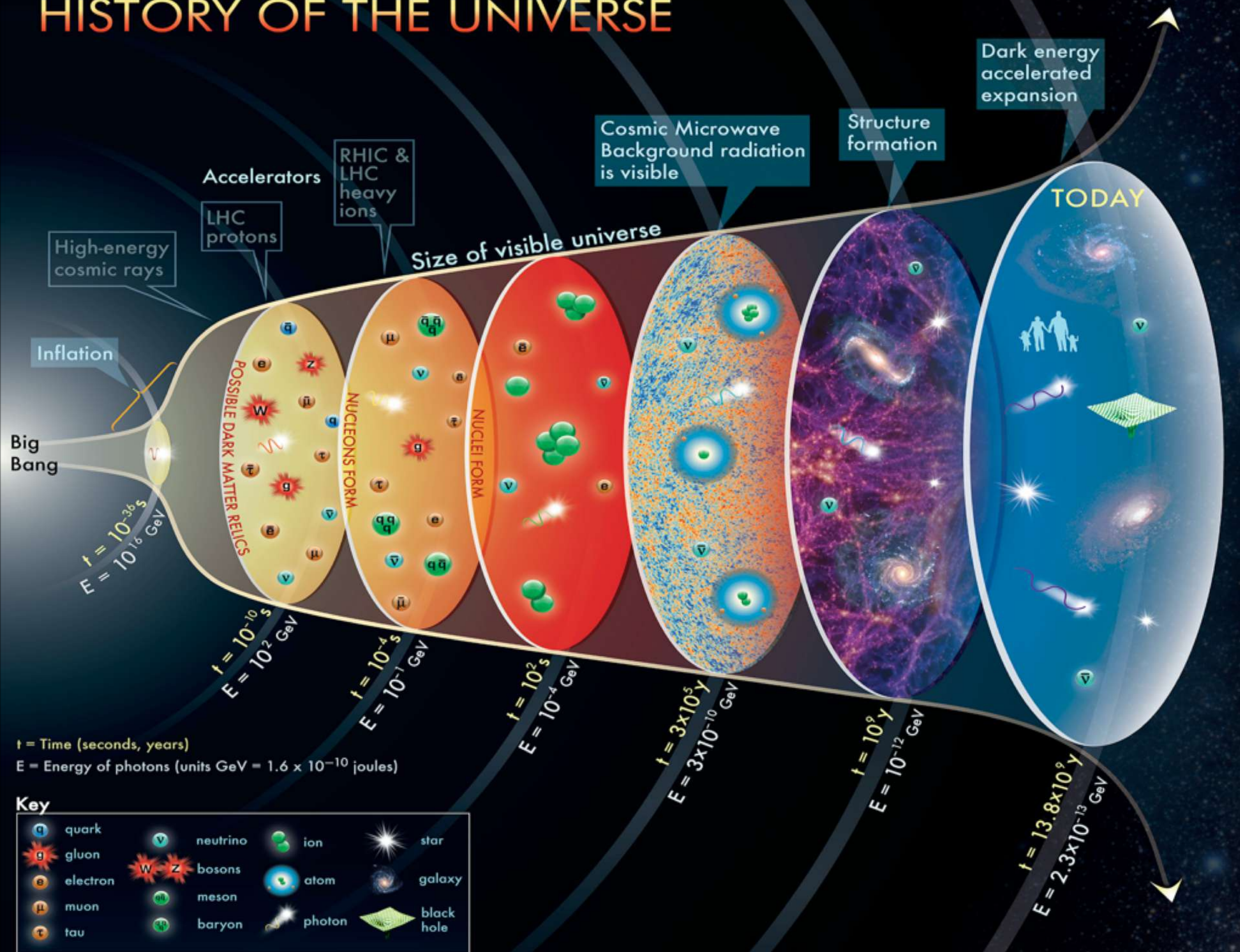
Amedeo Balbi

*Dipartimento di Fisica,
Università di Roma
«Tor Vergata»*



Drawing by Alan Dunn; © 1950, 1978
The New Yorker Magazine, Inc.

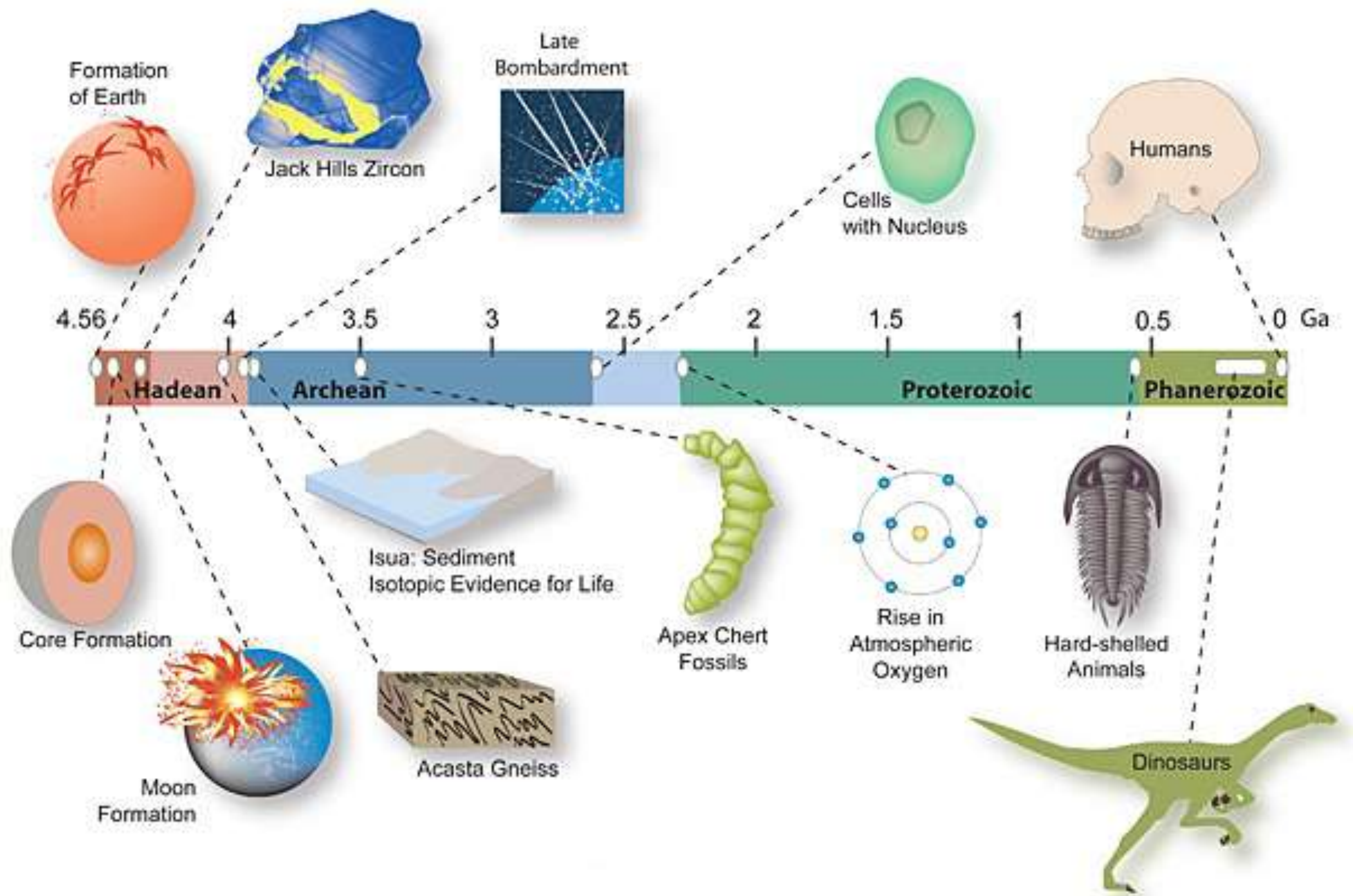
HISTORY OF THE UNIVERSE



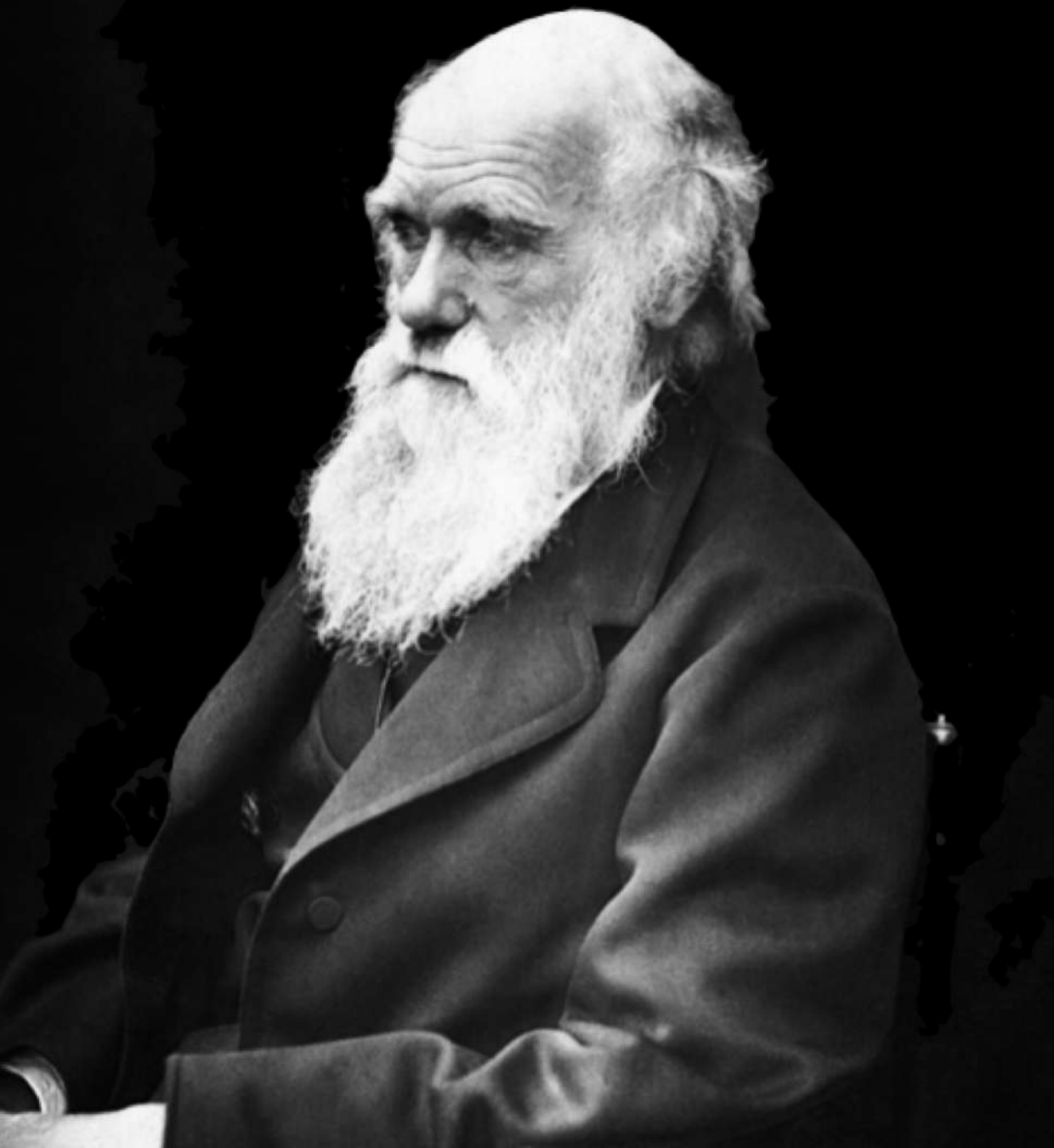
The concept for the above figure originated in a 1986 paper by Michael Turner.

Particle Data Group, LBNL © 2015

Supported by DOE





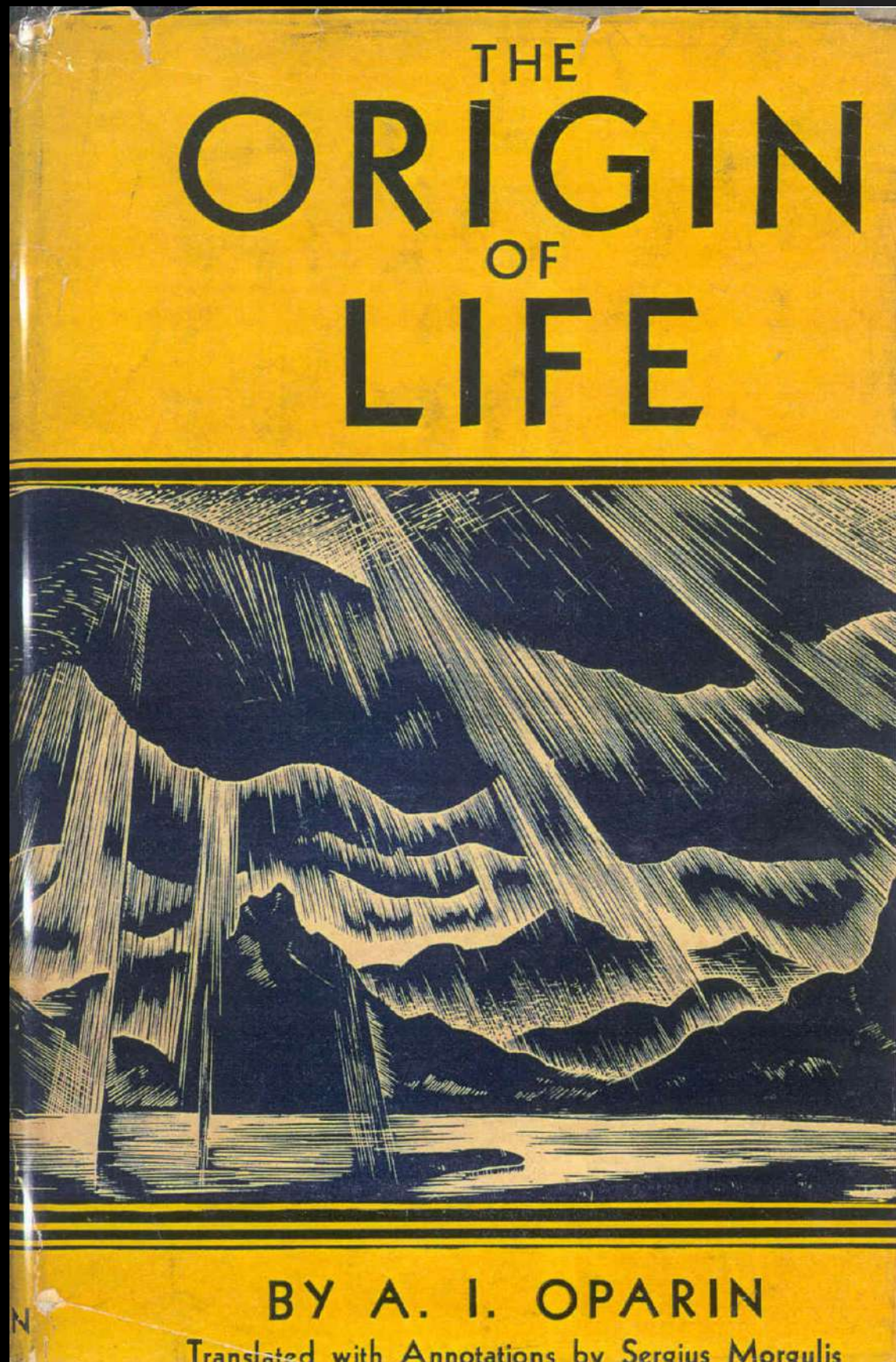


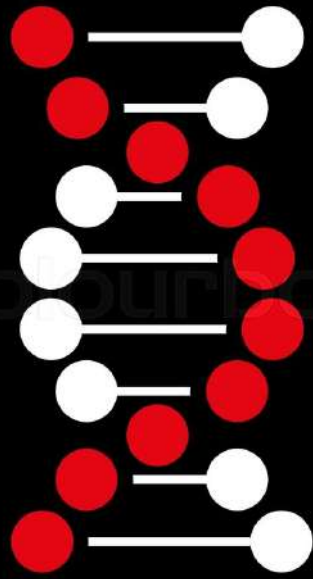
“It is **mere rubbish**, thinking
at present of the origin of life.”

— Charles Darwin, letter to Joseph
Hooker (1863)

“[...] in **some warm little
pond**, with all sorts of
ammonia and phosphoric salts,
light, heat, electricity, etc., [...]”

— Charles Darwin, letter to Joseph
Hooker (1871)





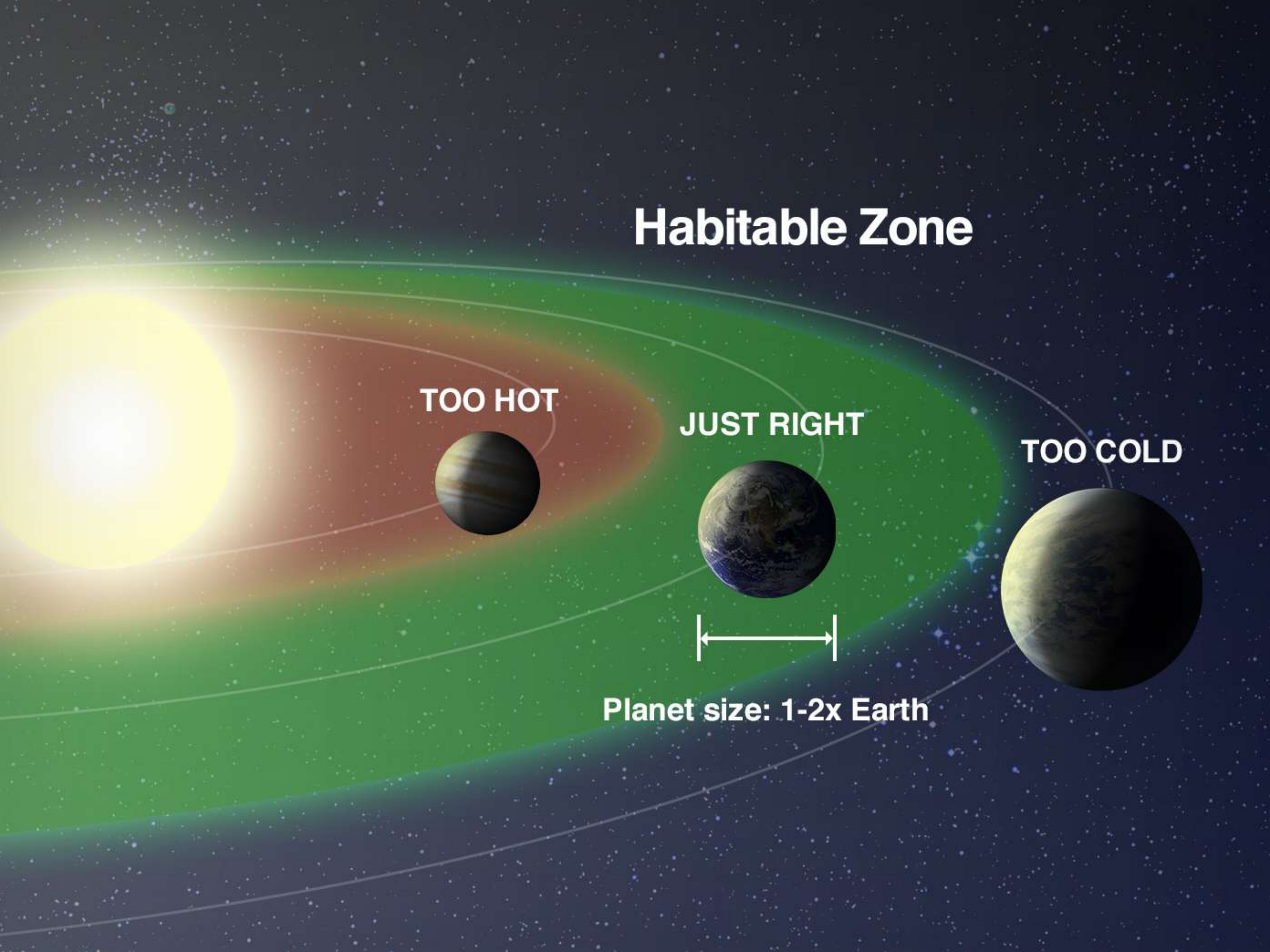
Habitable Zone

TOO HOT

JUST RIGHT

TOO COLD

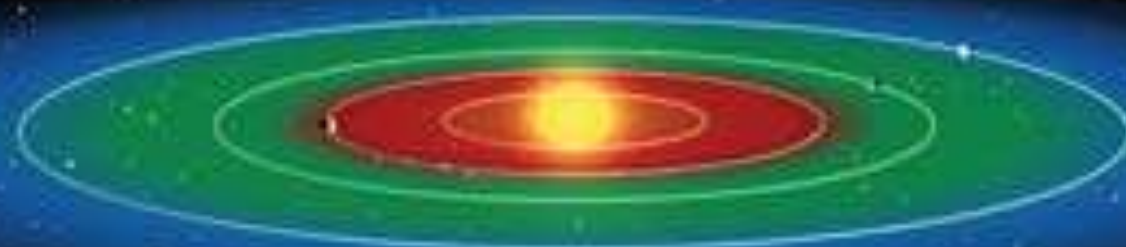
Planet size: 1-2x Earth



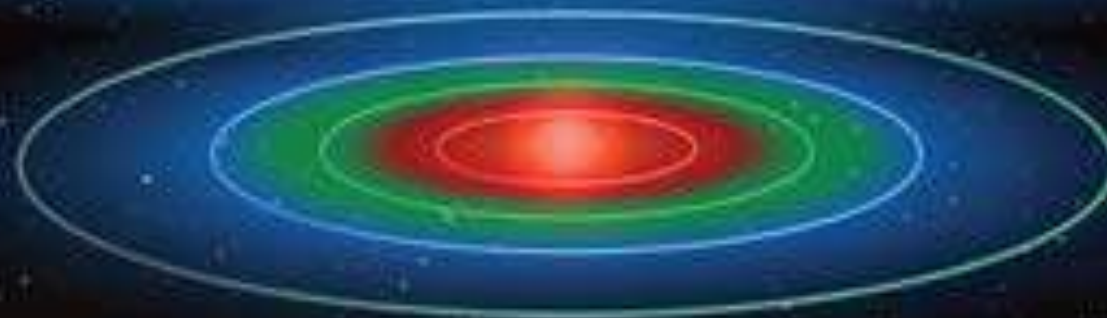
Hotter Stars



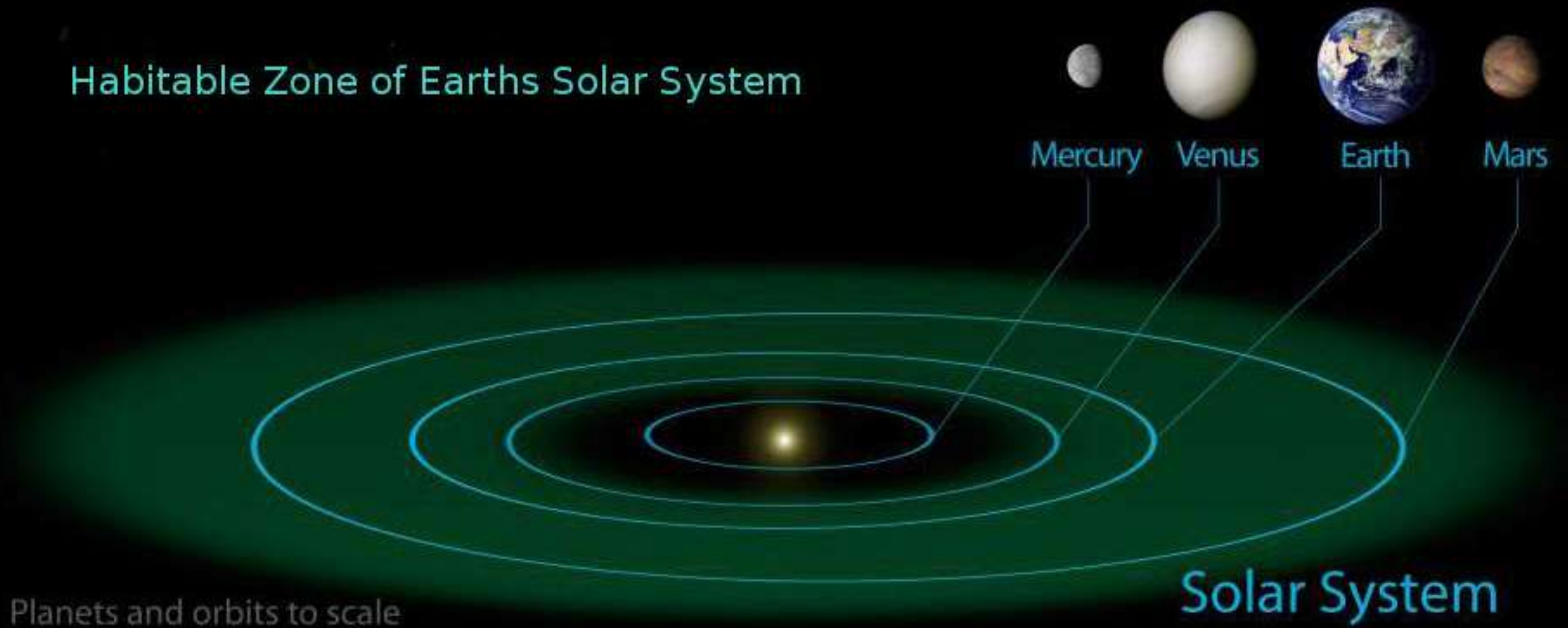
Sunlike Stars



Cooler Stars



Habitable Zone of Earths Solar System





Venus



Earth



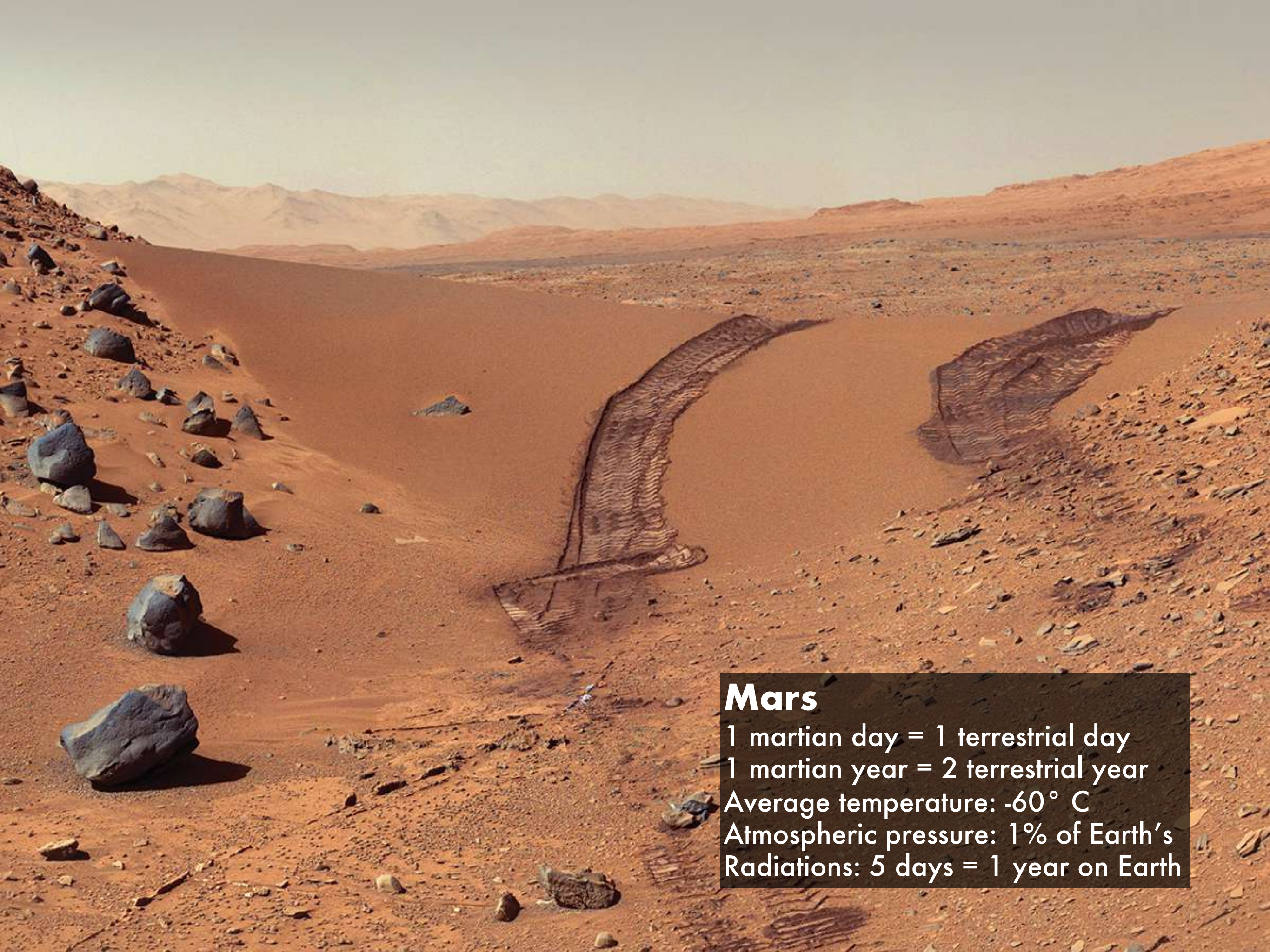
Mars



Venus

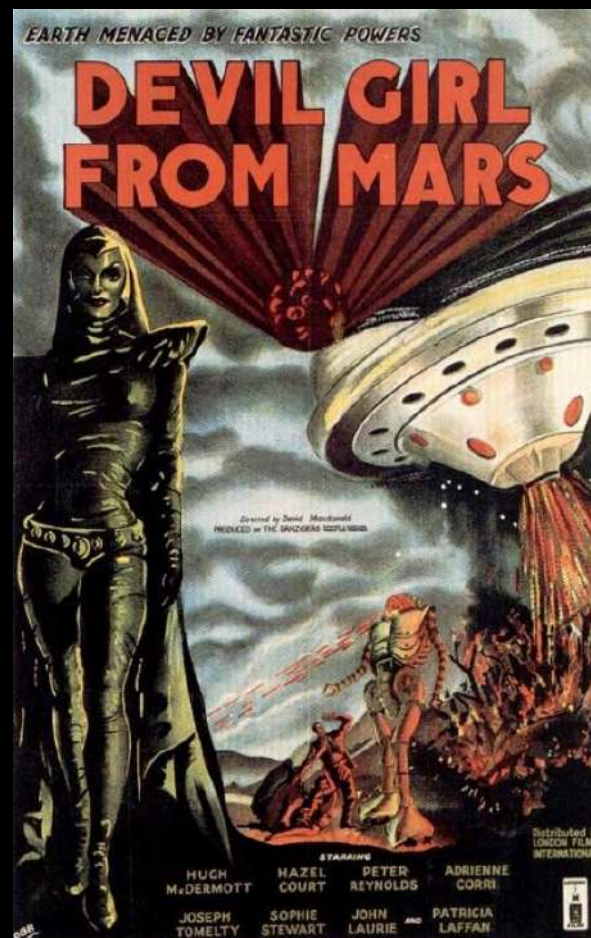
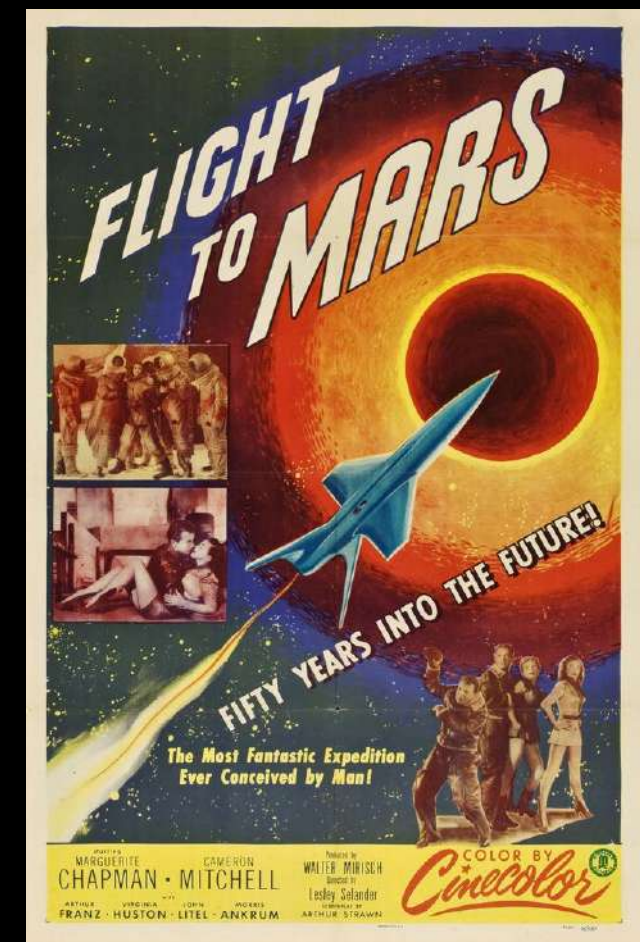
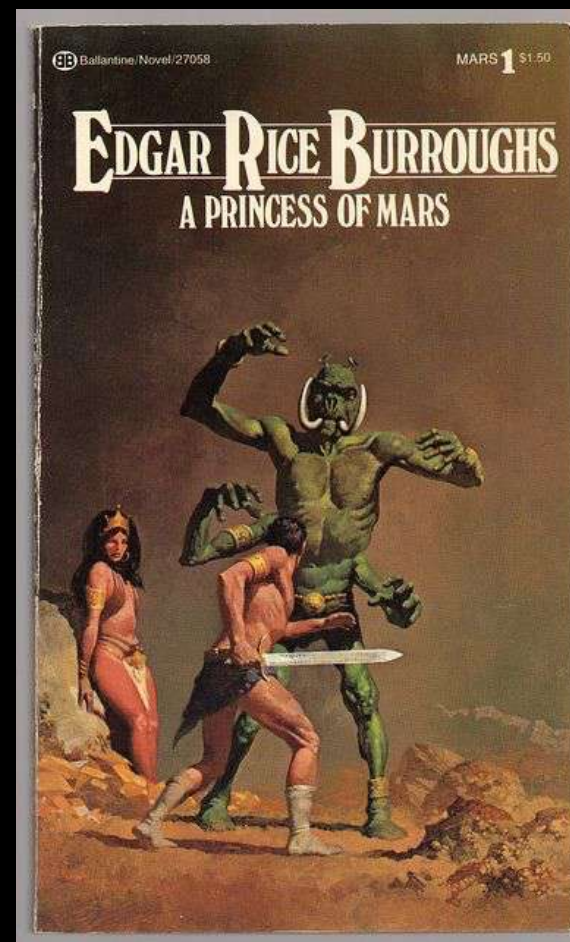
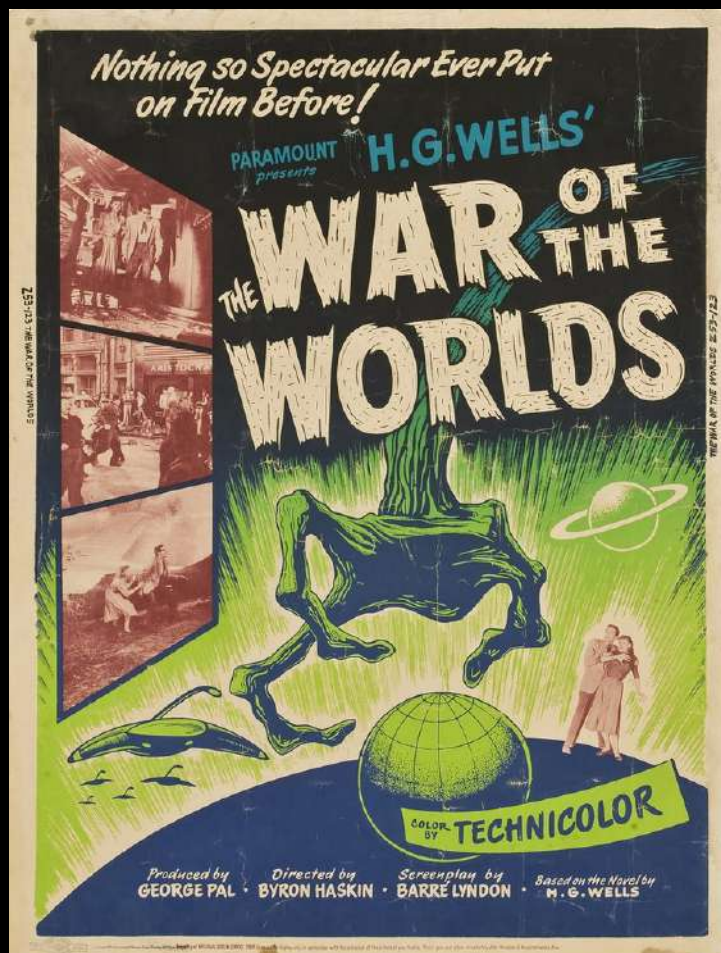
Average temperature: 500°C

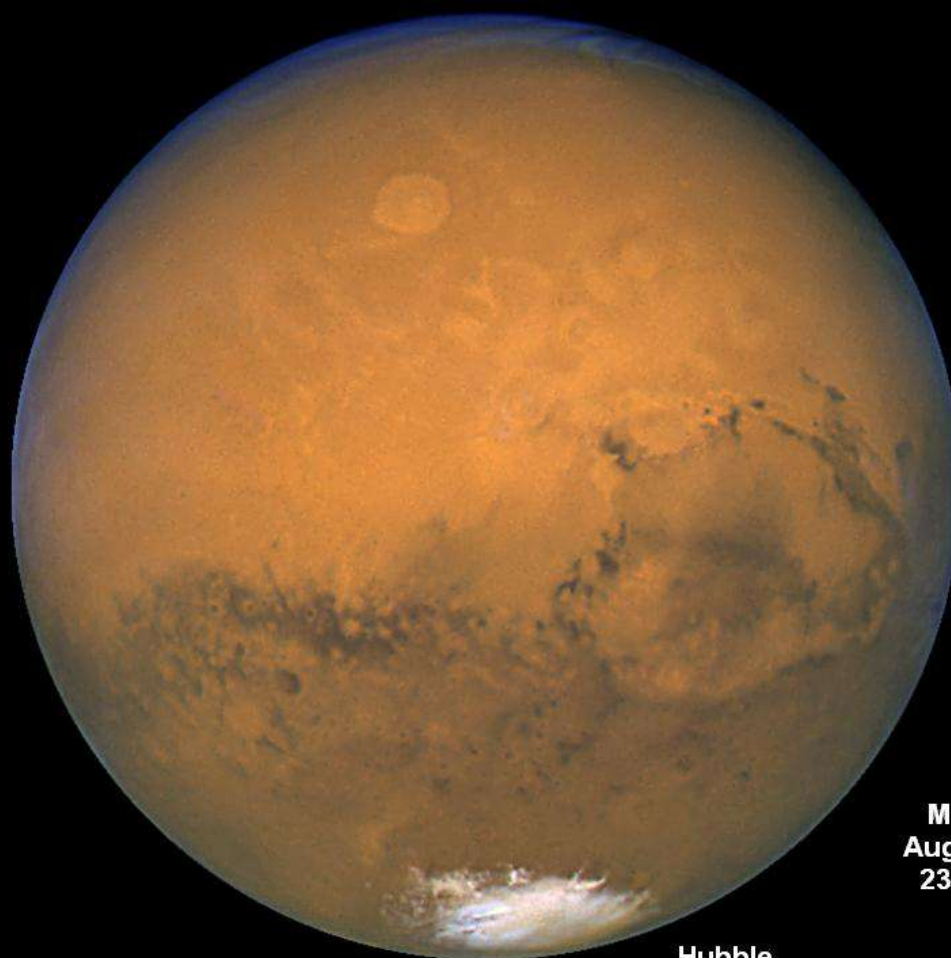
Atmospheric pressure: 90 atm



Mars

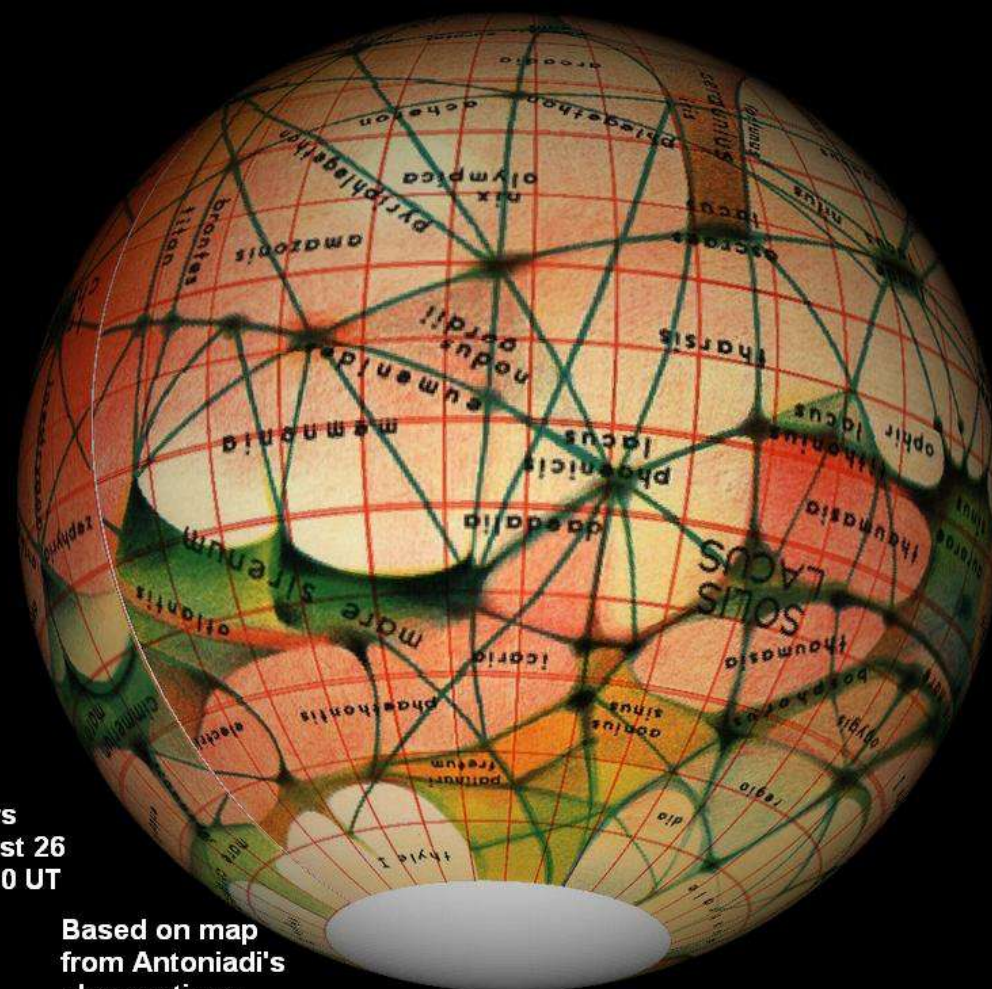
1 martian day = 1 terrestrial day
1 martian year = 2 terrestrial year
Average temperature: -60°C
Atmospheric pressure: 1% of Earth's
Radiations: 5 days = 1 year on Earth



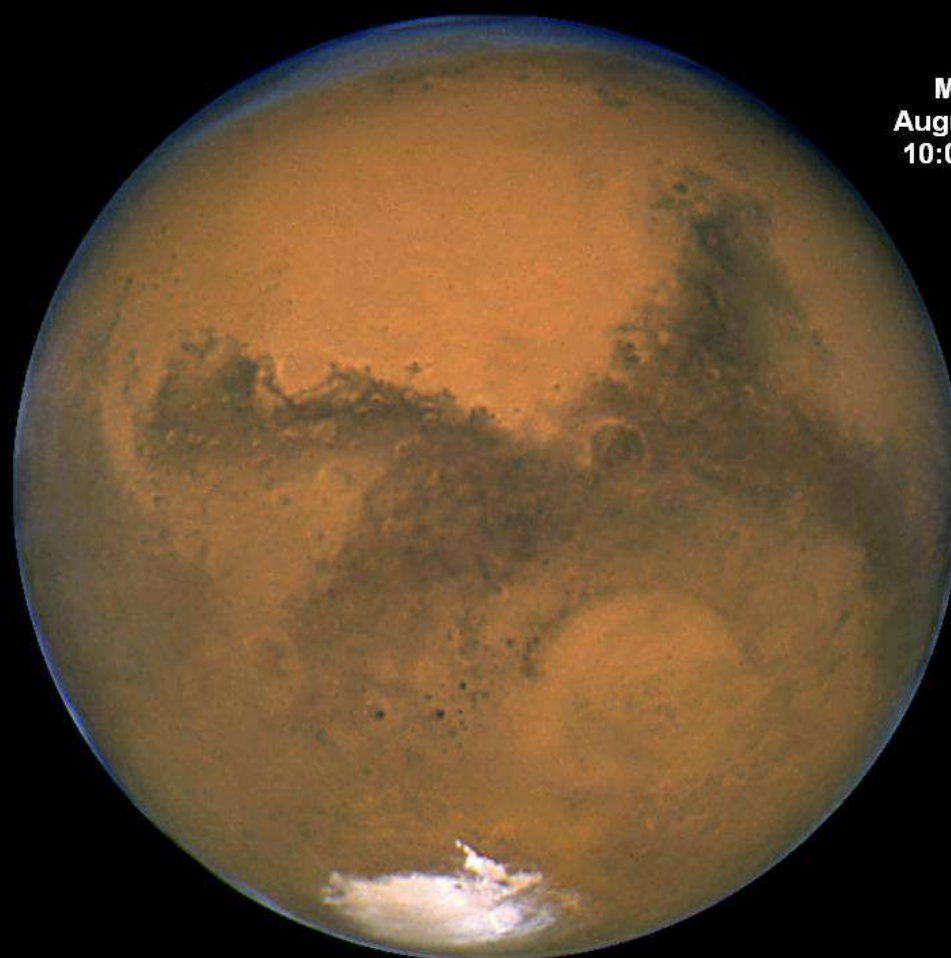


Hubble

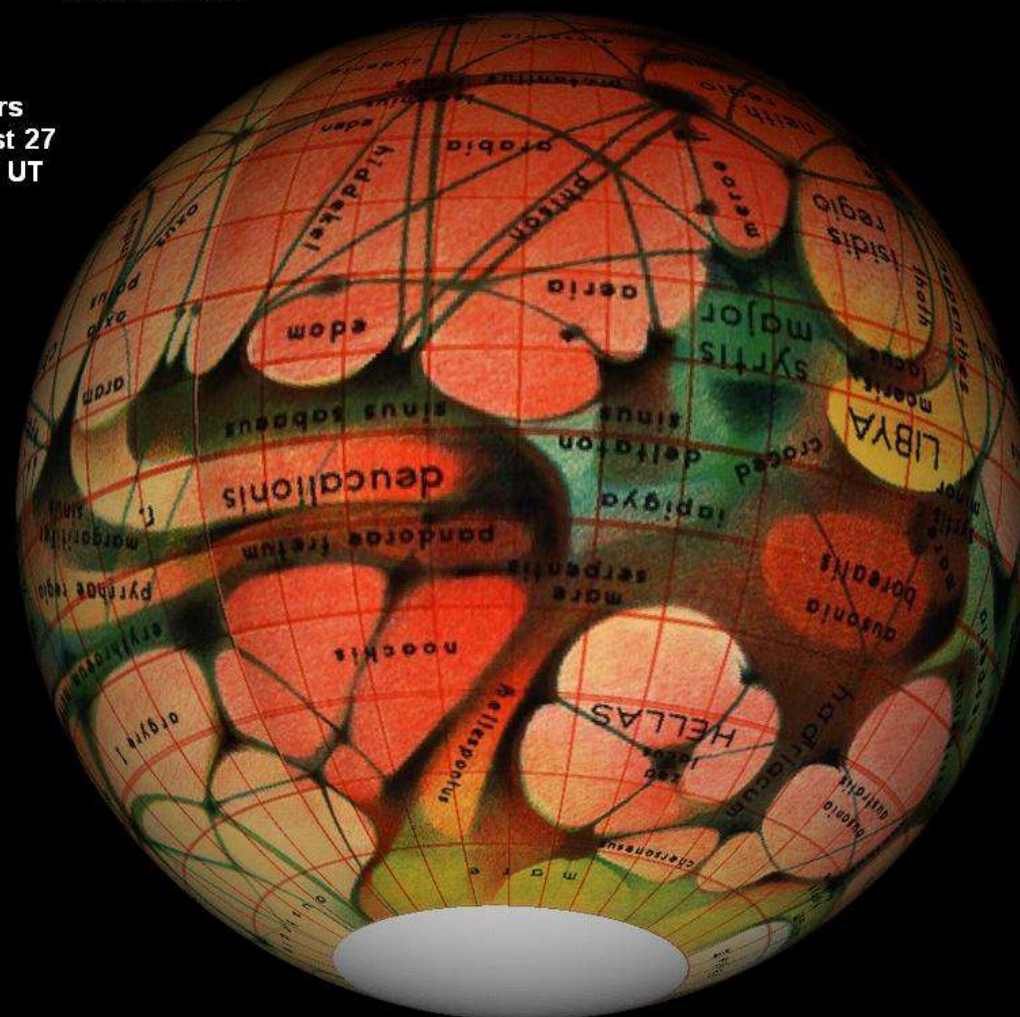
Mars
August 26
23:00 UT



Based on map
from Antoniadi's
observations



Mars
August 27
10:00 UT



THE ASTROPHYSICAL JOURNAL

AN INTERNATIONAL REVIEW OF SPECTROSCOPY AND
ASTRONOMICAL PHYSICS

VOLUME 126

SEPTEMBER 1957

NUMBER 2

SPECTROSCOPIC EVIDENCE FOR VEGETATION ON MARS

WILLIAM M. SINTON

Smithsonian Astrophysical Observatory

Received May 6, 1957

ABSTRACT

A new test for the presence of vegetation on Mars depends on the fact that all organic molecules have absorption bands in the vicinity of $3.4\ \mu$. These bands have been studied in the reflection spectrum of terrestrial plants, and it is found that for most plants a doublet band appears which has a separation of about $0.1\ \mu$ and is centered about $3.46\ \mu$. Spectra of Mars taken during the 1956 opposition indicate the probable presence of this band. This evidence and the well-known seasonal changes of the dark areas make it extremely probable that vegetation in some form is present.

Science

Vol 130, Issue 3384
06 November 1959

Further Evidence of Vegetation on Mars

The presence of large organic molecules is
indicated by recent infrared-spectroscopic tests.

William M. Sinton

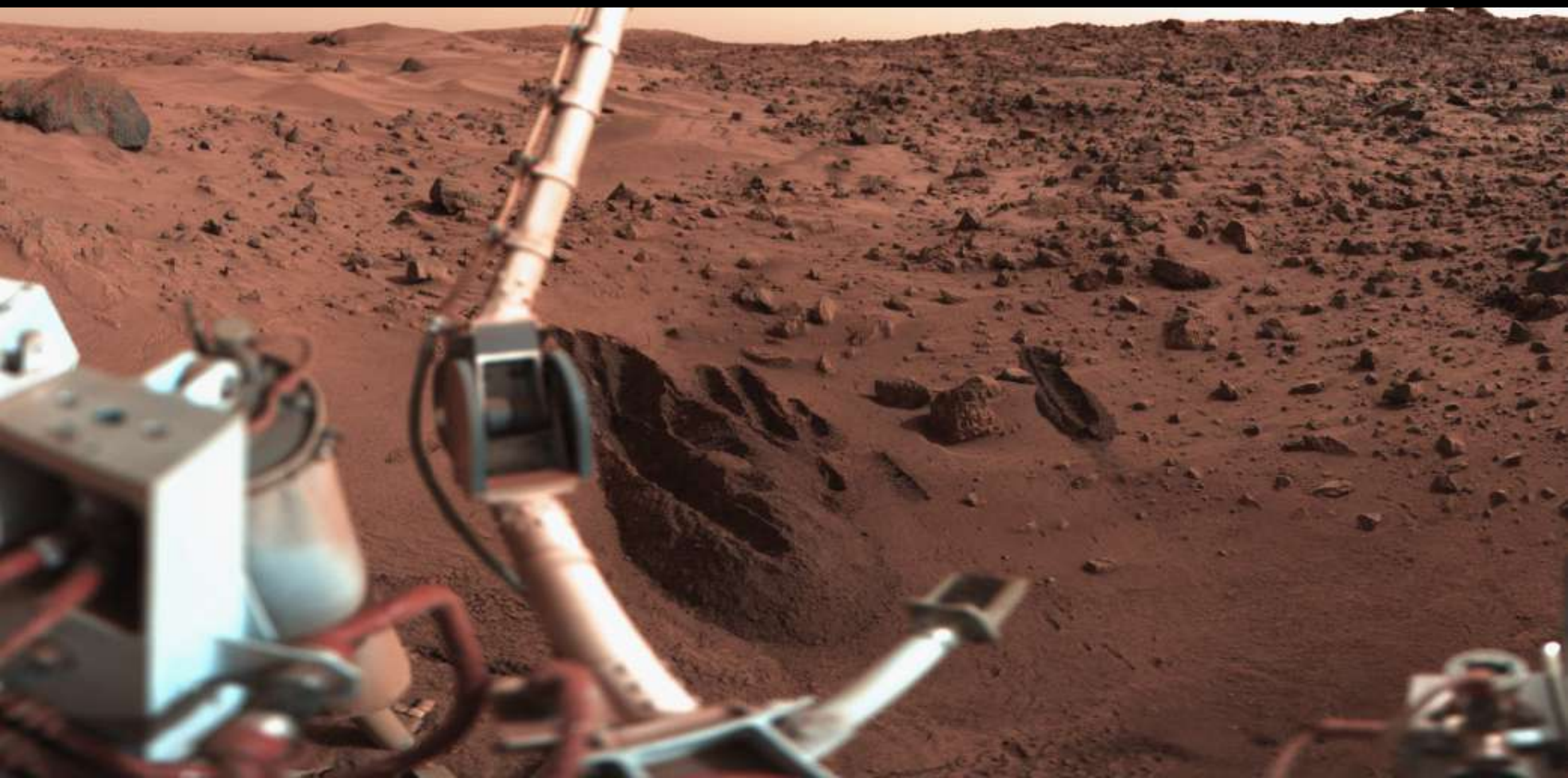
There has long been evidence pointing to the presence of vegetation on Mars. Photographs taken by E. C.

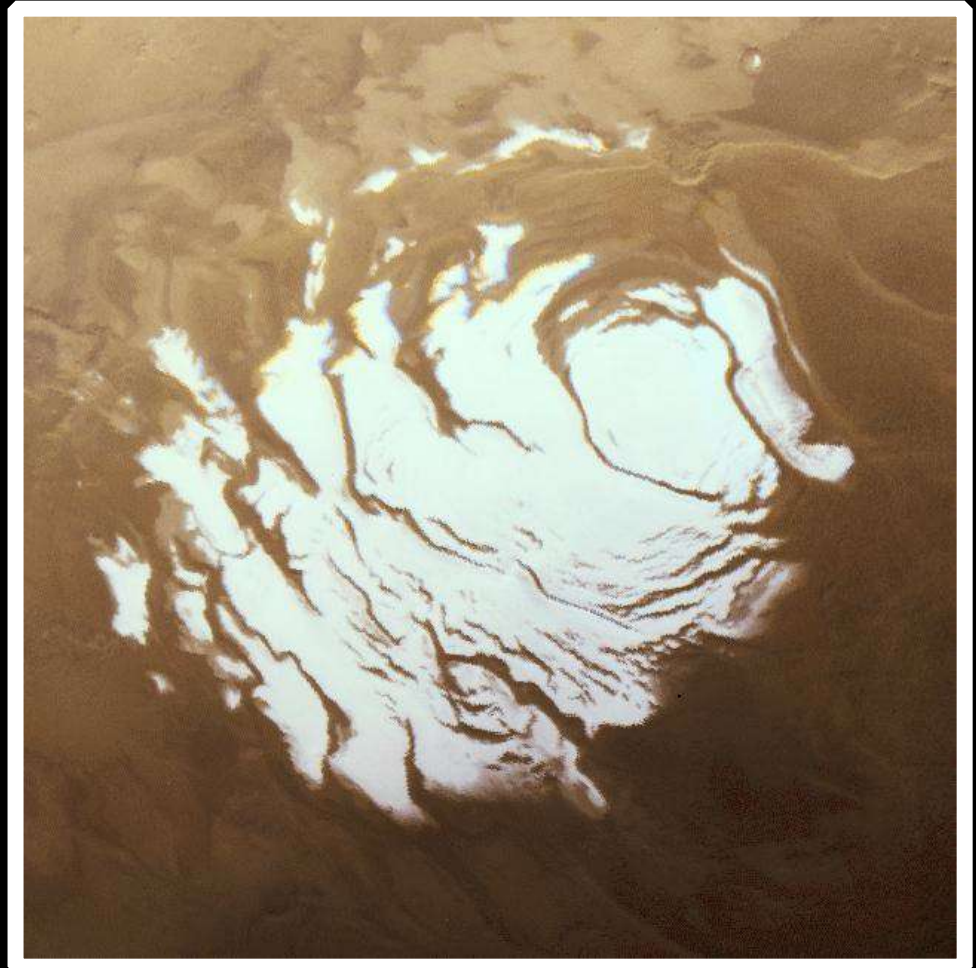
(2). The region in which it is situated has, however, been undergoing development for many years.

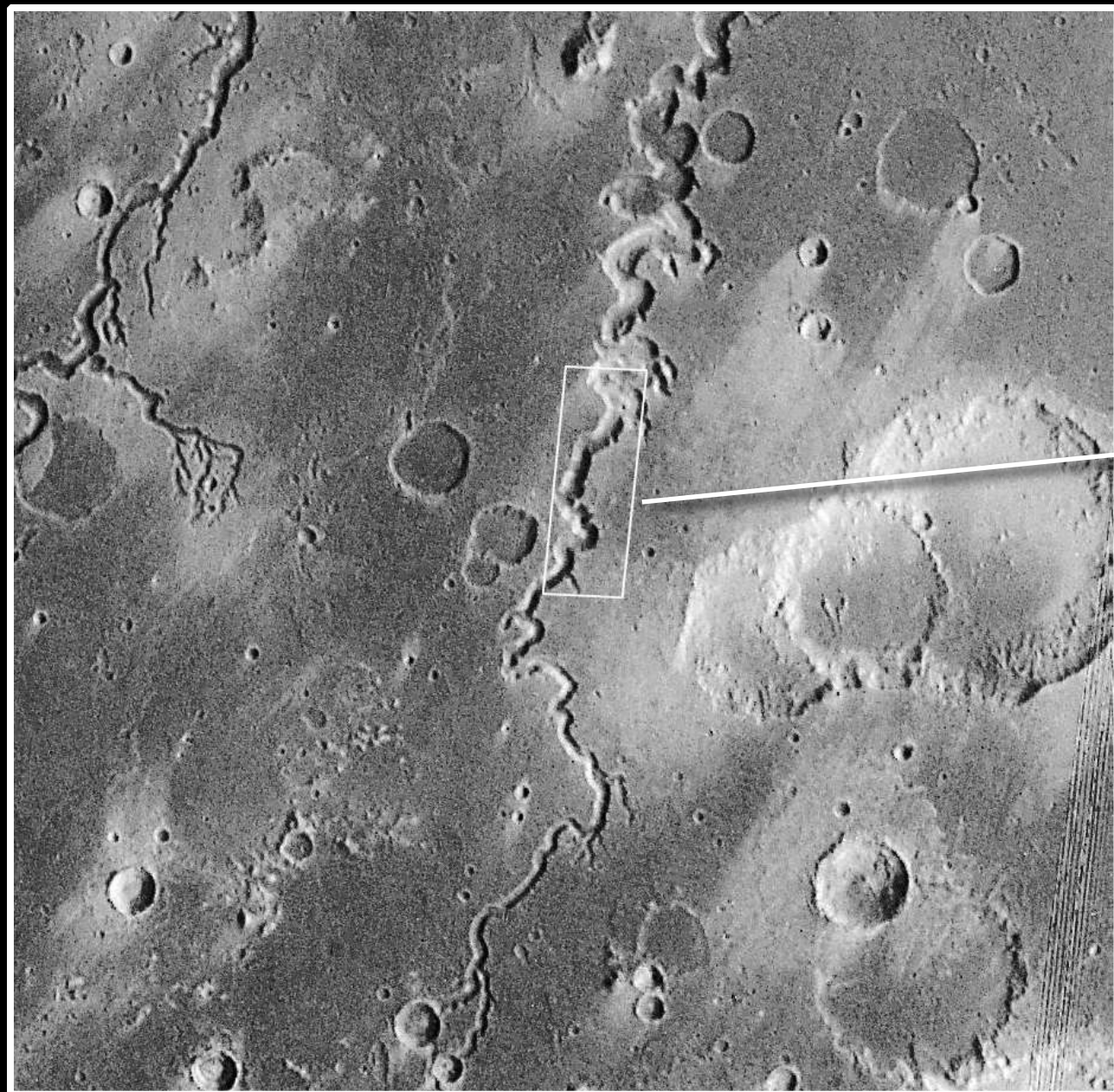
Martian plants possess a pigment that absorbs the near infrared.

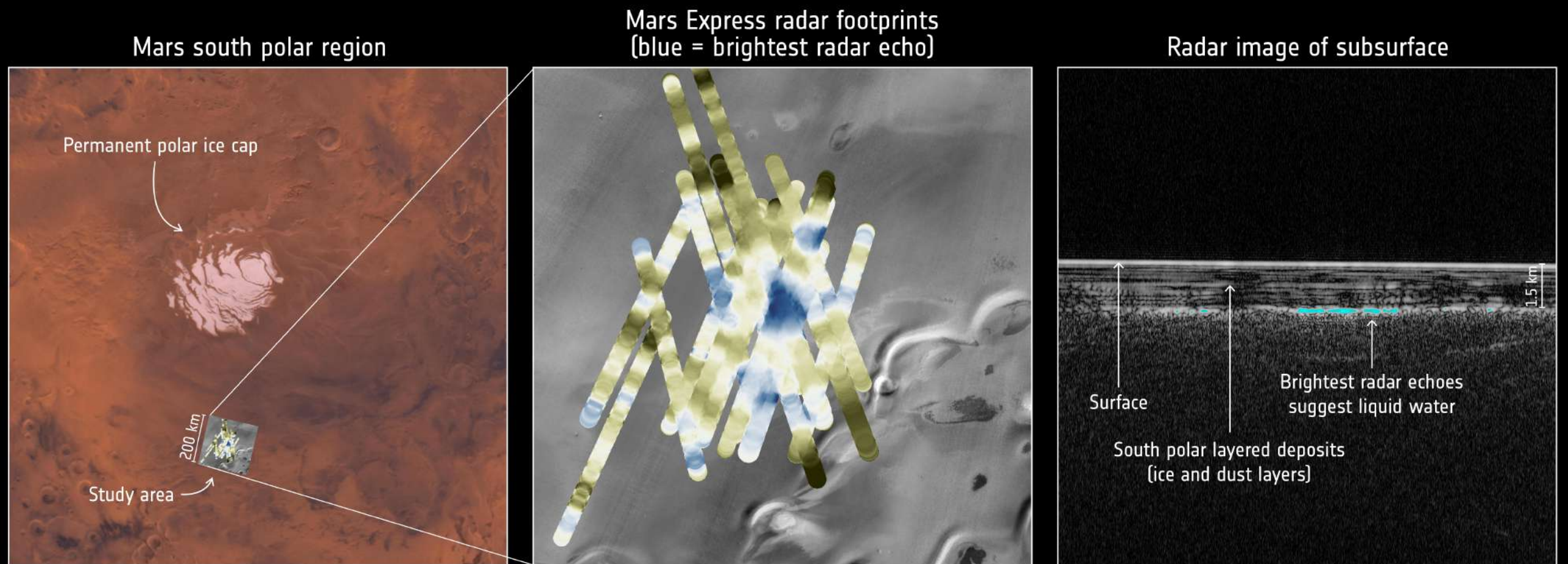
Using the 61-inch telescope of the Harvard College Observatory during the 1956 opposition, I made a new test for the presence of organic molecules on Mars (5). Organic molecules possess strong absorption bands at $3.5\ \mu$ as a result of the resonance of their carbon-hydrogen bonds. It was found that in the plants tested this band was double, most likely as a result of interaction between a pair of hydrogen atoms attached to the same carbon atom, as occurs in paraffin molecules.

The results of the 1956 observations indicated the presence of the band in the light reflected from Mars, but they left some doubt about the reality of the absorption. Furthermore, the regions of Mars which produced the absorption were not ascertained in this work. At

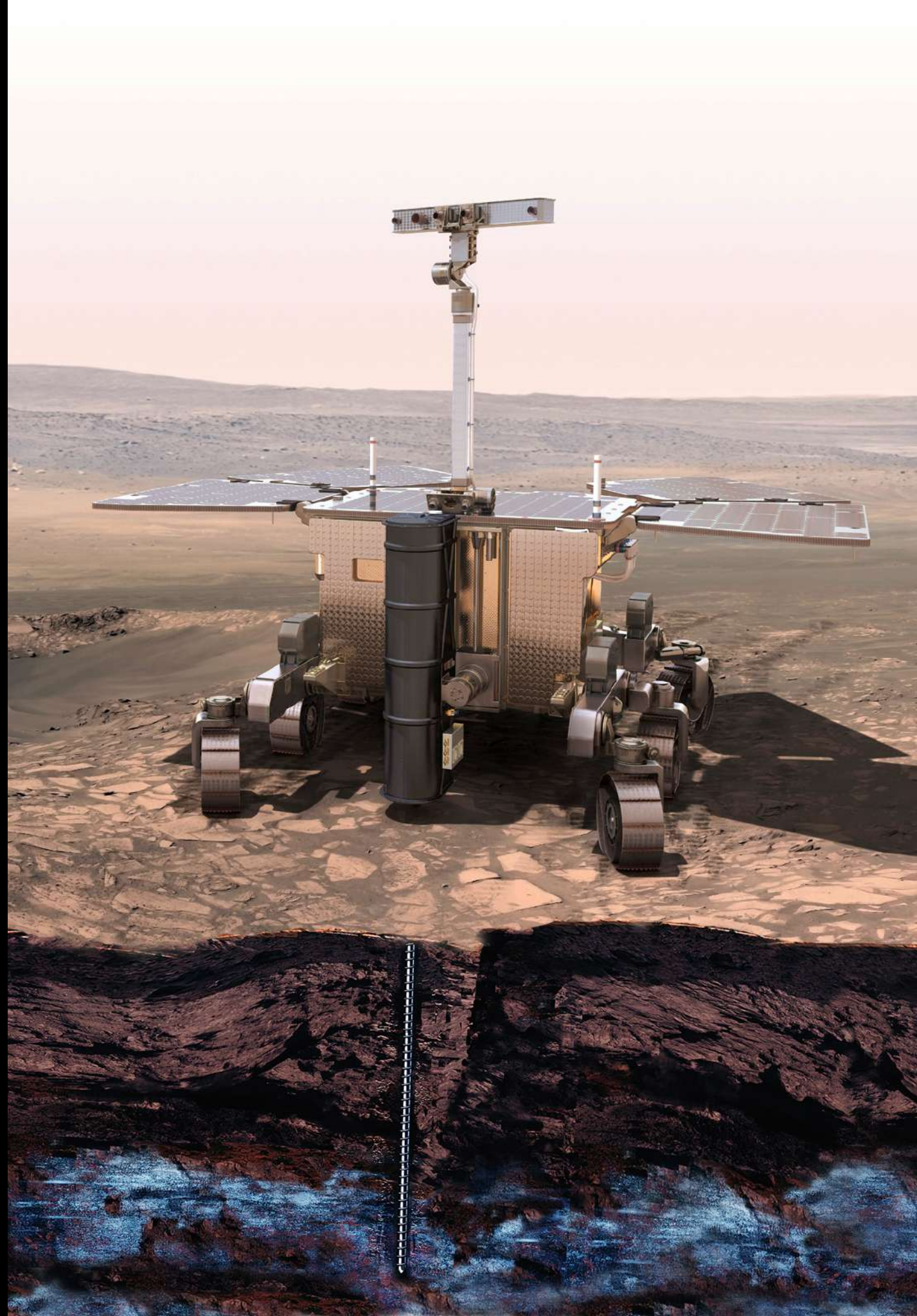








Orosei et al., *Science* (2018)

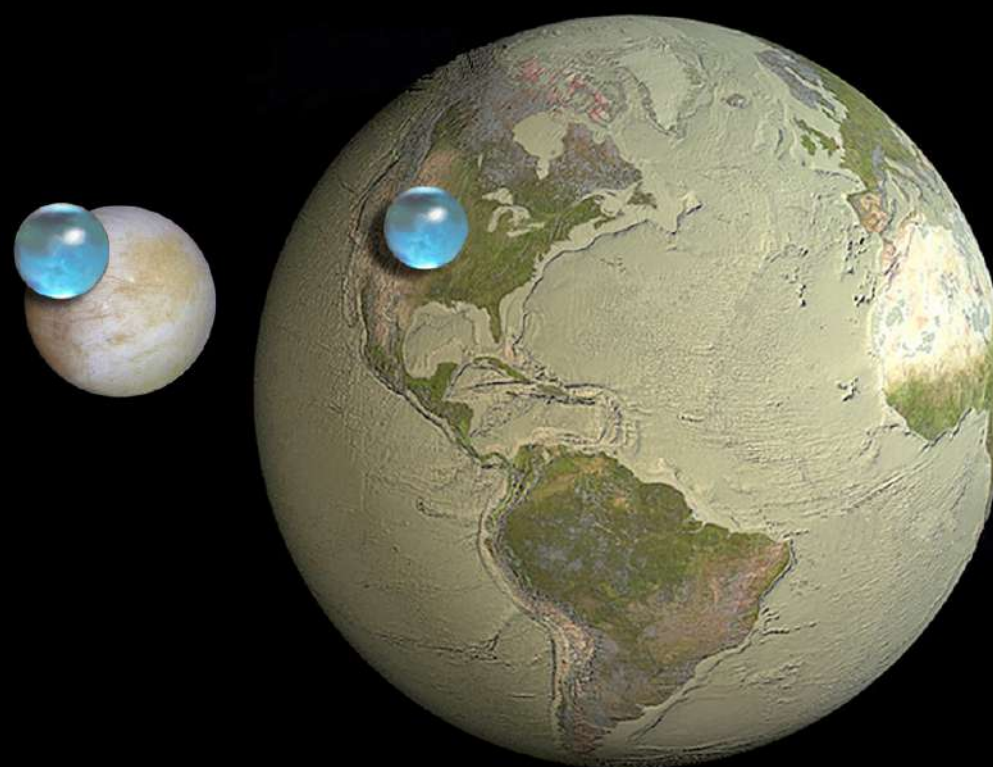
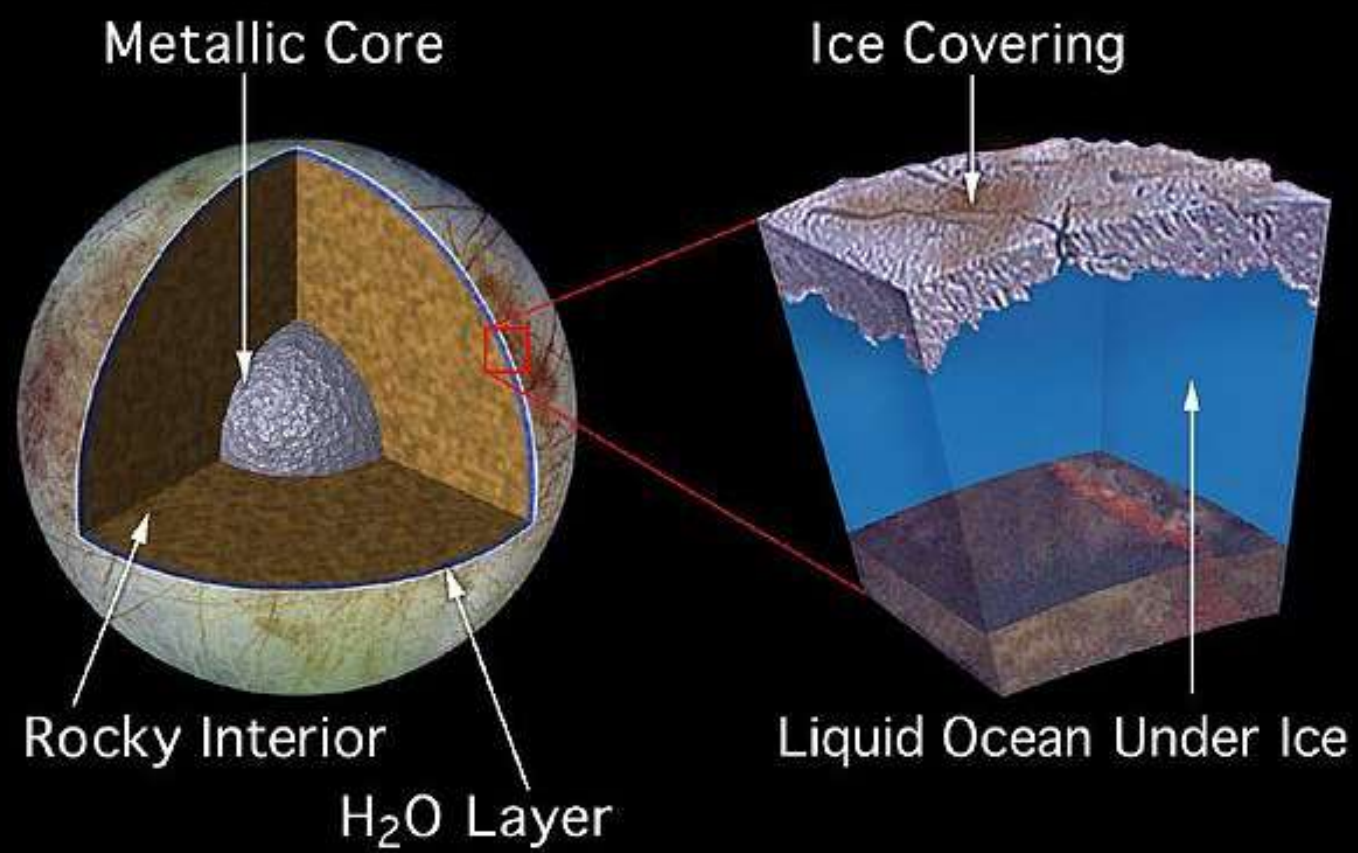


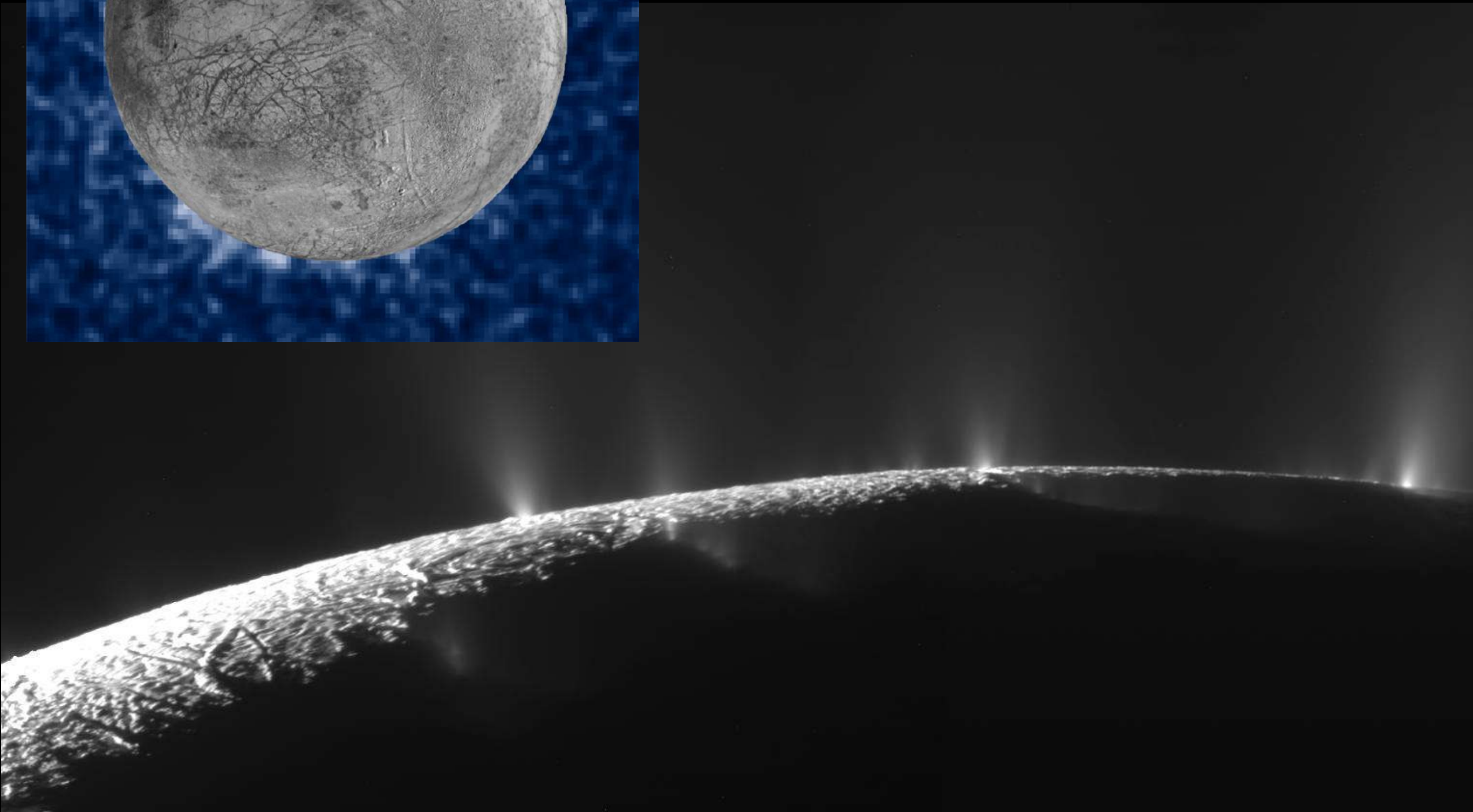
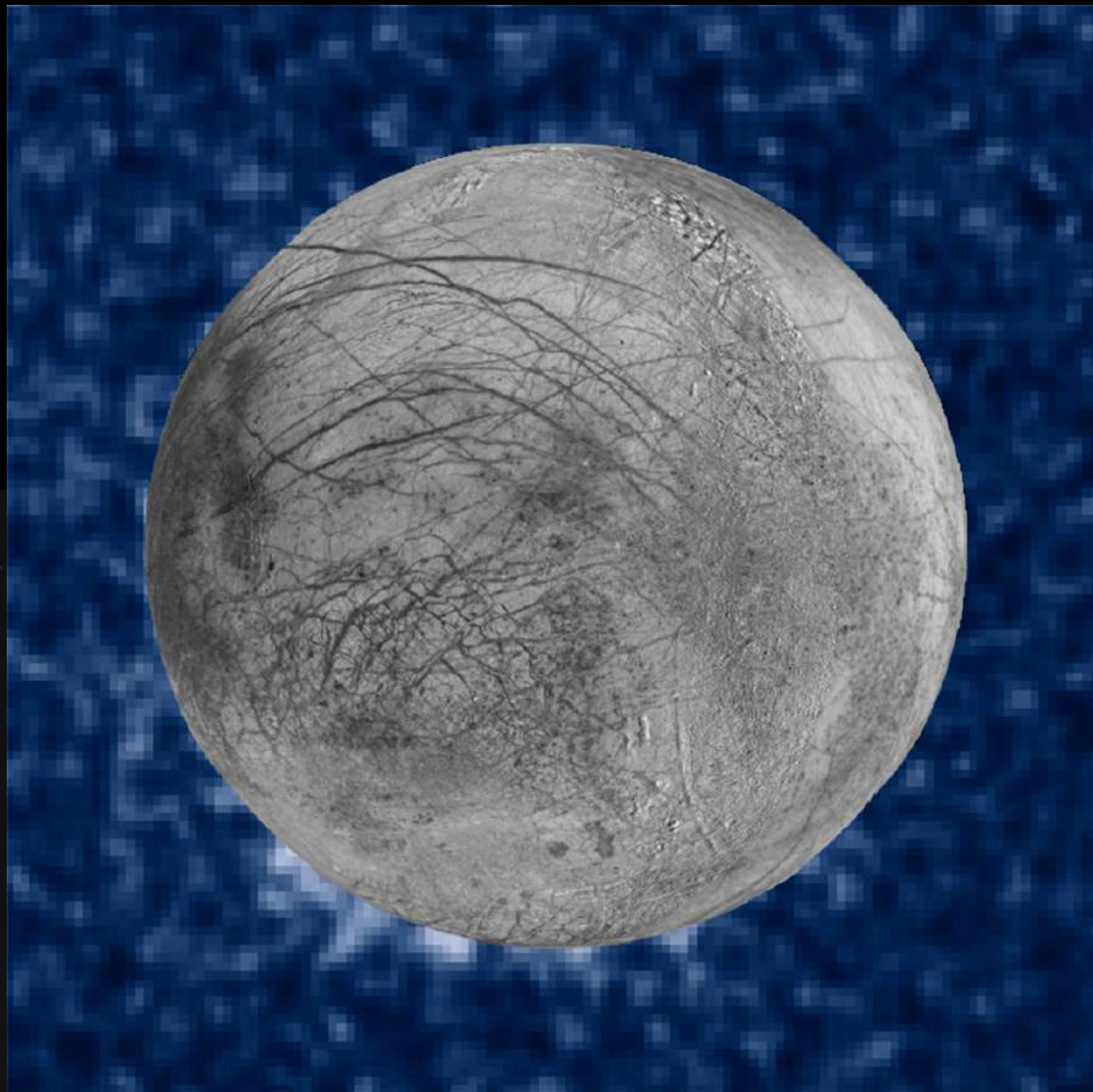
Black Smokers & Hot Vent Creatures at Endeavour Ridge (2190m)

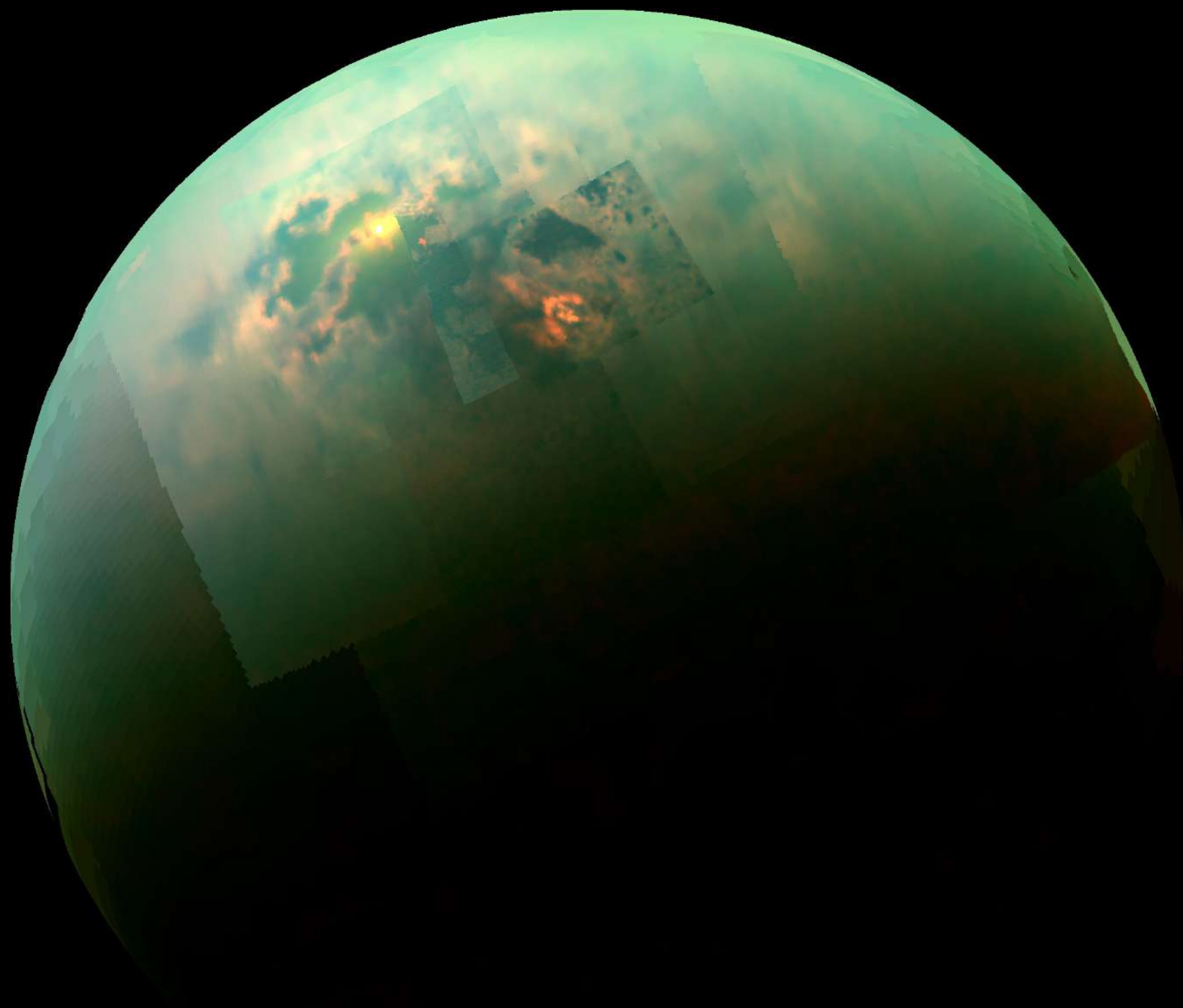
8 October 2010:

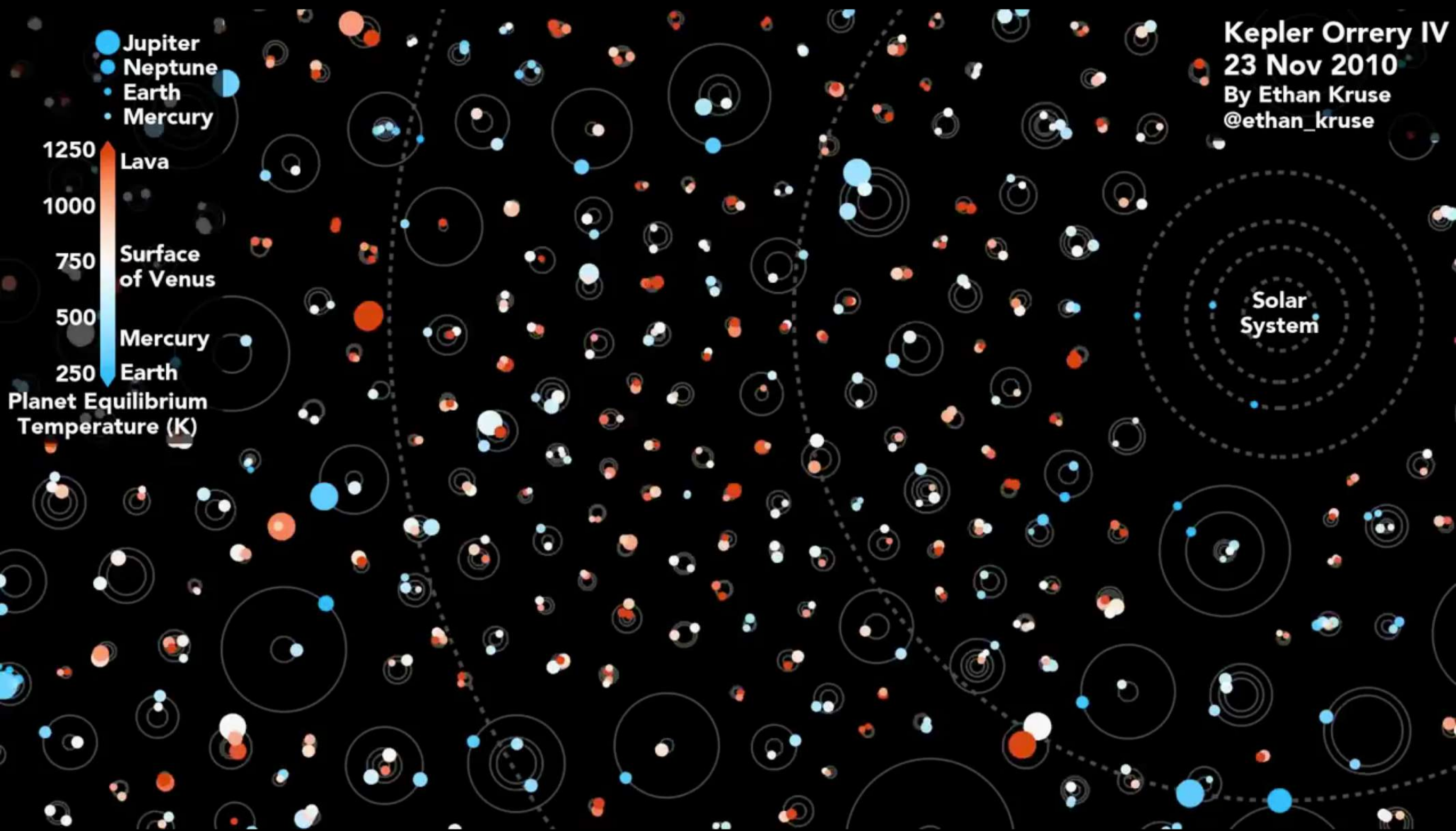
Specially adapted tubeworms, scale worms
and limpets thrive in the
hot sulfide-laced waters of Grotto
Hydrothermal Vent.











Kepler Orrery IV
23 Nov 2010
By Ethan Kruse
@ethan_kruse

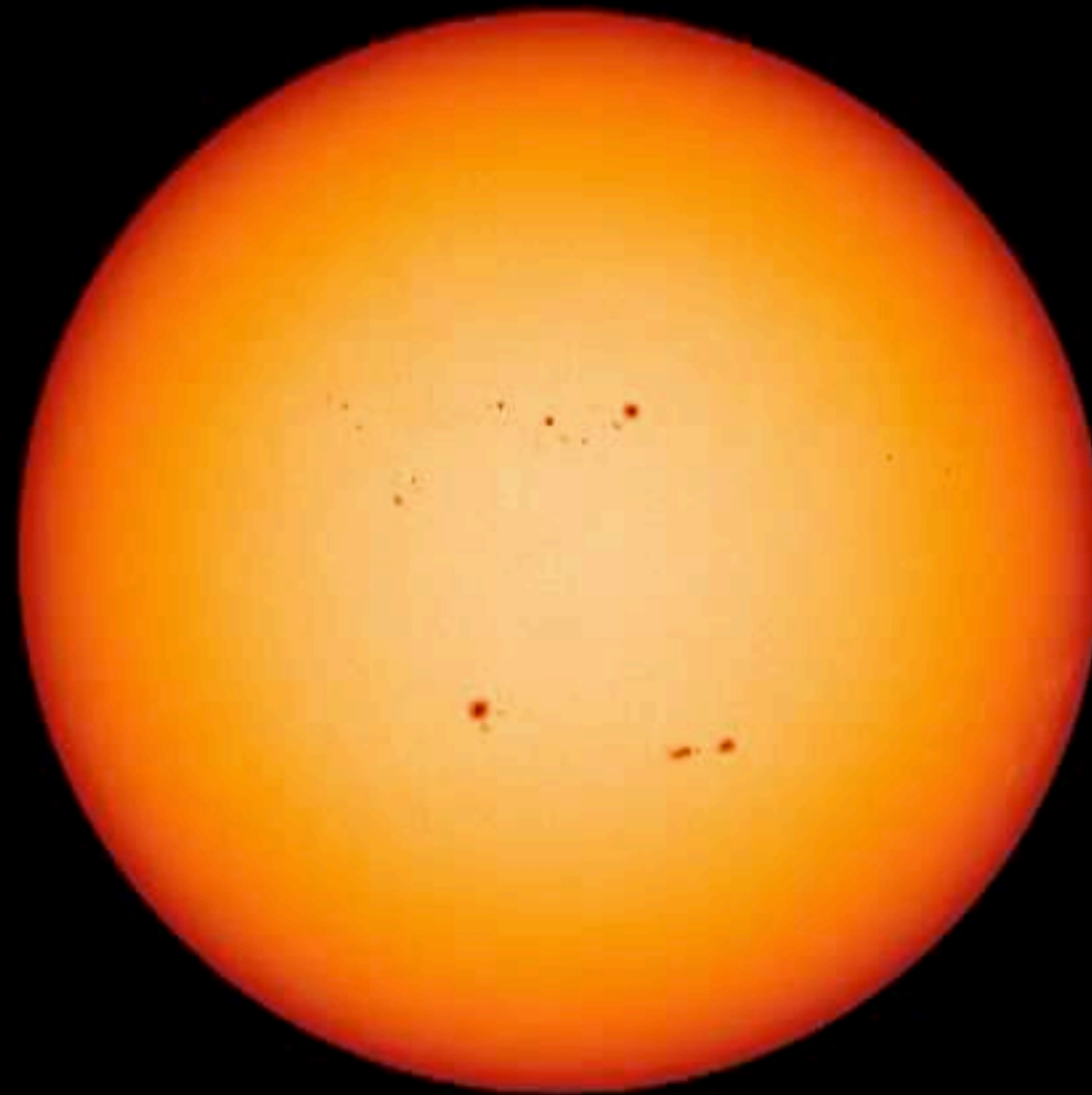


PLANET QUEST

THE SEARCH FOR ANOTHER EARTH



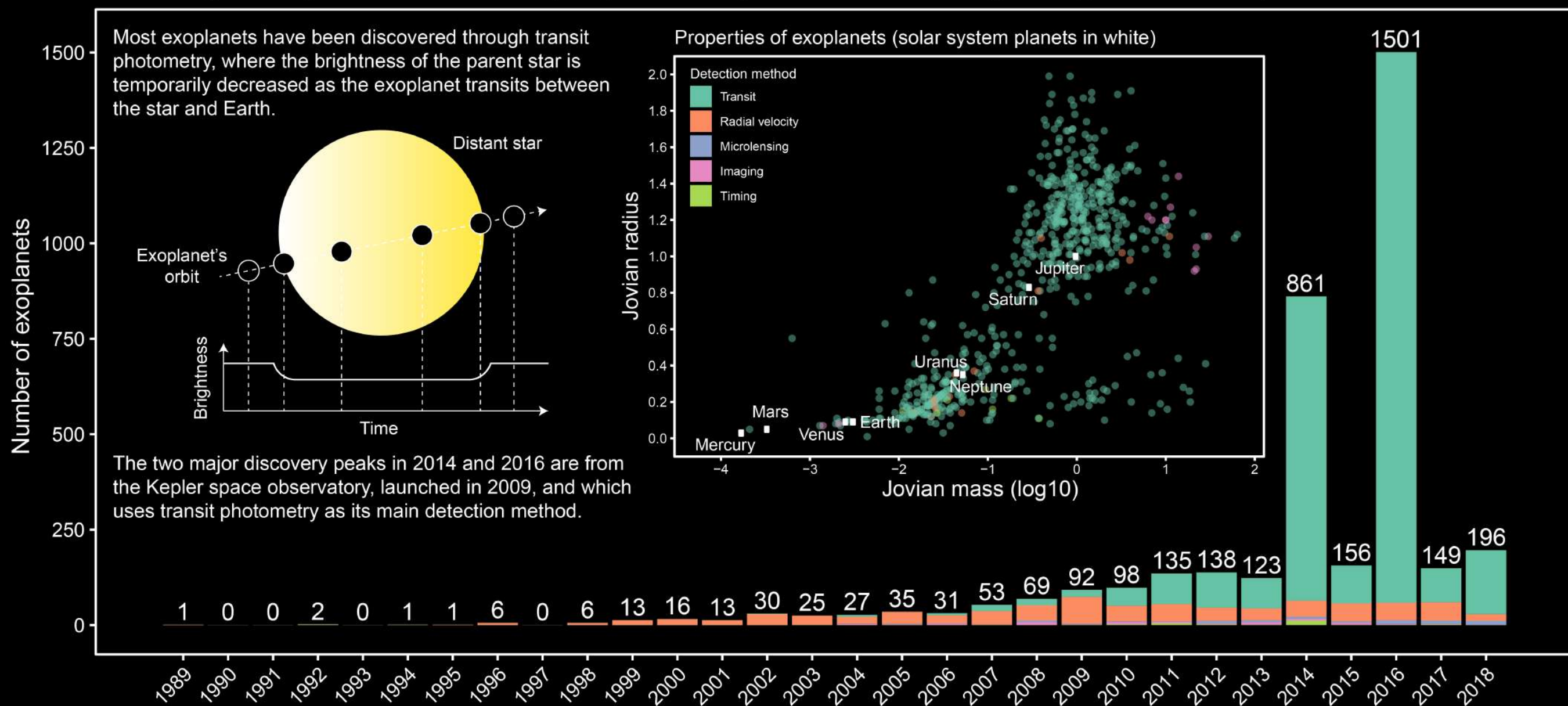
PLANETQUEST
THE SEARCH FOR ANOTHER EARTH



Time (GMT)

June 5

Discovery of exoplanets since 1989 per year and by detection method (per August 21st 2018)





The Periodic Table of Exoplanets

Over 3800 Exoplanets



Hot Zone



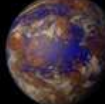



Warm 'Habitable' Zone

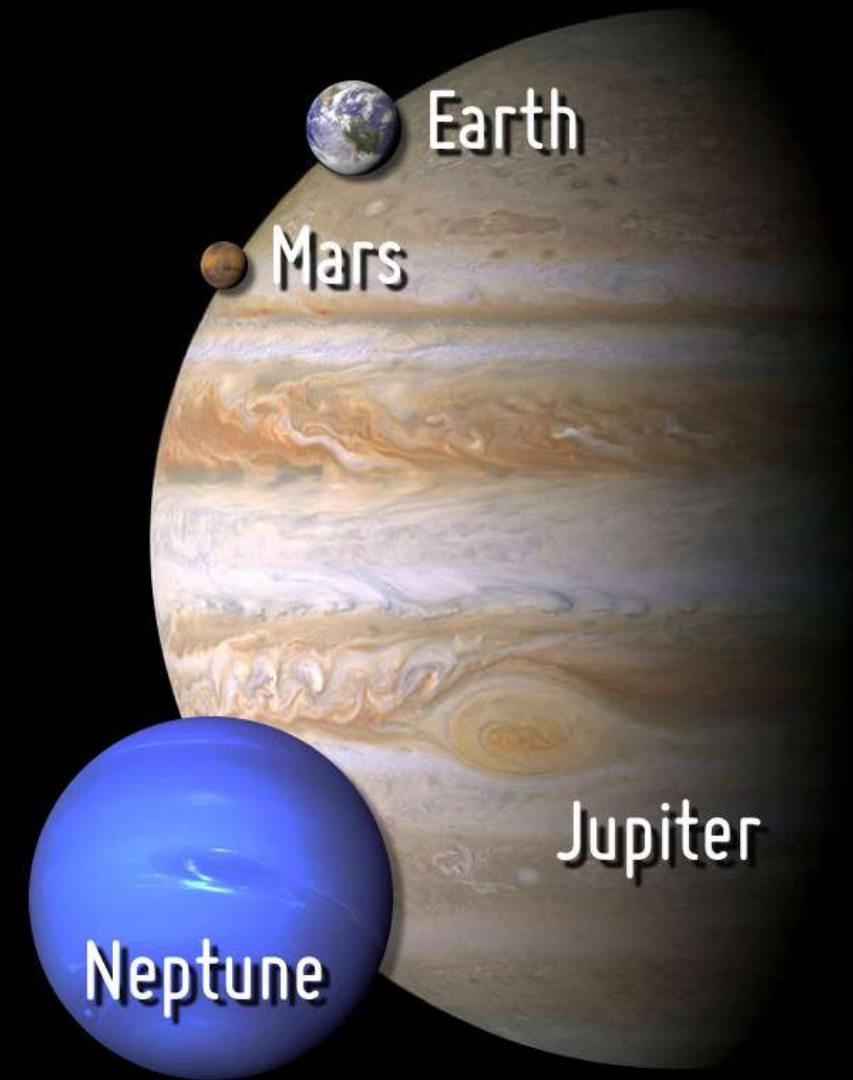
Cold Zone

Terrans Group				Giants Group	
Miniterrans (Mercury Size)	Subterrans (Mars Size)	Terrans (Earth Size)	Superterrans (Super-Earths & Mini-Neptunes)	Neptunians (Neptune Size)	Jovians (Jupiter Size)
$10^{-5} - 0.1 M_E$ or $0.03 - 0.4 R_E$	$0.1 - 0.5 M_E$ or $0.4 - 0.8 R_E$	$0.5 - 5 M_E$ or $0.8 - 1.5 R_E$	$5 - 10 M_E$ or $1.5 - 2.5 R_E$	$10 - 50 M_E$ or $2.5 - 6.0 R_E$	$> 50 M_E$ or $> 6 R_E$
Hot Zone	Hot Zone	Hot Zone	Hot Zone	Hot Zone	Hot Zone
Hot Miniterrans 5 0.1%	Hot Subterrans 54 1.4%	Hot Terrans 614 15.9%	Hot Superterrans 971 25.1%	Hot Neptunians 776 20.0%	Hot Jovians 827 21.4%
Warm 'Habitable' Zone	Potentially Habitable	Potentially Habitable	Potentially Habitable	Warm 'Habitable' Zone	Warm 'Habitable' Zone
Warm Miniterrans 0 0%	Warm Subterrans 1 <0.1%	Warm Terrans 22 0.6%	Warm Superterrans 32 0.8%	Warm Neptunians 39 1.0%	Warm Jovians 139 3.6%
Cold Zone	Cold Zone	Cold Zone	Cold Zone	Cold Zone	Cold Zone
Cold Miniterrans 0 0%	Cold Subterrans 1 <0.1%	Cold Terrans 8 0.2%	Cold Superterrans 10 0.3%	Cold Neptunians 25 0.6%	Cold Jovians 301 7.8%

Potentially Habitable Exoplanets

Ranked by Distance from Earth (light years)

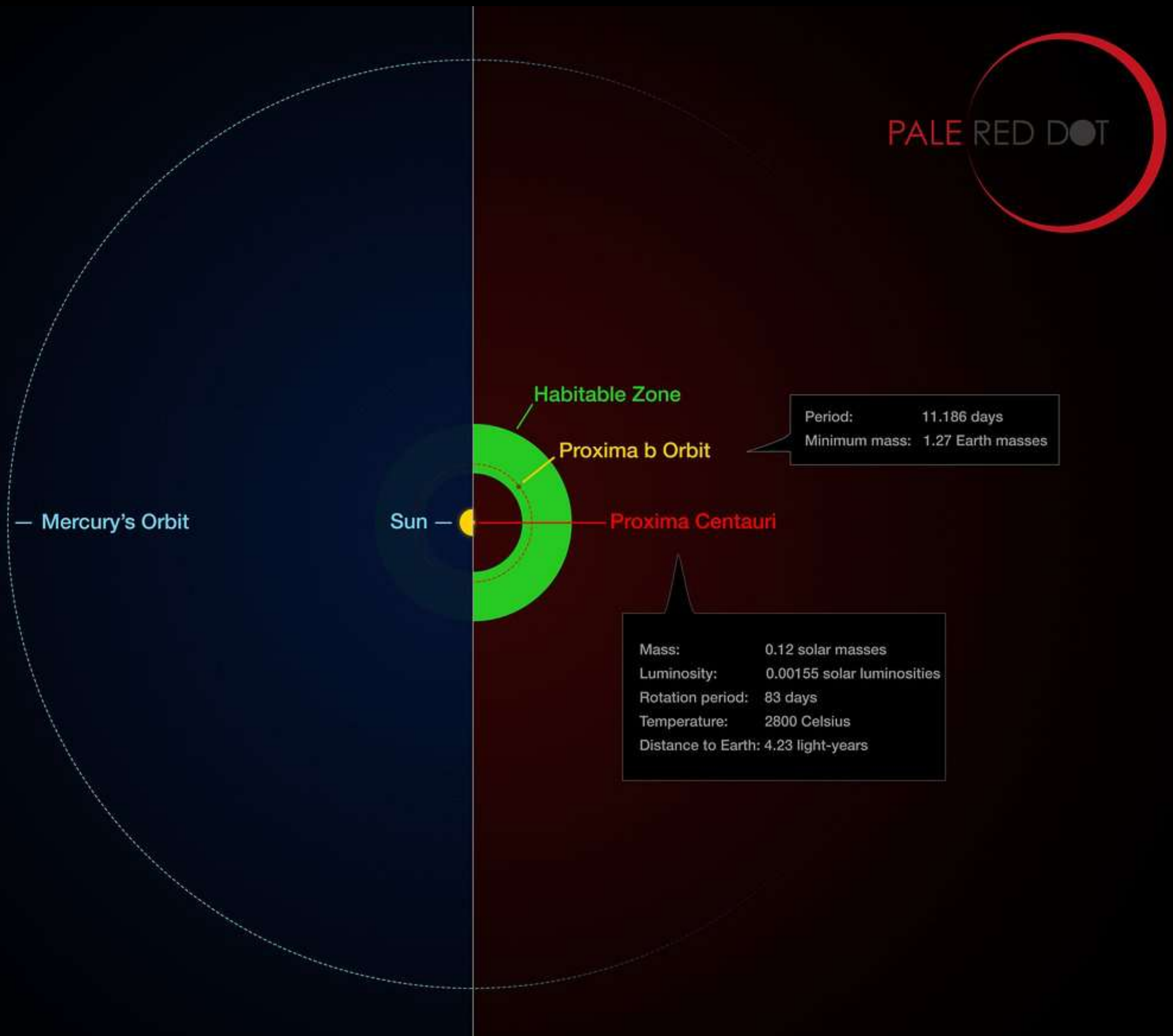
 [4.2 ly] Proxima Cen b	 [12 ly] tau Cet e	 [12 ly] GJ 1061 c	 [12 ly] GJ 1061 d	 [12 ly] GJ 273 b	 [12 ly] Teegarden's Star b
 [12 ly] Teegarden's Star c	 [14 ly] Wolf 1061 c	 [17 ly] GJ 3323 b	 [22 ly] GJ 667 C c	 [22 ly] GJ 667 C e	 [22 ly] GJ 667 C f
 [41 ly] TRAPPIST-1 d	 [41 ly] TRAPPIST-1 e	 [41 ly] TRAPPIST-1 f	 [41 ly] TRAPPIST-1 g	 [217 ly] K2-72 e	 [561 ly] Kepler-186 f
 [770 ly] Kepler-1229 b	 [1115 ly] Kepler-442 b	 [1200 ly] Kepler-62 f			



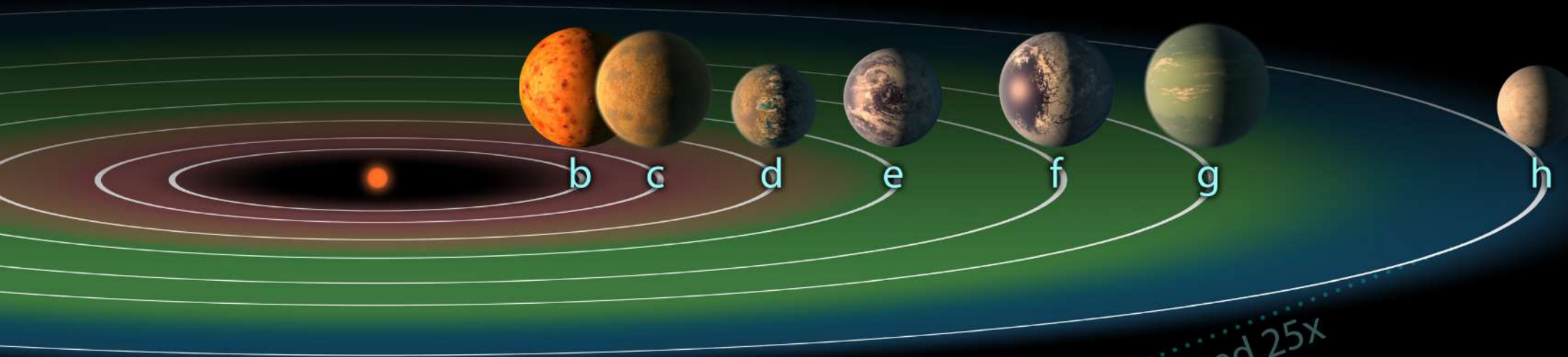
Artistic representations. Earth, Mars, Jupiter, and Neptune for scale. Distance from Earth is between brackets.

CREDIT: PHL @ UPR Arecibo (phl.upr.edu) Sep 4, 2019

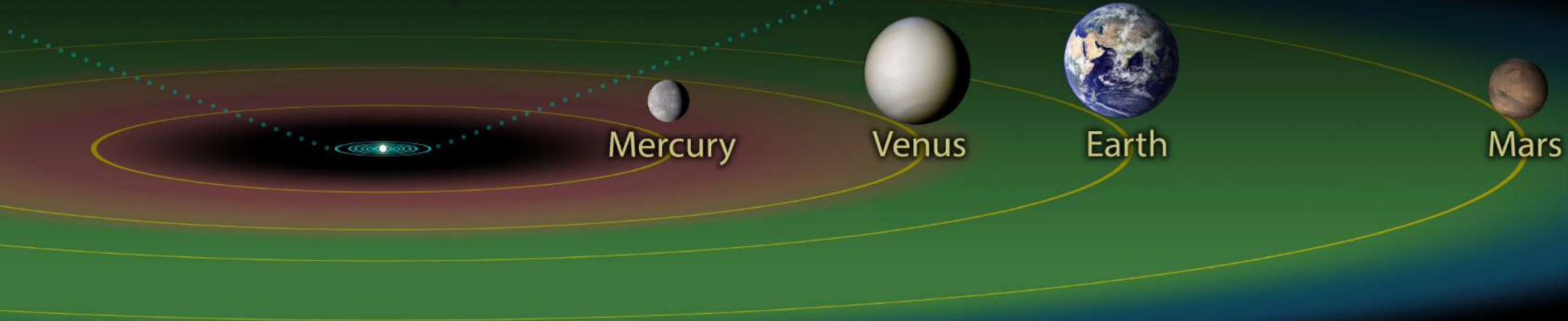
PALE RED DOT



TRAPPIST-1 System



Inner Solar System



Enlarged 25x

Illustration

A search for life on Earth from the Galileo spacecraft

**Carl Sagan^{*}, W. Reid Thompson^{*}, Robert Carlson[†], Donald Gurnett[‡]
& Charles Hord[§]**

^{*} Laboratory for Planetary Studies, Cornell University, Ithaca, New York 14853, USA

[†] Atmospheric and Cometary Sciences Section, Jet Propulsion Laboratory, Pasadena, California 91109, USA

[‡] Department of Physics and Astronomy, University of Iowa, Iowa City, Iowa 52242-1479, USA

[§] Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, Colorado 80309, USA

In its December 1990 fly-by of Earth, the Galileo spacecraft found evidence of abundant gaseous oxygen, a widely distributed surface pigment with a sharp absorption edge in the red part of the visible spectrum, and atmospheric methane in extreme thermodynamic disequilibrium; together, these are strongly suggestive of life on Earth. Moreover, the presence of narrow-band, pulsed, amplitude-modulated radio transmission seems uniquely attributable to intelligence. These observations constitute a control experiment for the search for extraterrestrial life by modern interplanetary spacecraft.

At ranges varying from ~100 km to ~100,000 km, spacecraft have now flown by more than 60 planets, satellites, comets and asteroids. They have been equipped variously with imaging systems, photometric and spectrometric instruments extending from ultraviolet to kilometre wavelengths, magnetometers and charged-particle detectors. In none of these encounters has compelling, or even strongly suggestive, evidence for extraterrestrial life been found. For the Moon, Venus and Mars, orbiter and lander observations confirm the conclusion from fly-by spacecraft. Still, extraterrestrial life, if it exists, might be quite unlike the forms of life

with which we are familiar, or present only marginally. The most elementary test of these techniques—the detection of life on Earth by such an instrumented fly-by spacecraft—had, until recently, never been attempted.

Galileo is a single-launch Jupiter orbiter and entry probe currently in interplanetary space and scheduled to arrive in the Jupiter system in December 1995. It could not be sent directly to Jupiter; instead, the mission incorporated two close gravitational assists at the Earth and one at Venus. This greatly lengthened the transit time, but it also permitted close observations of the Earth. The

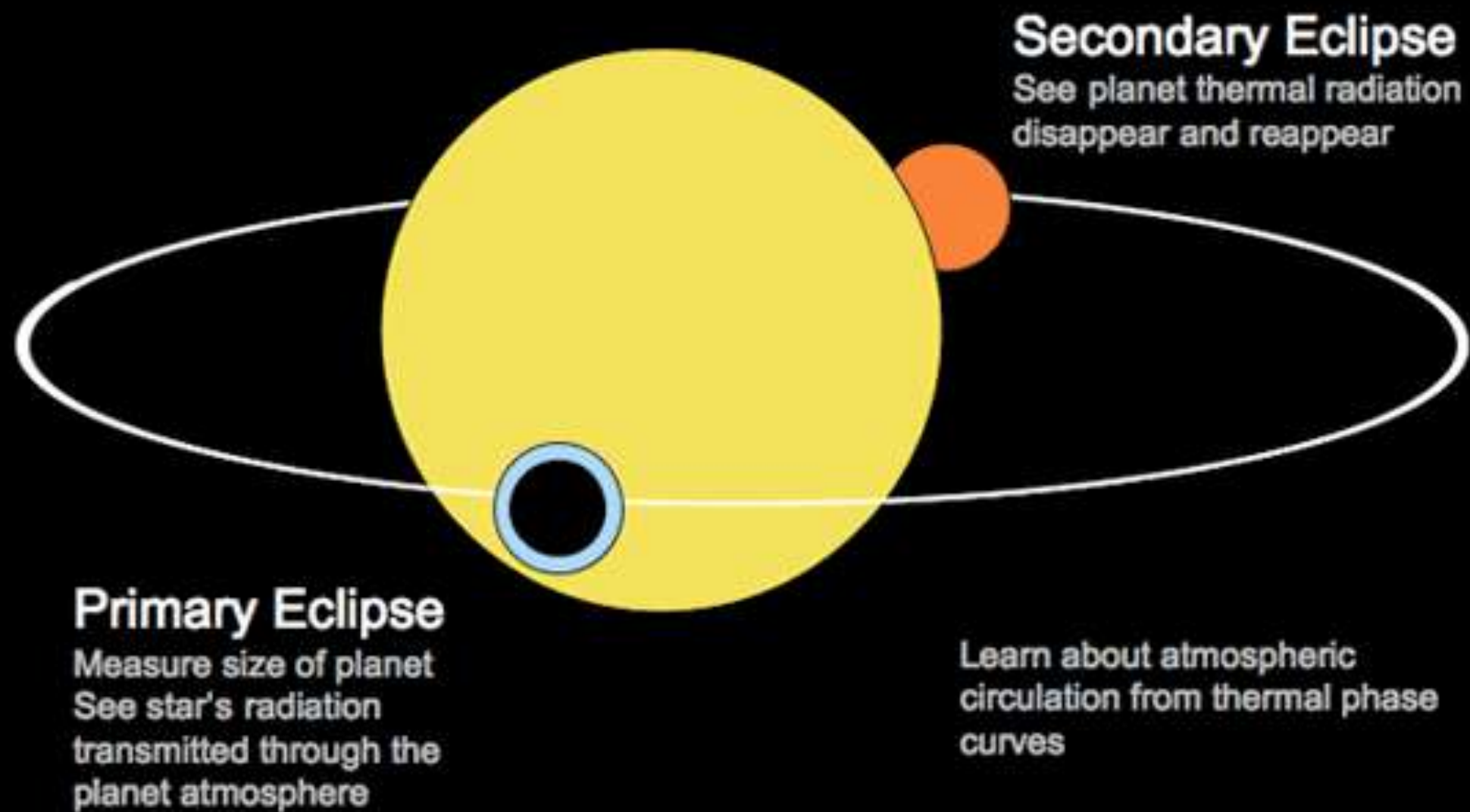
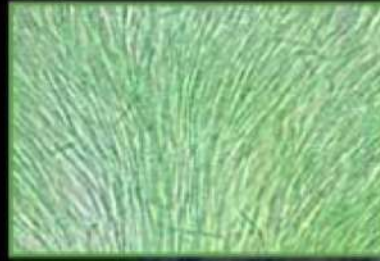
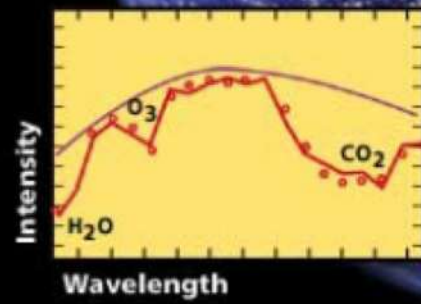


Figure by S. Seager

O₃ Ozone, produced by plants, algae



H₂O Liquid water



Methane produced by living organisms

